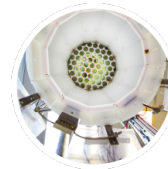


Bem vindos (remotamente ...) ao LIP !



**Experimental particle and
astroparticle physics**



**Development of new
instruments and methods**



Scientific computing



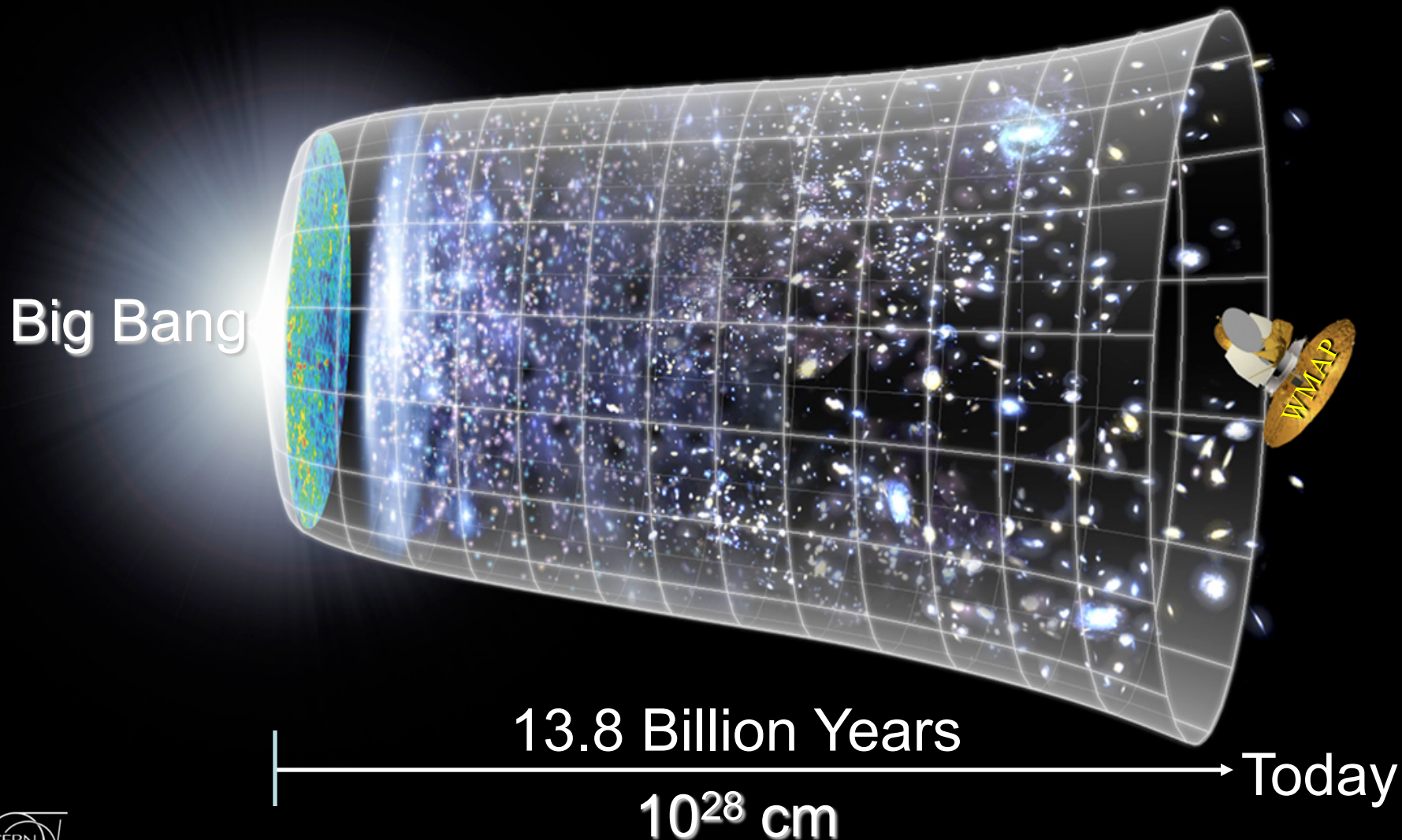
**Knowledge transfer,
education and outreach**

A aventura da Física de Partículas e Astropartículas

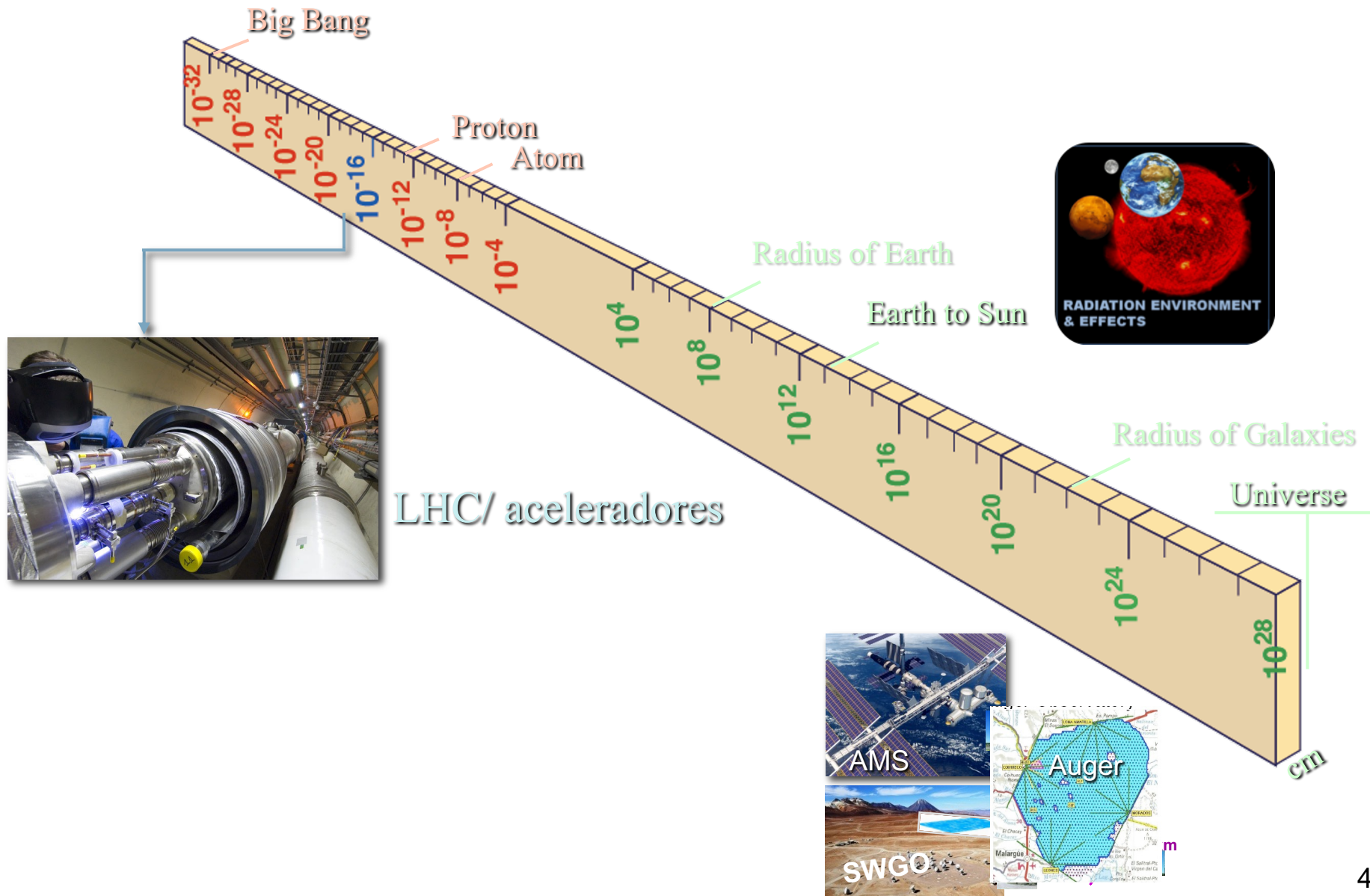


Mário Pimenta
Lisboa, Julho 2021

O Universo para compreender, ...



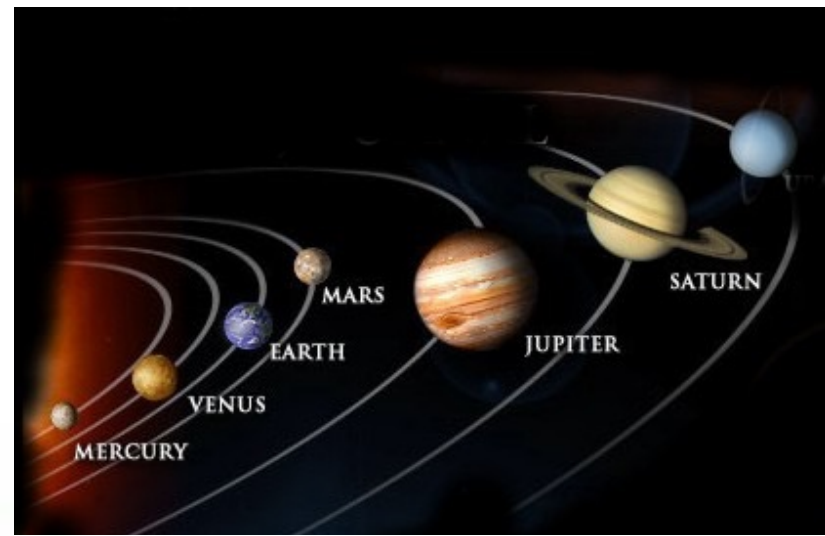
Escalas, ...



Da maçã ao Universo



I. Newton



Lei da atracção
Universal:

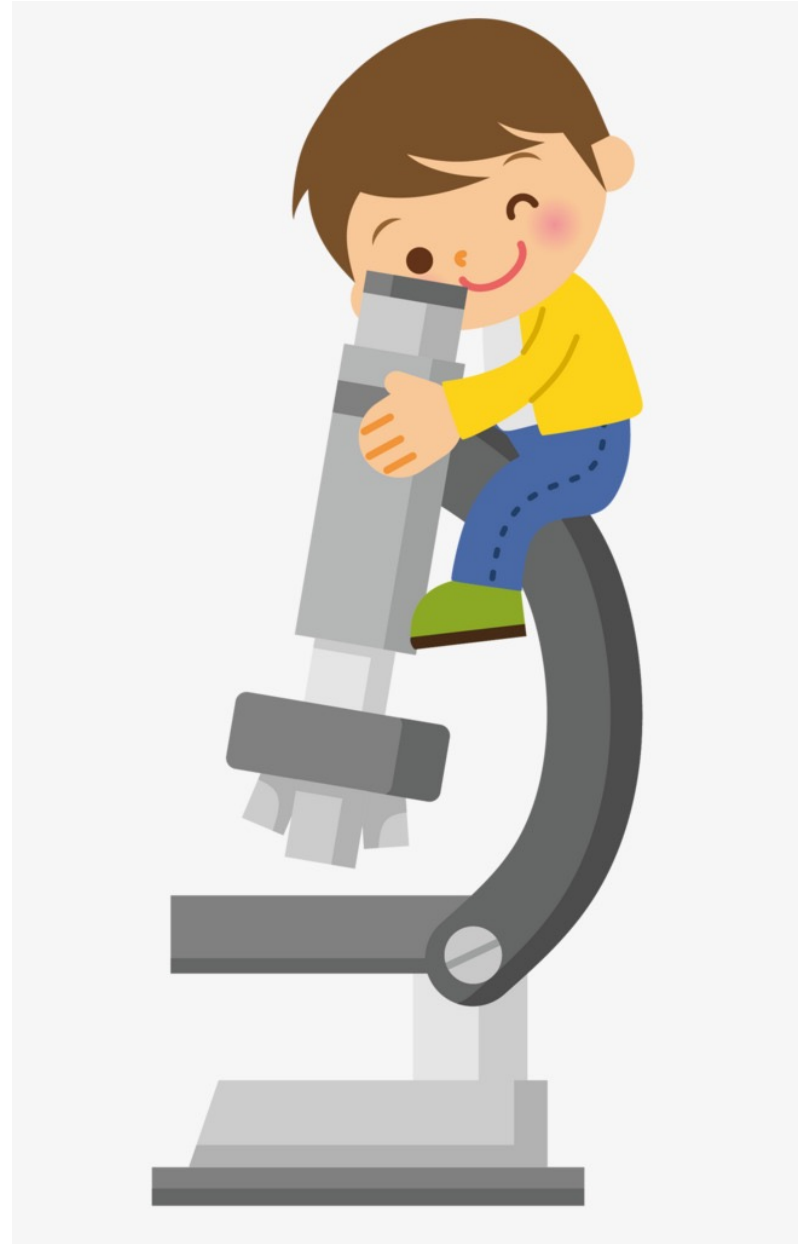
$$F = G \frac{m_{1g} m_{2g}}{r_{12}^2}$$

A gravidade !

Será que a maçã é
elementar?



Ver o interior da
maçã ???



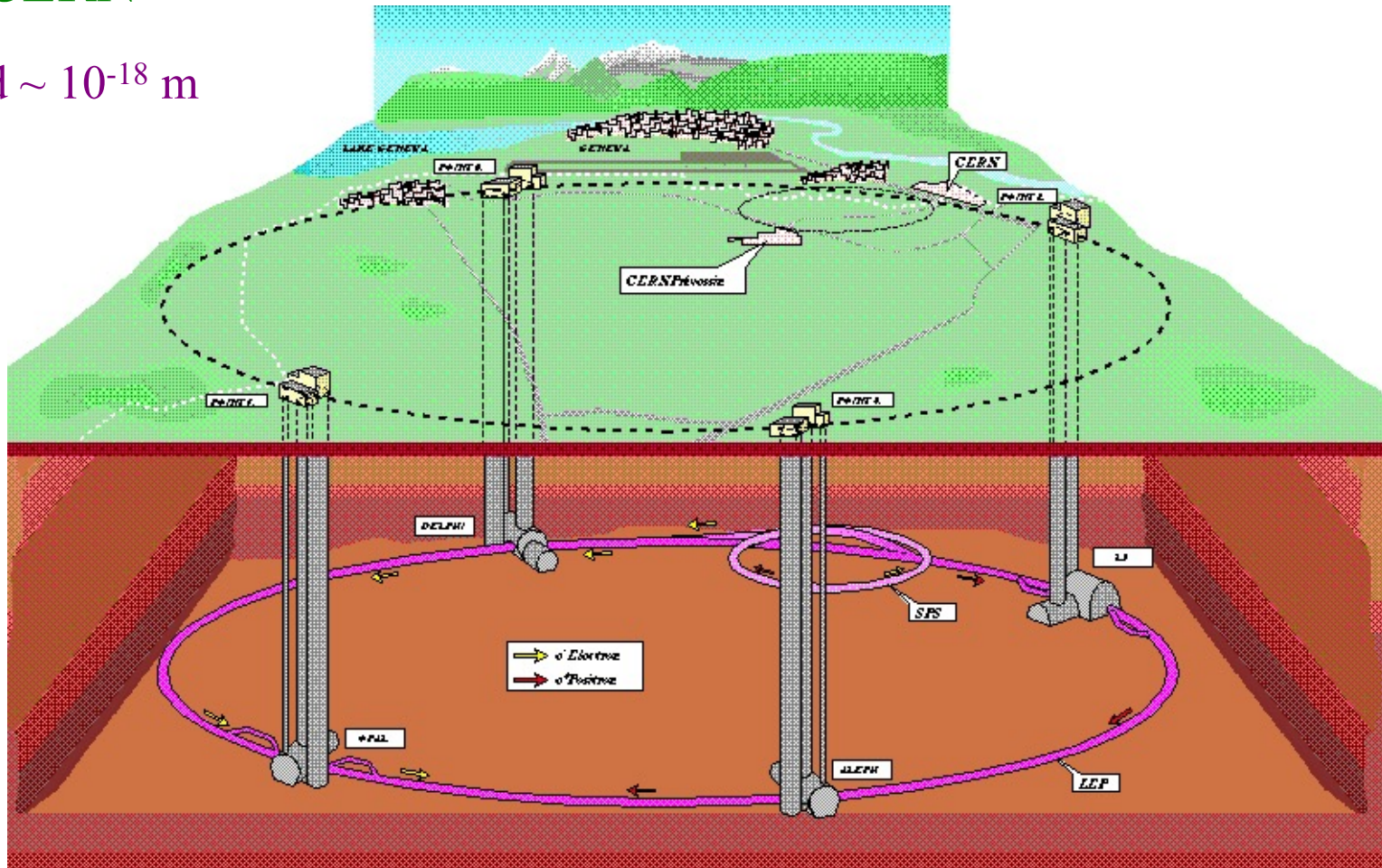
Ver com partículas

$$\lambda = \frac{h}{p}$$

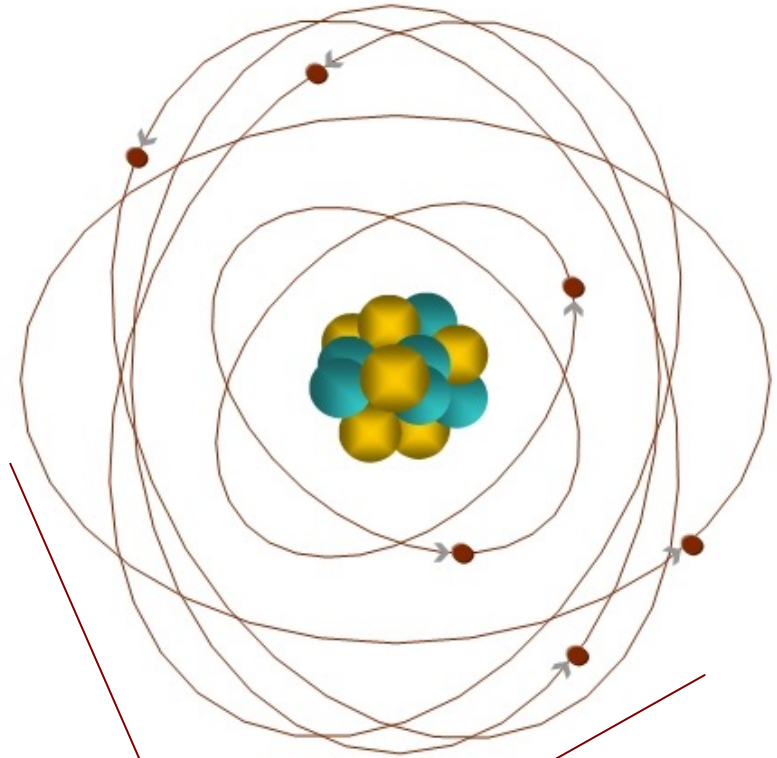
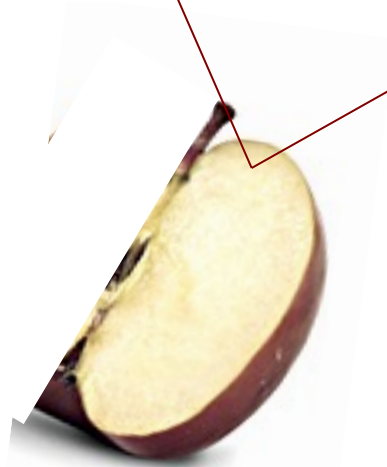


CERN

$d \sim 10^{-18}$ m

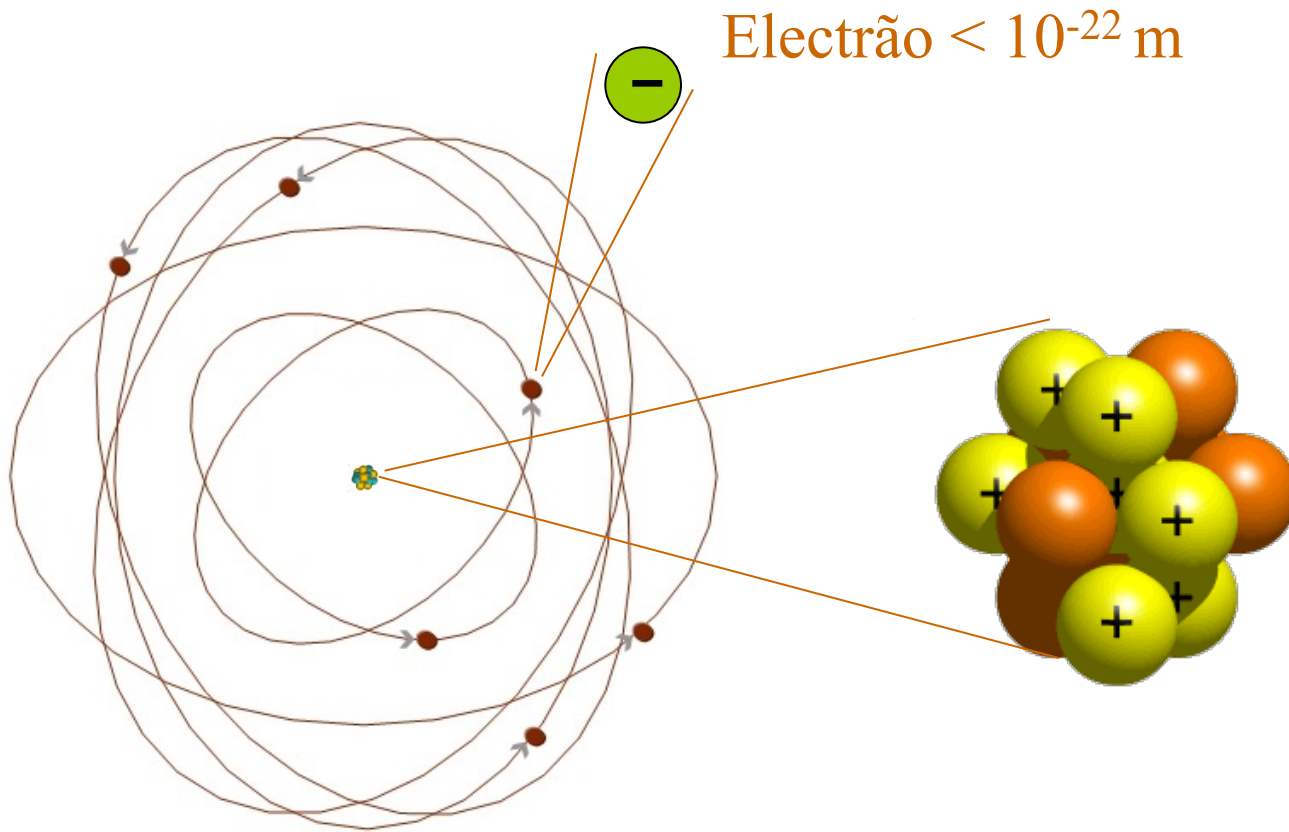


Será que a maçã é elementar?

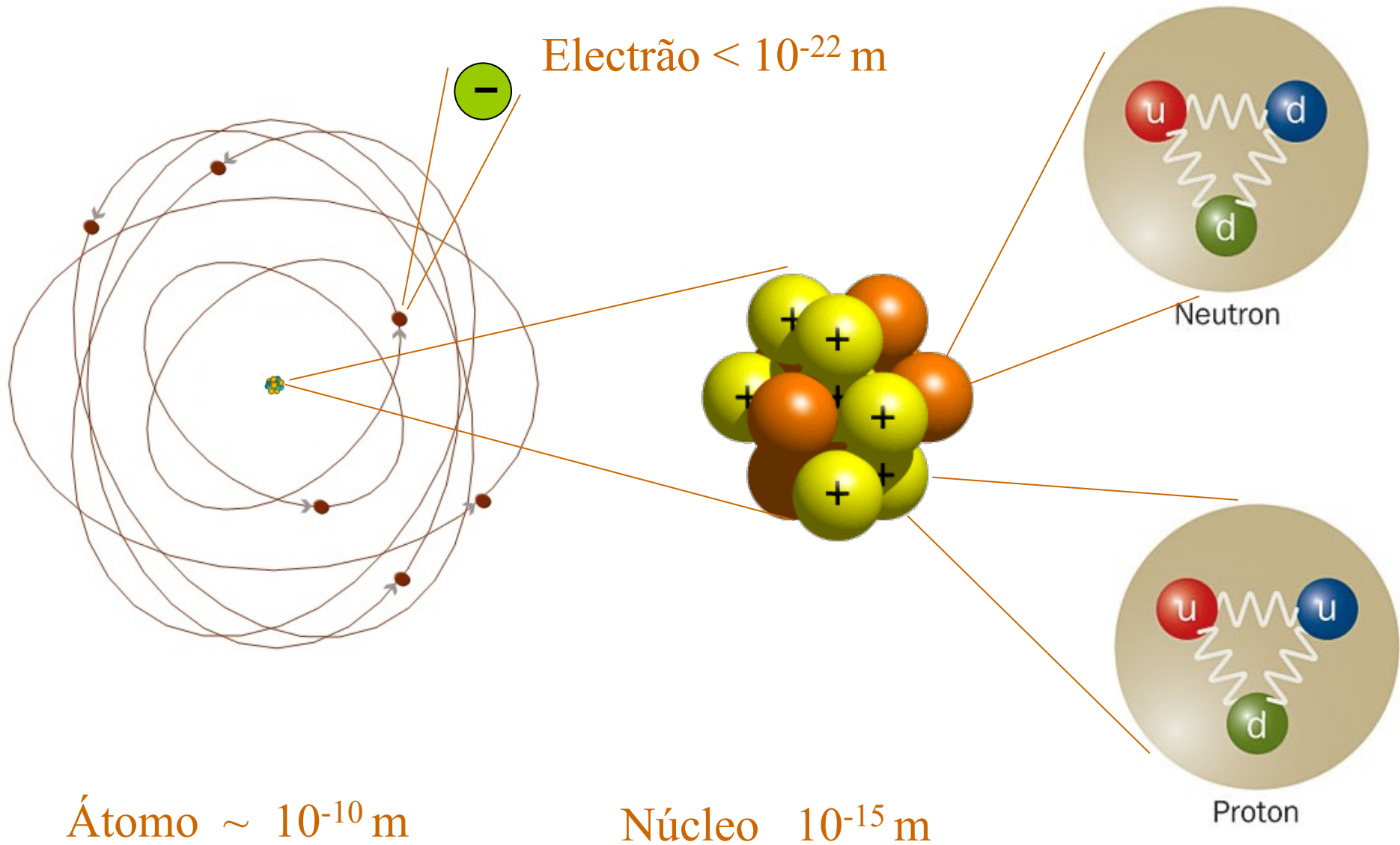


Átomo $\sim 10^{-10}$ m

No interior dos átomos



No interior dos átomos



Vários quarks, muitas partículas !!!

Quarks:



$$\begin{pmatrix} u \\ d \end{pmatrix}$$

$$Q = + 2/3$$

$$Q = - 1/3$$

Vários quarks, muitas partículas !!!

Quarks:

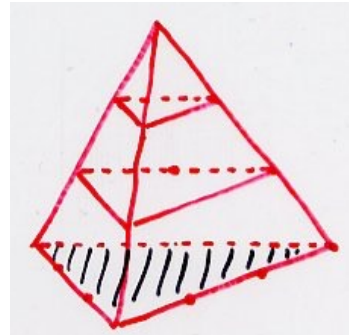


$$\begin{pmatrix} u \\ d \end{pmatrix} \quad \begin{pmatrix} c \\ s \end{pmatrix} \quad \begin{pmatrix} t \\ b \end{pmatrix} \quad \begin{pmatrix} ? \\ ? \end{pmatrix} \quad \begin{matrix} Q = + 2/3 \\ Q = - 1/3 \end{matrix}$$

Combinar



$$q_i q_j q_k$$



$$q_i \bar{q}_i$$

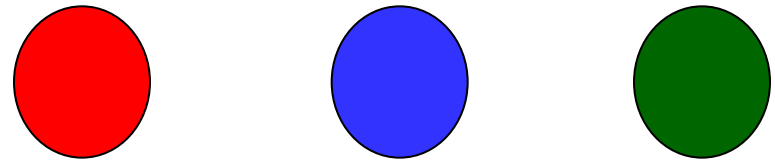


$$p, n, \Sigma, \lambda, \Xi, \Delta, \dots$$

$$\pi, k, \eta, \rho, \psi, \dots$$

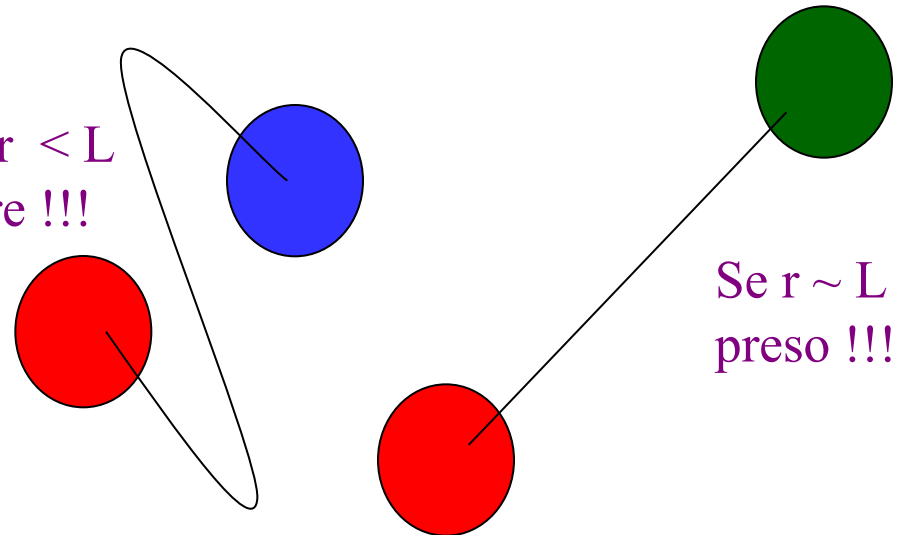
A força forte

Quarks de três cores:



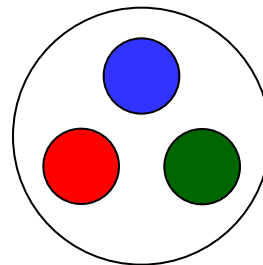
Ligados por “cordas”

Se $r < L$
livre !!!

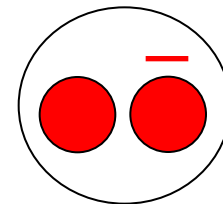


Se $r \sim L$
preso !!!

Partículas “sem cor”

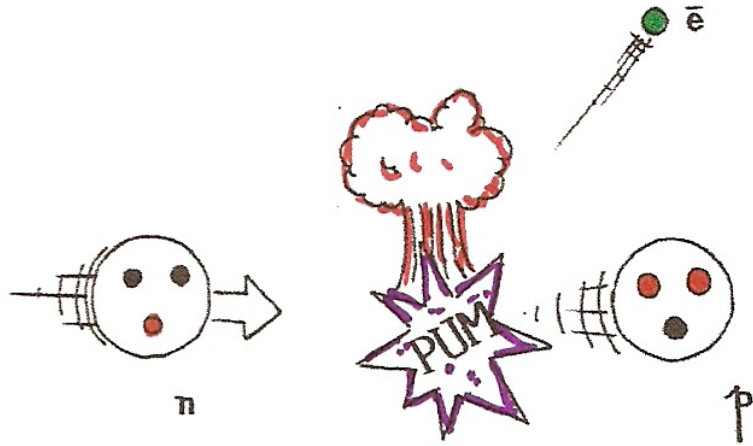


Bariões

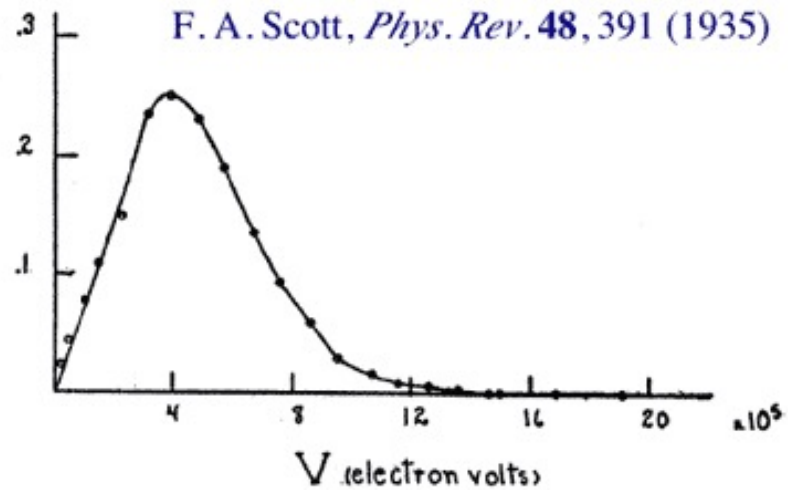


Mesões

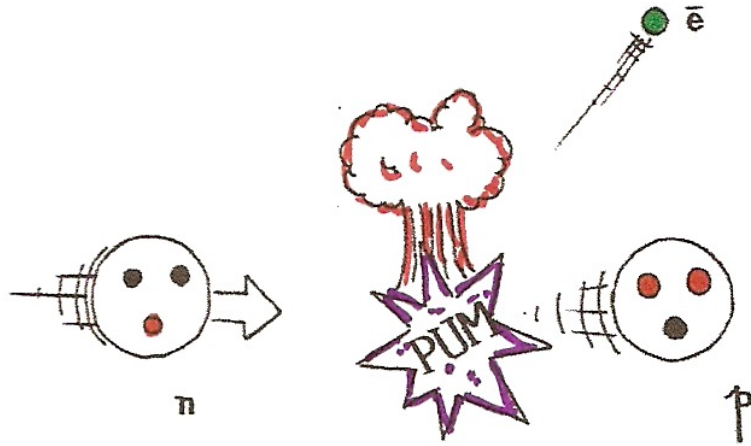
O neutrão decai...



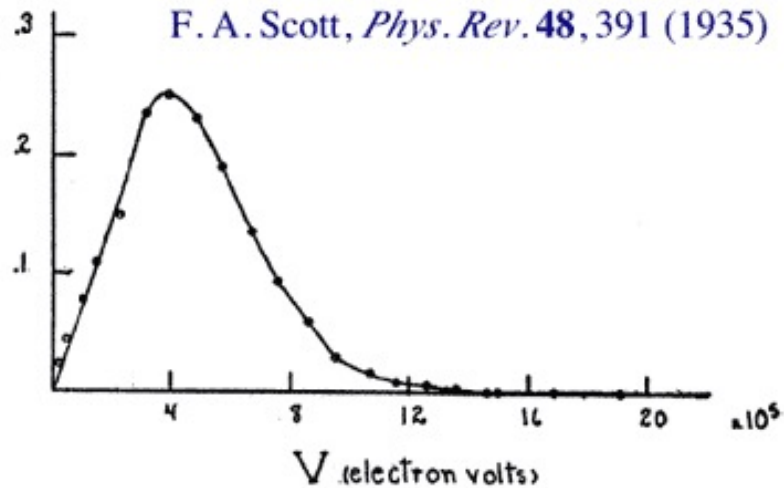
conservação E e \vec{P} ???



O neutrão decai...



conservação E e \vec{P} ???



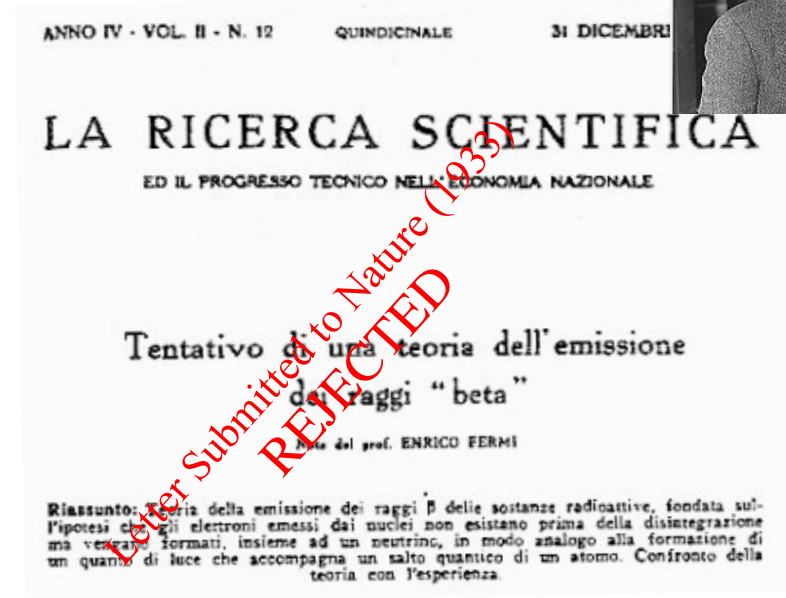
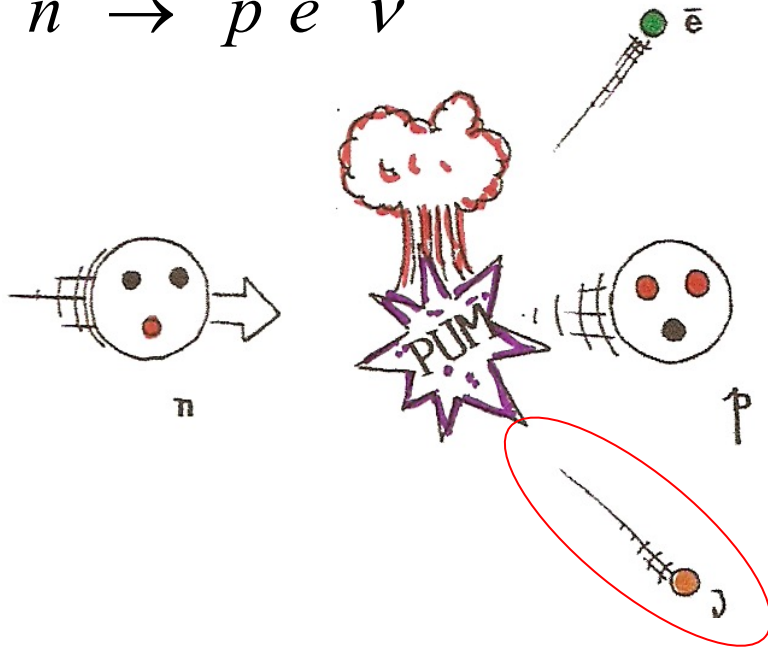
Pauli (1930)



O neutrino !!!

A força fraca

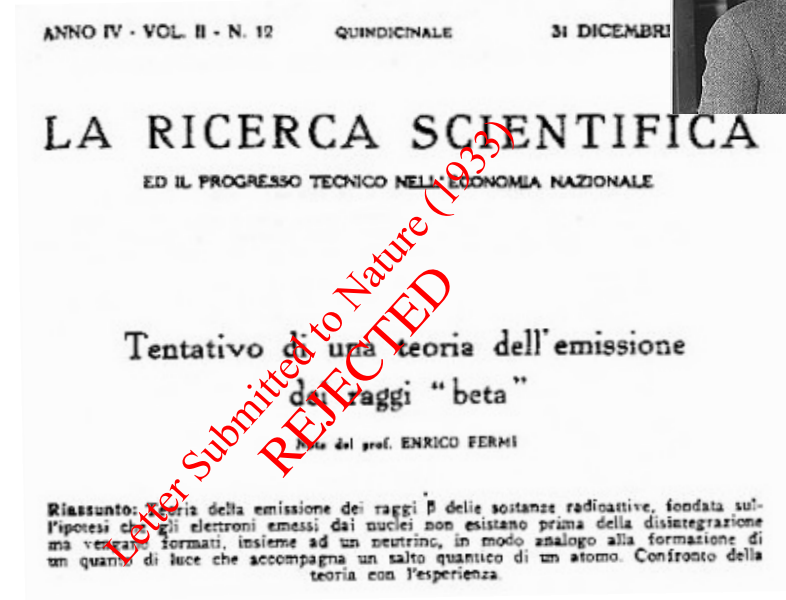
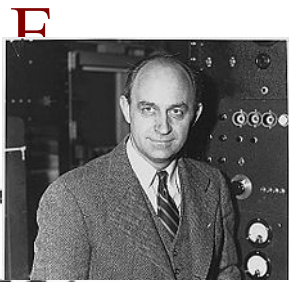
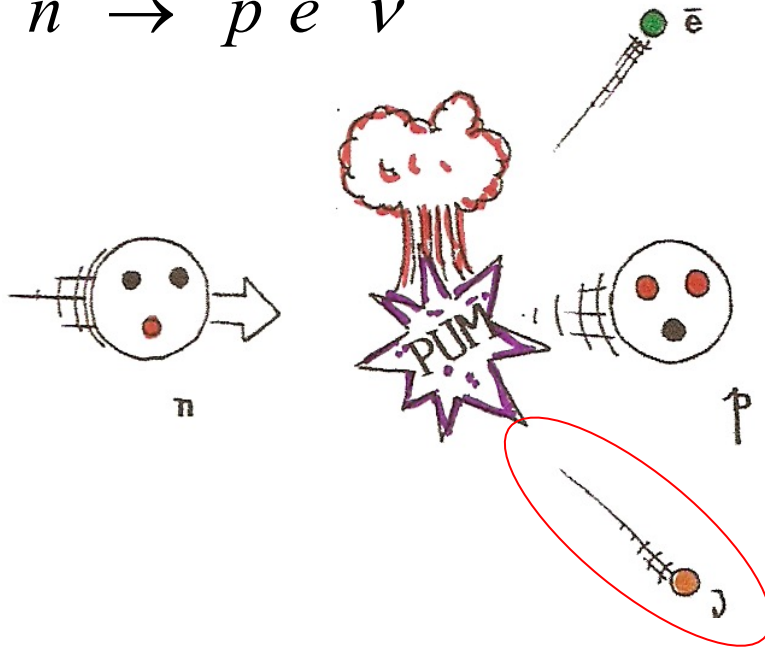
$$n \rightarrow p e^- \bar{\nu}$$



Nuovo Cimento and Zeitschrift fur Physik

A força fraca

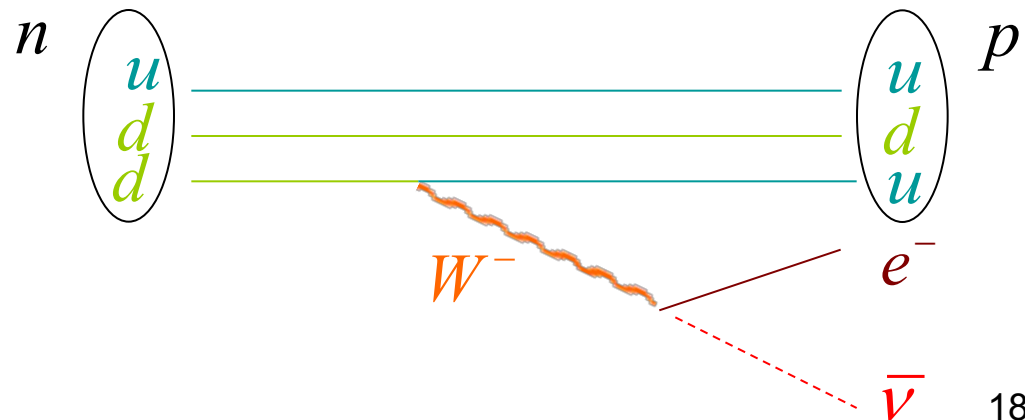
$$n \rightarrow p e^- \bar{\nu}$$



Nuovo Cimento and Zeitschrift fur Physik

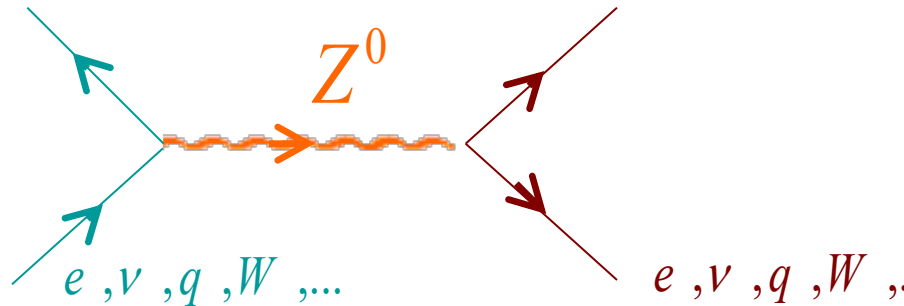
A visão moderna !

Um quark **d** transforma-se num quark **u** emitindo um bosão **W** que “decai” num par electrão, anti-neutrino

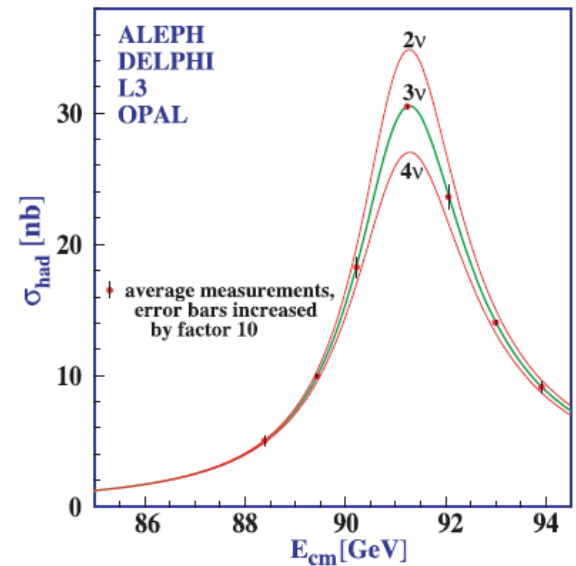
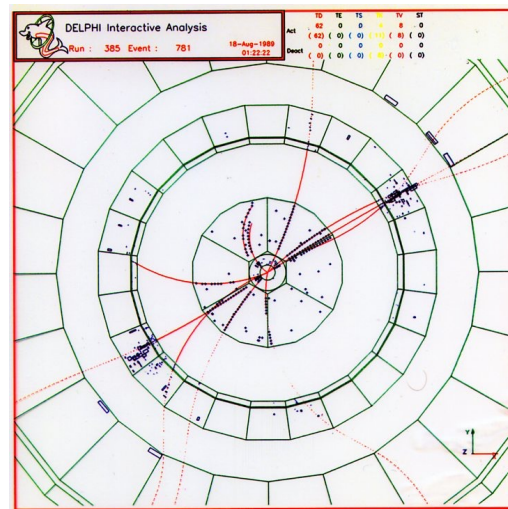
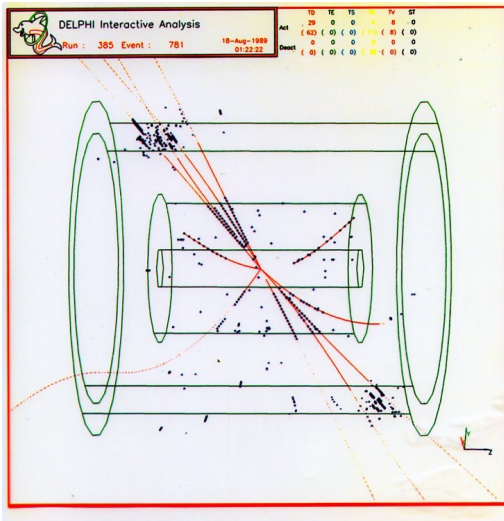
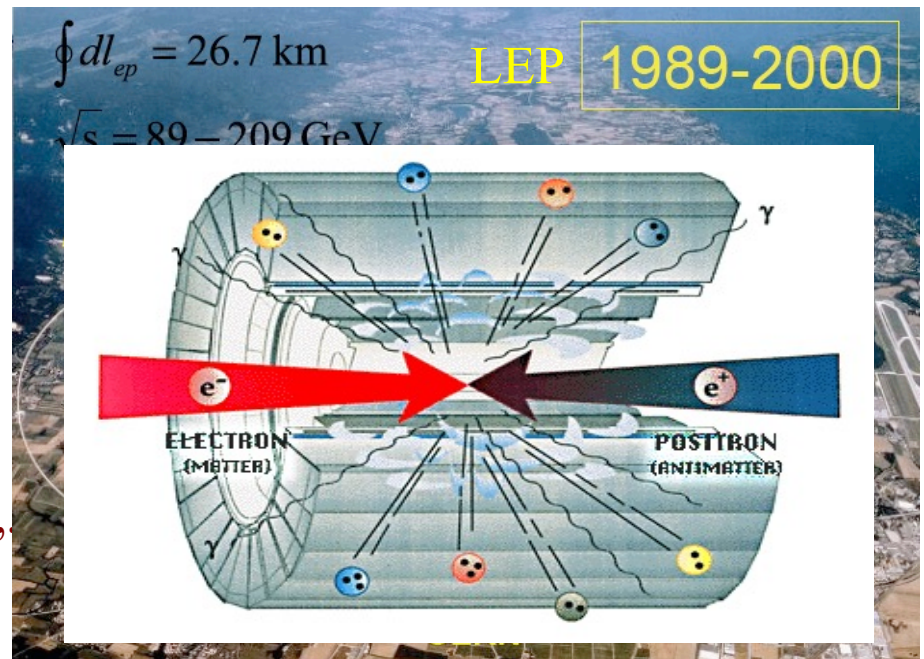


$O Z^0$

Parceiro neutro dos W 's



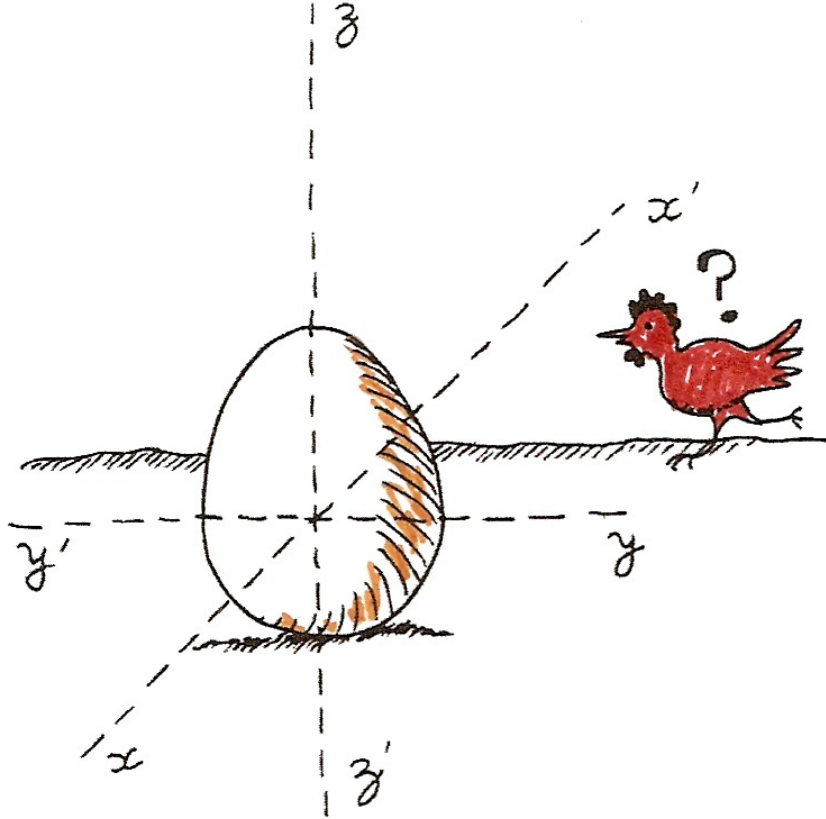
$$e^+e^- \rightarrow q\bar{q}$$



$$M_Z = 91.1875 \pm 0.0021 \text{ GeV}/c^2$$

Simetrias !!!

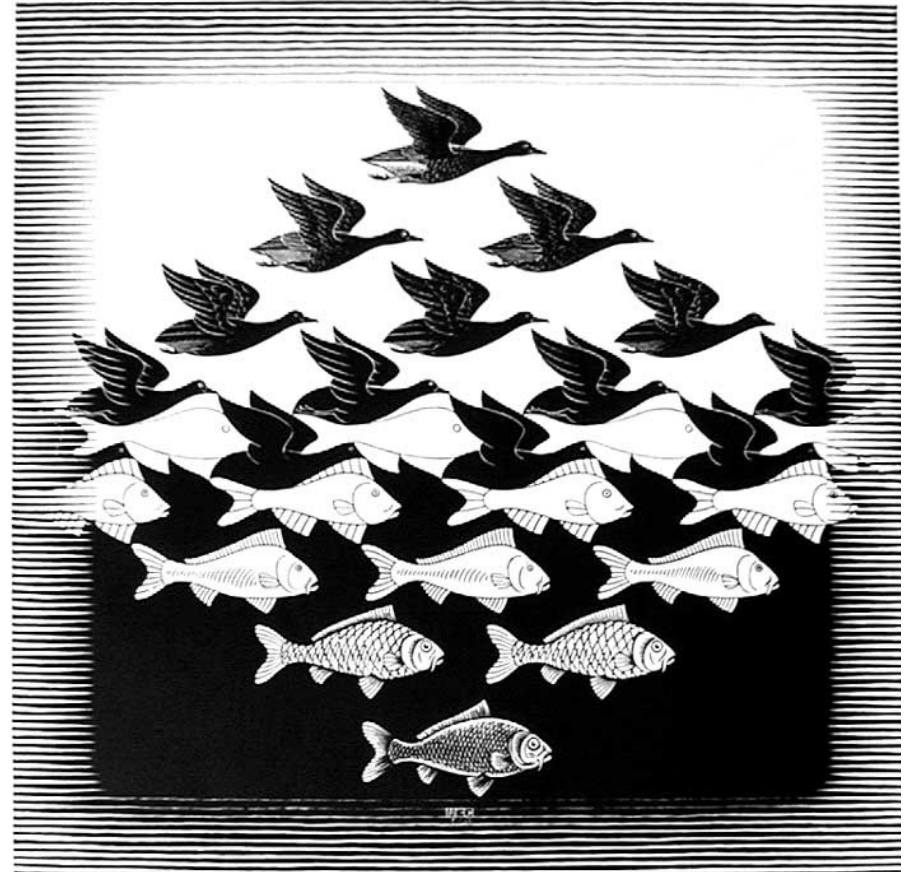
Globais



Leis de conservação:
Energia, Momento linear,
Momento Angular

Locais

M.C.Escher

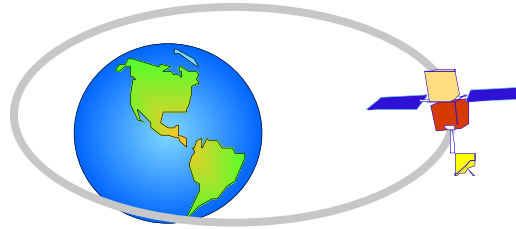


Campos de interacção:
Electromagnético, fraco, forte,
gravítico???,

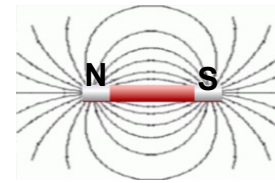
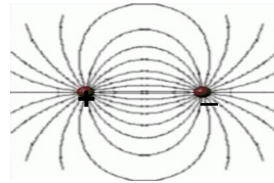
As interações

Unificação e Simetrias (locais)

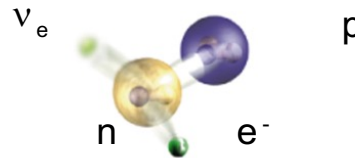
gravíticas



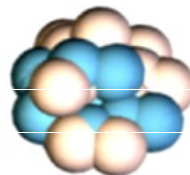
electromagneticas



fracas



fortes



Mas as Partículas têm massa

em eV/c^2

$$m_\nu \sim 10^{-1}$$

$$m_\gamma = 0$$

$$m_e \sim 500 \cdot 10^{-3}$$

$$m_W \sim 80 \cdot 10^9$$

$$m_u \sim 5 \cdot 10^{-6}$$

$$m_Z \sim 91 \cdot 10^9$$

$$m_t \sim 174 \cdot 10^9$$

$$m_H \sim 125 \cdot 10^9$$

...

Como é que as massas são geradas ???

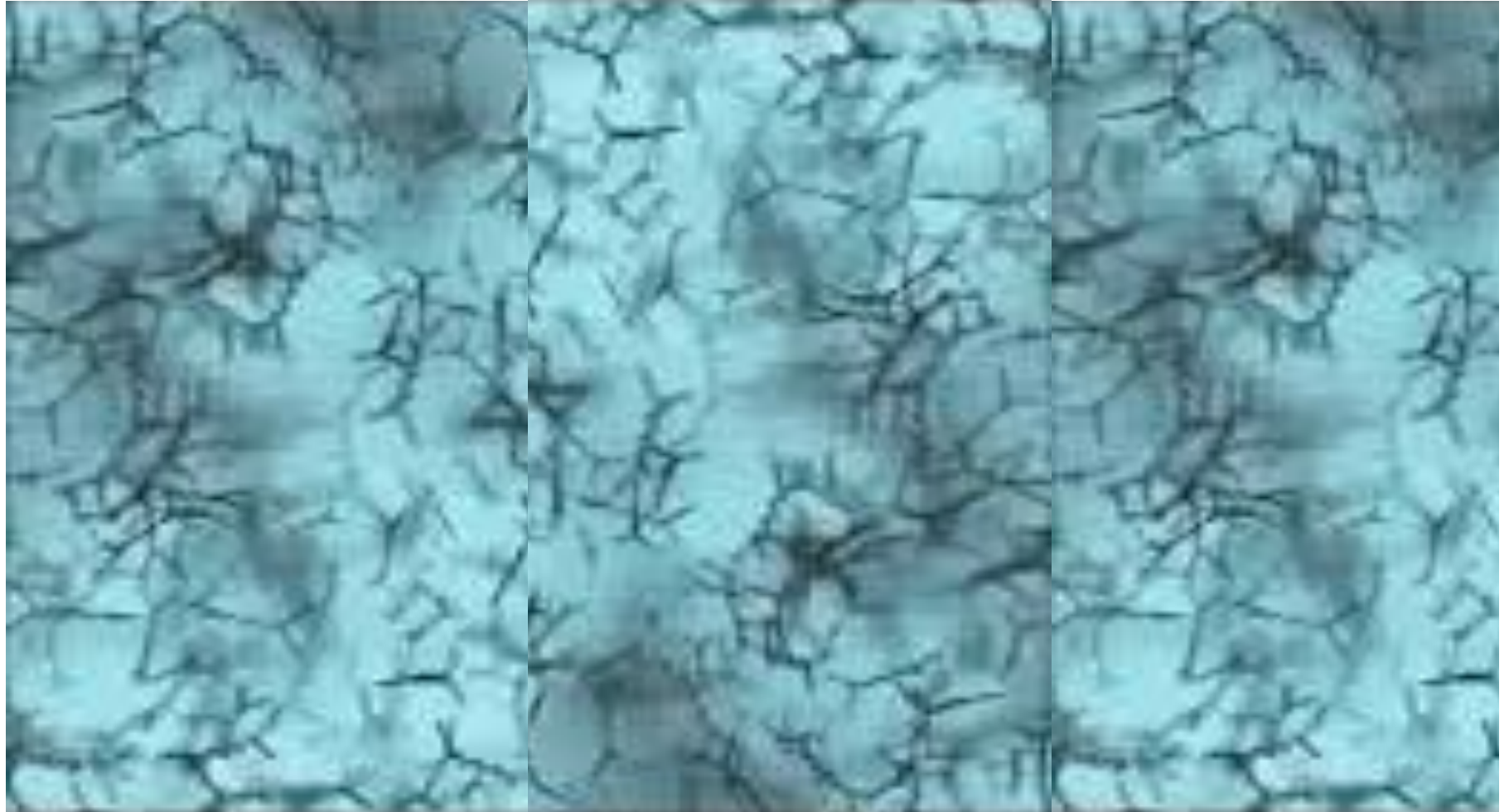
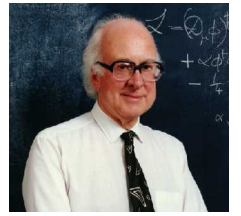
Por é que são tão diferentes ???

$$1 \text{ eV} \sim 1.8 \cdot 10^{-36} \text{ Kg}$$

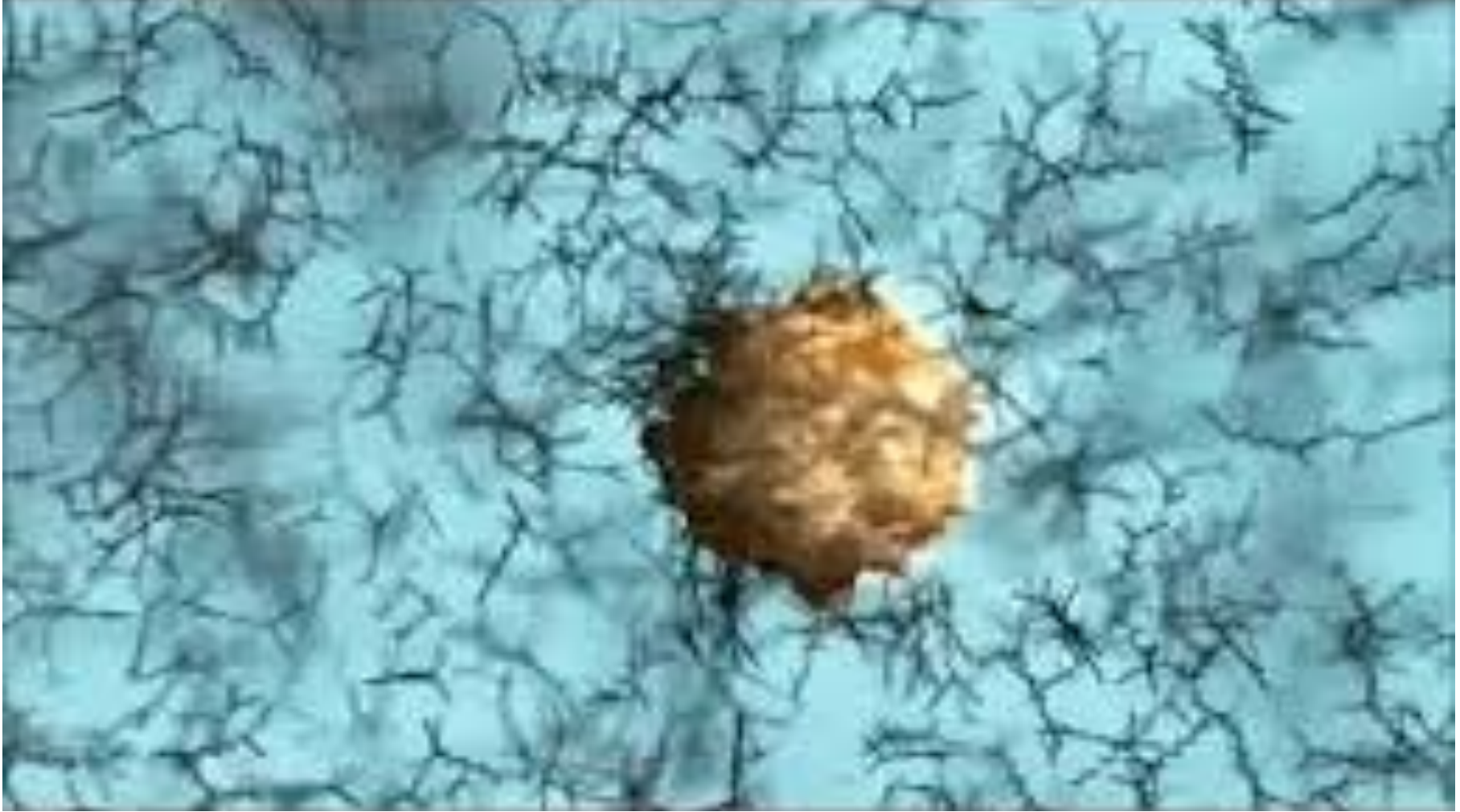
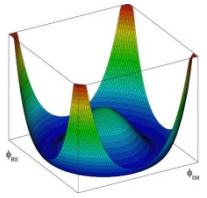
O campo de Higgs

Higgs, Englert, Brout - 1964

Peter Higgs

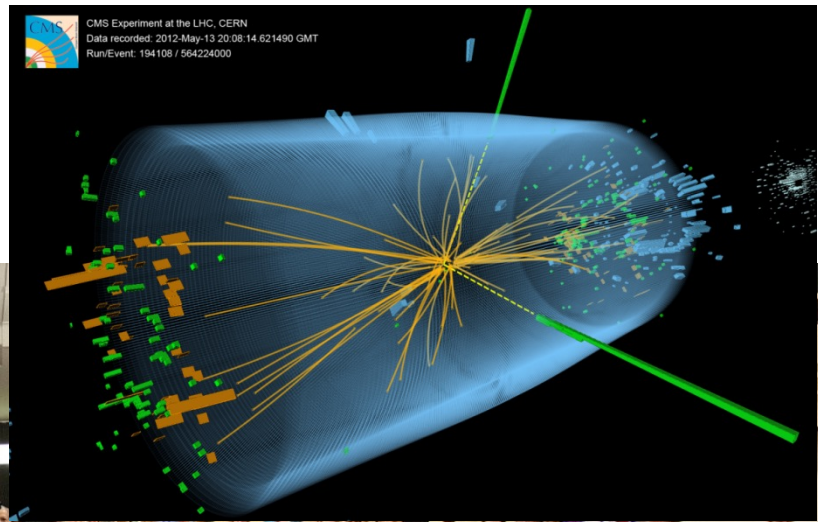


O mecanismo de Higgs



A massa é gerada na interacção com o campo

O bóson de Higgs



4 Julho 2012: CERN

As partículas elementares hoje!

Three Generations
of Matter (Fermions) spin $\frac{1}{2}$

| | I | II | III |
|----------|--|--|--|
| mass → | 2.4 MeV | 1.27 GeV | 173.2 GeV |
| charge → | $\frac{2}{3}$ | $\frac{2}{3}$ | $\frac{2}{3}$ |
| name → | u up | c charm | t top |
| Quarks | d down | s strange | b bottom |
| | ν_e electron neutrino | ν_μ muon neutrino | ν_τ tau neutrino |
| Leptons | e electron | μ muon | τ tau |
| | | | |

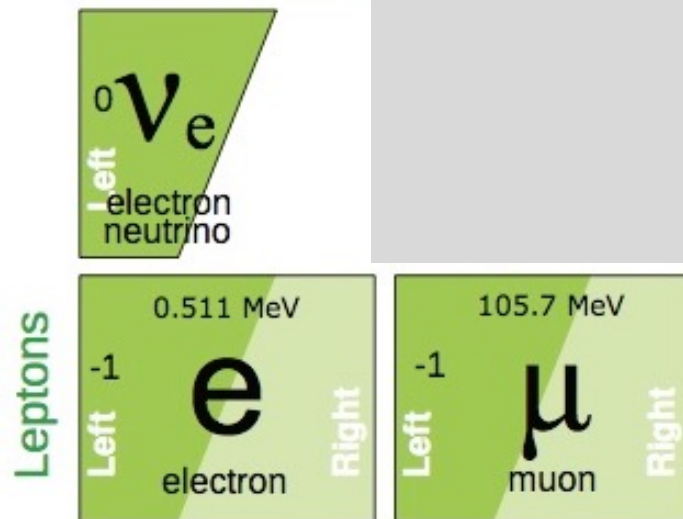
| | |
|------------------------|---|
| Bosons (Forces) spin 1 | g gluon |
| | γ photon |
| | Z^0 weak force |
| | W^\pm weak force |

| |
|-------------------------|
| H Higgs boson |
|-------------------------|

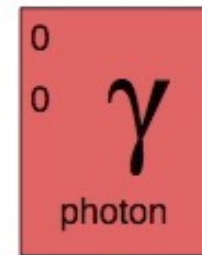
spin 0

As partículas elementares em 1957

Three Generations
of Matter (Fermions) spin $\frac{1}{2}$



Bosons (Forces) spin 1



As partículas elementares em 1980

Three Generations of Matter (Fermions) spin $\frac{1}{2}$

| | I | II | III |
|----------|--|--|--|
| mass → | 2.4 MeV | 1.27 GeV | |
| charge → | $\frac{2}{3}$ | $\frac{2}{3}$ | |
| name → | u up | c charm | |
| Quarks | Left Right | Left Right | |
| | 4.8 MeV $-\frac{1}{3}$ d down | 104 MeV $-\frac{1}{3}$ s strange | 4.2 GeV $-\frac{1}{3}$ b bottom |
| Leptons | Left Right | Left Right | |
| | 0 ν_e electron neutrino | 0 ν_μ muon neutrino | |
| | 0.511 MeV -1 e electron | 105.7 MeV -1 μ muon | 1.777 GeV -1 τ tau |

Bosons (Forces) spin 1

0
0
g
gluon

0
0
 γ
photon

As partículas elementares hoje!

Three Generations
of Matter (Fermions) spin $\frac{1}{2}$

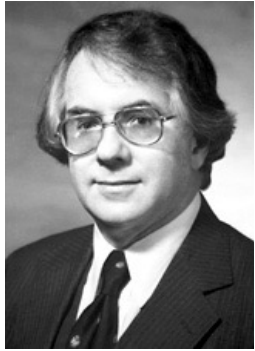
| | I | II | III |
|----------|--|--|--|
| mass → | 2.4 MeV | 1.27 GeV | 173.2 GeV |
| charge → | $\frac{2}{3}$ | $\frac{2}{3}$ | $\frac{2}{3}$ |
| name → | u up | c charm | t top |
| Quarks | Left Right | Left Right | Left Right |
| | d down | s strange | b bottom |
| | Left Right | Left Right | Left Right |
| | ν_e electron neutrino | ν_μ muon neutrino | ν_τ tau neutrino |
| Leptons | Left Right | Left Right | Left Right |
| | e electron | μ muon | τ tau |
| | Left Right | Left Right | Left Right |
| | 0.511 MeV | 105.7 MeV | 1.777 GeV |

| | |
|------------------------|---|
| Bosons (Forces) spin 1 | 0 0 g gluon |
| | 0 0 γ photon |
| | 91.2 GeV 0 Z⁰ weak force |
| | 80.4 GeV ± 1 W[±] weak force |

| |
|--|
| 126 GeV 0 0 H Higgs boson |
| spin 0 |

O Modelo Padrão: $SU(2)_L \otimes U_1$

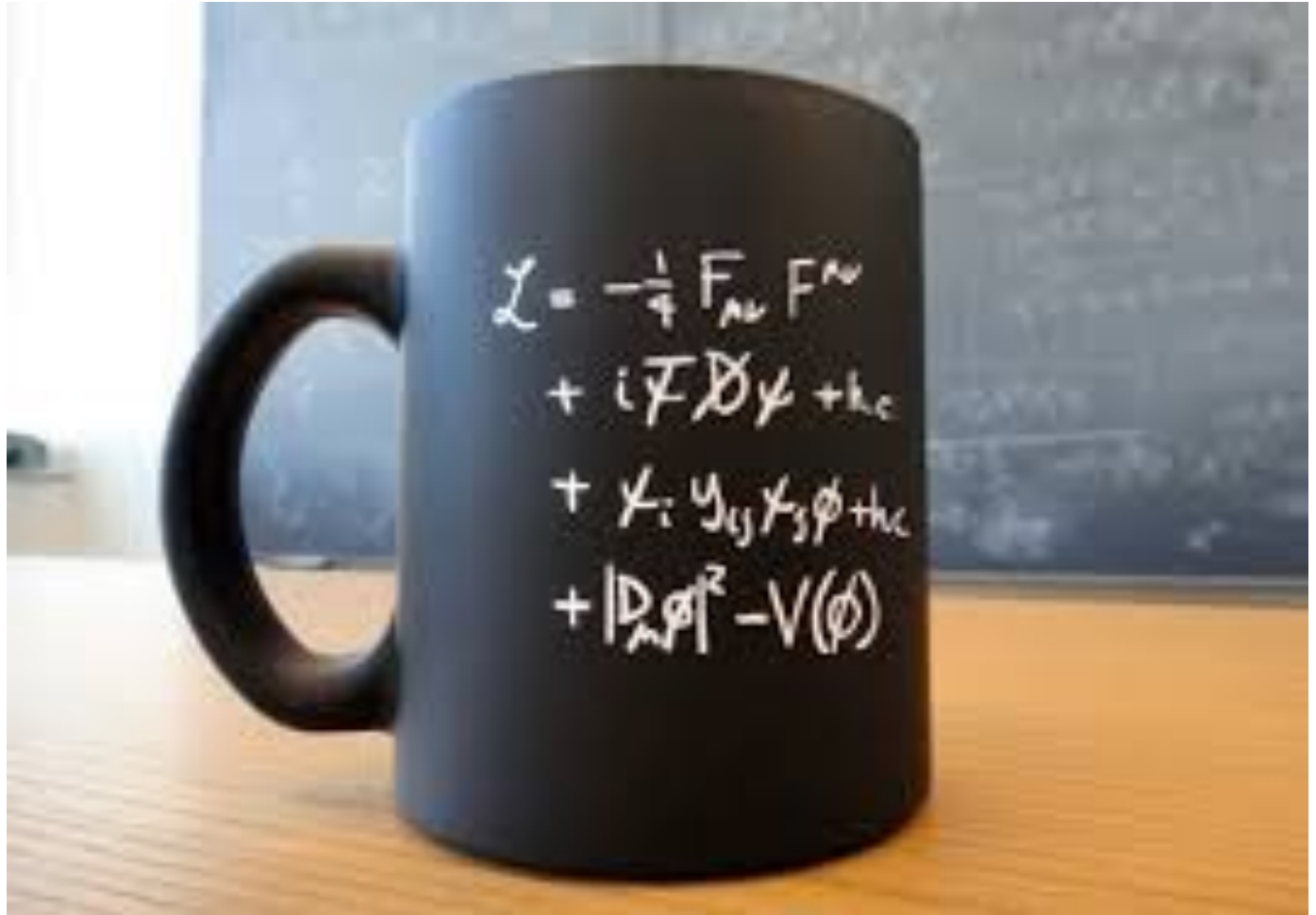
Glashow



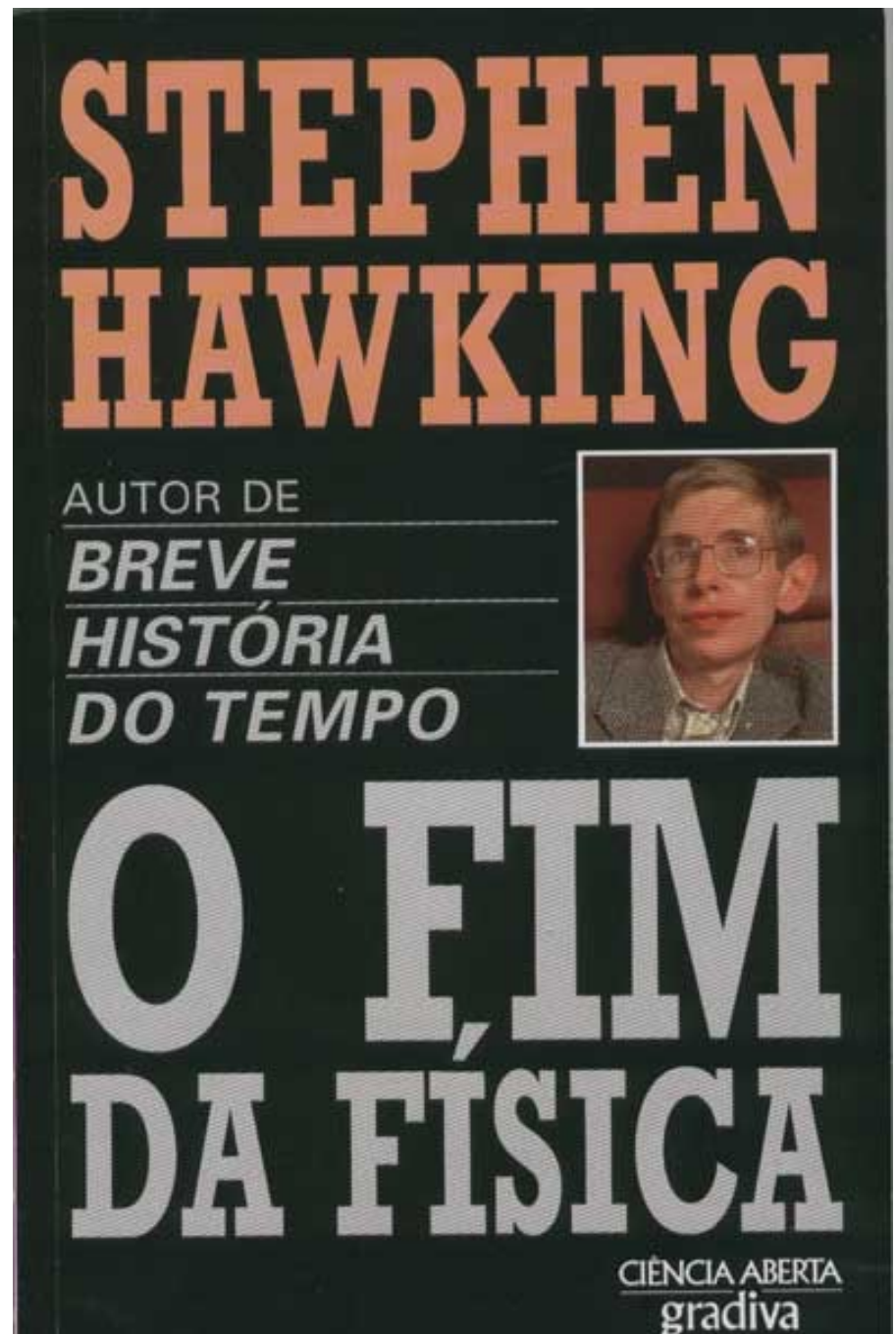
Salam



Weinberg



O fim da Física ?



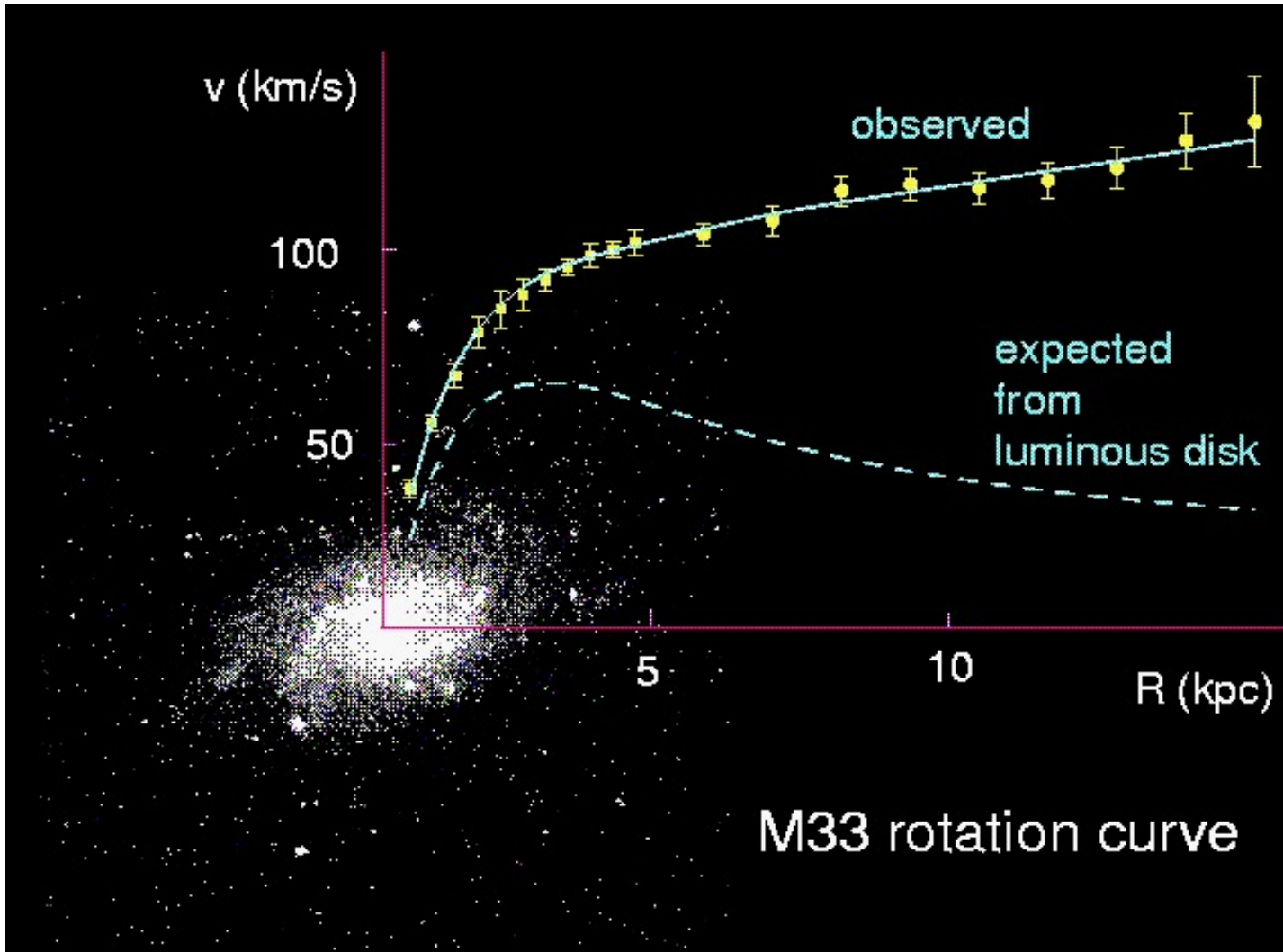
A matéria luminosa



A velocidade de rotação das estrelas periféricas

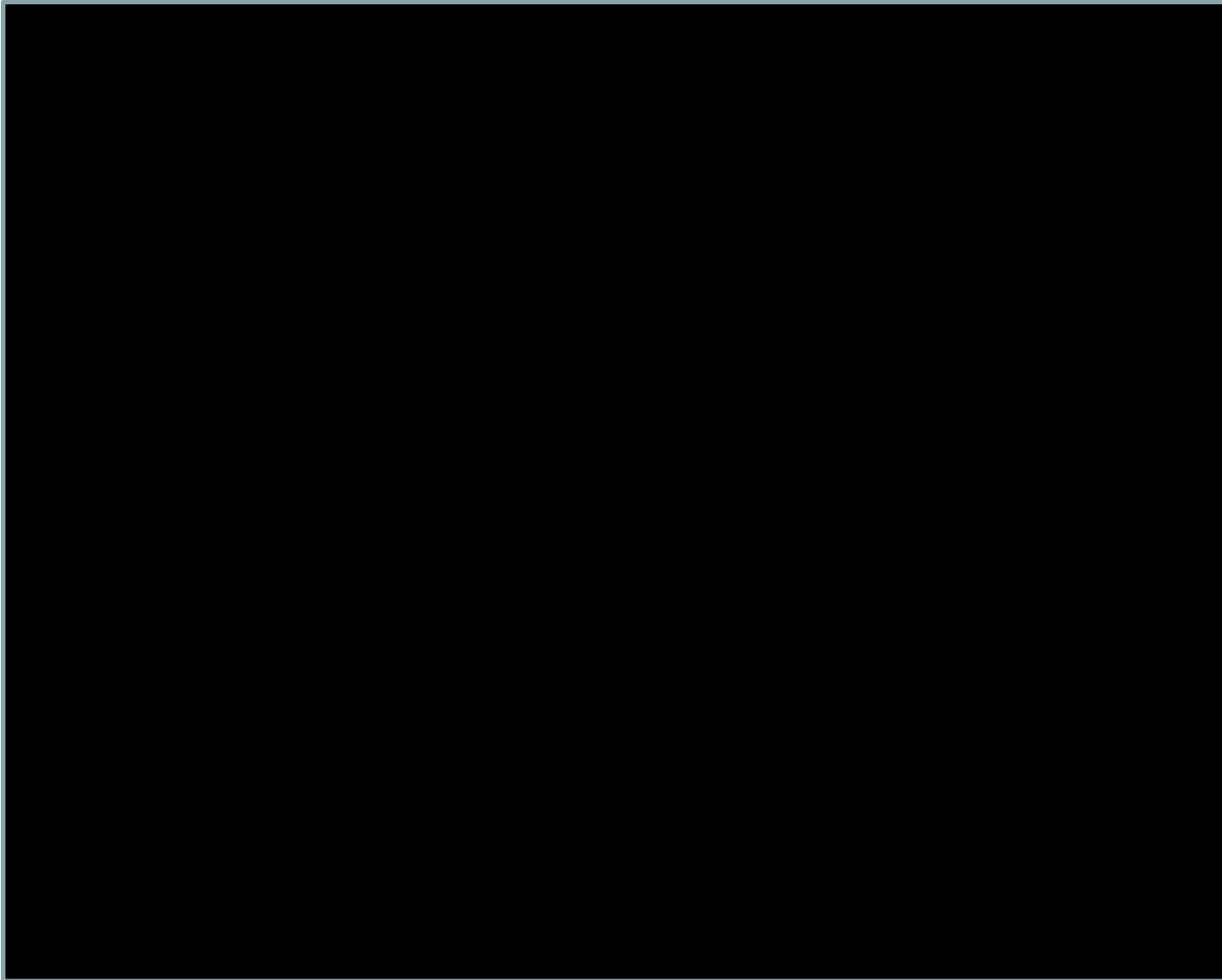


Rodar em torno do centro da galáxia

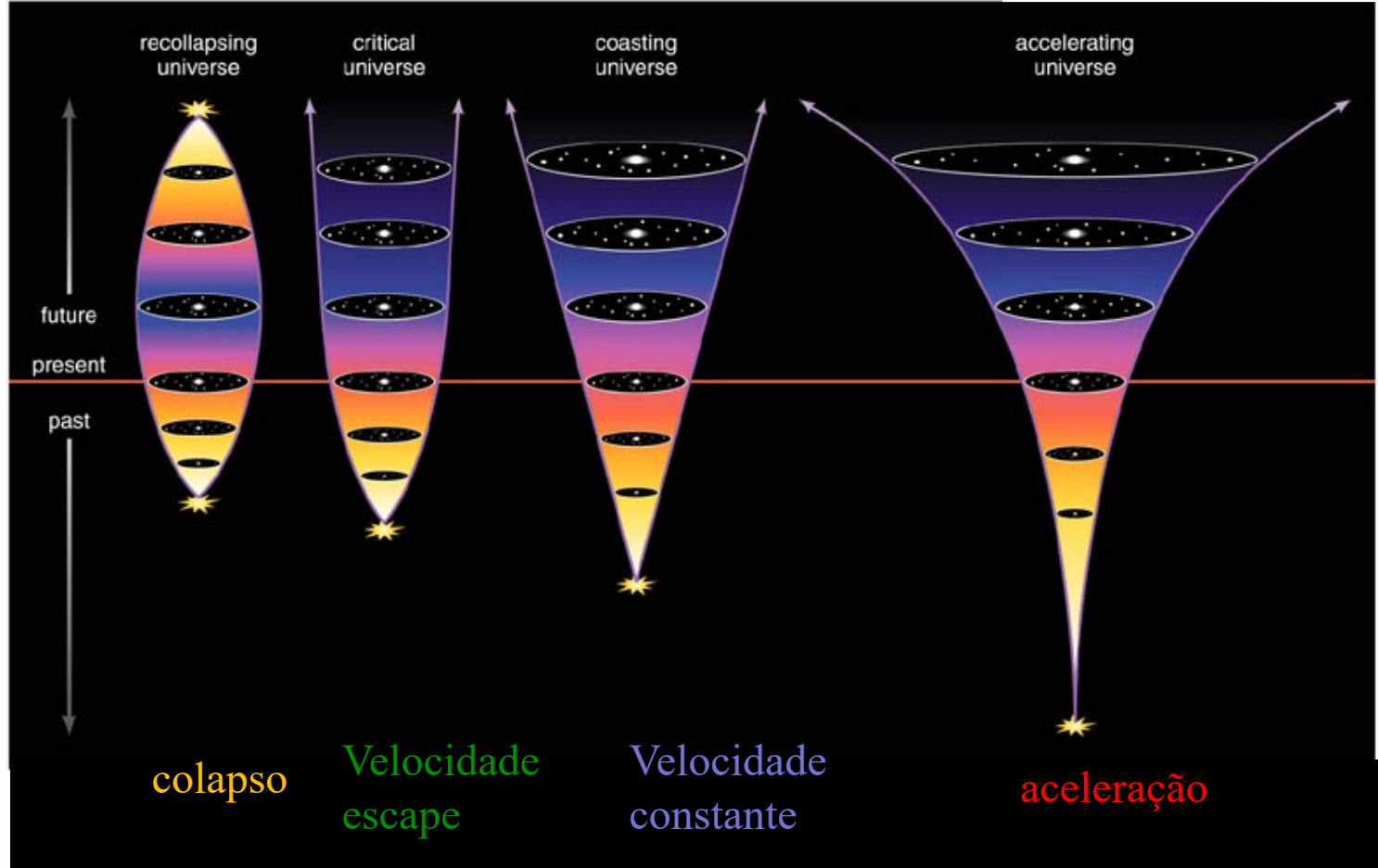
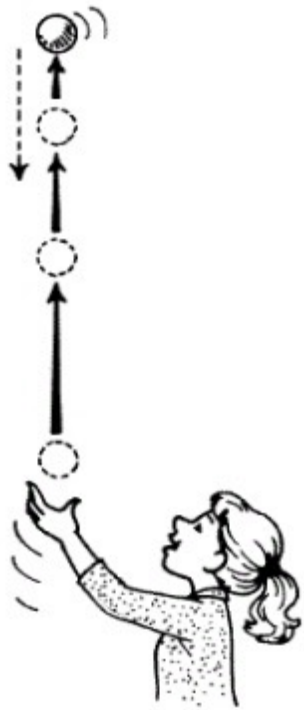
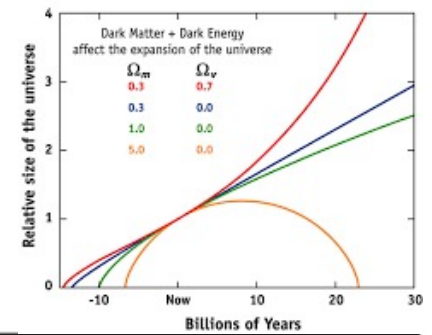


Um halo de matéria escura!

A matéria escura

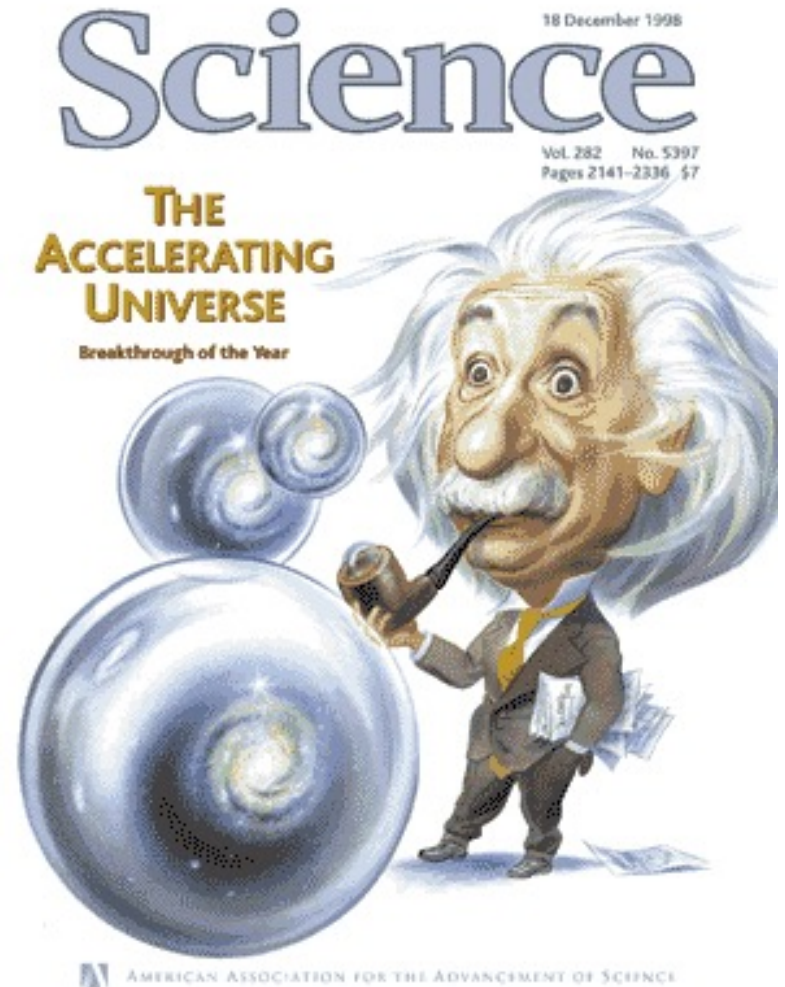
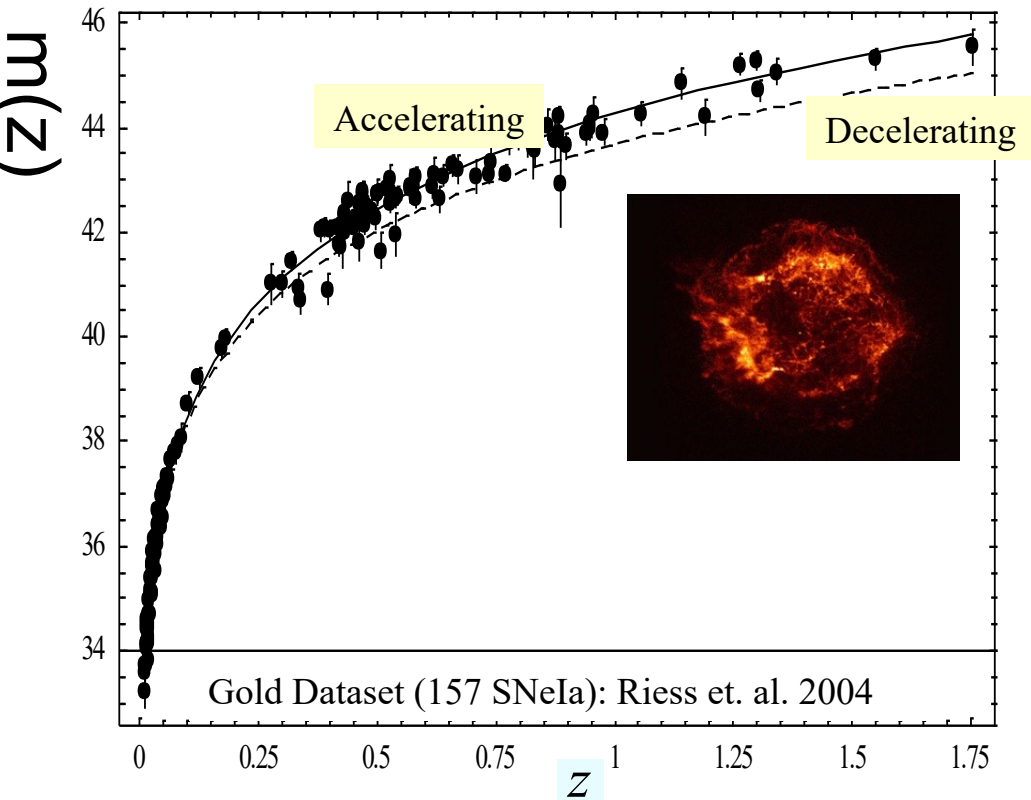


A expansão do Universo



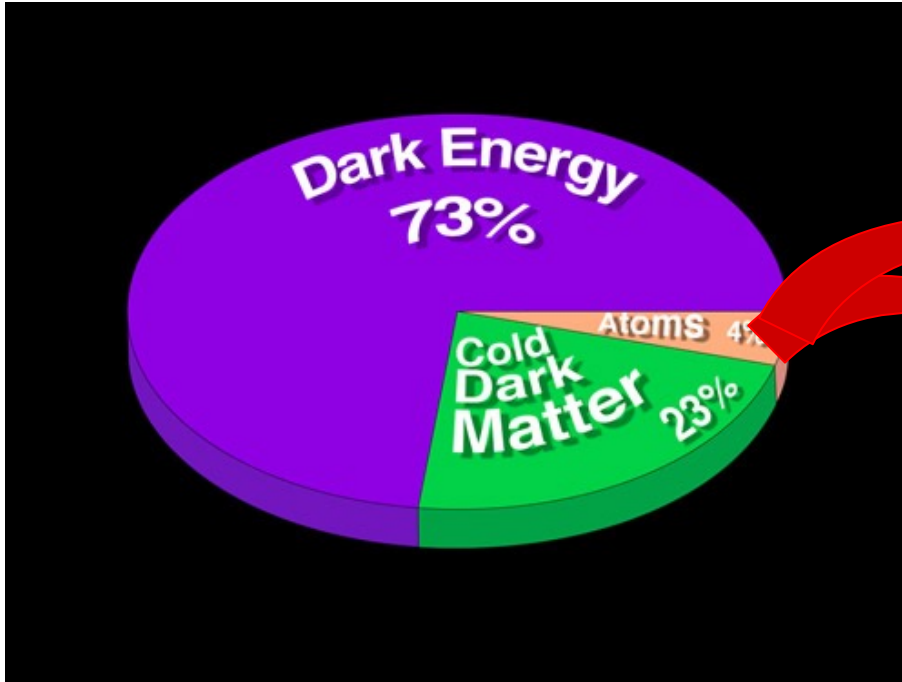
A energia escura !

Diagrama de Hubble



O Universo encontra-se numa expansão acelerada !!!
Energia escura – anti-gravítica

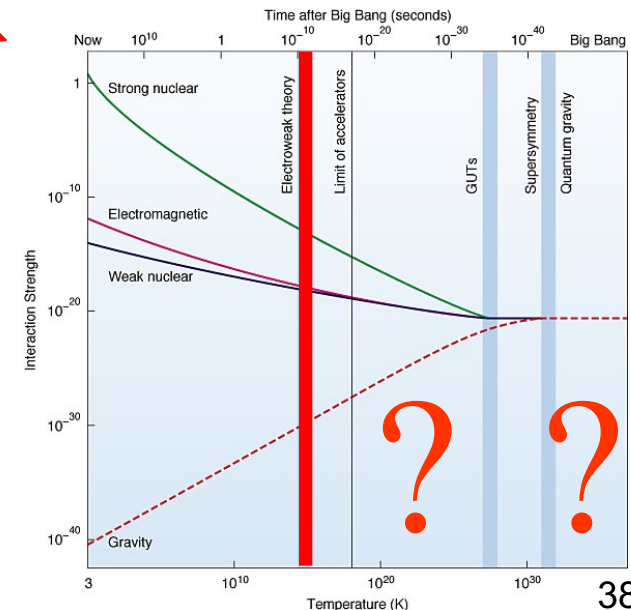
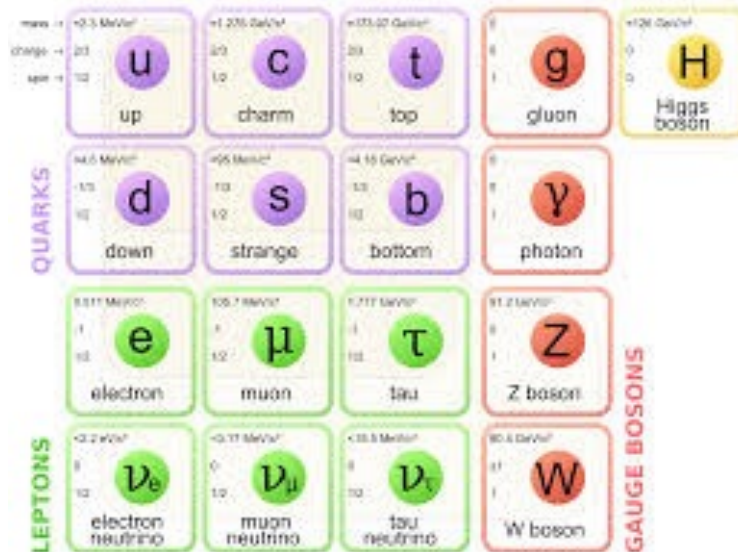
Fronteiras do nosso (des)conhecimento



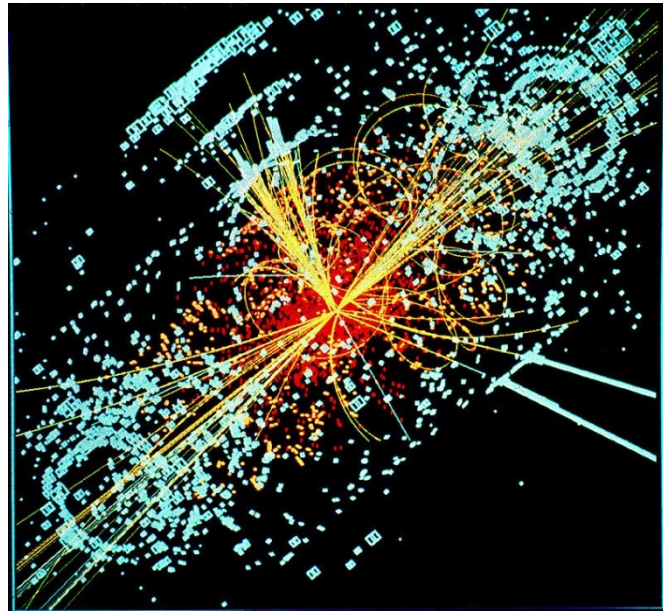
O modelo
padrão

$E_{CM} \sim 1 \text{ TeV}$

$R \sim 10^{-18} \text{ m}$

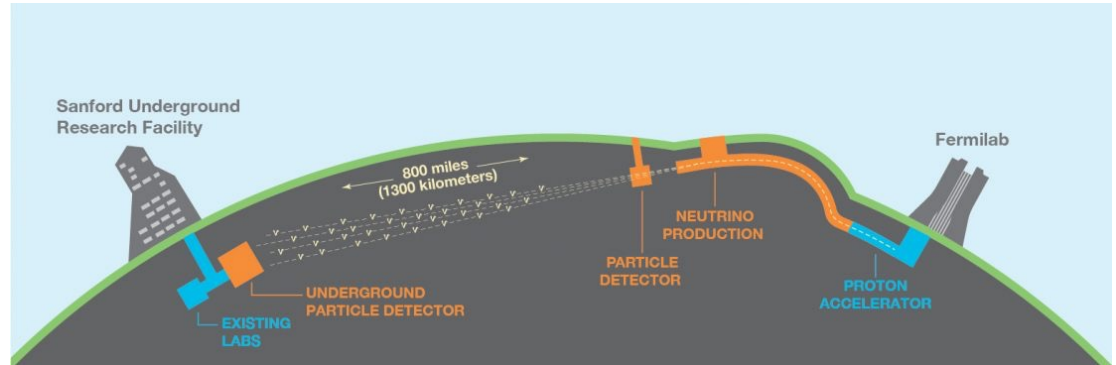


Olhar o Universo no sec. XXI

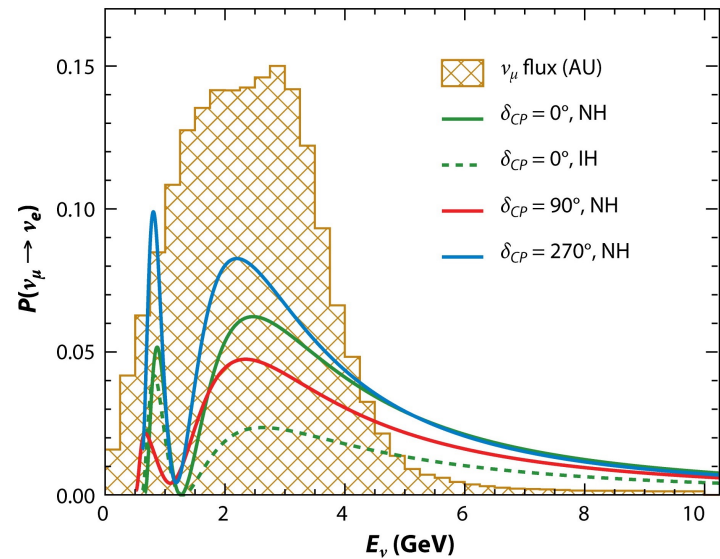


DUNE

violação de CP no
sector dos neutrinos ???



Padrões de Oscilação



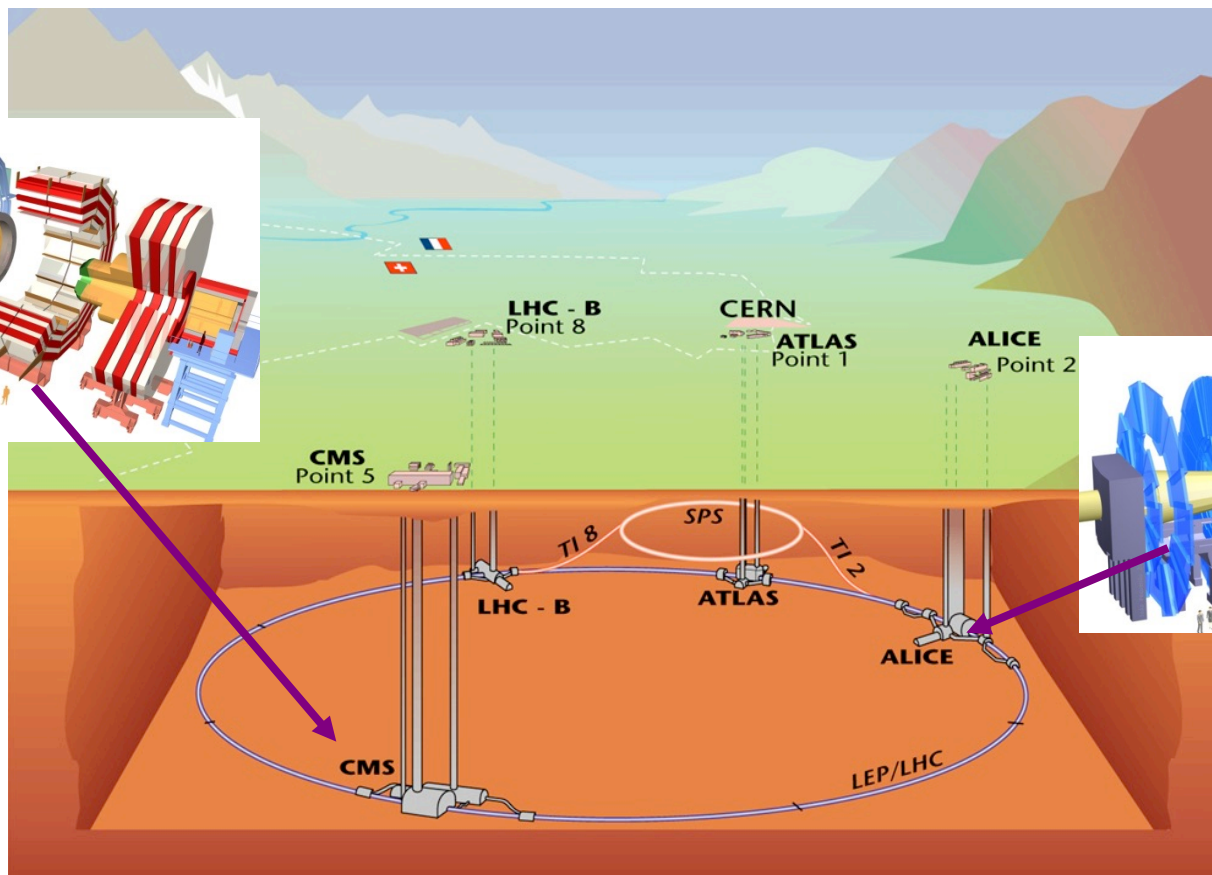
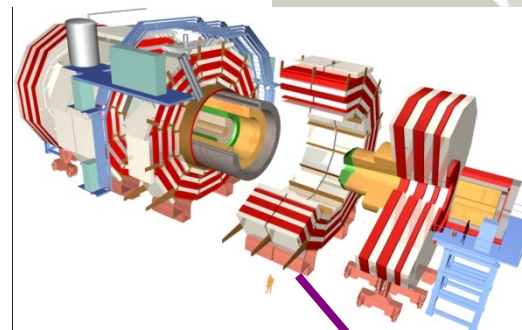
Proto-DUNE @ CERN



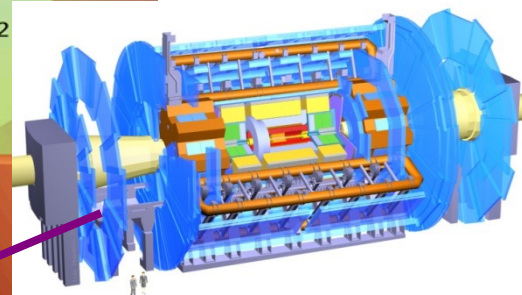
O LHC no CERN



CMS



ATLAS

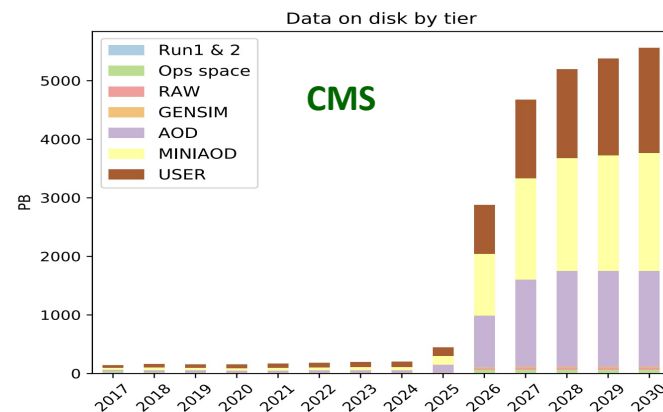
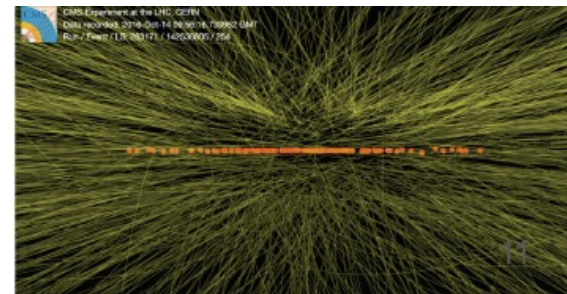


Os próximos 50 anos no CERN!

Até 2038 - LHC – alta intensidade

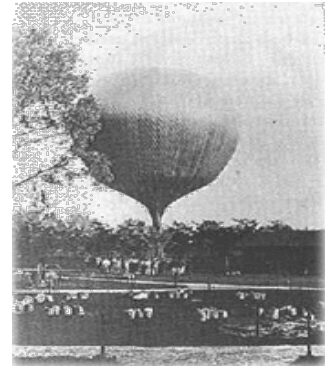
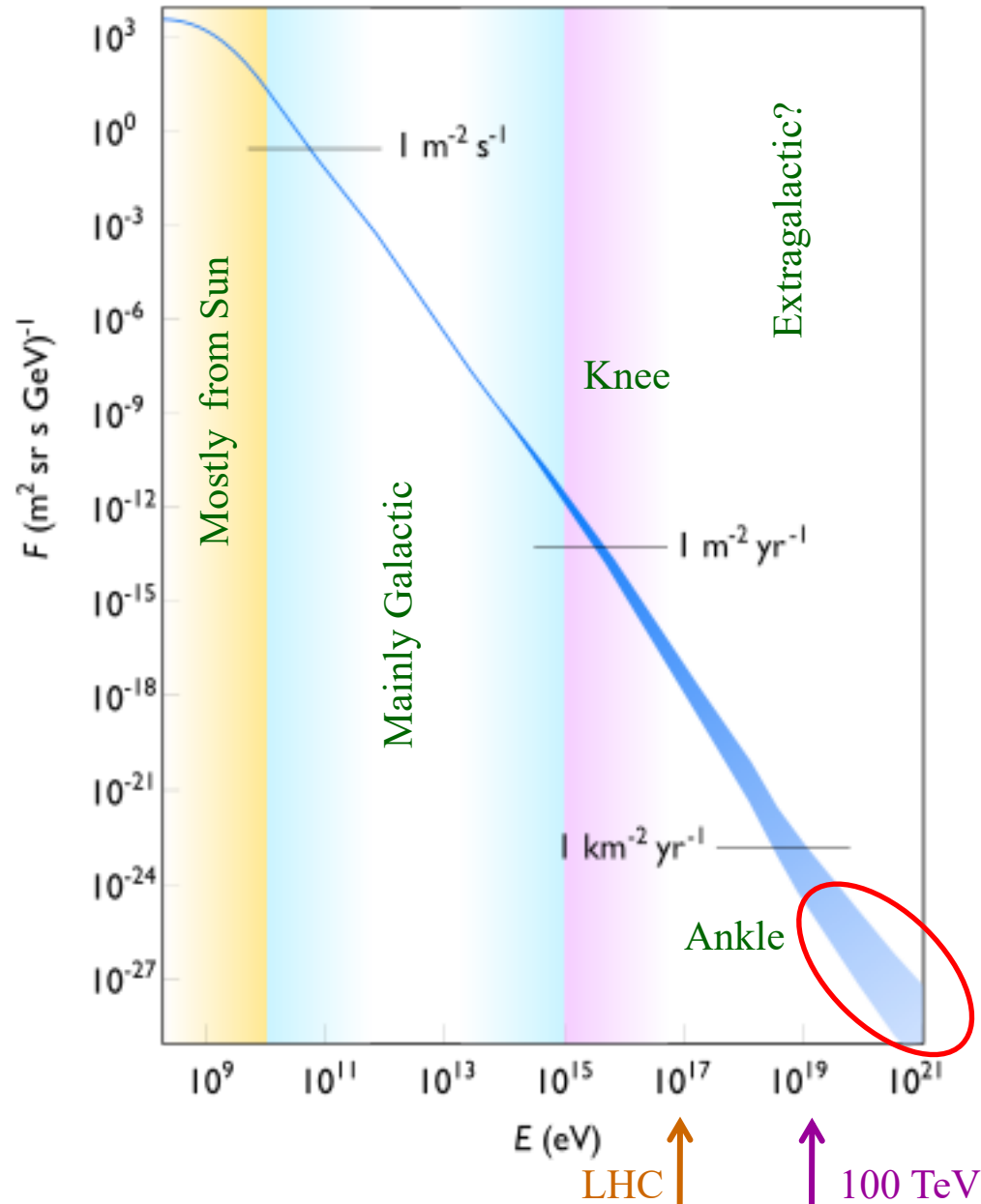
Inicia-se agora o caminho para novos aceleradores (em estudo o FCC)

E muito R&D e desafios tecnológicos a vencer !



Raios cósmicos carregados

Viktor Hess, 1912



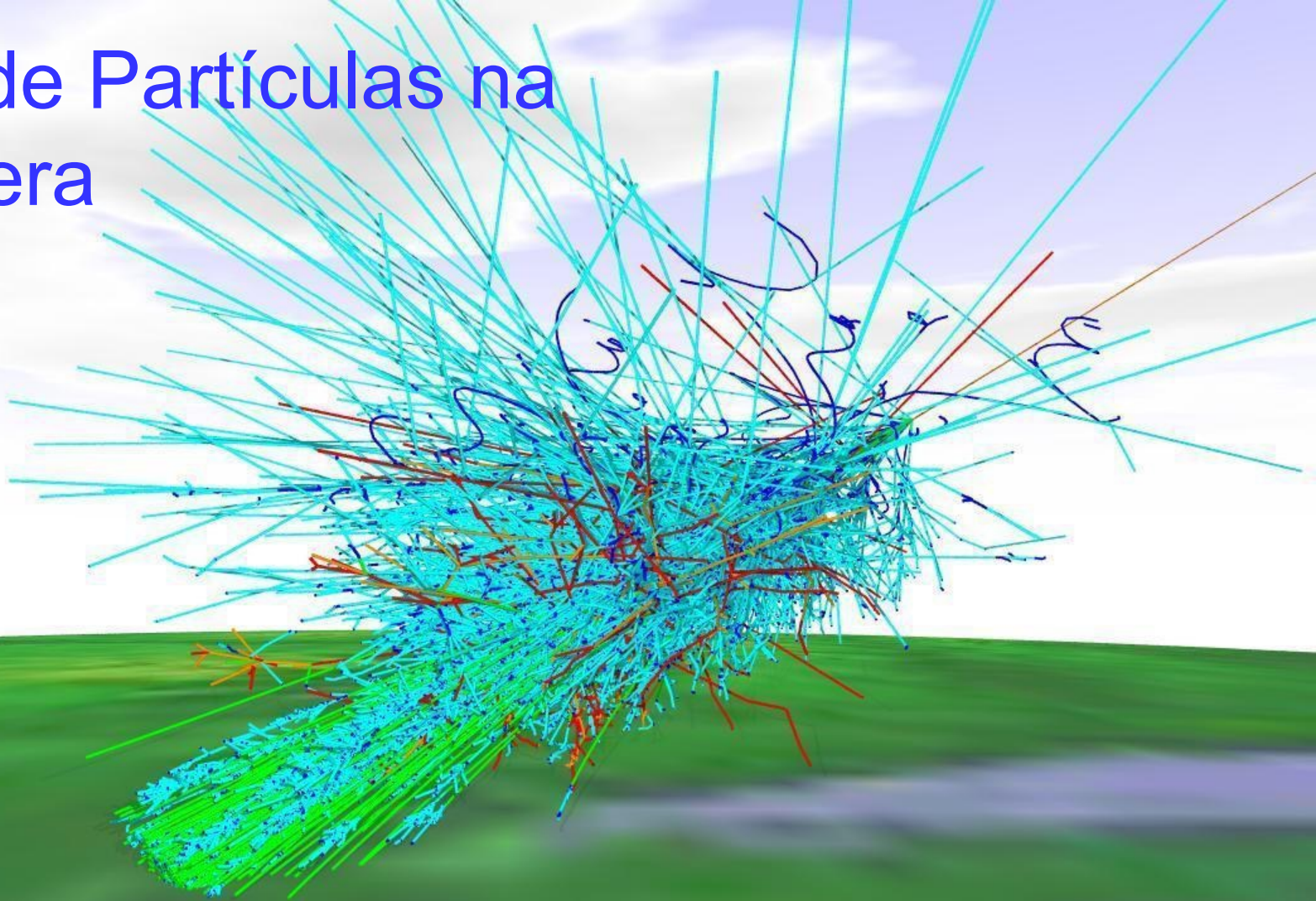
$$\frac{dN}{dE} \propto E^{-\alpha}$$

$$\alpha = \begin{cases} 2.7 & E < 10^{16} \\ 3.0 & 10^{16} < E < 10^{18} \\ 2.7? & E > 10^{18} \end{cases}$$

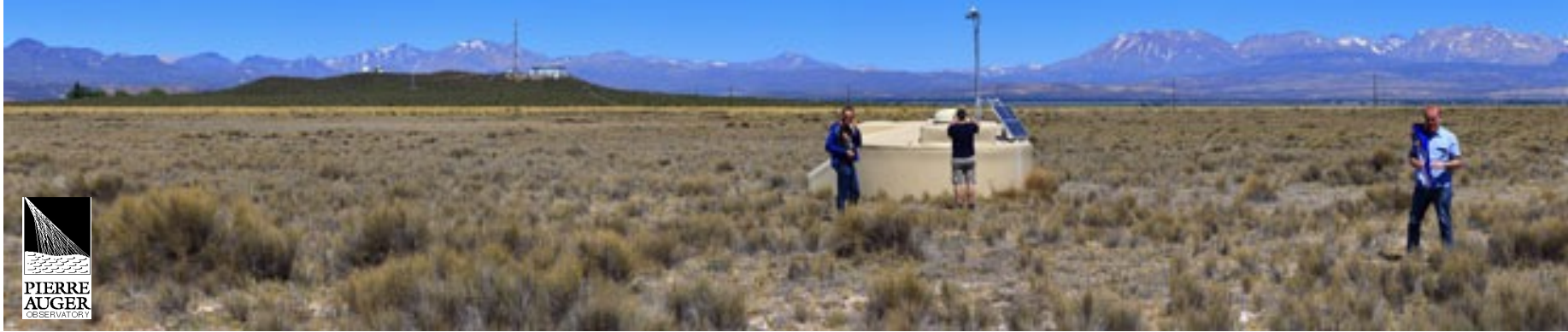
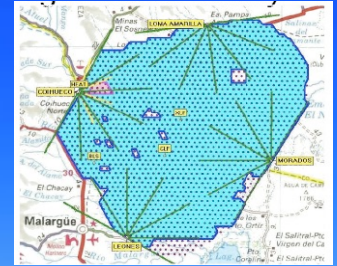
Um LHC do tamanho da órbita de Mercúrio !!!



Física de Partículas na atmosfera

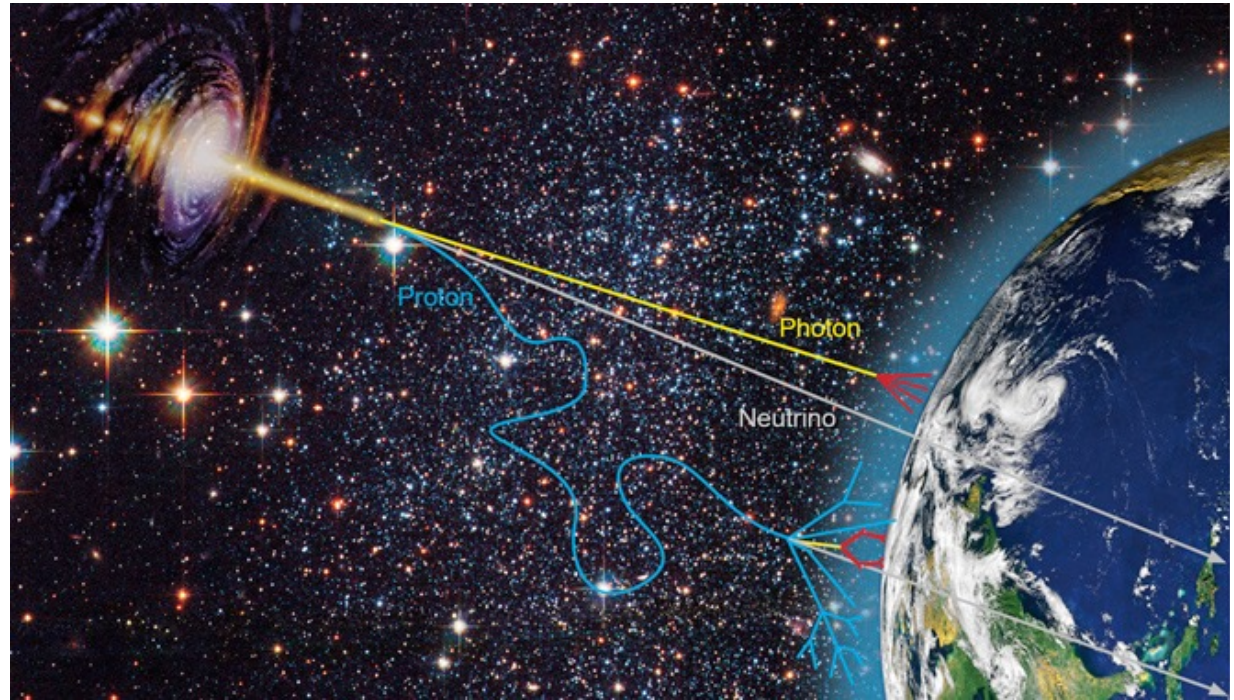


Na pampa argentina

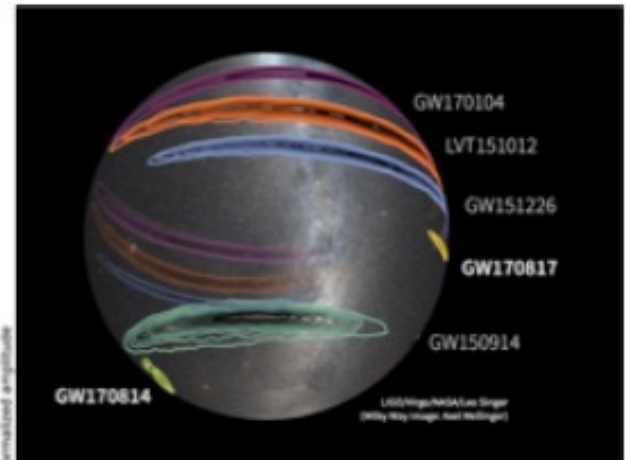
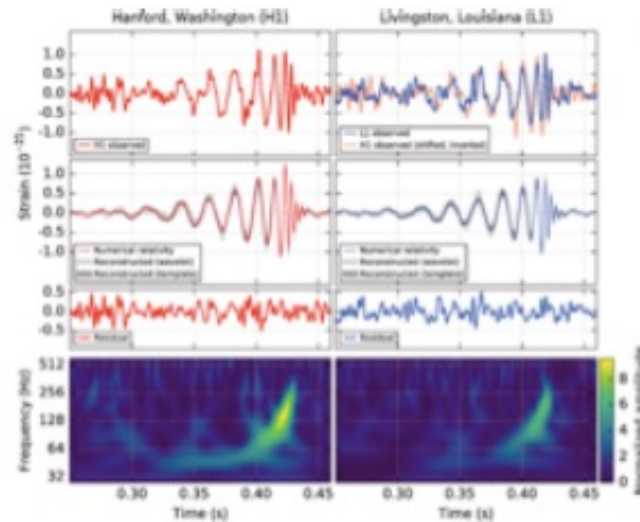


Mensageiros do Universo

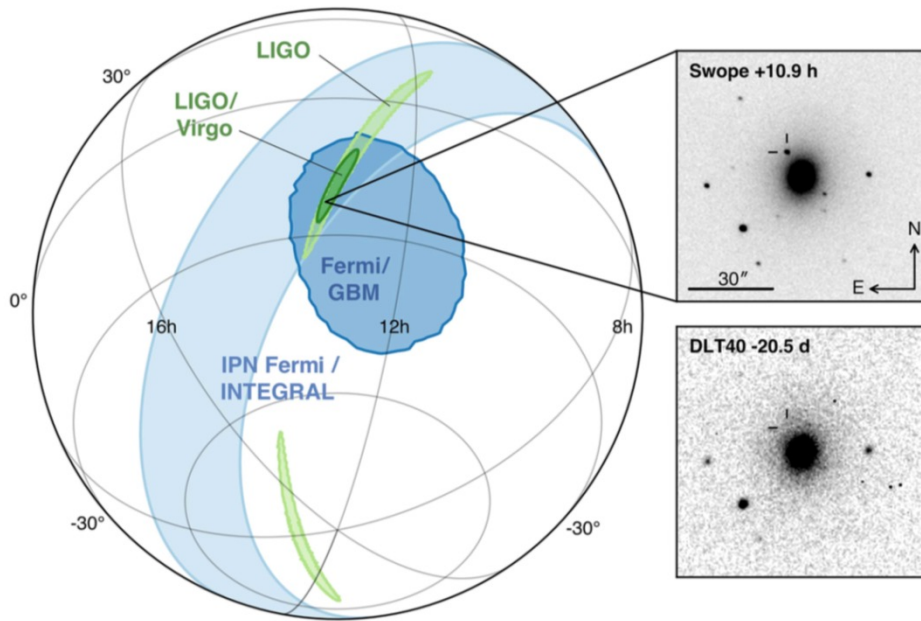
γ , p, nuclei, ν



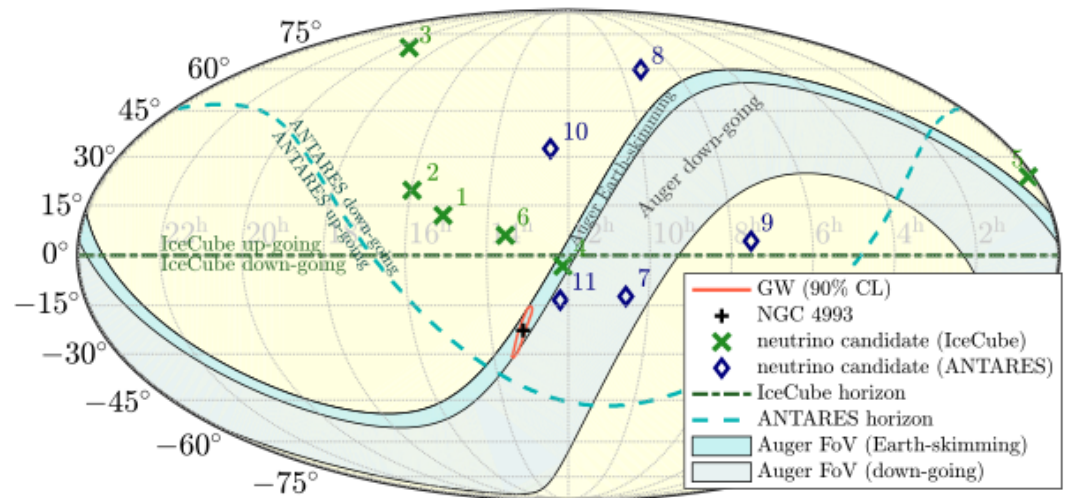
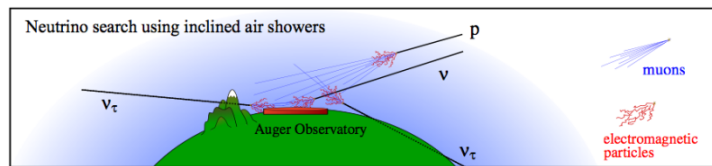
GW



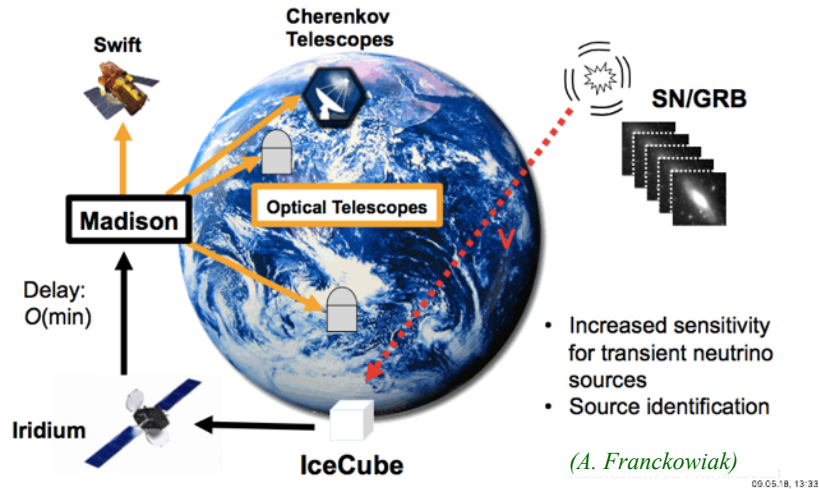
GW170817



A primeira observação
“multimessenger” da
fusão de duas estrelas
de neutrões



A primeira fonte de neutrinos astrofísicos de alta energia



TITLE: GCN CIRCULAR
 NUMBER: 21916
 SUBJECT: IceCube-170922A - IceCube observation of a high-energy neutrino candidate event
 DATE: 17/09/23 01:09:26 GMT
 FROM: Erik Blaufuss at U. Maryland/IceCube <blaufuss@icecube.umd.edu>

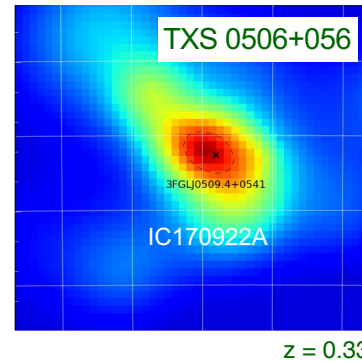
Claudio Kopfer (University of Alberta) and Erik Blaufuss (University of Maryland) report on behalf of the IceCube Collaboration (<http://icecube.wisc.edu/>).

On 22 Sep, 2017 IceCube detected a track-like, very-high-energy event with a high probability of being of astrophysical origin. The event was identified by the Extremely High Energy (EHE) track event selection. The IceCube detector was in a normal operating state. EHE events typically have a neutrino interaction vertex that is outside the detector, produce a muon that traverses the detector volume, and have a high light level (a proxy for energy).

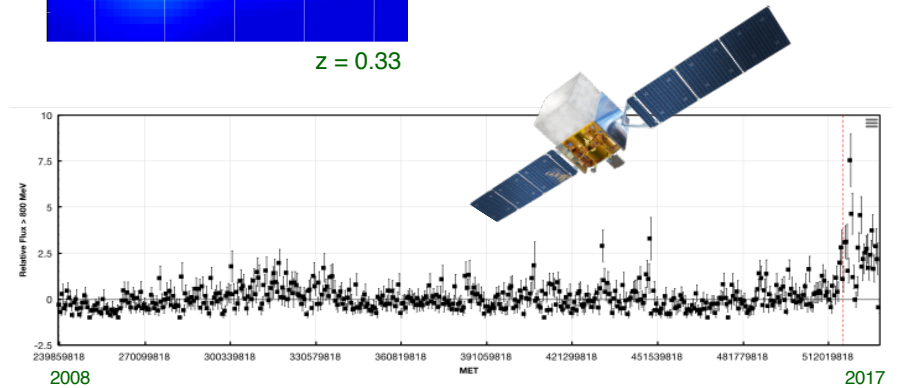
After the initial automated alert (https://gcn.gsfc.nasa.gov/notices_amon/50579430_130033.amon), more sophisticated reconstruction algorithms have been applied offline, with the direction refined to:

Date: 22 Sep, 2017
 Time: 20:54:30.43 UTC
 RA: 77.43 deg (+0.80 deg/+1.30 deg 90% PSF containment) J2000
 Dec: 5.72 deg (-0.40 deg/+0.70 deg 90% PSF containment) J2000

We encourage follow-up by ground and space-based instruments to help identify a possible astrophysical source for the candidate neutrino.



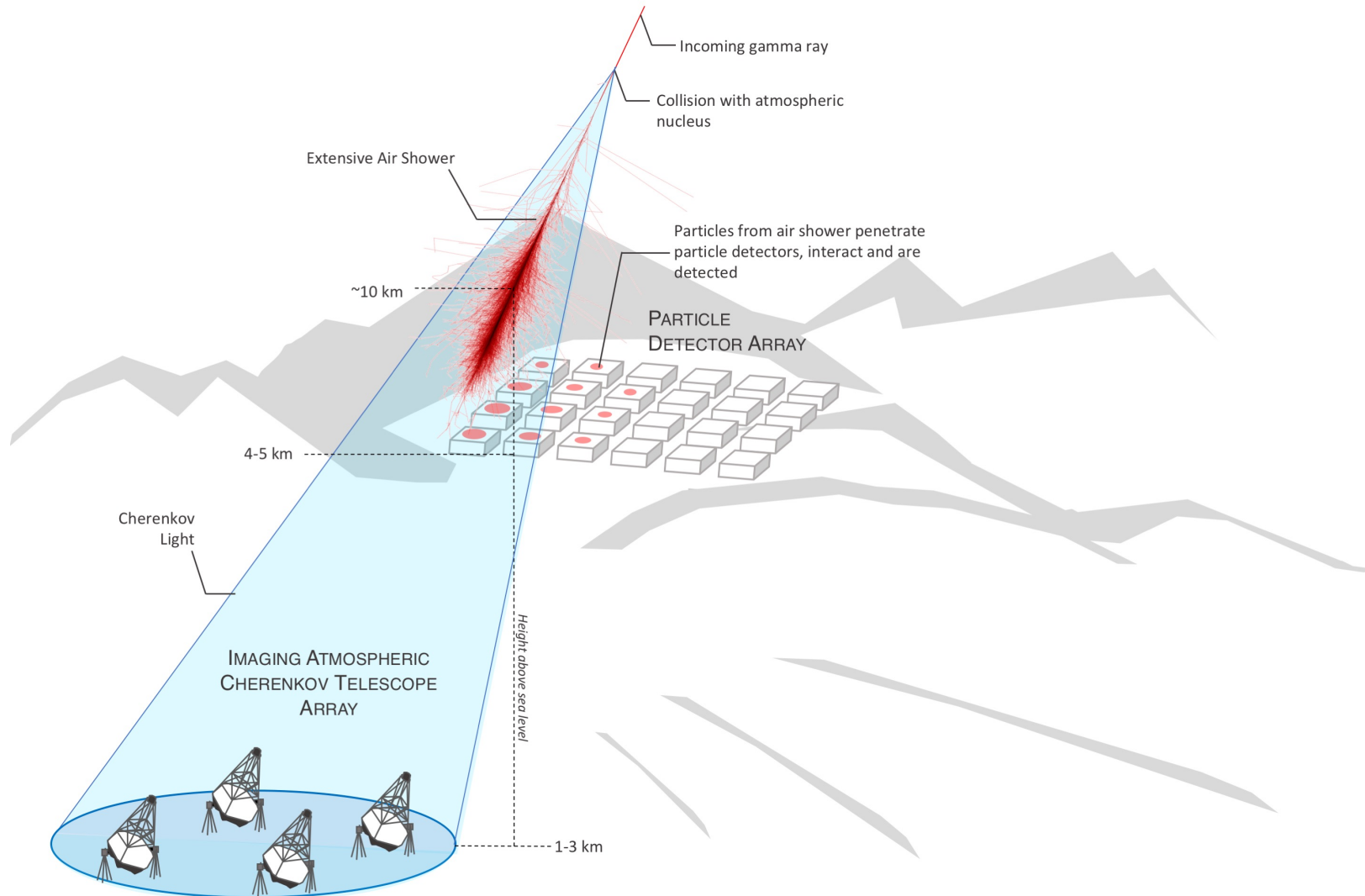
MAGIC: flare E > 100 GeV



FERMI: flare (found 6 days later)

IceCube 170922A, publications in preparation

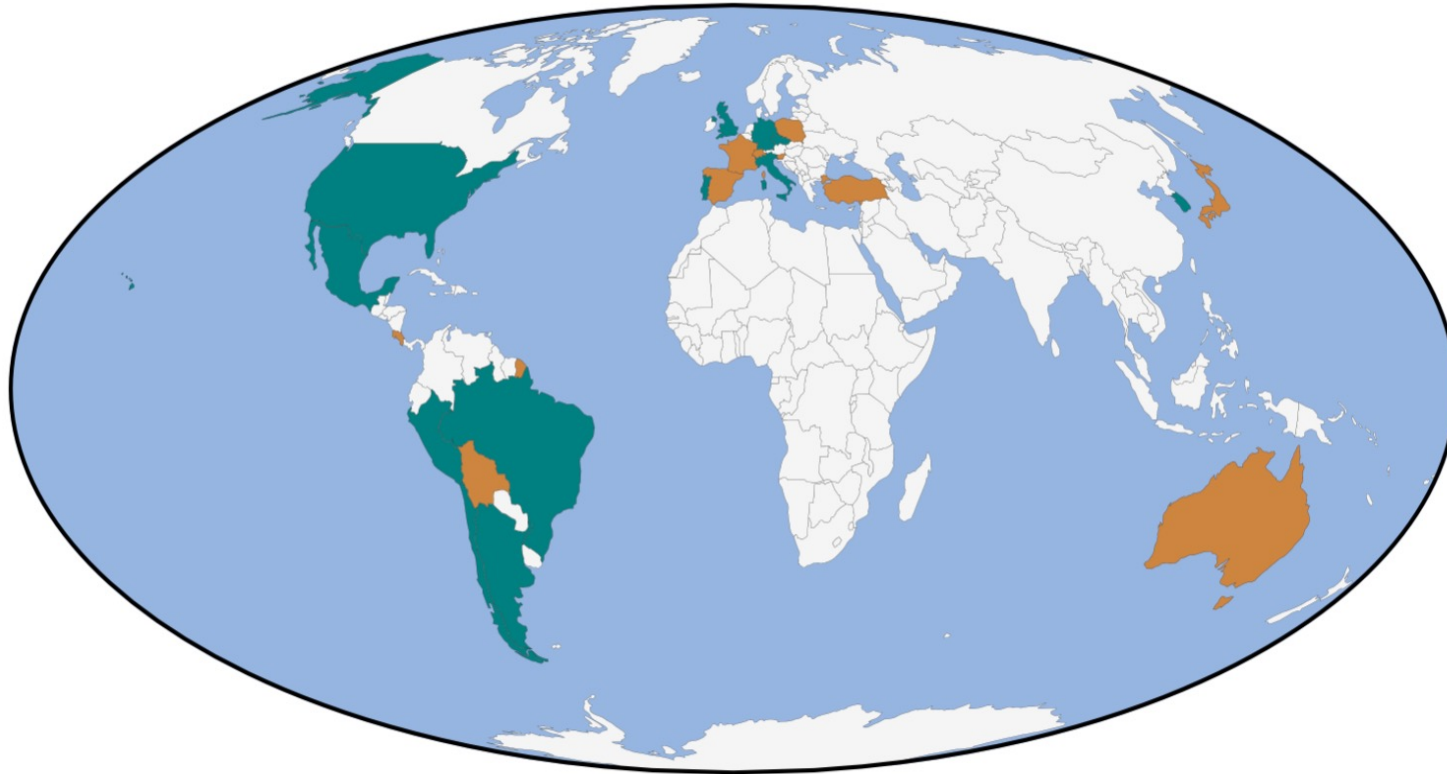
Detector fotões no alto das montanhas ..



Not to scale

SWGO

Southern Wide-field Gamma-ray Observatory



Countries in SWGO

Institutes

Argentina*, Brazil, Chile, Czech Republic, Germany*, Italy, Mexico, Peru, Portugal, South Korea, United Kingdom, United States*

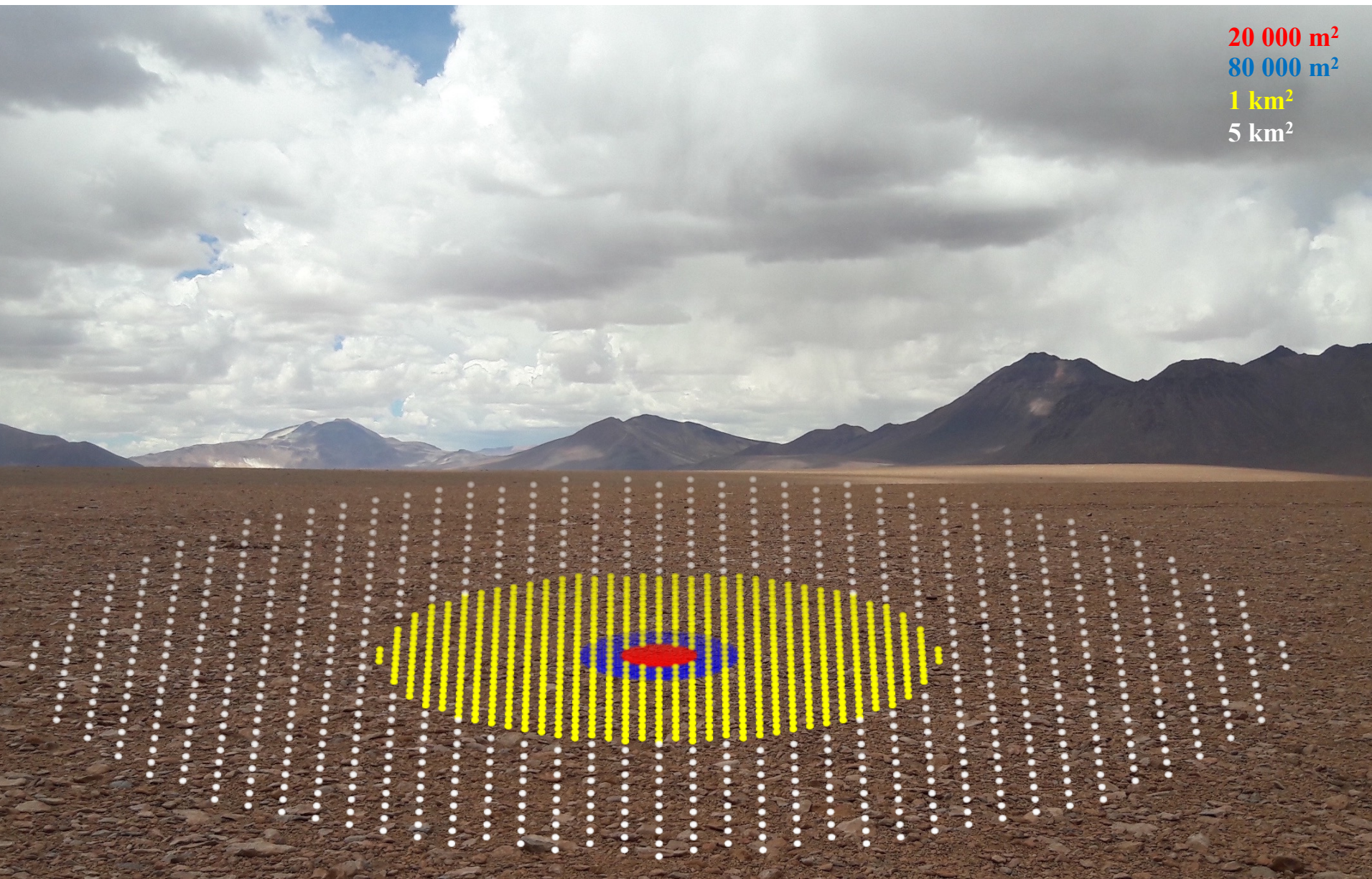
Supporting scientists

Australia, Bolivia, Costa Rica, France, Japan, Poland, Slovenia, Spain, Switzerland, Turkey

**also supporting scientists*

100 + 32 cientistas, 53 institutos, 13 países

Olhar para o centro da galáxia e para o Universo a 4500-5000 m de altitude na América do Sul com um detector de fótons energéticos



Da Terra aos Céus ...



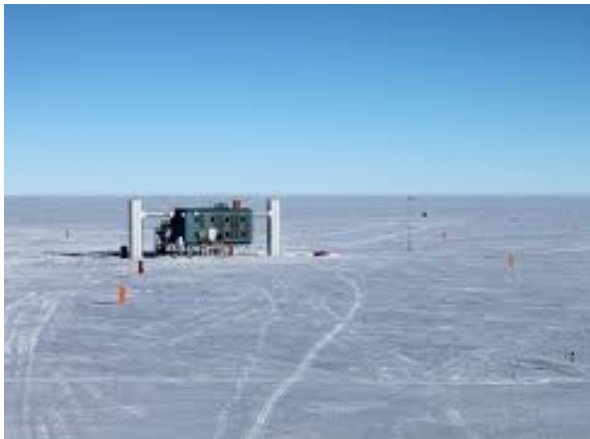
AMS



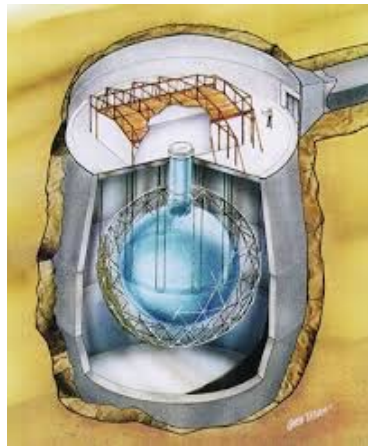
Fermi



LIGO



Ice Cube



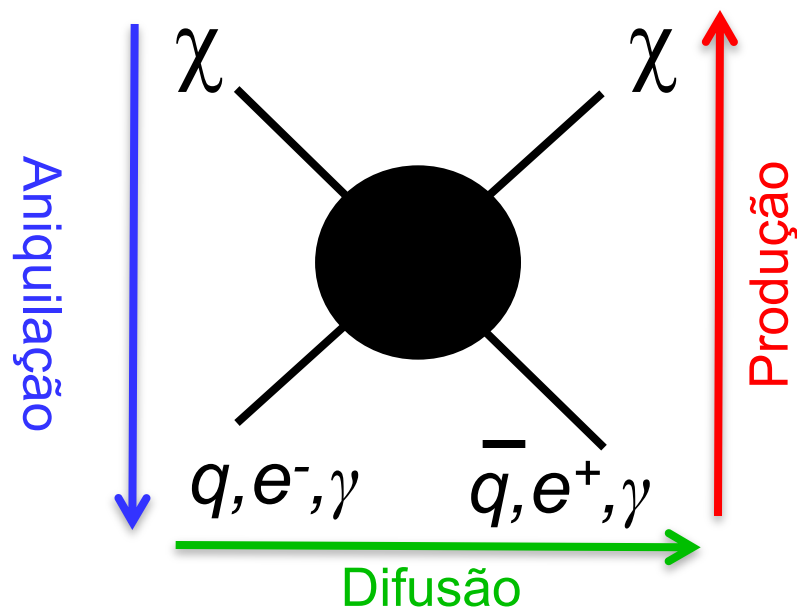
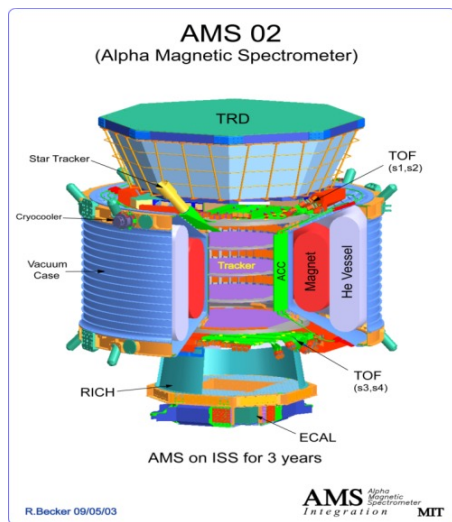
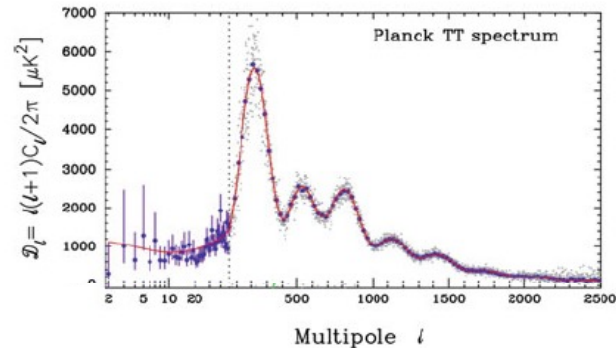
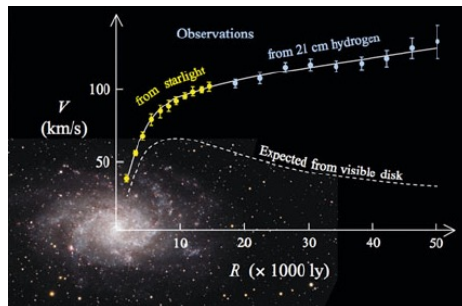
SNO⁺



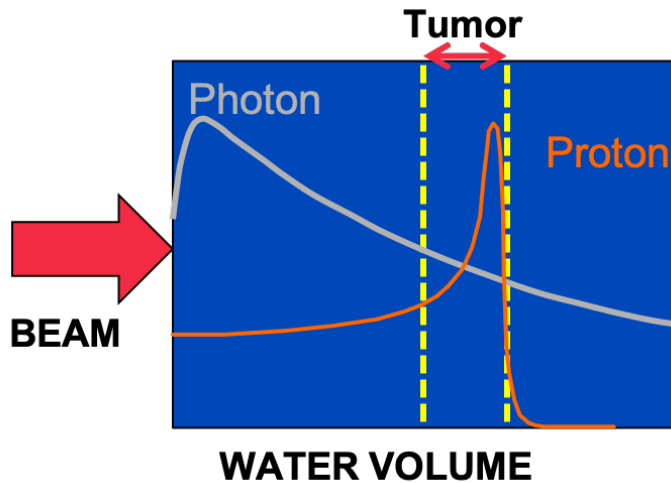
LST - CTA

E muitos mais

Matéria Escura



Terapia com Protões

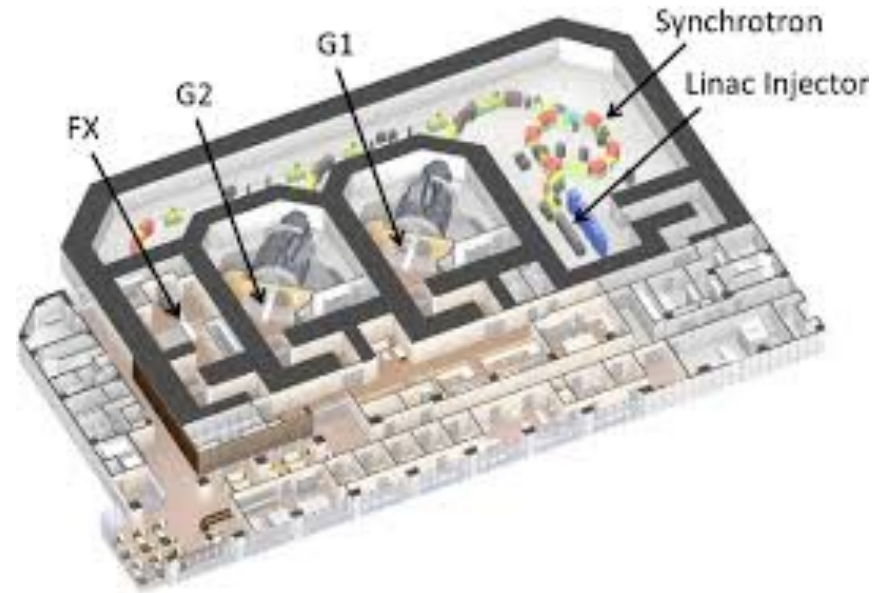


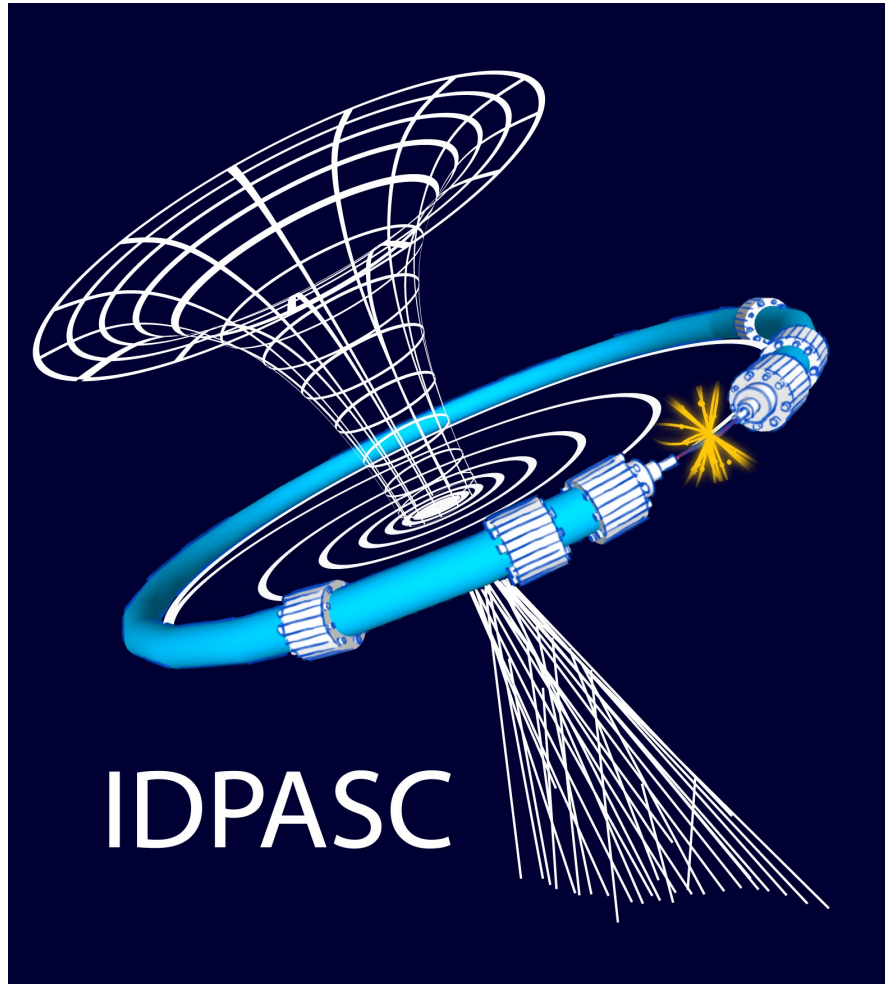
Grupo Hospitalar IPO, IST, LIP, ICNAS (UC)

Associação Portuguesa de Proto-Terapia e Tecnologias Avançadas para a Prevenção e Tratamento do Cancro (ProtoTera)

Loures (CTN): ciclotrão de 250 MeV com duas salas de tratamento e uma sala dedicada a desenvolvimento tecnológico, industrial e científico

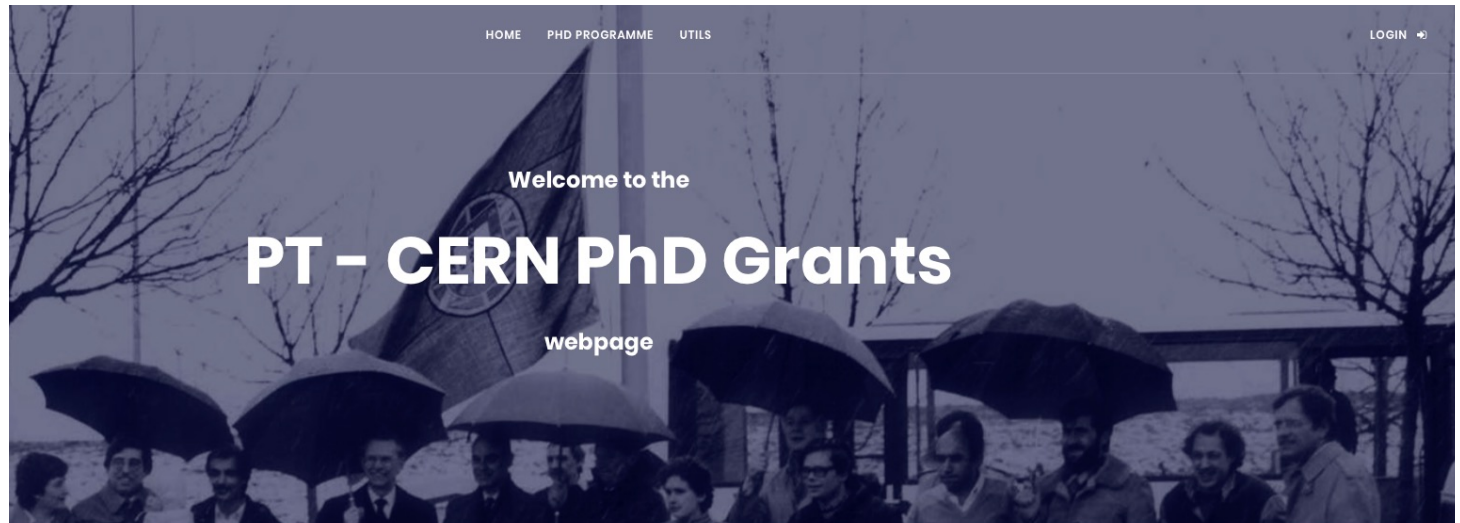
Coimbra (ICNAS): ciclotrão de 30 MeV + LINAC 70 MeV, com uma sala de tratamento para tumores oculares e produção de radioisótopos e radio fármacos para diagnóstico e terapia





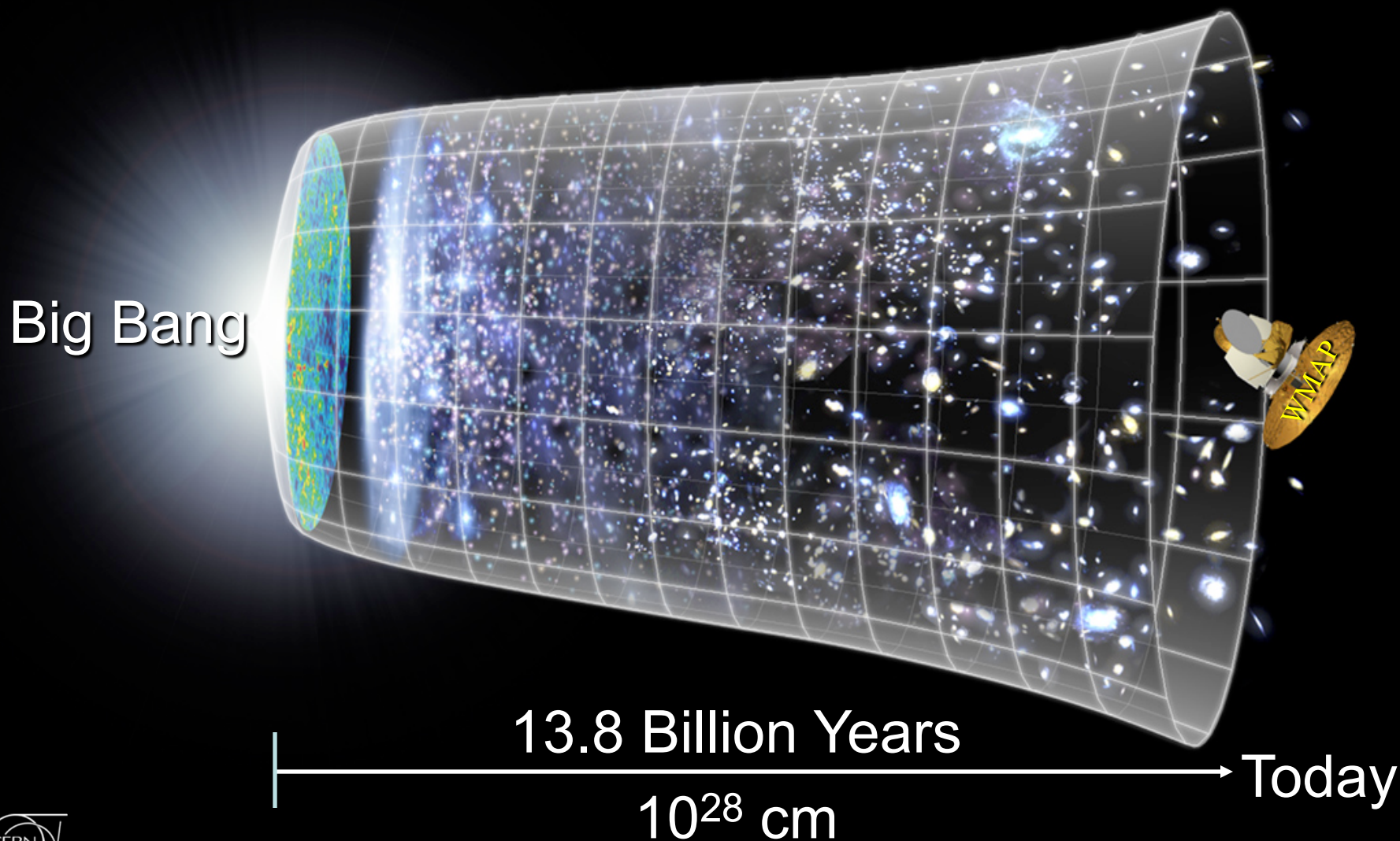
International Doctorate Network in Particle Physics, Astrophysics and Cosmology

Bolsas de Doutoramento



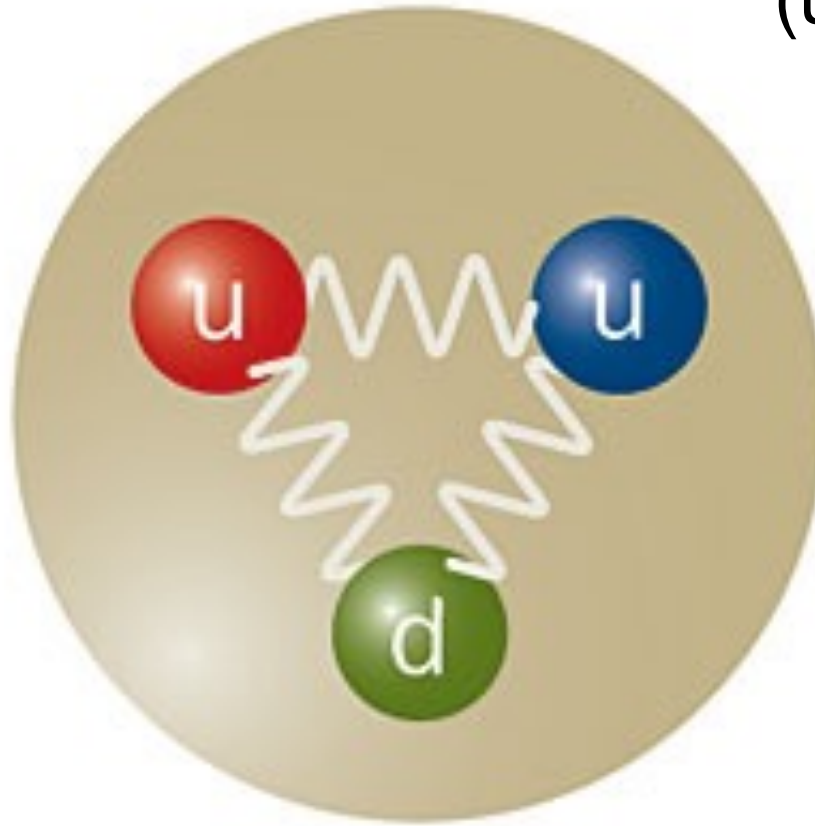
FCT – ProtoTera PhD Grants

O Universo para compreender, ...



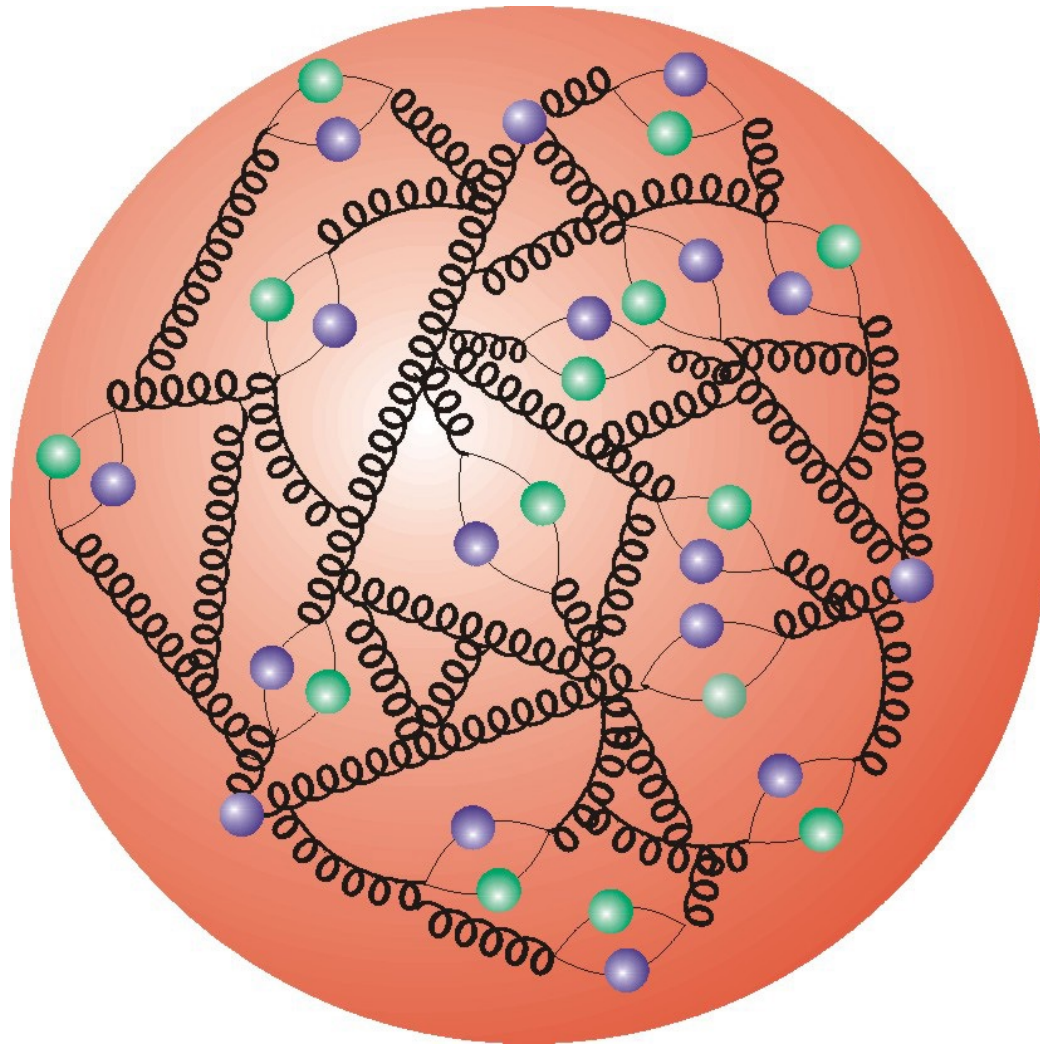
O Protão

(uud)



10^{-14} m

O Protão



A massa do Protão



$$M = \sum m_{\text{quarks}} + E_{\text{campo}}/c^2$$



A massa do Protão

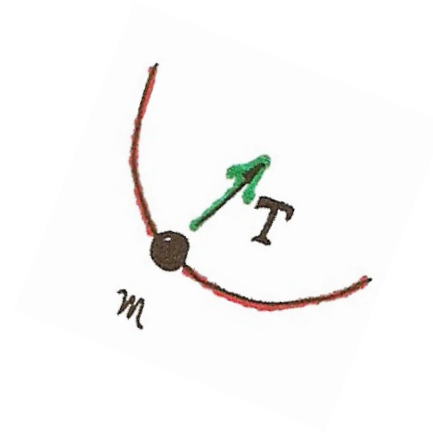
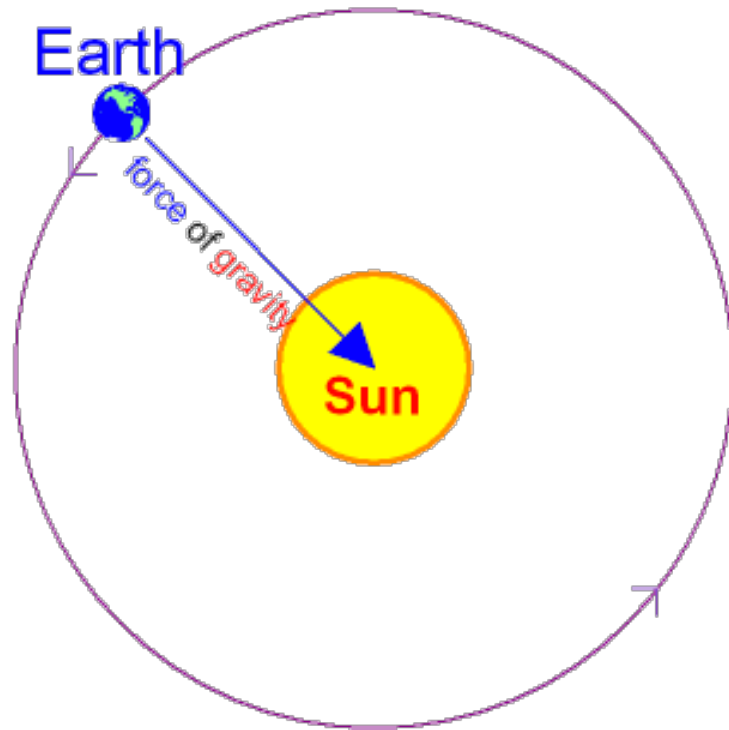


$$M = \sum m_{\text{quarks}} + E_{\text{campo}}/c^2$$

$$938 \sim (15 + 925) \text{ MeV}/c^2$$

$$1 \text{ MeV}/c^2 \sim 1.8 \cdot 10^{-30} \text{ Kg}$$

A massa do Sol

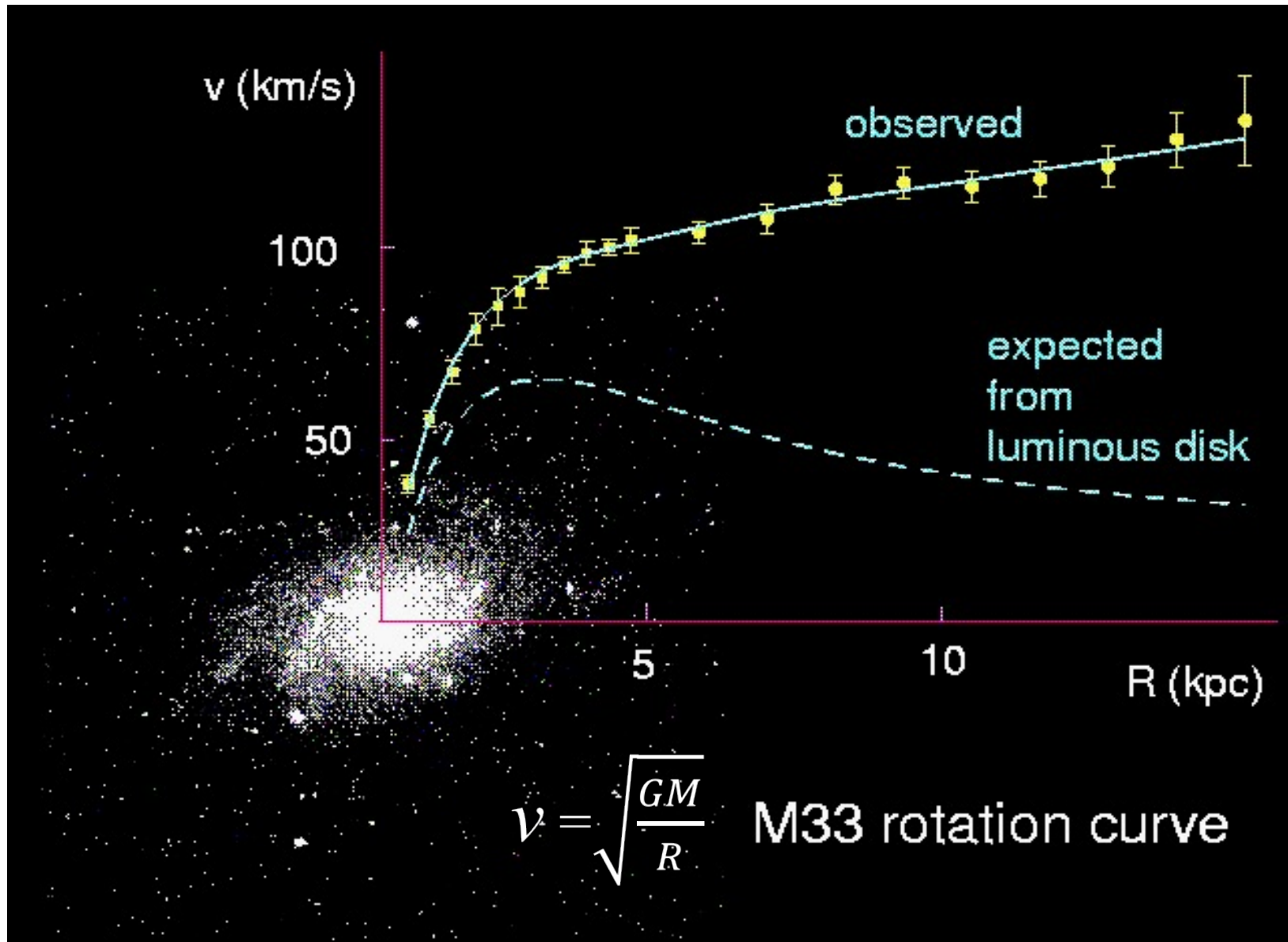


$$F_c = \frac{m v^2}{r}$$

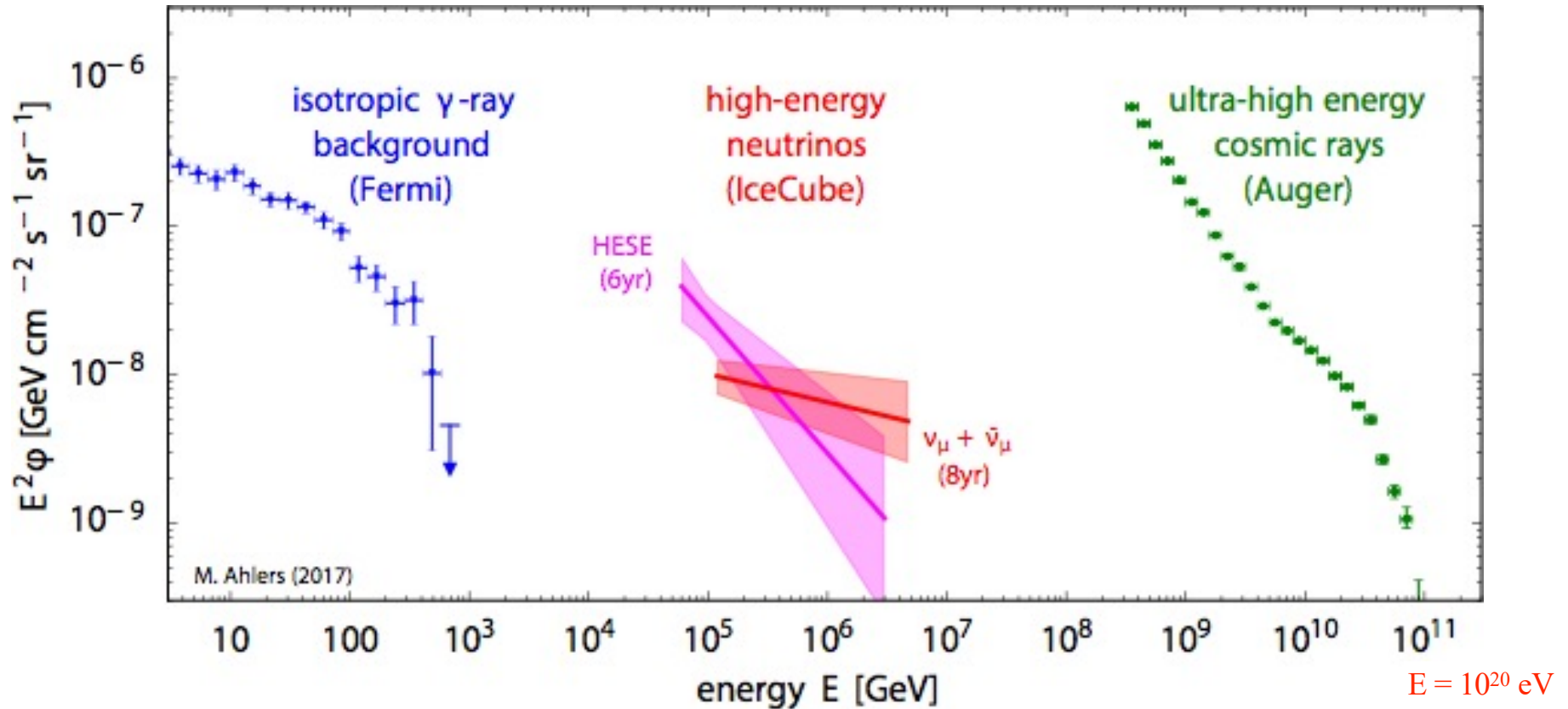
$$\frac{m_T v^2}{R} \approx G \frac{m_T M_S}{R^2}$$

$$M_S \sim 1.98 \cdot 10^{30} \text{ kg}$$

A velocidade de rotação das estrelas periféricas

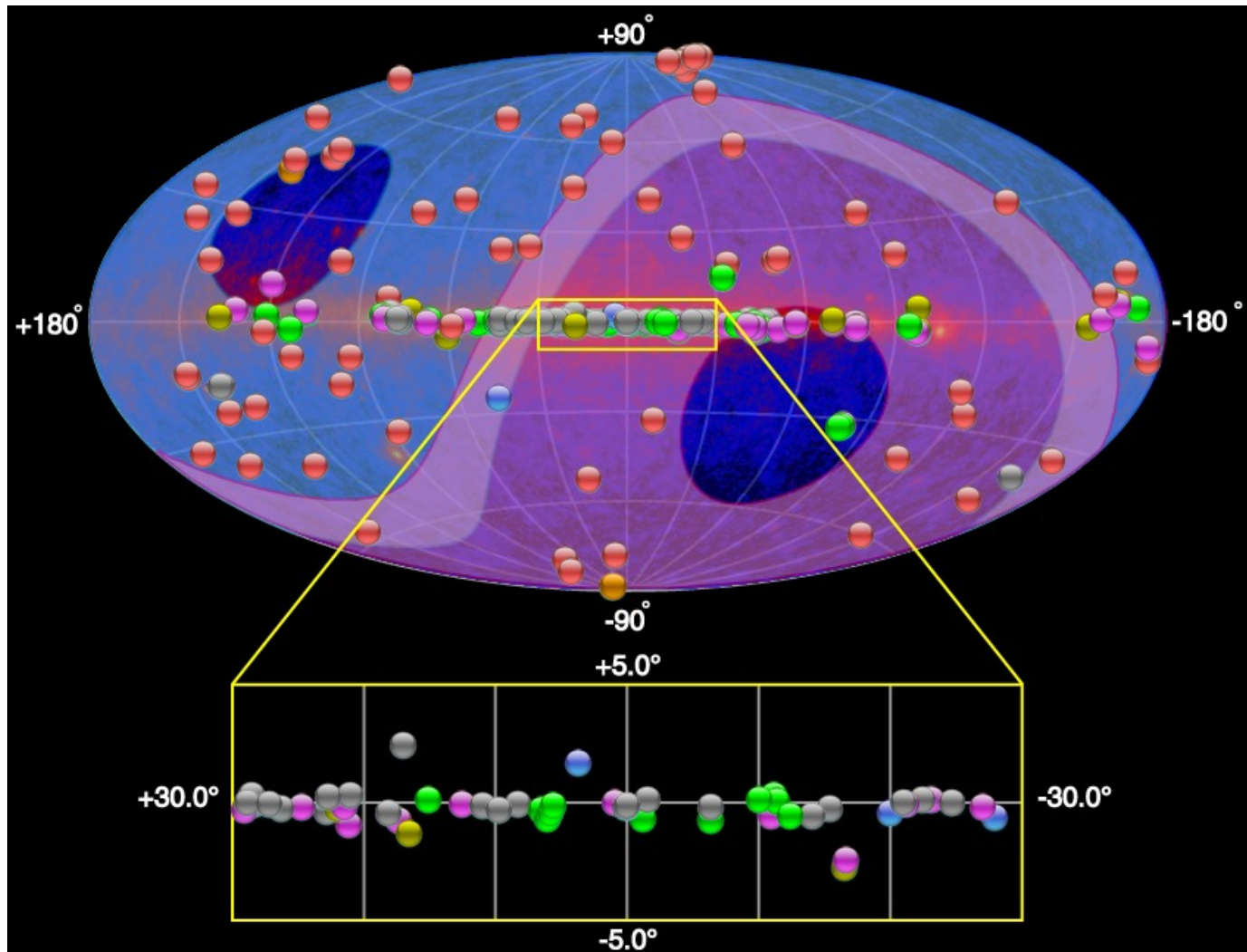


The Universe at the highest energies !



Energy density per decade similar in all three messenger particles

Fontes raios gamma de alta energia (TeV)



Source Types

- PWN
- Binary XRB PSR Gamma BIN
- HBL IBL FRI FSRQ
Blazar LBL AGN
(unknown type)
- Shell SNR/Molec. Cloud
Composite SNR
Superbubble
- Starburst
- DARK UNID Other
- uQuasar Star Forming
Region Globular Cluster
Cat. Var. Massive Star
Cluster BIN BL Lac
(class unclear) WR

(>200)

As bolhas de Fermi

