

Semilinear wave model for critical collapse

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The story begins

Weak Cosmic Censorship

Critical phenomena in General Relativity (GR) [1]

What are the critical phenomena?

Universal

Self-similar behaviour

Power law behaviour near the threshold

[1] Phys. Rev. Lett. 70,

What do we want?

Structural requirements on the non linearities

The extent to which nearby threshold solutions might display critical behaviour

How do we do it?

Construct *simple* models that qualitatively capture this behaviour near threshold blow-up solutions

How to obtain models

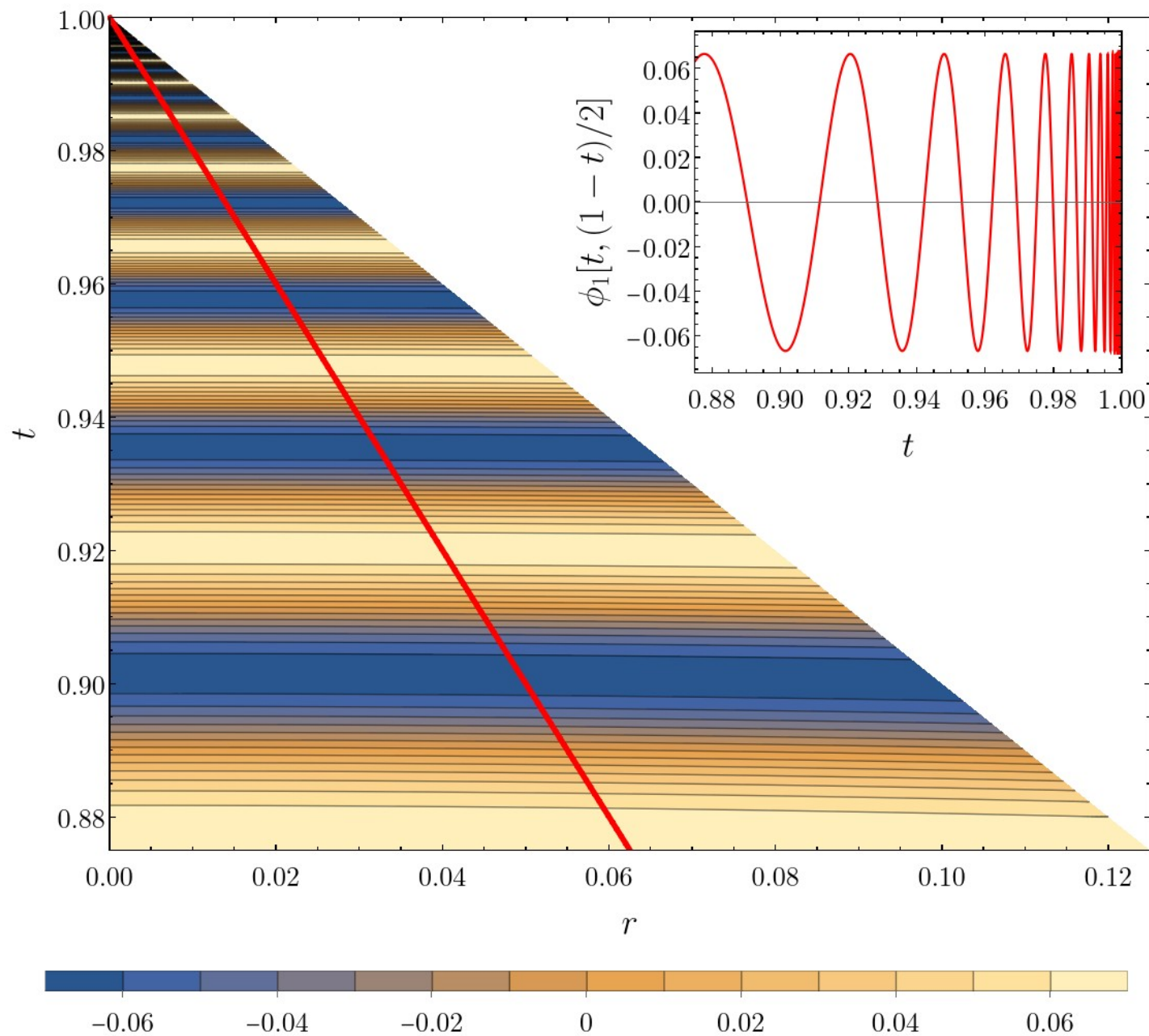
$$\square \varphi = 0 \qquad \phi = D[\varphi]$$

$$\square \phi - \frac{\partial_{\varphi}^2 D}{(\partial_{\varphi} D)^2} \nabla_a \phi \nabla^a \phi = 0$$

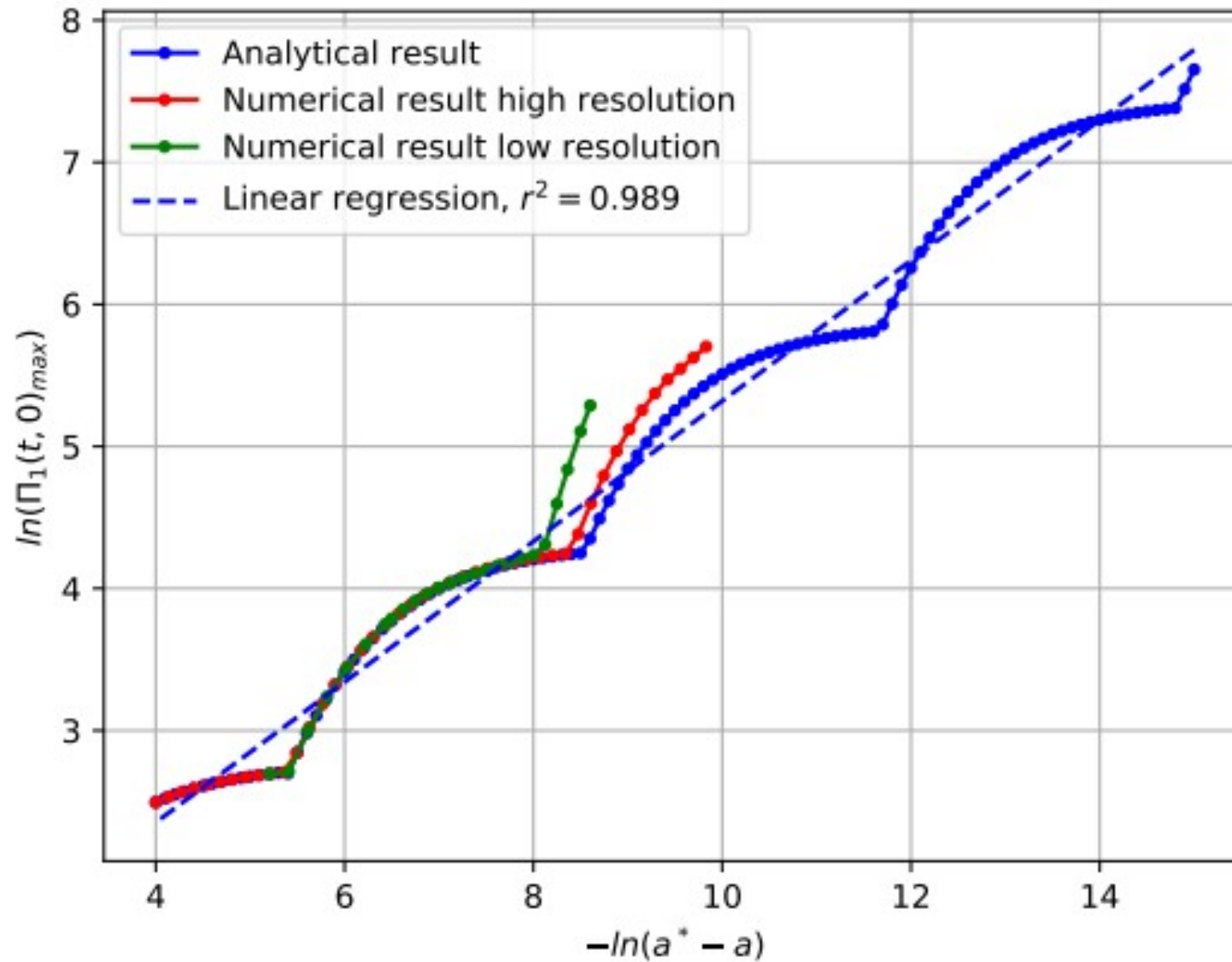
Model example

$$\square\phi + \frac{\phi - \sqrt{1 - \phi^2}}{1 - \phi^2} \nabla_a \phi \nabla^a \phi = 0$$

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What happens if we deviate from spherical symmetry?

$$\square\varphi = -\partial_t^2\varphi + \partial_r^2\varphi + \frac{2}{r}\partial_r\varphi + \Delta\varphi = 0$$

$$\varphi = \sum_{l=0}^{\infty} \sum_{m=-l}^l \varphi_{lm}(t, r) Y_{lm}(\theta^A)$$

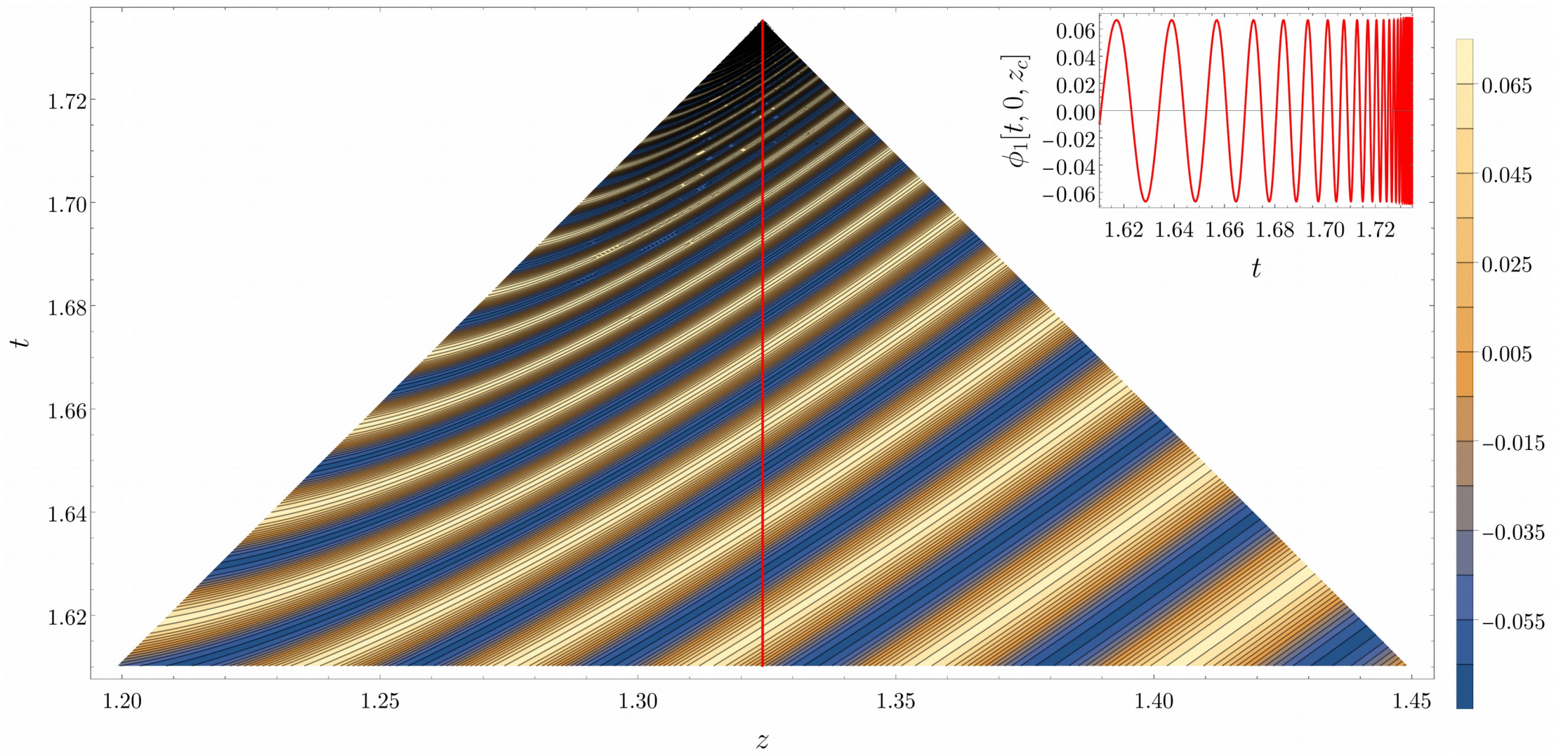
Partial waves

$$-\partial_t^2 \varphi_{lm} + \partial_r^2 \varphi_{lm} + \frac{2}{r} \varphi_{lm} - \frac{l(l+1)}{r^2} \varphi_{lm} = 0$$

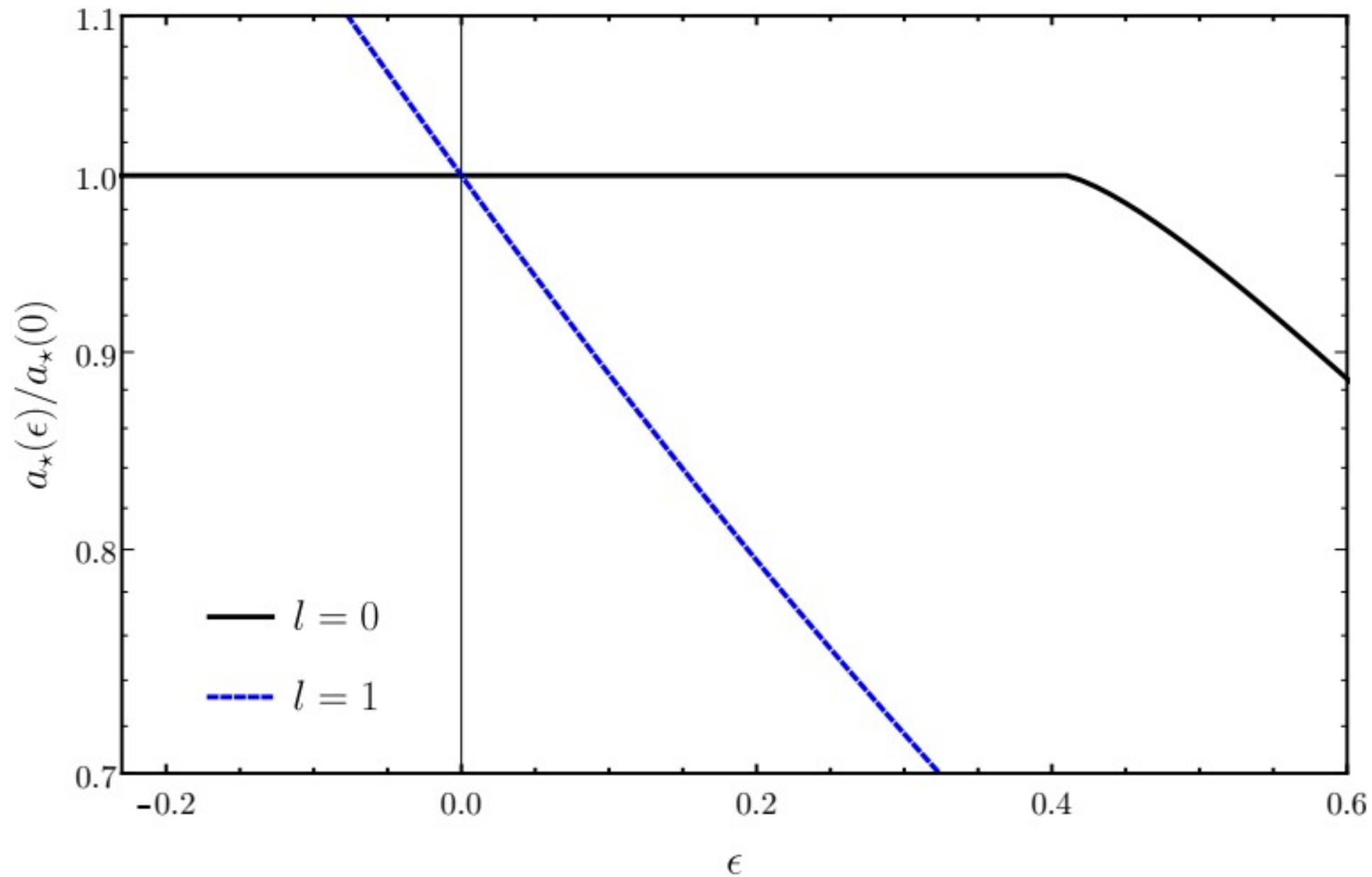
$$\sum_{k=0}^l \frac{(k+l)!}{2^k k! (l-k)!} \frac{1}{r^{k+1}} [F^{l-k}(u) - (-1)^{l-k} F^{l-k}(v)]$$

$$u = t - r$$
$$v = t + r$$

$$l = 1, m = 0, \varphi_{10}$$



$$\phi_a(t, r) = D[a \varphi_\star(t, r)] \quad \tilde{\phi}_a = D[a \varphi_\star + \epsilon \tilde{\varphi}]$$



Conclusions

First known simple model that shows critical phenomena (including DSS)

Spherical symmetry: Model that captures all the characteristics of critical phenomena

Code: We have a tool to study nonlinear models with non analytic solution

Aspherical deformations: Away from spherical symmetry we would lose universality

Predictions for GR, that are compatible up to now (Work in progress)