# QCD at non-zero isospin asymmetry and its physical relevance

# **Bastian Brandt**

#### Universität Bielefeld

#### in collaboration with Francesca Cuteri and Gergely Endrődi



Faculty of Physics





16.06.2021



### Contents

- 1. Introduction: (Lattice) QCD and isospin chemical potential
- 2. Phase diagram
- 3. Equation of State
- 4. Early Universe at large lepton flavour asymmetries
- 5. Conclusions

**QCD** at non-zero isospin asymmetry and its physical relevance Introduction: (Lattice) QCD and isospin chemical potential



#### 1. Introduction:

#### (Lattice) QCD and isospin chemical potential



## Motivation

theoretical first principles description of:

evolution of early Universe

(in particular: quark and hadron epochs)



[ Credit: BICEP2 collaboration/CERN/NASA\* ]

\* taken from Keck website: https://www.keckobservatory.org <sup>†</sup> Wikimedia Commons



### Motivation

theoretical first principles description of:

evolution of early Universe

(in particular: quark and hadron epochs)

compact stars: neutron stars



#### [ Credit: BICEP2 collaboration/CERN/NASA\* ]



[ Credit: Casey Reed - Penn St. U.<sup>†</sup> ]

\* taken from Keck website: https://www.keckobservatory.org <sup>†</sup> Wikimedia Commons



## Motivation

theoretical first principles description of:

evolution of early Universe

(in particular: quark and hadron epochs)

compact stars: neutron stars

mass radius relation (solving TOV equation)



[ Credit: BICEP2 collaboration/CERN/NASA\* ]



[ Credit: Casey Reed - Penn St. U.<sup>†</sup> ]



# Motivation

theoretical first principles description of:

evolution of early Universe

(in particular: quark and hadron epochs)

- compact stars: neutron stars mass radius relation (solving TOV equation)
- fireball in heavy-ion collisions

hydrodynamical treatment of equilibrated quark gluon plasma





#### [ Credit: BICEP2 collaboration/CERN/NASA\* ]



[ Credit: Casey Reed - Penn St. U.<sup>†</sup> ]

\* taken from Keck website: https://www.keckobservatory.org † V

† Wikimedia Commons



## Motivation

theoretical first principles description of:

evolution of early Universe

(in particular: quark and hadron epochs)

- compact stars: neutron stars mass radius relation (solving TOV equation)
- fireball in heavy-ion collisions

hydrodynamical treatment of equilibrated quark gluon plasma

 $\Rightarrow$  dominated by strong force

fundamental input:

- phase diagram
- equation of state (EoS)





#### [ Credit: BICEP2 collaboration/CERN/NASA\* ]



[ Credit: Casey Reed - Penn St. U.<sup>†</sup> ]

\* taken from Keck website: https://www.keckobservatory.org

† Wikimedia Commons



# Introduction – Quantum Chromodynamics (QCD)

theory of the strong interactions

#### **Standard Model of Elementary Particles**





# Introduction – Quantum Chromodynamics (QCD)

- theory of the strong interactions
- governs dynamics of smallest to largest boundstates

(mesons & baryons to neutron star cores)







#### Standard Model of Elementary Particles



# Introduction – Quantum Chromodynamics (QCD)

- theory of the strong interactions
- governs dynamics of smallest to largest boundstates

(mesons & baryons to neutron star cores)

responsible for 99.9% of the mass of visible matter in the Universe

mass [MeV] pion nucleon





#### Standard Model of Elementary Particles







### Introduction – Quantum Chromodynamics

u d s c t b

- QCD is a quantum field theory: particles  $\longleftrightarrow$  field excitations
- QCD particles and fields:

quarks

g gluons

$$\psi_f(x) \in \text{fundamental repr. of } SU(3)$$

 $A_{\mu}(x) \in \text{Lie algebra (su(3)) of } SU(3)$ 



#### Introduction – Quantum Chromodynamics

- $\blacktriangleright \ \mathsf{QCD} \ \mathsf{is a quantum field theory:} \quad \mathsf{particles} \longleftrightarrow \mathsf{field excitations}$
- QCD particles and fields:

quarks $\boldsymbol{U}$  $\boldsymbol{d}$  $\boldsymbol{s}$  $\boldsymbol{c}$  $\boldsymbol{t}$  $\boldsymbol{b}$  $\psi_f(x)$  $\in$  fundamental repr. of SU(3)gluons $\boldsymbol{e}$  $\boldsymbol{A}_{\mu}(x)$  $\in$  Lie algebra (su(3)) of SU(3)

► QCD Lagrangian: (Euclidean spacetime – after Wick rotation)  $\mathcal{L} = \sum_{f} \bar{\psi}_{f} \{ \gamma_{\mu} (\partial_{\mu} + ig_{s}A_{\mu}) + m_{f} \} \psi_{f} + \frac{1}{4} \text{Tr}(F_{\mu\nu}F_{\mu\nu}[A_{\mu}, g_{s}])$ fermionic Yang-Mills (gluonic)



### Introduction – Quantum Chromodynamics

- QCD is a quantum field theory: particles  $\longleftrightarrow$  field excitations
- QCD particles and fields:

u d s c t b quarks  $\psi_f(x) \in \text{fundamental repr. of } SU(3)$ gluons g  $A_{\mu}(x) \in \text{Lie algebra (su(3)) of } SU(3)$ 

QCD Lagrangian: (Euclidean spacetime – after Wick rotation)  $\mathcal{L} = \sum_{f} \bar{\psi}_{f} \{ \gamma_{\mu} (\partial_{\mu} + ig_{s}A_{\mu}) + m_{f} \} \psi_{f} + \frac{1}{4} \operatorname{Tr} (F_{\mu\nu} F_{\mu\nu} [A_{\mu}, g_{s}]) \}$ fermionic Yang-Mills (gluonic)

Euclidean path integral (PI) quantisation: [Feynman '48]

$$\mathcal{Z} = \int \mathcal{D} A \mathcal{D} \psi \mathcal{D} ar{\psi} \, \exp \Big( - \int d^4 x \mathcal{L}(x) \Big)$$

finite temperature partition function: (grand canonical – zero density)

$$\mathcal{Z} = \mathsf{Tr}\big[e^{-H/T}\big] = \int_{\mathrm{BC}} DAD\psi D\bar{\psi} \, \exp\Big(-\int_0^{1/T} \int d^3 x \mathcal{L}(x)\Big)$$



natural regulator for PI: discrete spacetime

[Wilson '74]





- natural regulator for PI: discrete spacetime
- gluonic link variables:  $U_{\mu}(x) = e^{iaA_{\mu}(x)} \hat{=} \longrightarrow$





- natural regulator for PI: discrete spacetime
- gluonic link variables:  $U_{\mu}(x) = e^{iaA_{\mu}(x)} \hat{=} \longrightarrow$
- ▶ PI: fermions integrated out  $Z = \int [dU] \det (D[U]) e^{-S_{gluon}[U]}$

gluon action and Dirac operator (staggered):

$$S_{\text{gluon}} = \frac{\beta}{N_c} \sum_{p} \text{ReTr}\left\{ \square \right\} \qquad D = \sum_{x,f} \left\{ m_f + \frac{\eta_{\mu}}{2a} \left( \longrightarrow - \longleftarrow \right)_{\mu} \right\}$$







- natural regulator for PI: discrete spacetime
- gluonic link variables:  $U_{\mu}(x) = e^{iaA_{\mu}(x)} \hat{=} \longrightarrow$
- ▶ PI: fermions integrated out  $\mathcal{Z} = \int [dU] \det (D[U]) e^{-S_{gluon}[U]}$

gluon action and Dirac operator (staggered):

$$S_{\text{gluon}} = \frac{\beta}{N_c} \sum_{p} \text{ReTr}\left\{ \square \right\} \qquad D = \sum_{x,f} \left\{ m_f + \frac{\eta_{\mu}}{2a} \left( \longrightarrow - \longleftarrow \right)_{\mu} \right\}$$

• in finite volume: finite dimensional integral  
(use a 
$$N_t \times N_s^3$$
 lattice – periodic spatial BC)

- $\Rightarrow$  Monte-Carlo methods importance sampling
- $\mathcal{O}(10^9)$  degrees of freedom







- natural regulator for PI: discrete spacetime
- gluonic link variables:  $U_{\mu}(x) = e^{iaA_{\mu}(x)} \hat{=} \longrightarrow$
- PI: fermions integrated out  $\mathcal{Z} = \int [dU] \det (D[U]) e^{-S_{gluon}[U]}$

gluon action and Dirac operator (staggered):

$$S_{\text{gluon}} = \frac{\beta}{N_c} \sum_{p} \text{ReTr}\left\{ \square \right\} \qquad D = \sum_{x,f} \left\{ m_f + \frac{\eta_{\mu}}{2a} \left( \longrightarrow - \longleftarrow \right)_{\mu} \right\}$$





[Copyright: Forschungszentrum Jülich]



- natural regulator for PI: discrete spacetime
- gluonic link variables:  $U_{\mu}(x) = e^{iaA_{\mu}(x)} \hat{=} \longrightarrow$
- PI: fermions integrated out  $\mathcal{Z} = \int [dU]$

gluon action and Dirac operator (staggered):

$$S_{\text{gluon}} = \frac{\beta}{N_c} \sum_{p} \text{ReTr}\left\{ \square \right\} \qquad D = \sum_{x,f} \left\{ m_f + \frac{\eta_{\mu}}{2a} \left( \longrightarrow - \longleftarrow \right)_{\mu} \right\}$$

 in finite volume: finite dimensional integral (use a N<sub>t</sub> × N<sub>s</sub><sup>3</sup> lattice – periodic spatial BC)
 ⇒ Monte-Carlo methods – importance sampling

 $\mathcal{O}(10^9)$  degrees of freedom  $\rightarrow$  use GPUs



[ Copyright: NVIDIA (from nvidia.com) ]

[Wilson '74]





## Introduction – Running coupling and Confinement

- QFT: physical quantities are renormalised quantities
  - $\begin{array}{l} \Rightarrow \quad \mbox{physical coupling } \alpha_s = g_s^R \mbox{ depends on the energy scale } \mu_E \\ \alpha_s \rightarrow 0 \mbox{ for } \mu_E \rightarrow \infty \qquad \mbox{ asymptotic freedom } [ \mbox{ Gross, Wilczek '73; Politzer '73 } ] \\ \alpha_s \rightarrow \infty \mbox{ for } \mu_E \rightarrow 0 \end{array}$

typical energy scales:  $g_s^R = \mathcal{O}(1)$ 

 $\Rightarrow$  non-perturbative methods are needed



# Introduction – Running coupling and Confinement

- QFT: physical quantities are renormalised quantities
  - $\Rightarrow \mbox{ physical coupling } \alpha_s = g_s^R \mbox{ depends on the energy scale } \mu_E \\ \alpha_s \rightarrow 0 \mbox{ for } \mu_E \rightarrow \infty \mbox{ asymptotic freedom [Gross, Wilczek '73; Politzer '73]}$

 $\alpha_{s} \rightarrow \infty$  for  $\mu_{E} \rightarrow 0$ 

typical energy scales:  $g_s^R = \mathcal{O}(1)$ 

 $\Rightarrow$  non-perturbative methods are needed

#### Confinement

- no direct evidence for QCD particles in colliders indirect evidence: jet events
  - $\rightarrow~$  quarks and gluons are confined in hadrons



[ TASSO DESY (Image: Oxford PPU) ]



# Introduction – Running coupling and Confinement

- QFT: physical quantities are renormalised quantities
  - $\label{eq:asymptotic freedom} \begin{array}{l} \Rightarrow \quad \mbox{physical coupling } \alpha_s = g_s^R \mbox{ depends on the energy scale } \mu_E \\ \alpha_s \rightarrow 0 \mbox{ for } \mu_E \rightarrow \infty \qquad \mbox{asymptotic freedom} \quad \begin{tabular}{l} $$ Gross, Wilczek '73; Politzer '73 end{tabular} \end{tabular}$

 $\alpha_{\rm s} \rightarrow \infty ~{\rm for}~\mu_{\rm E} \rightarrow {\rm 0}$ 

typical energy scales:  $g_s^R = \mathcal{O}(1)$ 

 $\Rightarrow$  non-perturbative methods are needed

#### Confinement

- no direct evidence for QCD particles in colliders indirect evidence: jet events
  - $\rightarrow~$  quarks and gluons are confined in hadrons
- heuristic explanation for quark confinement: flux tube between quark and antiquark

(evidence from pure gauge theory simulations [Lang, Rebbi '82; ... ] )



[ TASSO DESY (Image: Oxford PPU) ]





#### Introduction – chiral symmetry

fundamental global symmetry for most of the phenomena observed in QCD:

chiral symmetry 
$$SU_V(2) \times U_V(1) \times SU_A(2) \times U_A(1)$$
  $(N_f = 2)$   
intact intact broken spontaneously broken  
Baryon number  $\Rightarrow \langle \bar{\psi}\psi \rangle \neq 0$  anomalously

chiral limit  $(m_{u/d} = 0)$ 

 $\Rightarrow$  3 Goldstone bosons – pions



### Introduction – chiral symmetry

fundamental global symmetry for most of the phenomena observed in QCD:

chiral symmetry
$$SU_V(2) \times U_V(1) \times SU_A(2) \times U_A(1)$$
 $(N_f = 2)$ broken explicitlyintactbroken explicitly $m_u - m_d \neq 0$ Baryon number $m_{u/d} \neq 0 \Rightarrow \langle \bar{\psi}\psi \rangle \neq 0$ 

#### in nature (physical point)

 $\Rightarrow$  3 Goldstone bosons – pions  $\rightarrow$  pseudo Goldstone bosons –  $m_\pi 
eq 0$ 



### Introduction – chiral symmetry

fundamental global symmetry for most of the phenomena observed in QCD:

 $\begin{array}{c} \text{chiral symmetry} \\ \text{broken explicitly} \\ m_u - m_d \neq 0 \end{array} \begin{array}{c} \text{SU}_V(2) \times \text{U}_V(1) \times \text{SU}_A(2) \times \text{U}_A(1) \\ \text{intact} \\ \text{broken explicitly} \\ m_{u/d} \neq 0 \end{array} \begin{array}{c} (N_f = 2) \\ \text{intact} \\ \text{broken explicitly} \\ \text{order } \\ m_{u/d} \neq 0 \end{array}$ 

#### in nature (physical point)

- $\Rightarrow$  3 Goldstone bosons pions  $\rightarrow$  pseudo Goldstone bosons  $m_\pi 
  eq 0$
- Iow energy effective theory for Goldstone bosons:

chiral perturbation theory ( $\chi PT$ )

- systematic expansion in quark masses (external parameters (*T*, *B*, μ, ...) can be included)
- valid as long as  $m_f, p, T, \ldots < \Lambda_{\chi}$



At large T: quarks and gluons are liberated

 $\Rightarrow$  plasma of quarks and gluons



[ Homepage of the CMB experiment, GSI (FAIR) ]



At large T: quarks and gluons are liberated

- $\Rightarrow$  plasma of quarks and gluons
- transition at n = 0: (first principles LQCD)
  - transition is a crossover [Aoki et al '06]
  - $T_c pprox 157~{
    m MeV}~(\sim 10^{12}~{
    m K})$

[ Borsanyi et al '10; Bazavov et al '19 ]



[ Homepage of the CMB experiment, GSI (FAIR) ]



At large T: quarks and gluons are liberated

- $\Rightarrow$  plasma of quarks and gluons
- transition at n = 0: (first principles LQCD)
  - transition is a crossover [Aoki et al '06]
  - $T_c \approx 157 \text{ MeV} (\sim 10^{12} \text{ K})$ [Borsanyi *et al* '10; Bazavov *et al* '19]

driven by:

- effective restoration of SU<sub>A</sub>(2)
- deconfinement of quarks and gluons



Homepage of the CMB experiment, GSI (FAIR) ]





At large T: quarks and gluons are liberated

- $\Rightarrow$  plasma of quarks and gluons
- transition at n = 0: (first principles LQCD)
  - transition is a crossover [Aoki et al '06]
  - $T_c \approx 157$  MeV ( $\sim 10^{12}$  K)

[ Borsanyi et al '10; Bazavov et al '19 ]

driven by:

- effective restoration of SU<sub>A</sub>(2)
- deconfinement of quarks and gluons
- transition at  $n \neq 0$ ?

models predict a number of phases possible 1st order phase trans. + critical endpoint(s)

What about first principles information?









Lattice QCD: grand canonical ensemble density  $n \rightarrow$  chemical potential  $\mu$ "Physical" basis for QCD at  $N_f = 3$ : (strangeness not conserved in SM)

$$\mu_{u} = \frac{\mu_{B}}{3} + \frac{2\mu_{Q}}{3} \qquad \mu_{d} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} \qquad \mu_{s} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} - \mu_{s}$$



Lattice QCD: grand canonical ensemble density  $n \rightarrow$  chemical potential  $\mu$ 

"Physical" basis for QCD at  $N_f = 3$ : (strangeness not conserved in SM)

$$\mu_{u} = \frac{\mu_{B}}{3} + \frac{2\mu_{Q}}{3} \qquad \mu_{d} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} \qquad \mu_{s} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} - \mu_{s}$$

Convenient basis for simulations: ("isospin" basis)

$$\mu_u = \mu_L + \mu_I \qquad \mu_d = \mu_L - \mu_I \qquad \mu_s$$

relation between bases:

$$\mu_B = 3\mu_L - \mu_I$$
  $\mu_Q = 2\mu_I$   $\mu_S = \mu_L - \mu_I - \mu_s$ 



Lattice QCD: grand canonical ensemble density  $n \rightarrow$  chemical potential  $\mu$ 

"Physical" basis for QCD at  $N_f = 3$ : (strangeness not conserved in SM)

$$\mu_{u} = \frac{\mu_{B}}{3} + \frac{2\mu_{Q}}{3} \qquad \mu_{d} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} \qquad \mu_{s} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} - \mu_{s}$$

Convenient basis for simulations: ("isospin" basis)

$$\mu_u = \mu_L + \mu_I \qquad \mu_d = \mu_L - \mu_I \qquad \mu_s$$

▶  $\mu_L \neq 0 \neq \mu_s$ : complex action (sign) problem

$$\mathcal{Z} = \int [dU] \det \left( D[U](\mu_u, \mu_d, \mu_s) 
ight) e^{-S_{ ext{gluon}}[U]} \in \mathbb{C}$$

relation between bases:

$$\mu_B = 3\mu_L - \mu_I$$
  $\mu_Q = 2\mu_I$   $\mu_S = \mu_L - \mu_I - \mu_s$ 



Lattice QCD: grand canonical ensemble density  $n \rightarrow$  chemical potential  $\mu$ 

"Physical" basis for QCD at  $N_f = 3$ : (strangeness not conserved in SM)

$$\mu_{u} = \frac{\mu_{B}}{3} + \frac{2\mu_{Q}}{3} \qquad \mu_{d} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} \qquad \mu_{s} = \frac{\mu_{B}}{3} - \frac{\mu_{Q}}{3} - \mu_{s}$$

Convenient basis for simulations: ("isospin" basis)

$$\mu_u = \mu_L + \mu_I \qquad \qquad \mu_d = \mu_L - \mu_I \qquad \qquad \mu_s$$

•  $\mu_L \neq 0 \neq \mu_s$ : complex action (sign) problem

$$\mathcal{Z} = \int [dU] \det \left( D[U](\mu_u, \mu_d, \mu_s) 
ight) e^{-S_{ ext{gluon}}[U]} \in \mathbb{C}$$

• pure isospin chemical potential:  $\mu_L = \mu_s = 0$ 

det (D) is real, positive definite  $\longrightarrow$  suitable for importance sampling relation between bases:

 $\mu_B = 3\mu_L - \mu_I$   $\mu_Q = 2\mu_I$   $\mu_S = \mu_L - \mu_I - \mu_s$ 



10<sup>14</sup>vr

10<sup>12</sup>vr

10<sup>10</sup>yr

10<sup>8</sup>vi 10<sup>6</sup>vi

10<sup>4</sup>yr 100 yr

1 yr 10<sup>6</sup>s 10<sup>4</sup>s 100 s

1 s 10-25

10-4 10-65 10-85

#### Isospin asymmetry – physical significance

isospin asymmetry:  $n_l = n_u - n_d \neq 0$ 

- stable isotopes
- heavy-ion collisions (Pb or Au)  $N_n/N_n \approx 2/3$
- neutron star cores [Steiner et al '05]  $N_{p}/N_{p} \gtrsim 1/39$
- early universe?
  - $n_l \neq 0$  possible for large lepton flavour asymmetries [ Abuki, Brauner, Warringa '09; Wygas, Oldengott, Bödeker, Schwarz '18 ] discussed in more detail later  $\Rightarrow$



160

140



#### lsospin asymmetry – physical significance

isospin asymmetry:  $n_l = n_u - n_d \neq 0$ 

- stable isotopes
- ► heavy-ion collisions (Pb or Au)  $N_p/N_n \approx 2/3$
- neutron star cores [Steiner *et al* '05]  $N_p/N_n \gtrsim 1/39$

#### early universe?

 $n_l \neq 0$  possible for large lepton flavour asymmetries [ Abuki, Brauner, Warringa '09; Wygas, Oldengott, Bödeker, Schwarz '18 ]  $\Rightarrow$  discussed in more detail later

typically:  $n_B$  dominates (exception **eU**)



however:  $n_l$  important ingredient


#### 2. Phase diagram

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \square$  Phase diagram

#### UNIVERSITÄT BIELEFELD Faculty of Physics

# Phase diagram

 $\chi \mathsf{PT}$  [Son, Stephanov '01]

expected phase diagram:

- hadronic phase (white)
- quark-gluon plasma



[ Brandt, Endrődi, Schmalzbauer '18 ]



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \square$  Phase diagram

# Phase diagram

 $\chi$ PT [Son, Stephanov '01]

expected phase diagram:

- hadronic phase (white)
- quark-gluon plasma
- T = 0 and  $\mu_I \ge m_\pi/2$ :  $(\mu_Q \ge m_\pi)$

condensation of charged pions (Bose-Einstein Condensation – BEC)





UNIVERSITÄT

Faculty of Physics

QCD at non-zero isospin asymmetry and its physical relevance Phase diagram

# Phase diagram

 $\chi PT$ [ Son, Stephanov '01 ]

expected phase diagram:

- hadronic phase (white)
- guark-gluon plasma
- $\blacktriangleright$  T = 0 and  $\mu_l \ge m_\pi/2$ :  $(\mu_Q \ge m_\pi)$ condensation of charged pions (Bose-Einstein Condensation – BEC)
- $\blacktriangleright \mu_I \gg m_{\pi}/2$ : BCS phase!?

perturbation theory: 1-gluon exchange is attractive







[Brandt, Endrődi, Schmalzbauer '18]



$$D = \gamma_{\mu} D_{\mu} + m_{ud} + \gamma_{0} \tau_{3} \mu_{I}$$
  
SU<sub>V</sub>(2)   
explicit  
$$\mu_{I} \neq 0$$



$$D = \gamma_{\mu} D_{\mu} + m_{ud} + \gamma_{0} \tau_{3} \mu_{I}$$
  
SU<sub>V</sub>(2)  $\longrightarrow_{I}$  U<sub>Q</sub>(1)  $\xrightarrow{I}$  Ø  
explicit spontaneous  
 $\mu_{I} \neq 0$   $\mu_{I} \ge m_{\pi}/2$ 





$$D = \gamma_\mu D_\mu + m_{ud} + \gamma_0 \tau_3 \mu_\mu$$

$$SU_{V}(2) \xrightarrow{\bullet} U_{Q}(1) \xrightarrow{\bullet} \varnothing$$
explicit spontaneous
$$\mu_{I} \neq 0 \qquad \mu_{I} \geq m_{\pi}/2$$

- cannot observe spontaneous symmetry breaking in finite V
- low mode in simulations





$$D = \gamma_{\mu} D_{\mu} + m_{ud} + \gamma_0 \tau_3 \mu_I + i \gamma_5 \tau_2 \lambda$$

Ø

explicit

$$\begin{array}{ccc} \mathsf{SU}_V(2) & \longrightarrow & \mathsf{U}_Q(1) & \longrightarrow & \varnothing \\ & & & & & \\ & & & \\ & & & & \\$$

- cannot observe spontaneous symmetry breaking in finite V
- low mode in simulations

need to break symmetry explicitly

 $\Rightarrow$  introduce regulator:  $\sim \lambda$ pionic source [Kogut, Sinclair '02] physical results: extrapolate  $\lambda \rightarrow 0$ reliable extrapolations: main task for analysis





$$D = \gamma_{\mu}D_{\mu} + m_{ud} + \gamma_{0} au_{3} \mu_{I} + i\gamma_{5} au_{2} \lambda$$

Ø

$$U_{V}(2) \xrightarrow{\bullet} U_{Q}(1) \xrightarrow{\bullet} \varnothing$$
explicit
explicit
$$\mu_{I} \neq 0$$
pionic source  $\lambda$ 

- cannot observe spontaneous symmetry breaking in finite V
- low mode in simulations

#### need to break symmetry explicitly

 $\Rightarrow$  introduce regulator:  $\sim \lambda$ pionic source [Kogut, Sinclair '02] physical results: extrapolate  $\lambda \rightarrow 0$ reliable extrapolations: main task for analysis

improvement program:

S

[Brandt, Endrődi, Schmalzbauer '18]

- valence quark improvement
- leading order reweighting





# Status and setup

- model/EFT results: (recent review: [Mannarelli '19])
  - $\chi$ PT (P)NJL phase diagram in hep-ph/0508117
  - linear sigma model HRG

(for detailled references see [Brandt, Endrődi, Schmalzbauer '18])

first results from lattice QCD:

 $N_t = 4$ , unphysical masses, unimproved:

- $N_f = 2$  [Kogut, Sinclair, '02; '04]
- $N_f = 8$  [de Forcrand, et al, '07]

canonical approach, unph. masses, T = 0:

•  $N_f=2+1$  [Detmold, Orginas, Shi '12]

here: • improved actions (N<sub>f</sub> = 2 + 1)

- physical quark masses
- well controlled  $\lambda$ -extrapolations (from now on everything  $\lambda = 0$ )



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \square$  Phase diagram



0.6 0.8

0.2 0.4

μ / m.

### Phase diagram from the lattice

continuum phase diagram

[ Brandt, Endrődi, Schmalzbauer '18 ]

• BEC phase boundary



(Mer)

order parameter:

renormalised pion condensate



## Phase diagram from the lattice

- continuum phase diagram
   [ Brandt, Endrődi, Schmalzbauer '18 ]
  - BEC phase boundary
     O(2) universality class



- order parameter: renormalised pion condensate
- finite size scaling: consistency with O(2) scaling

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \square$  Phase diagram



## Phase diagram from the lattice

- continuum phase diagram

   [Brandt, Endrödi, Schmalzbauer '18 ]
  - BEC phase boundary
     O(2) universality class
  - chiral crossover



relevant observable:

renormalised chiral condensate



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \square$  Phase diagram



## Phase diagram from the lattice

- continuum phase diagram
   [ Brandt, Endrödi, Schmalzbauer '18 ]
  - BEC phase boundary
     O(2) universality class
  - chiral crossover
  - pseudo-triple point



meeting point:

crossover and BEC phase boundary:

pseudo coexistence of three phases

from this point on:

phase boundaries coincide



#### Comparison to model and EFT results



lattice results as model input at  $\mu \neq 0$ :

- calibrate models
- improve understanding of model capabilities





## Phase diagram in the $n_l - T$ plane

using a model independent 2d interpolation of  $n_l$ 

(model independent: all possible "good" spline fits via Monte-Carlo)

[ S. Borsanyi, private comm.; Brandt, Endrődi '16 ]

determine the  $n_I - T$  phase diagram





## Phase diagram in the $n_l - T$ plane

using a model independent 2d interpolation of  $n_l$ 

(model independent: all possible "good" spline fits via Monte-Carlo)

[ S. Borsanyi, private comm.; Brandt, Endrődi '16 ]

determine the  $n_I - T$  phase diagram



... work in progress



#### 3. Equation of state

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Equation of state



### Equation of state at T = 0

Grand canonical ensemble T = 0:

$$p(T = 0, \mu_I = 0) = 0$$
  $n_I = \frac{\partial p}{\partial \mu_I}$   $\epsilon = -p + n_I \mu_I$ 

 $\Rightarrow \quad p(0,\mu_{I}) = \int_{0}^{\mu_{I}} d\mu \, n_{I}(0,\mu_{I}) \qquad I = -4p + n_{I}(0,\mu_{I})\mu_{I}$ 

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Equation of state



## Equation of state at T = 0

Grand canonical ensemble T = 0:

 $p(T = 0, \mu_I = 0) = 0$   $n_I = \frac{\partial p}{\partial \mu_I}$   $\epsilon = -p + n_I \mu_I$ 

 $\Rightarrow \quad p(0,\mu_{I}) = \int_{0}^{\mu_{I}} d\mu \, n_{I}(0,\mu_{I}) \qquad I = -4p + n_{I}(0,\mu_{I})\mu_{I}$ 

- lattice simulation: [Brandt et al '18]
   T = 0 never exactly fulfilled
  - $\Rightarrow$  use large  $N_t$  so that  $T \approx 0$



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \vdash$  Equation of state



### Equation of state at T = 0

Grand canonical ensemble T = 0:

 $p(T = 0, \mu_I = 0) = 0$   $n_I = \frac{\partial p}{\partial \mu_I}$   $\epsilon = -p + n_I \mu_I$ 

 $\Rightarrow \quad p(0,\mu_I) = \int_0^{\mu_I} d\mu \, n_I(0,\mu_I) \qquad I = -4p + n_I(0,\mu_I)\mu_I$ 

- lattice simulation: [Brandt *et al* '18] T = 0 never exactly fulfilled
  - $\Rightarrow$  use large  $N_t$  so that  $T \approx 0$
- correct for  $T \neq 0$  effects with  $\chi$ PT particularly relevant for  $\mu_l \approx m_\pi/2$



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \vdash$  Equation of state



## Equation of state at T = 0

Grand canonical ensemble T = 0:

 $p(T = 0, \mu_I = 0) = 0$   $n_I = \frac{\partial p}{\partial \mu_I}$   $\epsilon = -p + n_I \mu_I$ 

 $\Rightarrow \quad p(0,\mu_I) = \int_0^{\mu_I} d\mu \, n_I(0,\mu_I) \qquad I = -4p + n_I(0,\mu_I)\mu_I$ 

- Introduction: [Brandt et al '18] T = 0 never exactly fulfilled
  - $\Rightarrow$  use large  $N_t$  so that  $T \approx 0$
- correct for  $T \neq 0$  effects with  $\chi PT$ particularly relevant for  $\mu_l \approx m_{\pi}/2$ 
  - $\Rightarrow$  fit at LO gives  $f_{\pi} = 133(4)$  MeV





## Equation of state at T = 0

Resulting equation of state:



- in good agreement with NLO χPT [Adhikari, Andersen '19]
- currently: only a single lattice spacing

continuum limit: ... work in progress



## Extracting the EoS at non-zero T

p(T,0) and I(T,0) are known [Borsanyi et al '13, Bazavov et al '14]

 $\Rightarrow$  need to determine:

 $\Delta p(T,\mu_I) \equiv p(T,\mu_I) - p(T,0) \quad \text{and} \quad \Delta I(T,\mu_I) \equiv I(T,\mu_I) - I(T,0)$ 



## Extracting the EoS at non-zero T

p(T, 0) and I(T, 0) are known [Borsanyi et al '13, Bazavov et al '14]

 $\Rightarrow$  need to determine:

 $\Delta p(T,\mu_I) \equiv p(T,\mu_I) - p(T,0)$  and  $\Delta I(T,\mu_I) \equiv I(T,\mu_I) - I(T,0)$ 

• computation of  $\Delta p$ :

As before:  $\Delta p(T, \mu_l) = \int_0^{\mu_l} d\mu \, n_l(T, \mu)$ 

• computation of  $\Delta I$ :

Starting point: 
$$\frac{\Delta I(T,\mu_l)}{T^4} = T \frac{\partial}{\partial T} \left( \frac{\Delta p(T,\mu_l)}{T^4} \right) + \frac{\mu_l n_l(T,\mu_l)}{T^4}$$

$$\Rightarrow \Delta I(T,\mu_l) = -4 \int_0^{\mu_l} d\mu'_l n_l(T,\mu'_l) + \int_0^{\mu_l} d\mu'_l T \frac{\partial}{\partial T} n_l(T,\mu'_l) + \mu_l n_l(T,\mu_l)$$

use 2d interpolation of  $n_I(T, \mu_I)$ 

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Equation of state





QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Equation of state







QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Early Universe at large lepton flavour asymmetries



#### 4. Early Universe at large lepton flavour asymmetries



QCD at non-zero isospin asymmetry and its physical relevance  $\Box$  Early Universe at large lepton flavour asymmetries



#### Evolution of the early Universe



for 100  ${\rm GeV}>$   $T\gtrsim$  10  ${\rm MeV}:$ 

 $n_B, n_Q, n_{L_e}, n_{L_\mu}, n_{L_\tau}$  conserved in comoving volume

Cosmic trajectory: isentropic expansion with parameters  $T, \mu_B, \mu_Q, \mu_{L_\ell}$  with  $\ell \in (e, \mu, \tau)$  (grand canonical ensemble)

conservation equations:

 $\frac{n_B}{s} = b \qquad \frac{n_Q}{s} = 0 \qquad \frac{n_{L_\ell}}{s} = l_\ell$ 

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Early Universe at large lepton flavour asymmetries



#### Evolution of the early Universe



for 100  ${\rm GeV}>$   $T\gtrsim$  10  ${\rm MeV}:$ 

 $n_B, n_Q, n_{L_e}, n_{L_{\mu}}, n_{L_{\tau}}$  conserved in comoving volume

Cosmic trajectory: isentropic expansion with parameters
 *T*, μ<sub>B</sub>, μ<sub>Q</sub>, μ<sub>Lℓ</sub> with ℓ ∈ (e, μ, τ)
 (grand canonical ensemble)

conservation equations:

 $\frac{n_B}{s} = b \qquad \frac{n_Q}{s} = 0 \qquad \frac{n_{L_\ell}}{s} = l_\ell$ 

- empirical constraints (CMB):
  - $b = 8.6(0.06) \cdot 10^{-11}$  [Planck collaboration '15]
  - $|\mathit{I}_e + \mathit{I}_\mu + \mathit{I}_ au| < 0.012$  [ Oldengott, Schwarz '17 ]



QCD at non-zero isospin asymmetry and its physical relevance  $\hfill \sqsubseteq$  Early Universe at large lepton flavour asymmetries



#### Evolution of the early Universe



for 100  ${\rm GeV}>$   $T\gtrsim$  10  ${\rm MeV}:$ 

 $n_B, n_Q, n_{L_e}, n_{L_{\mu}}, n_{L_{\tau}}$  conserved in comoving volume

Cosmic trajectory: isentropic expansion with parameters *T*, μ<sub>B</sub>, μ<sub>Q</sub>, μ<sub>Lℓ</sub> with ℓ ∈ (e, μ, τ) (grand canonical ensemble)

#### conservation equations:

 $\frac{n_B}{s} = b \qquad \frac{n_Q}{s} = 0 \qquad \frac{n_{L_\ell}}{s} = l_\ell$ 

empirical constraints (CMB):

- $b = 8.6(0.06) \cdot 10^{-11}$  [ Planck collaboration '15 ]
- $|\mathit{I}_e + \mathit{I}_\mu + \mathit{I}_ au| < 0.012$  [ Oldengott, Schwarz '17 ]

weak equilibrium & charge neutrality:

 $\Rightarrow$  standard scenario:  $\mu_B \approx \mu_Q \approx 0$ 

QCD at non-zero isospin asymmetry and its physical relevance  $\hfill {\mbox{ Learly Universe at large lepton flavour asymmetries}}$ 



#### Evolution of the early Universe



for 100  ${\rm GeV}>$   $T\gtrsim$  10  ${\rm MeV}:$ 

 $n_B, n_Q, n_{L_e}, n_{L_{\mu}}, n_{L_{\tau}}$  conserved in comoving volume

Cosmic trajectory: isentropic expansion with parameters  $T, \mu_B, \mu_Q, \mu_{L_\ell}$  with  $\ell \in (e, \mu, \tau)$  EoS for (grand canonical ensemble)  $p \approx p_Q$ 

conservation equations:

 $\frac{n_B}{s} = b \qquad \frac{n_Q}{s} = 0 \qquad \frac{n_{L_\ell}}{s} = l_\ell$ 

- empirical constraints (CMB):
  - $b=8.6(0.06)\cdot10^{-11}$  [ Planck collaboration '15 ]
  - $|\mathit{I}_{e}+\mathit{I}_{\mu}+\mathit{I}_{ au}| < 0.012$  [ Oldengott, Schwarz '17 ]

weak equilibrium & charge neutrality:

 $\Rightarrow$  standard scenario:  $\mu_B \approx \mu_Q \approx 0$ 

EoS for cosmic trajectory:

#### $p_{\rm QCD}$ from the lattice

• to first approximation:

EoS for  $\mu_B = \mu_Q = 0$ 

• effects of  $\mu_B \neq 0 \neq \mu_Q$ : Taylor expansion QCD at non-zero isospin asymmetry and its physical relevance  $\Box$  Early Universe at large lepton flavour asymmetries



## Cosmic trajectory with lepton flavour asymmetries



empirical constraints (CMB):

- $b = 8.6(0.06) \cdot 10^{-11}$  [Planck collaboration '15]
- $|I_e+I_\mu+I_ au| < 0.012$  [Oldengott, Schwarz '17]

▶ sum of  $I_{\ell}$  constrained – individual  $I_{\ell}$  not

 $I_{\ell} - I_{\ell'} \gtrsim 0.1$  possible in models:

- explanation of  ${}^{4}\mathrm{He}$  and light element abundances [Ichikawa et al '04, ...]
- from beyond standard model (BSM) physics [McDonald '00, ... ]

QCD at non-zero isospin asymmetry and its physical relevance  $\Box$  Early Universe at large lepton flavour asymmetries



## Cosmic trajectory with lepton flavour asymmetries



empirical constraints (CMB):

- $b=8.6(0.06)\cdot10^{-11}$  [Planck collaboration '15]
- $|I_e+I_\mu+I_ au| < 0.012$  [Oldengott, Schwarz '17]

▶ sum of  $I_{\ell}$  constrained – individual  $I_{\ell}$  not

 $I_{\ell} - I_{\ell'} \gtrsim 0.1$  possible in models:

• explanation of  ${}^{4}\mathrm{He}$  and light element abundances [Ichikawa et al '04, ... ]

• from beyond standard model (BSM) physics [McDonald '00, ... ]

• cosmic trajectories for  $|l_{\ell}| \gg 0$  (but  $|l_e + l_{\mu} + l_{\tau}| = 0$ )

- $\Rightarrow$  trajectories at large  $\mu_Q \gg |\mu_B| > 0$
- EoS from Taylor expansion [ Middeldorf-Wygas, Oldengott, Bödeker, Schwarz '20 ]
- EoS from model [Vovchenko, Brandt et al '20]



### Cosmic trajectories and pion condensation



#### First test: simple model with pion condensation

- HRG & effective mass model for pions (quasiparticle picture & rearrangement term)
- match the model to lattice QCD at T = 0

[ Vovchenko, Schaffner-Bielich, Hajkarim, Brandt et al '20 ]



### Cosmic trajectories and pion condensation



#### First test: simple model with pion condensation

- HRG & effective mass model for pions (quasiparticle picture & rearrangement term)
- match the model to lattice QCD at T = 0[Vovchenko, Schaffner-Bielich, Hajkarim, Brandt *et al* '20]

 $T \lesssim 160 \text{ MeV}$   $\mu_I \lesssim 0.8 - 0.9 m_{\pi}$ 

To determine range of applicability: compare EoS to lattice data ( $N_t = 10, 12$ )

expect the model to work reliably for:




#### Cosmic trajectories and pion condensation



First test: simple model with pion condensation

- HRG & effective mass model for pions (quasiparticle picture & rearrangement term)
- match the model to lattice QCD at T = 0[Vovchenko, Schaffner-Bielich, Hajkarim, Brandt *et al* '20]

To determine range of applicability: compare EoS to lattice data ( $N_t = 10, 12$ )

expect the model to work reliably for:

Cosmic trajectories with large  $|I_e + I_\mu|$ 

constraints:

- $|I_e + I_\mu + I_\tau| = 0$
- $I_e I_\mu = 0$
- dependence on  $I_e I_\mu$  very mild



 $T \lesssim 160 \text{ MeV}$   $\mu_I \lesssim 0.8 - 0.9 m_{\pi}$ 



#### Cosmic trajectories and pion condensation



Effects of pion condensation:

- pion condensation strongly affects EoS
- enhances relic density of primordial gravitational waves
- modifies fraction of primordial black holes heavier than solar masses





First test: simple model with pion condensation

- HRG & effective mass model for pions (quasiparticle picture & rearrangement term)
- match the model to lattice QCD at T = 0

[ Vovchenko, Schaffner-Bielich, Hajkarim, Brandt et al '20 ]

## Conclusions

• QCD phase diagram at  $\mu_I \neq 0$ 

open question:

existence/location of BCS phase?





### Conclusions

• QCD phase diagram at  $\mu_I \neq 0$ 

open question:

existence/location of BCS phase?

Equation of state

work in progress: continuum limit





### Conclusions

• QCD phase diagram at  $\mu_I \neq 0$ 

open question: existence/location of BCS phase?

Equation of state

work in progress: continuum limit

- early universe with lepton flavour asymmetry
   [Vovchenko, Schaffner-Bielich, Hajkarim, Brandt et al '20]
  - ⇒ conceivable lepton flavour asymmetries: pion condensation can occur possible environment for pion stars? [Brandt, Endrödi, Fraga, Hippert *et al* '18]





### Outlook: extension to non-zero baryon density

Physical systems discussed:

all feature  $\mu_L \neq 0 \neq \mu_s$ 

- early universe with large lepton asymm.
- compact stars
- heavy ion-collisions



eventually: need to overcome sign problem

for small  $\mu_L \neq 0 \neq \mu_s$ :

indirect methods: (starting from  $\mu = 0$ )

Taylor expansion; analytic continuation; reweighting

(only possible in absence of phase transition)

here: use results at  $\mu_{I} \neq 0$  as a novel starting point for Taylor expansion

 $\Rightarrow$  Work in progress . . .



# Hunting the BCS phase

First evidence:

coexistence of BEC and deconfinement

⇒ look at Polyakov loop (associated with deconfinement)

However: need a more clear-cut criterion





## Hunting the BCS phase

First evidence:

coexistence of BEC and deconfinement

⇒ look at Polyakov loop (associated with deconfinement)

However: need a more clear-cut criterion

• extension of Banks-Casher relation to  $\mu_I \neq 0$ :

$$\Delta^2 = rac{2\pi^3}{9} \left< 
ho(0) \right> \quad (V o \infty, \, m_{ud} o 0, \, \mu_I \, ext{ large})$$

[Kanazawa, Wettig, Yamamoto '12]





## Hunting the BCS phase

First evidence:

coexistence of BEC and deconfinement

⇒ look at Polyakov loop (associated with deconfinement)

However: need a more clear-cut criterion

• extension of Banks-Casher relation to  $\mu_I \neq 0$ :

$$\Delta^2 = rac{2\pi^3}{9} \left< 
ho(0) \right> \quad (V o \infty, \, m_{ud} o 0, \, \mu_I \, ext{large})$$

[Kanazawa, Wettig, Yamamoto '12]

$$\begin{array}{ll} \text{for } m_{ud} \neq 0 & \rho(0) \to \rho(m_{ud}) \\ \text{search plateau for } \rho(m_{ud}) \text{ at large } \mu_I \end{array}$$





# Hunting the BCS phase

First evidence:

coexistence of BEC and deconfinement

⇒ look at Polyakov loop (associated with deconfinement)

However: need a more clear-cut criterion

• extension of Banks-Casher relation to  $\mu_I \neq 0$ :

$$\Delta^2 = rac{2\pi^3}{9} \langle 
ho(0) 
angle \quad (V o \infty, \, m_{ud} o 0, \, \mu_I \, ext{large})$$

[Kanazawa, Wettig, Yamamoto '12]

for 
$$m_{\scriptscriptstyle ud} 
eq 0$$
:  $ho(0) 
ightarrow 
ho(m_{\scriptscriptstyle ud})$ 

search plateau for  $\rho(m_{ud})$  at large  $\mu_I$ 

Need: • finer lattices • larger  $\mu_l$ •  $\lambda \rightarrow 0$  extrap. • finite V study



