



LABORATÓRIO DE INSTRUMENTAÇÃO
E FÍSICA EXPERIMENTAL DE PARTÍCULAS
partículas e tecnologia

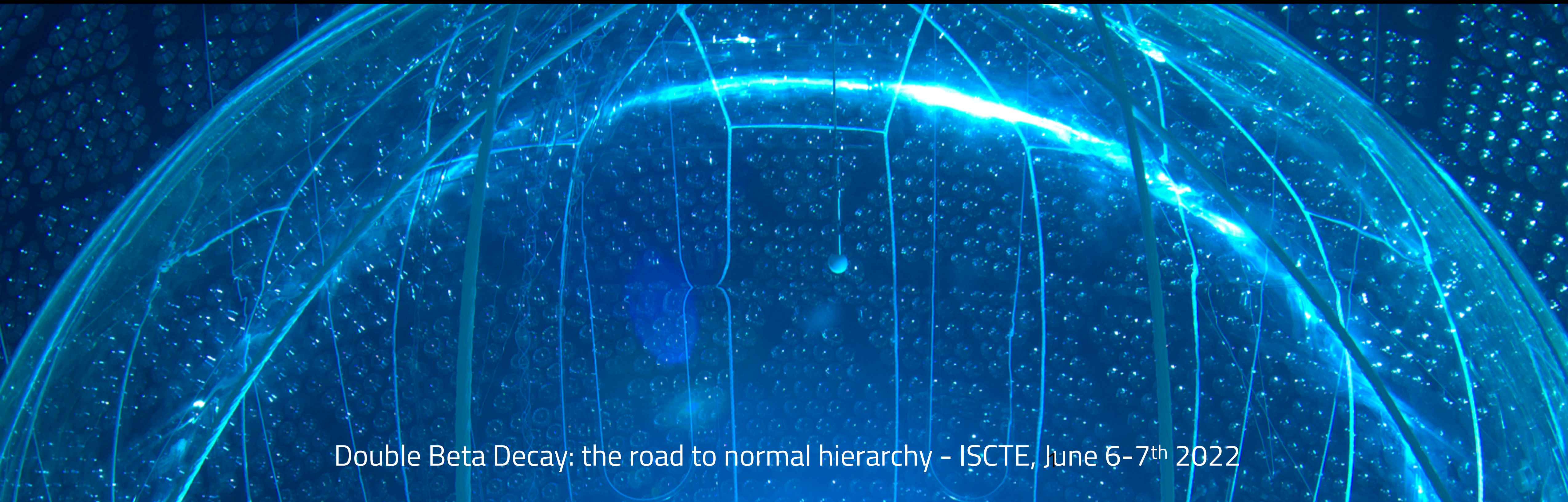


UNIVERSIDADE D
COIMBRA

The SNO+ Experiment



Nuno Barros, LIP/UC



Double Beta Decay: the road to normal hierarchy - ISCTE, June 6-7th 2022

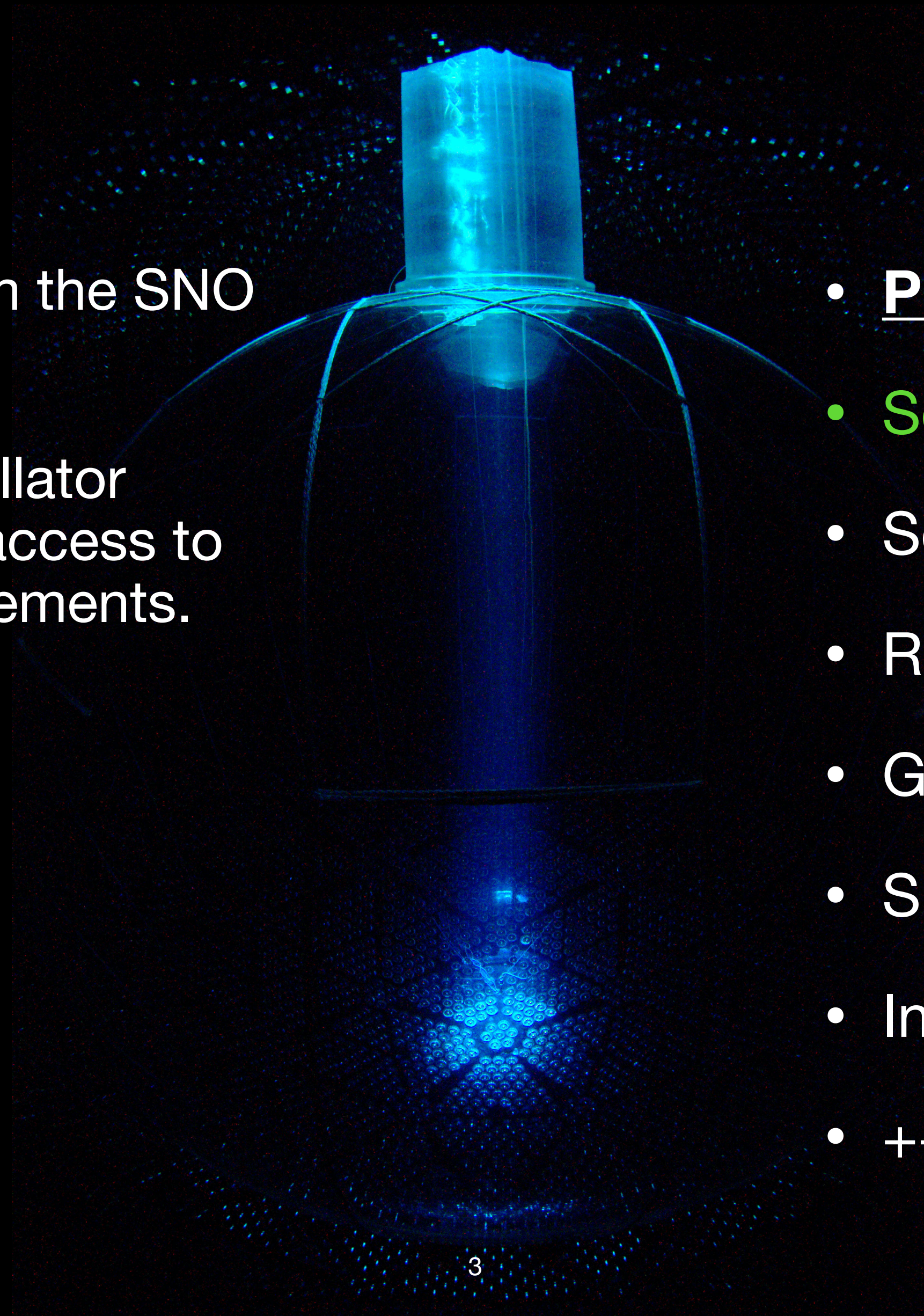
The SNO+ Collaboration



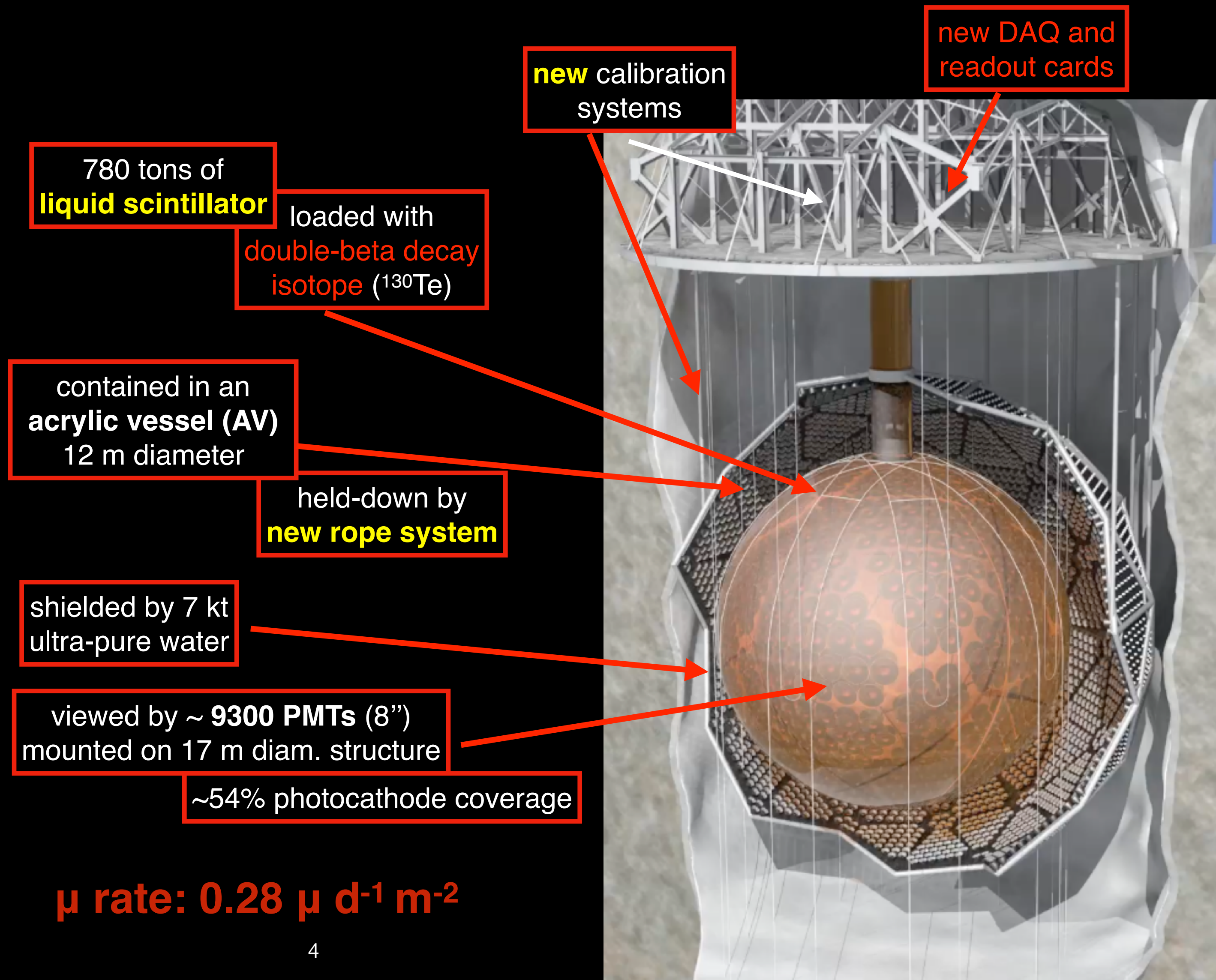
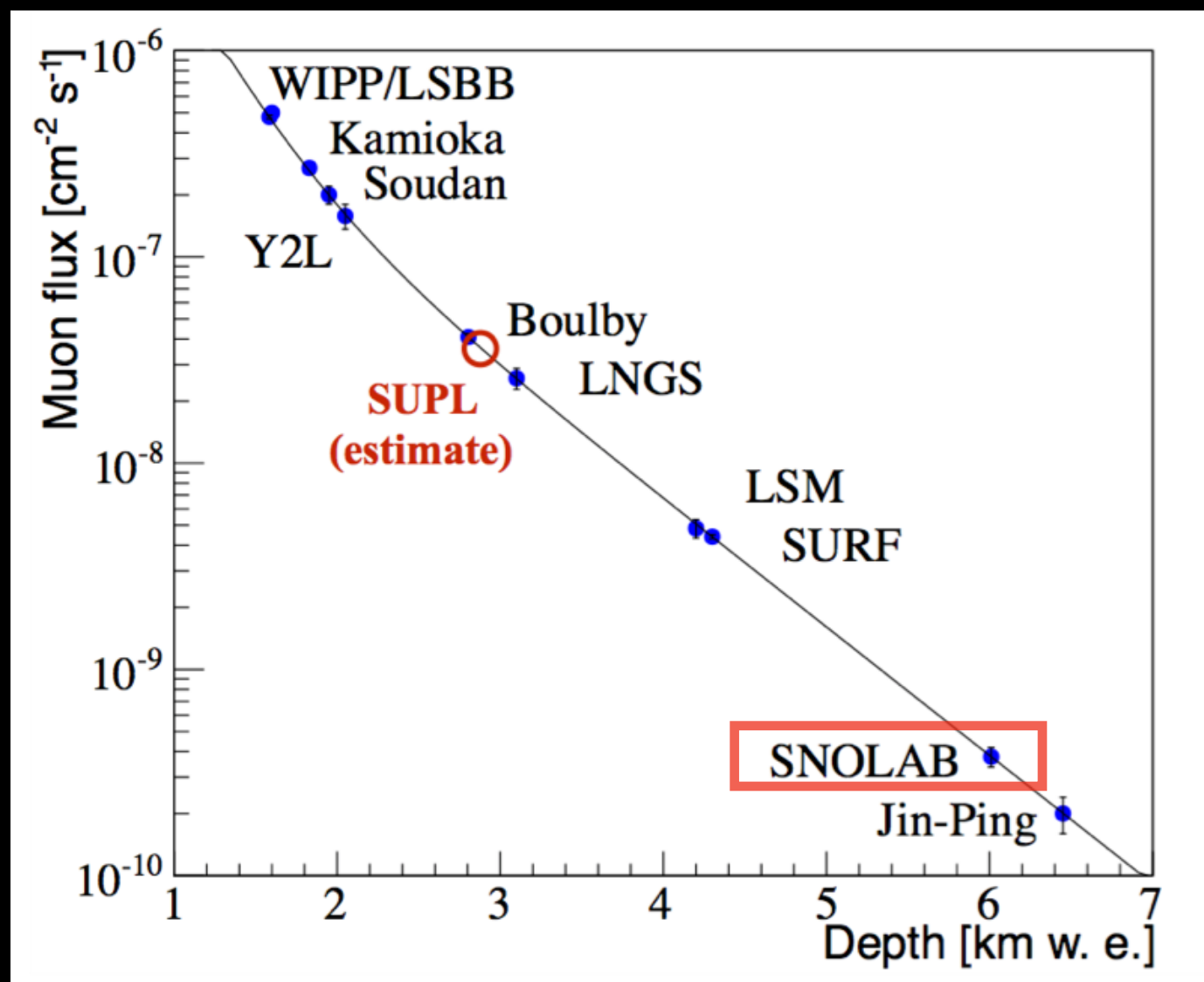
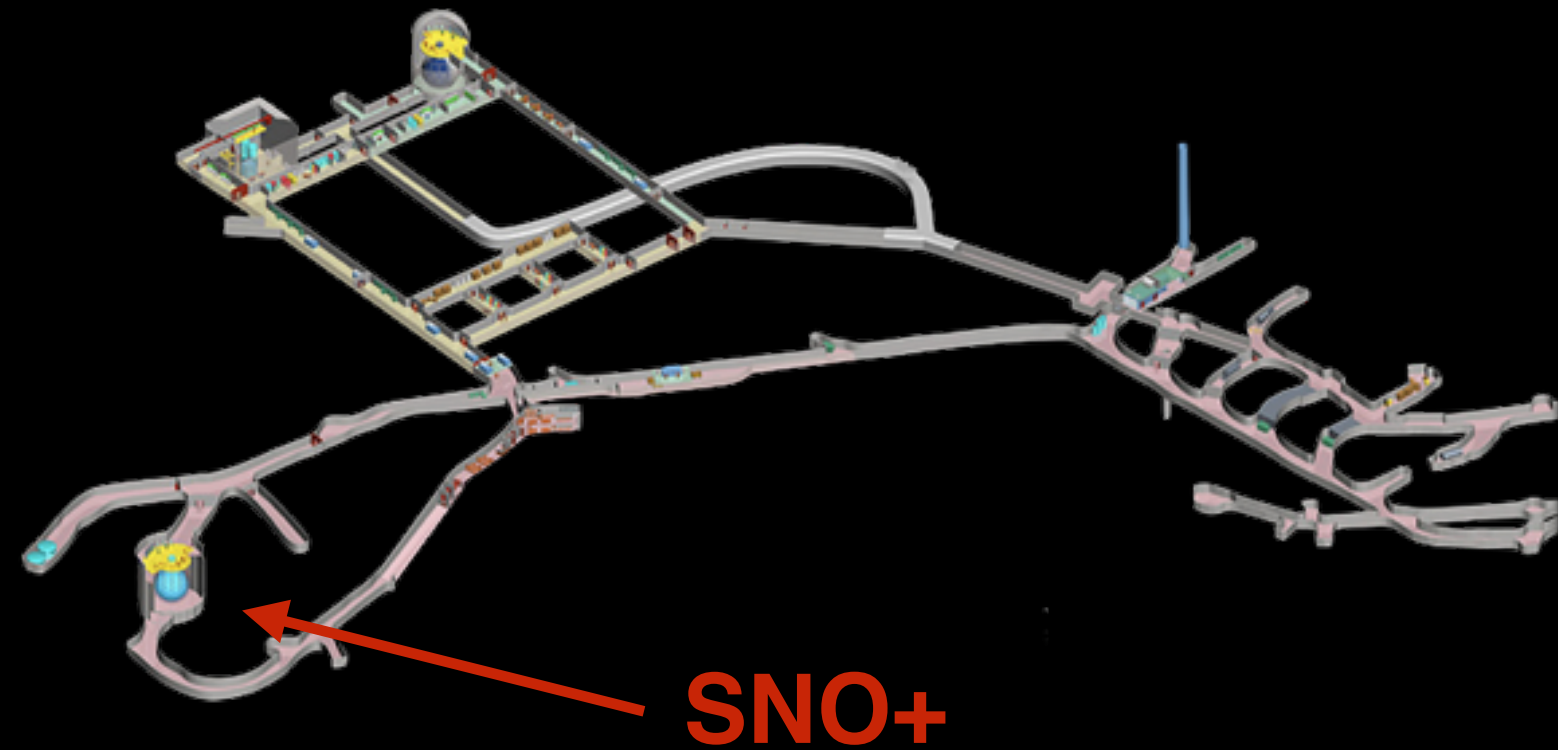
A Multi-purpose Liquid Scintillator Detector

- Infrastructure inherited from the SNO experiment.
- Adapted from D₂O to scintillator detecting medium to give access to low energy (~MeV) measurements.

- Physics Programme:
- Search for $0\nu\beta\beta$ in ^{130}Te
- Solar Neutrinos
- Reactor Anti-neutrinos
- Geo-neutrinos
- Supernova bursts
- Invisible nucleon decay
- +++



The SNO+ Detector



SNO+ Upgrades

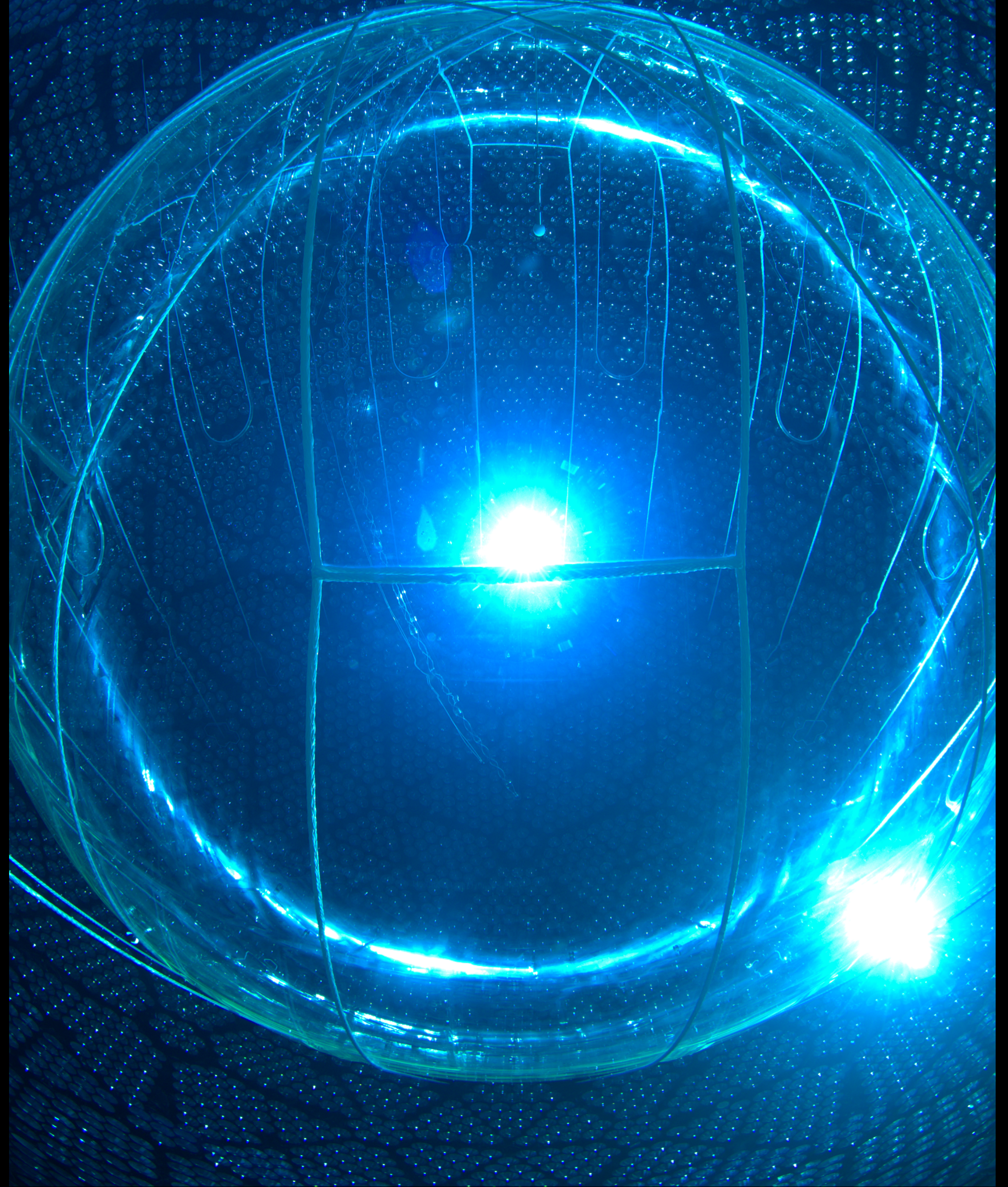
Besides ropes and purification plant

- Refurbishing the electronics
- Repair of many “dead” PMT bases
- All-new DAQ, adapted to the higher trigger rate
- New cover gas system
- New calibration systems capable of deploying in LAB scintillator
- New in-situ injected LED/laser light calibration system
- Calibration system cameras (for photogrammetry)

SNO+ Timeline

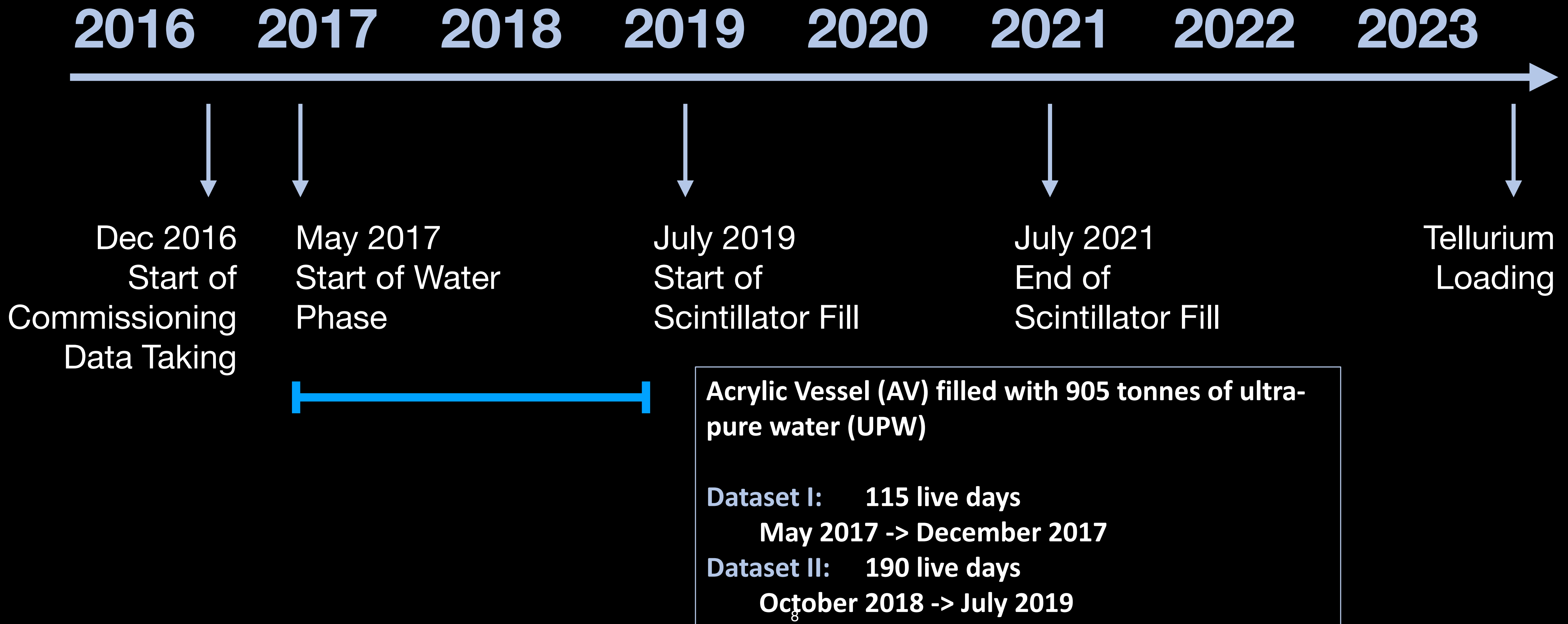


SNO+ Water Phase



SNO+ Timeline

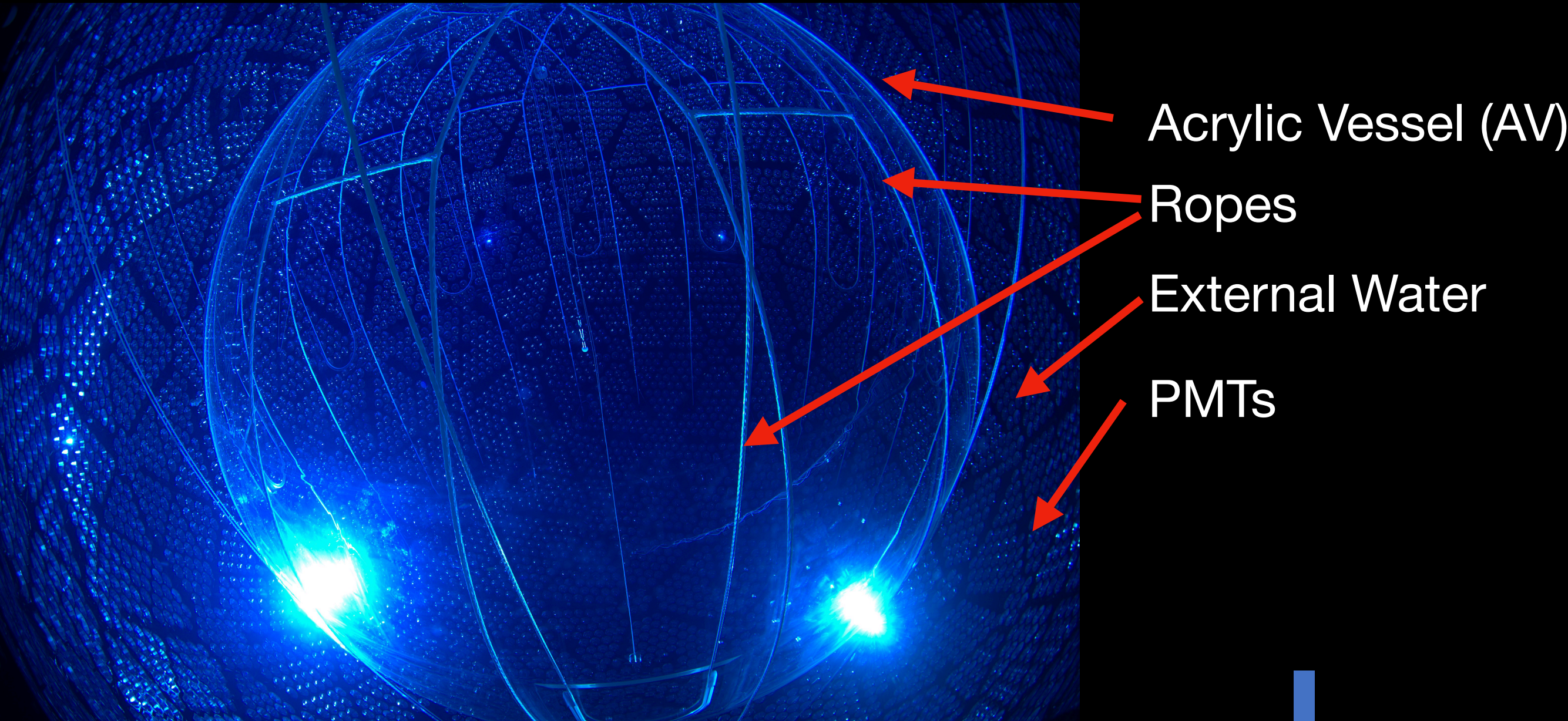
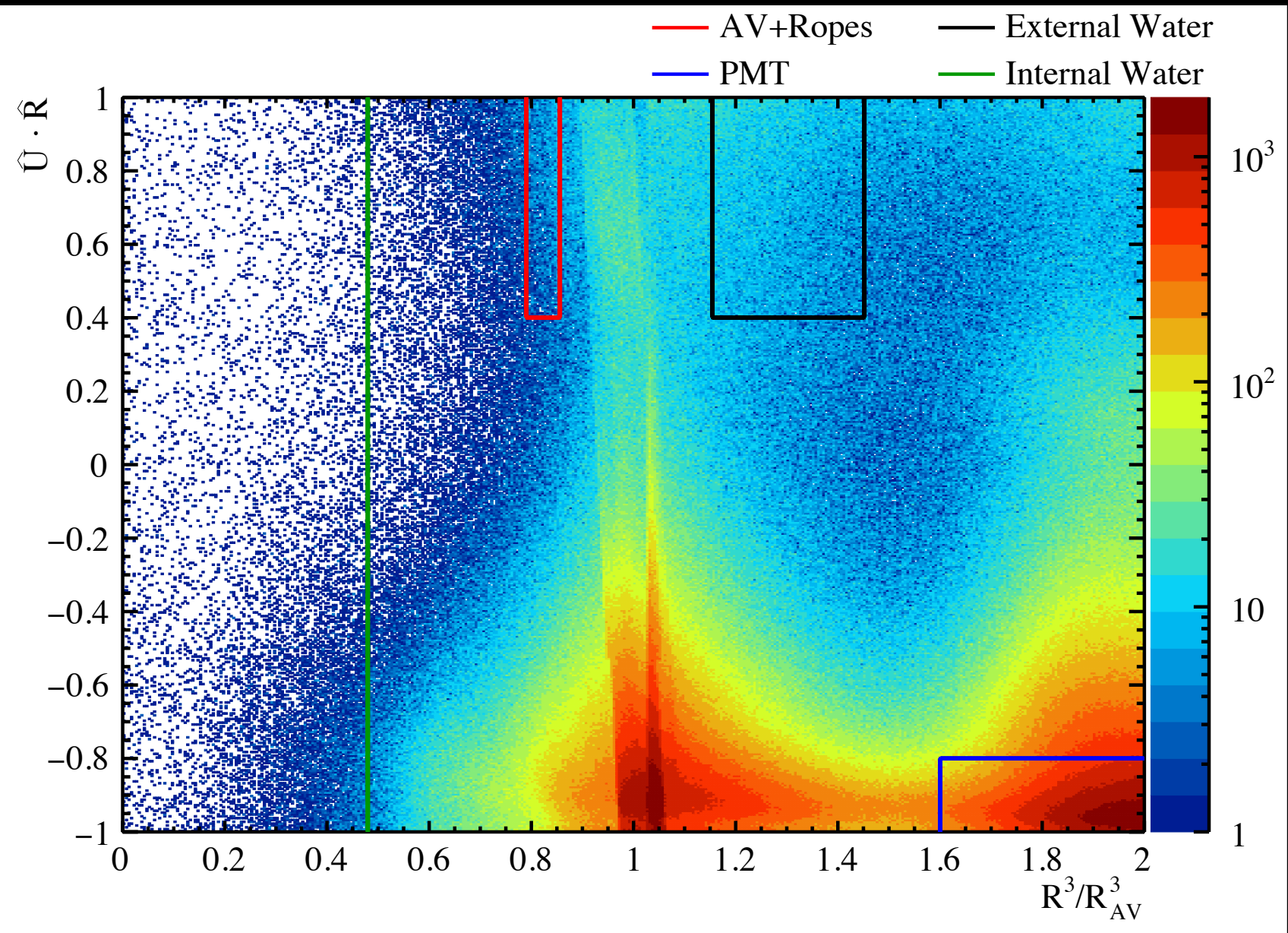
Water Phase



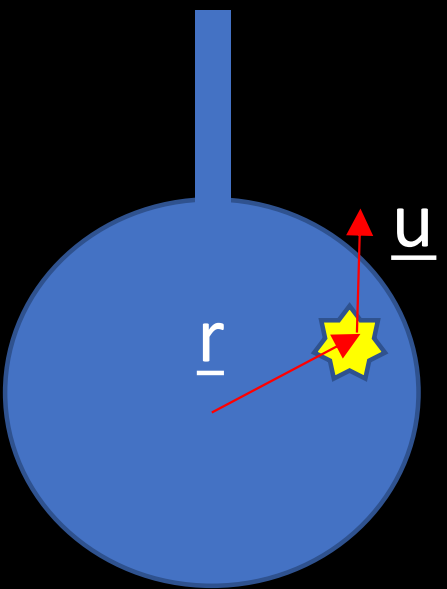
Water Phase Backgrounds

External backgrounds measurement

Simple detector configuration.
Measure components that don't change with detection medium.



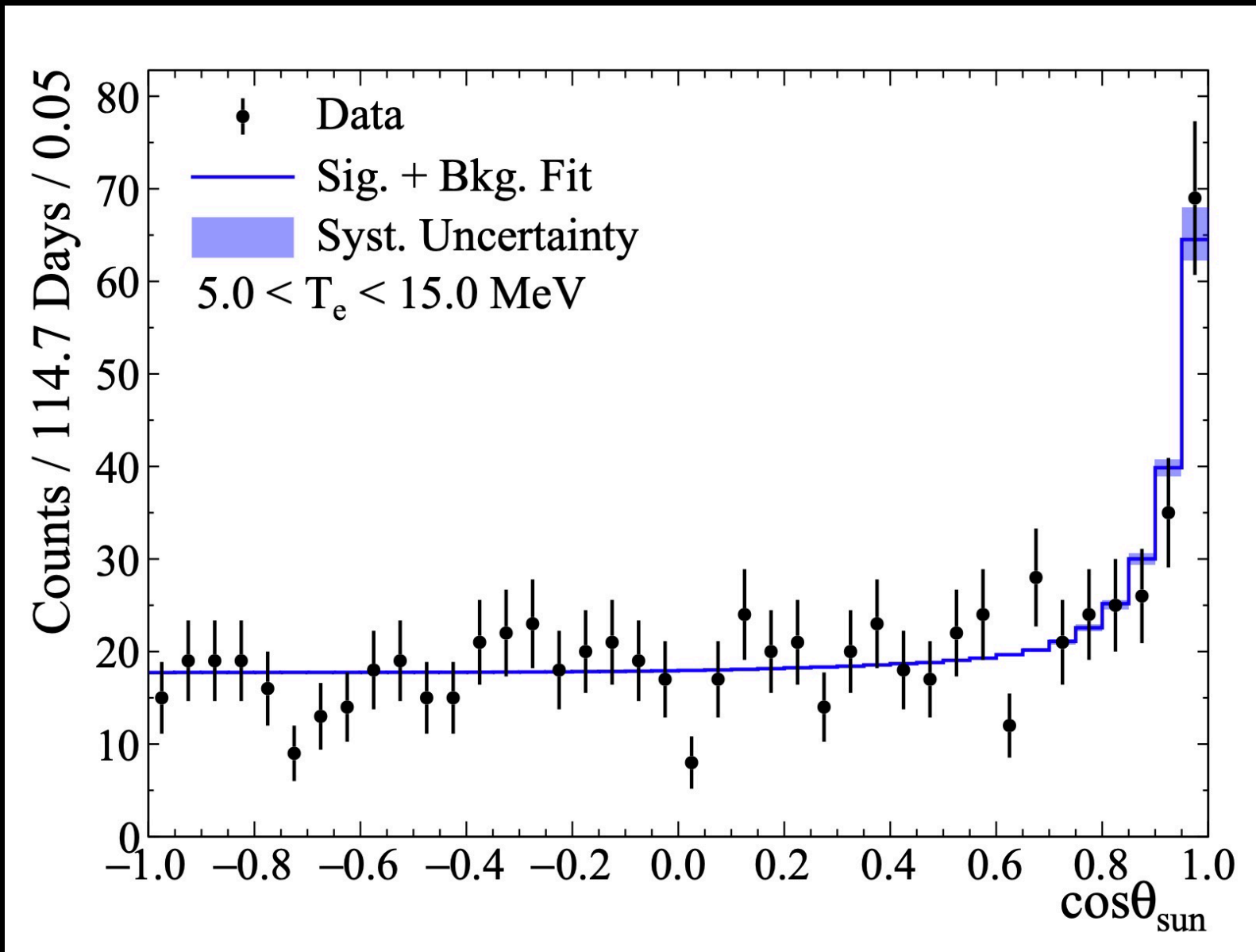
Background	Rate (Fraction of Nominal)
AV+Ropes	$0.52 \pm 0.02^{+0.39}_{-0.28}$
External Water	$0.03 \pm 0.01^{+0.61}_{-0.03}$
PMT	$2.04 \pm 0.04^{+3.69}_{-1.20}$



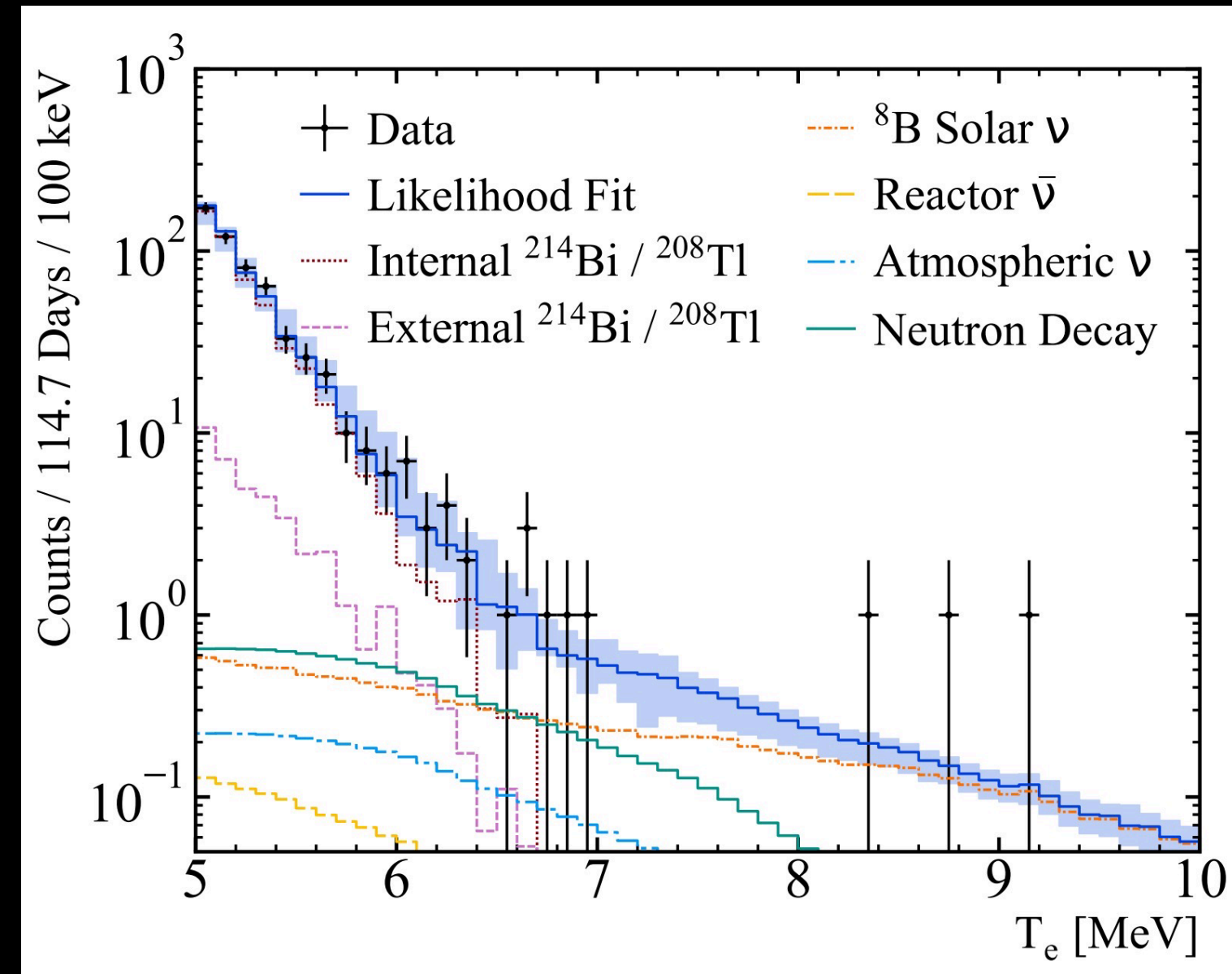
Box analysis performed
Fitted scale factors reduce $0\nu\beta\beta$ ROI backgrounds
from 1.21 to 0.93 events / year
ie >20% below goals

Water Phase Physics Results

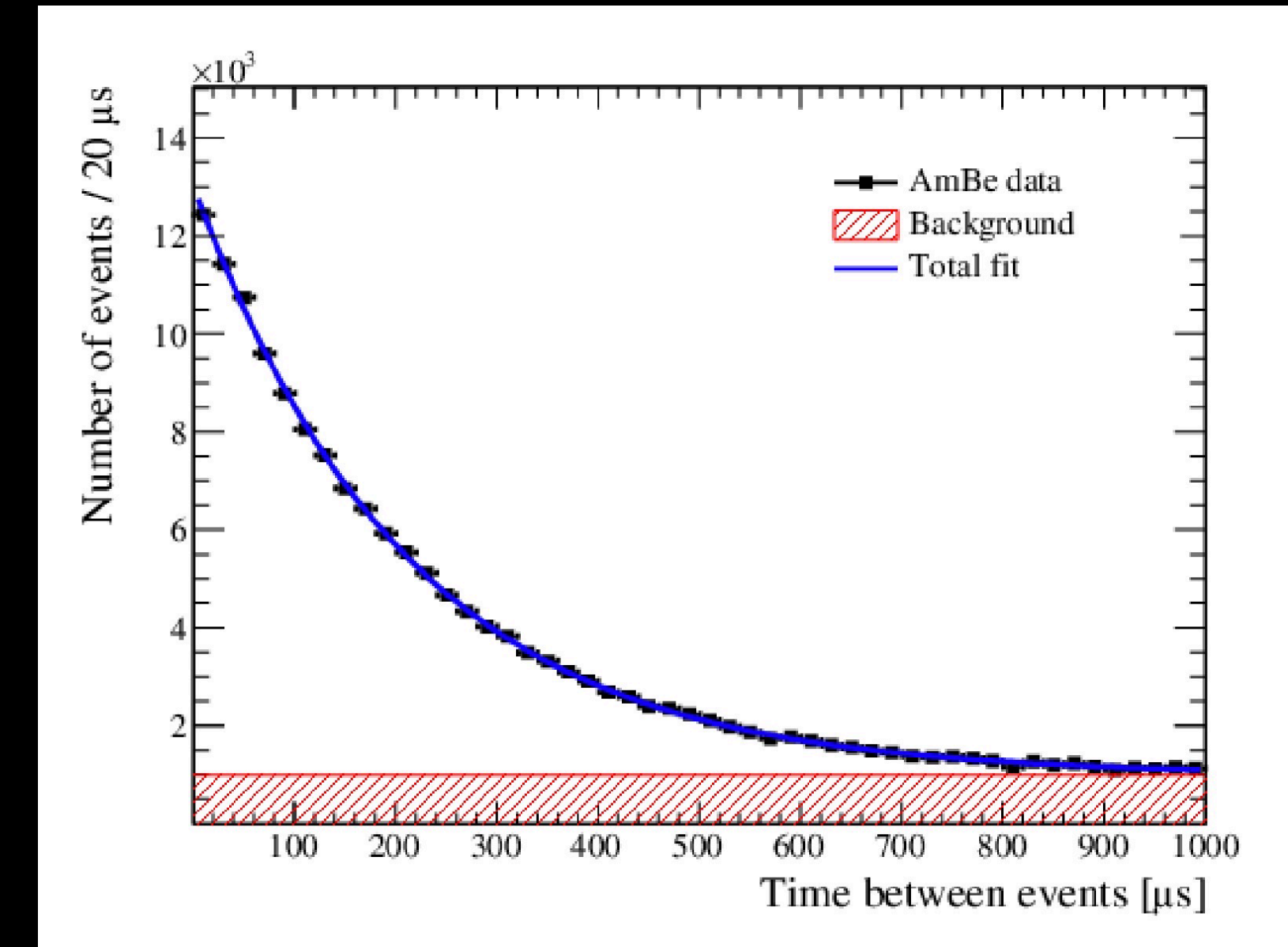
Dataset I (115 days)



^8B Solar neutrino flux
Phys. Rev. D 99, 012012 (2019)



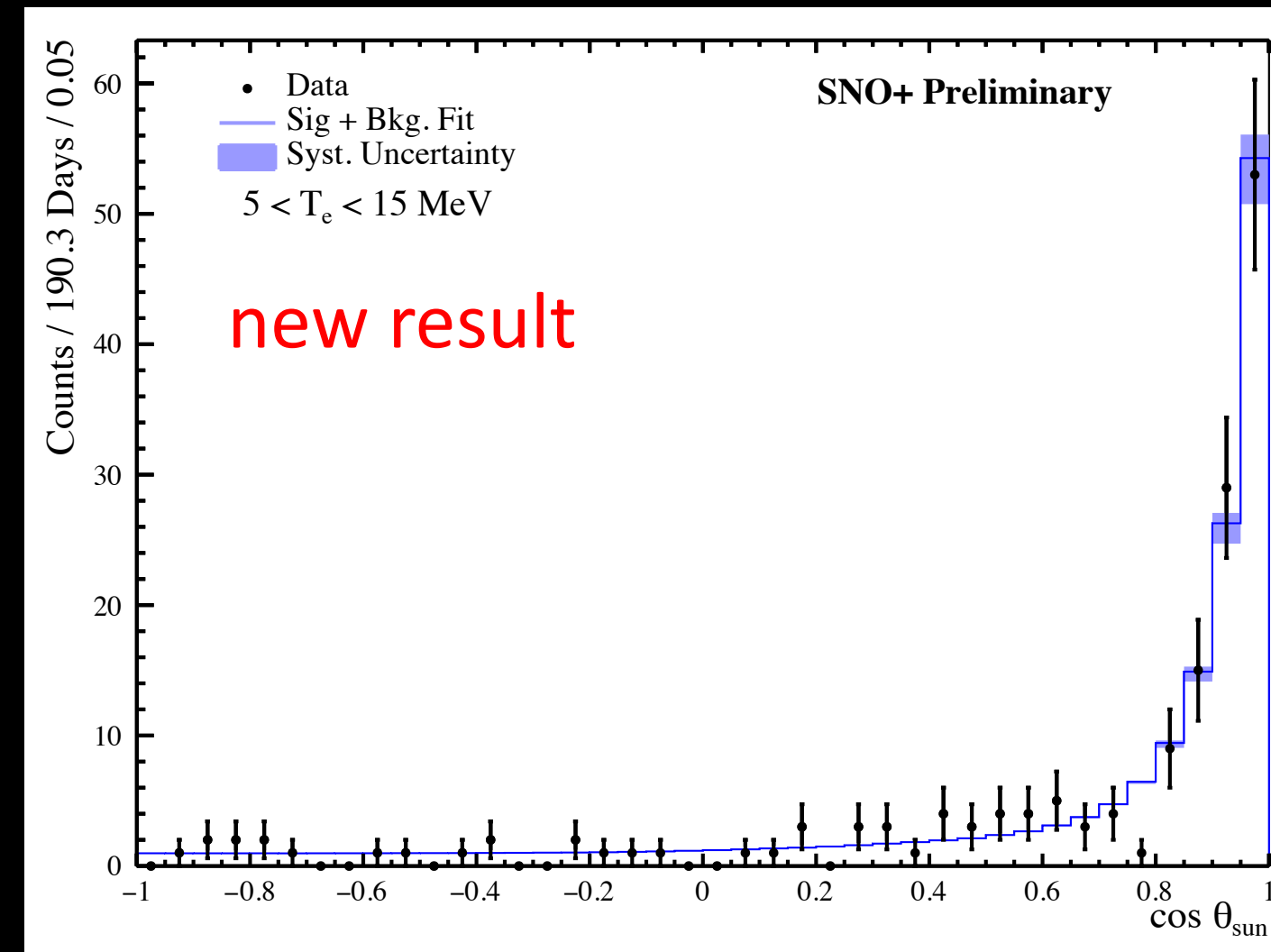
New limits for p, pp and pn
invisible nucleon decay modes
Phys. Rev. D 99, 032008 (2019)



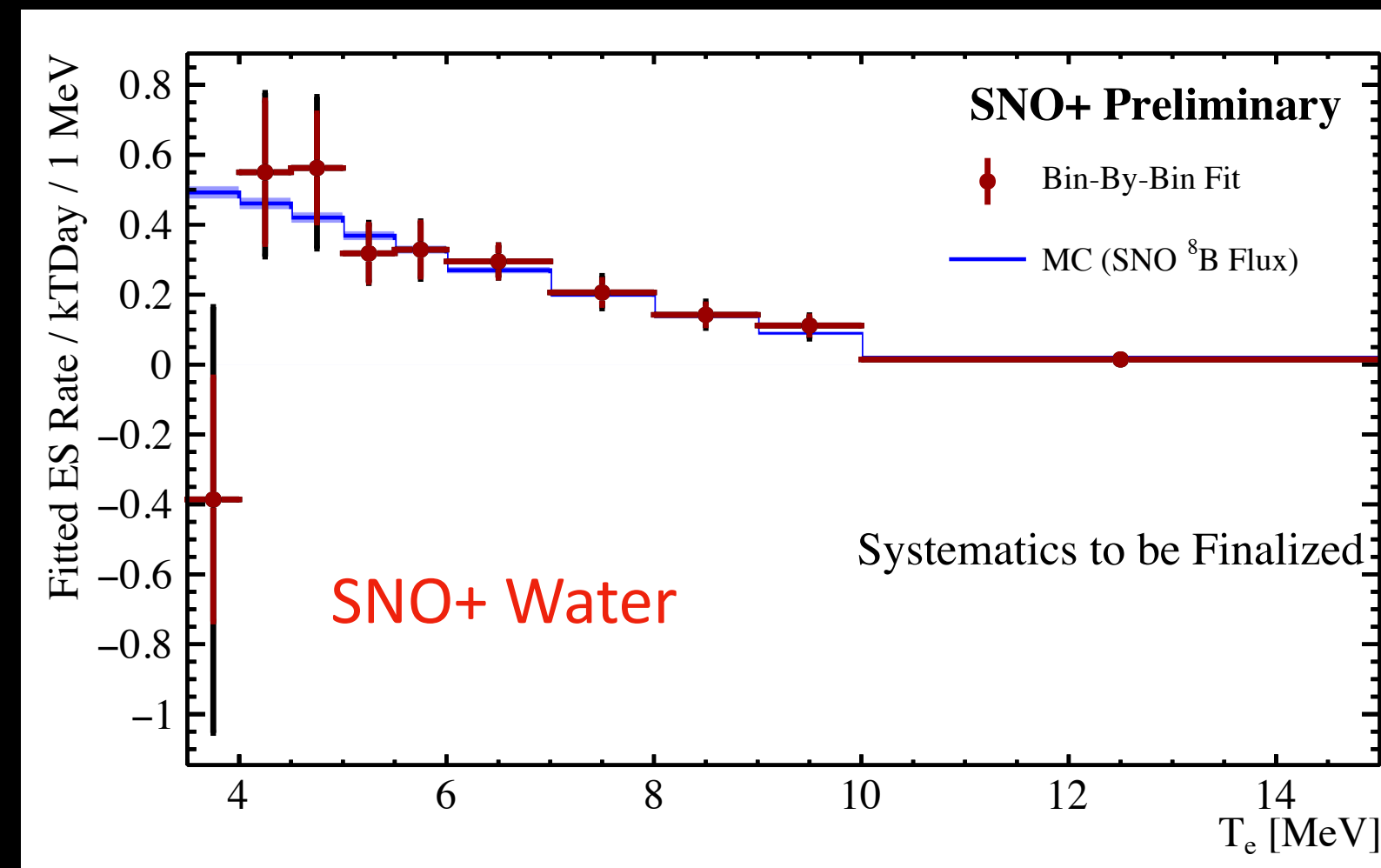
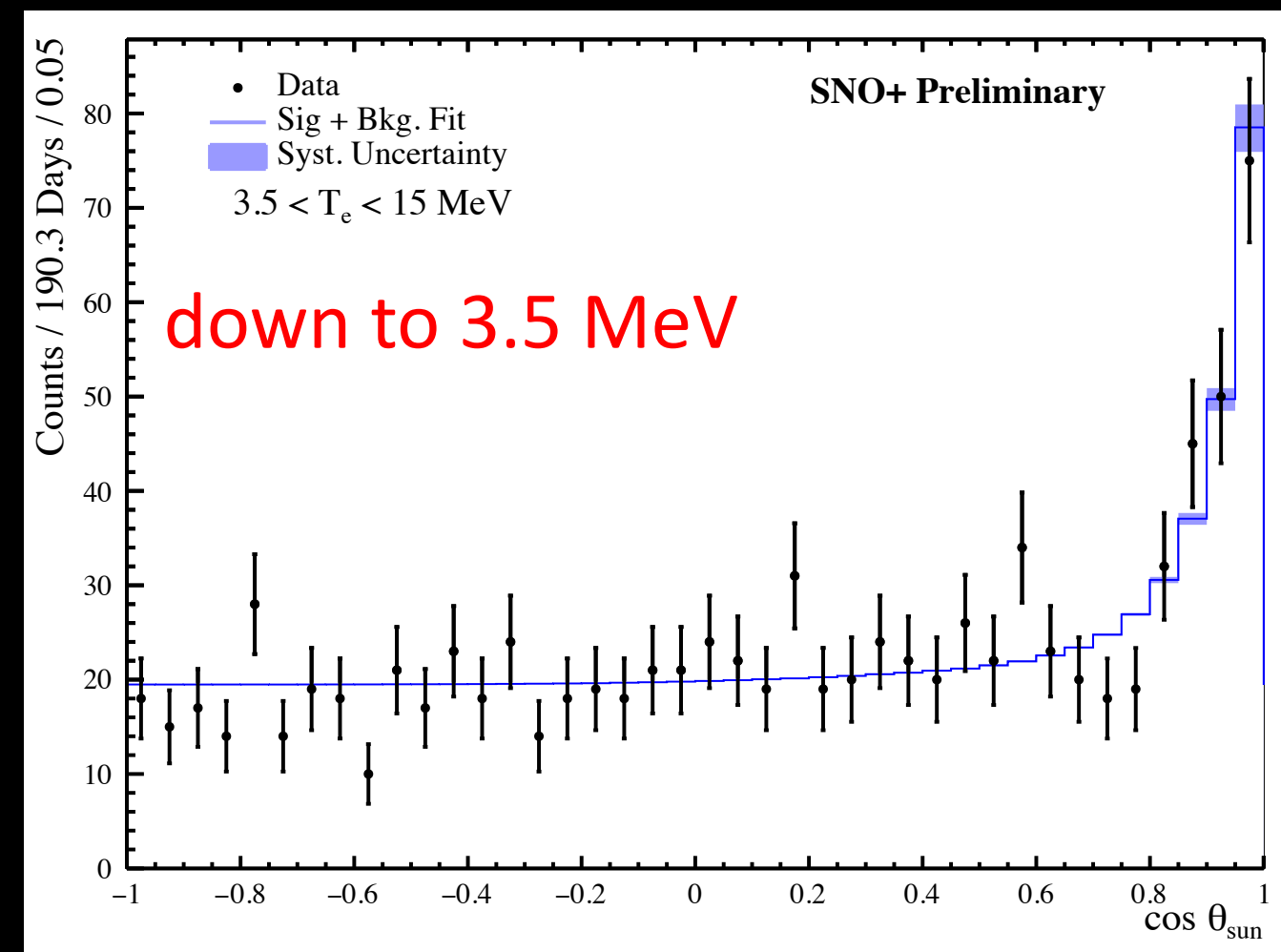
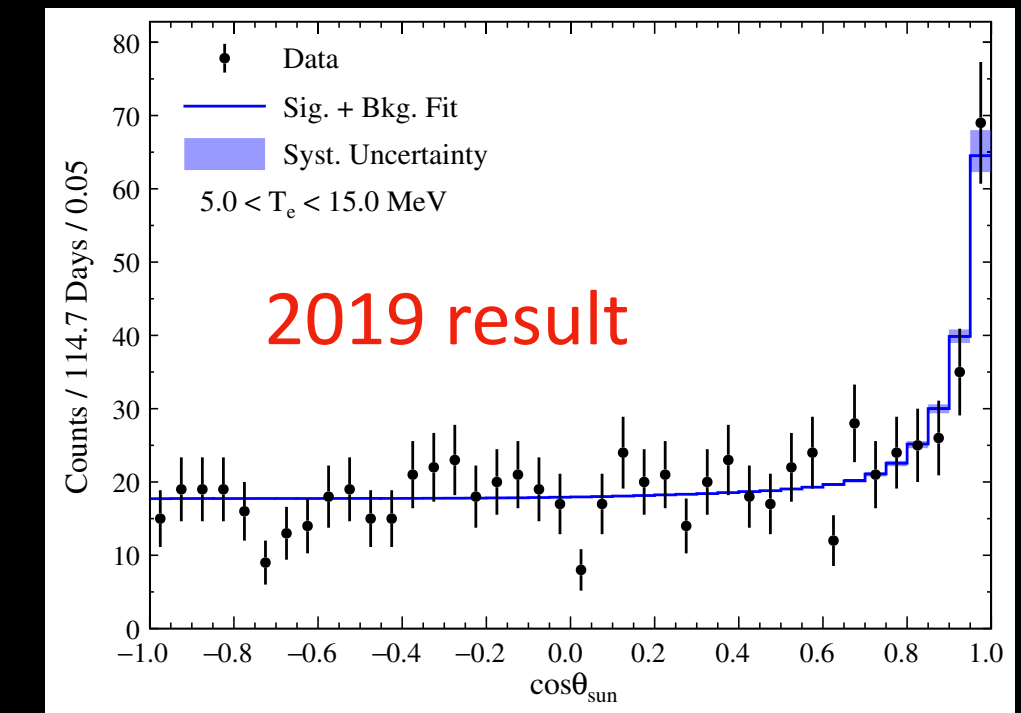
~50% efficiency for triggering on
a neutron in pure water
Phys. Rev. C 102, 014002 (2020)

Water Phase Physics Results

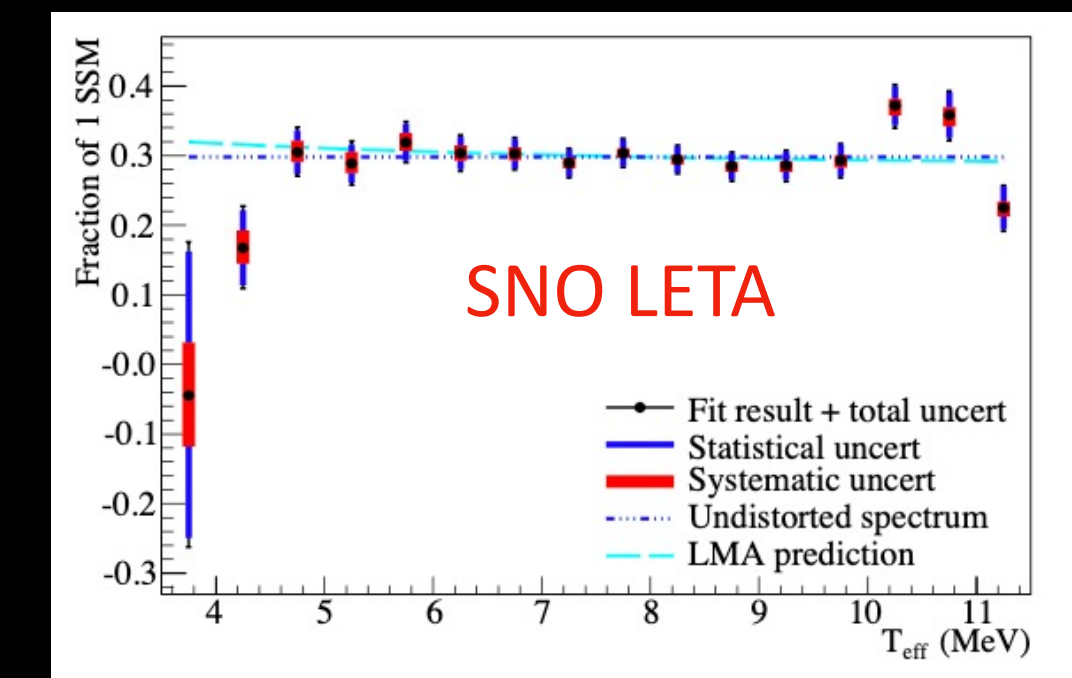
Updated Solar ν results - Dataset II (~190 days)



← Same threshold. →



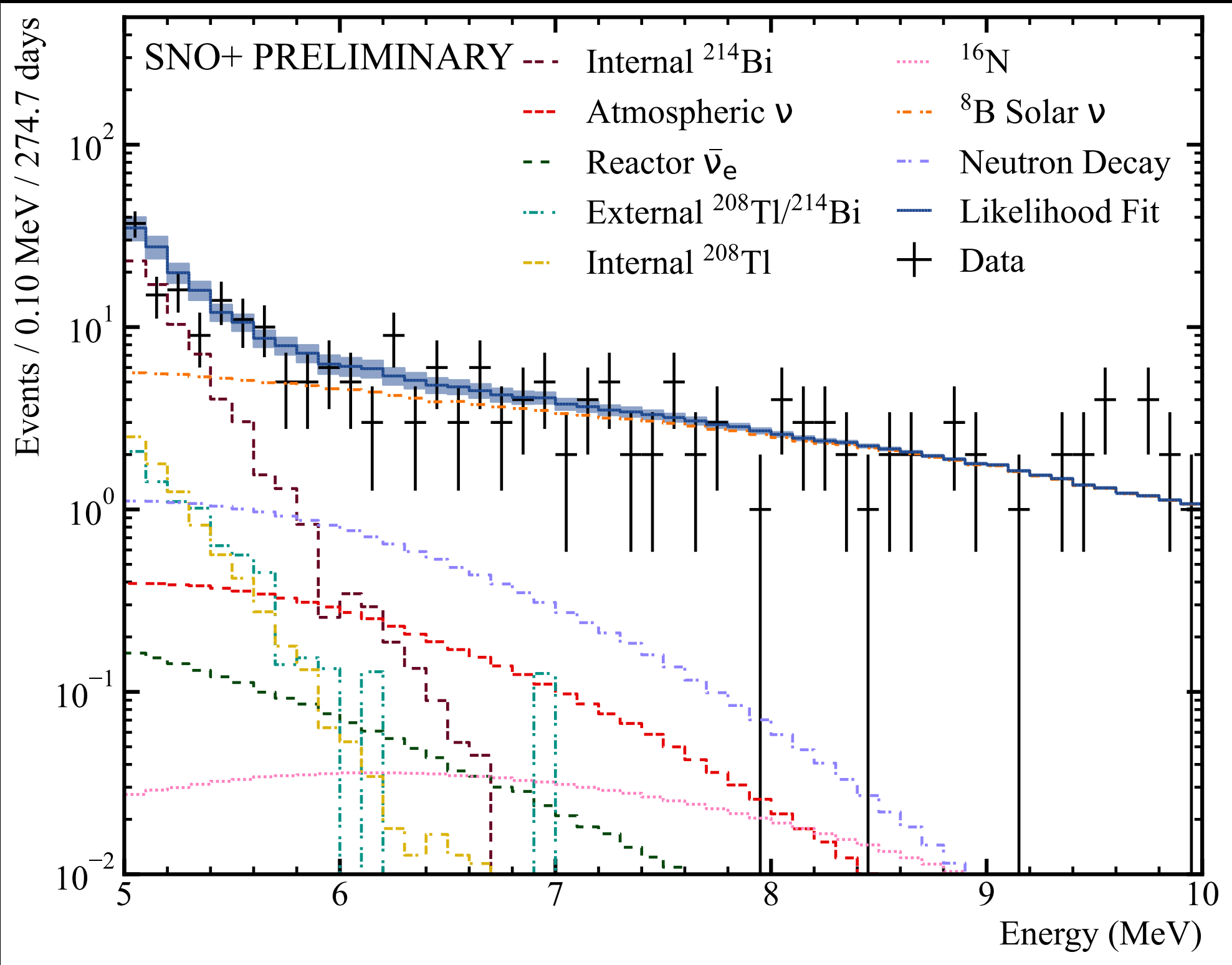
← Same threshold. →



Water Phase Physics Results

Updated Nucleon Decay results - Dataset II (~190 days)

- World's best limits on invisible modes of nucleon decay
- Recently updated in [arXiv:2205.06400](https://arxiv.org/abs/2205.06400)



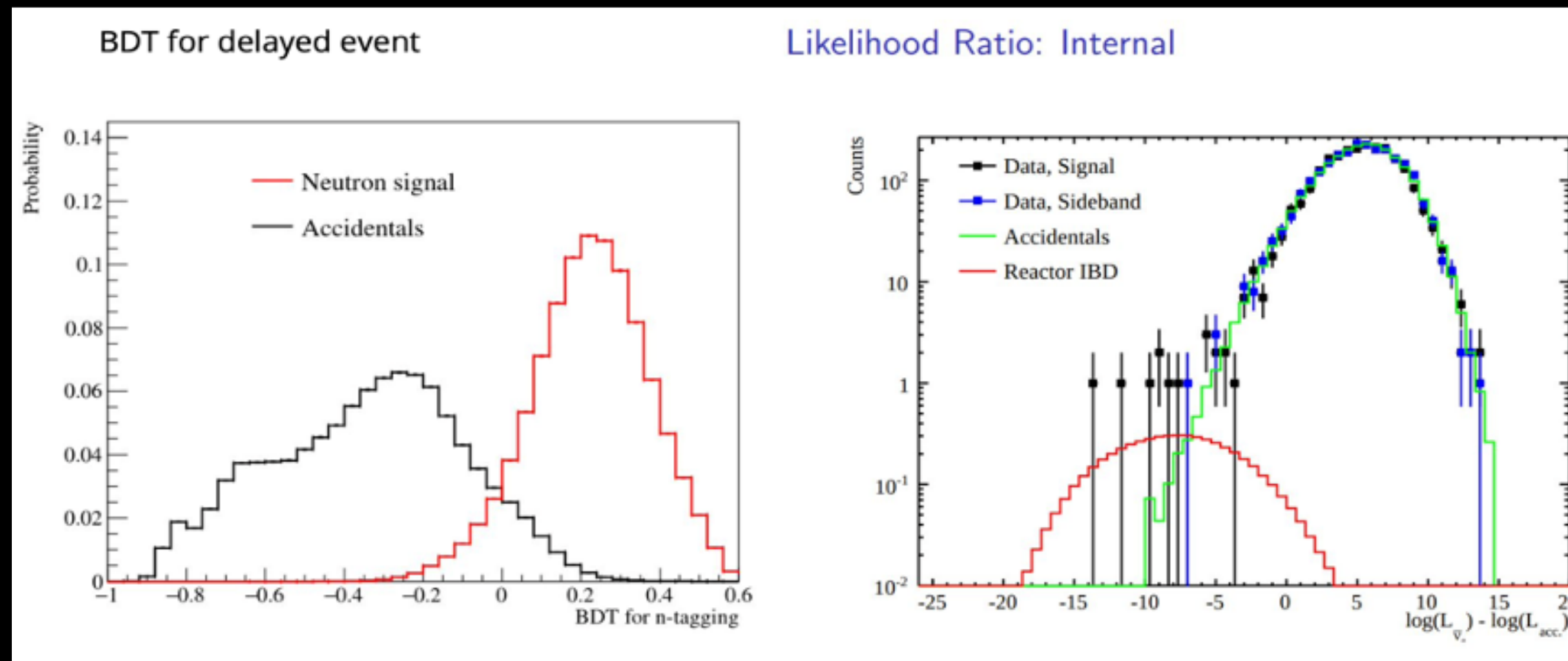
Decay Mode	Partial Lifetime Limit	Existing Limits
n	$9.0 \times 10^{29} \text{ y}$	$5.8 \times 10^{29} \text{ y}$ [5]
p	$9.6 \times 10^{29} \text{ y}$	$3.6 \times 10^{29} \text{ y}$ [6]
pp	$1.1 \times 10^{29} \text{ y}$	$4.7 \times 10^{28} \text{ y}$ [6]
np	$6.0 \times 10^{28} \text{ y}$	$2.6 \times 10^{28} \text{ y}$ [6]
nn	$1.5 \times 10^{28} \text{ y}$	$1.4 \times 10^{30} \text{ y}$ [5]

Improvement of a factor of 3 over existing limits

Water Phase Physics Results

Antineutrino Observation

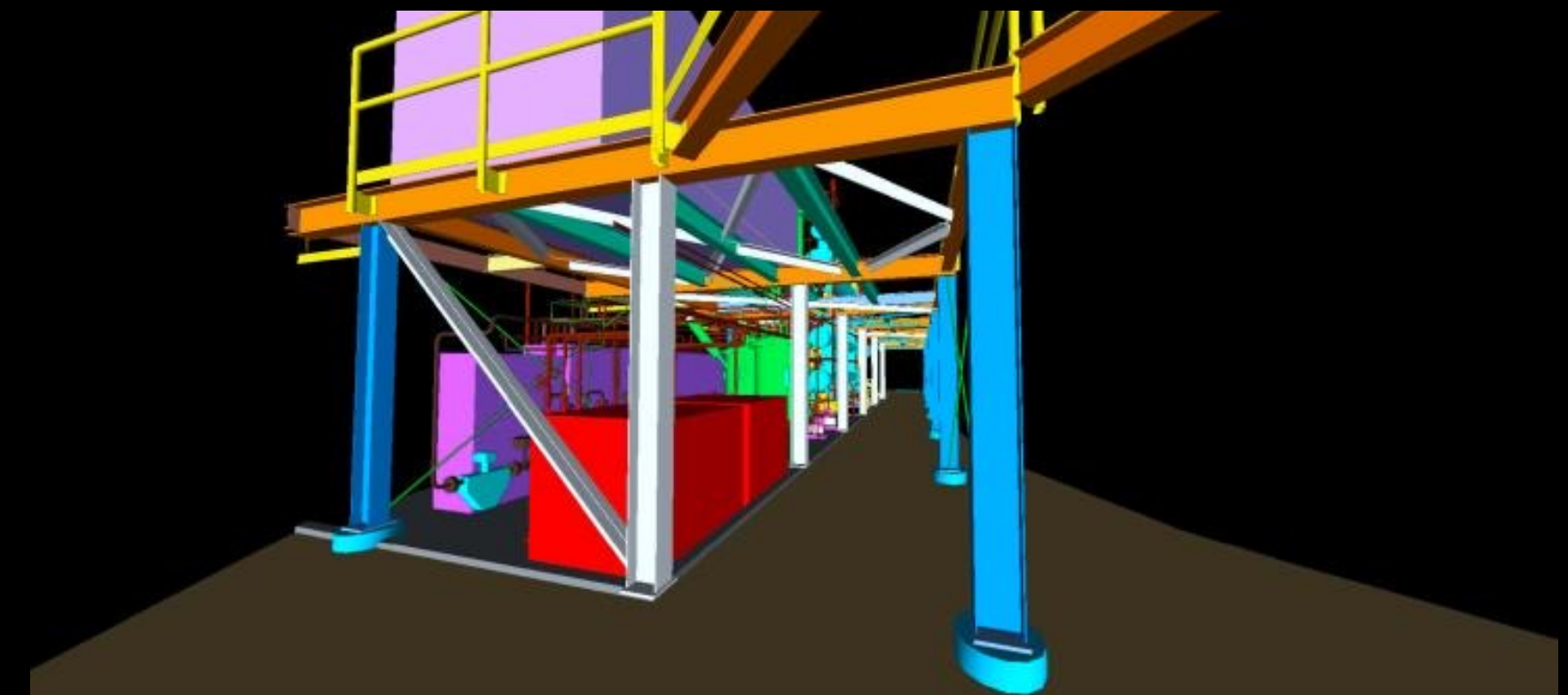
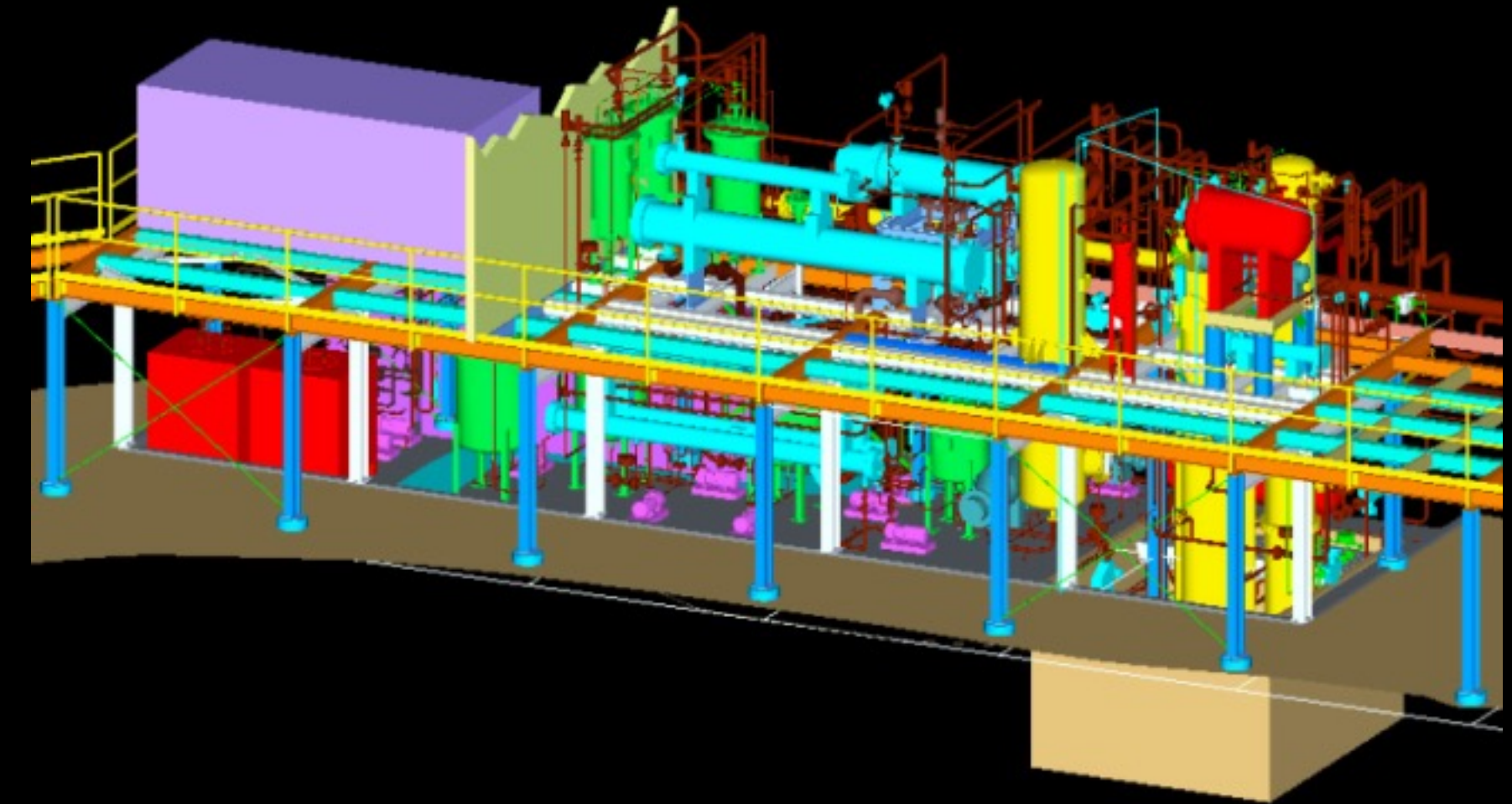
- First observation of reactor $\bar{\nu} + p \rightarrow e^+ + n$ events using pure water (undoped)
- publication being prepared; detection of 9 and 10 events in two distinct analyses (BDT and likelihood) with >3 sigma significance
- made possible by $\sim 50\%$ neutron detection efficiency (highest in a water Cherenkov detector)



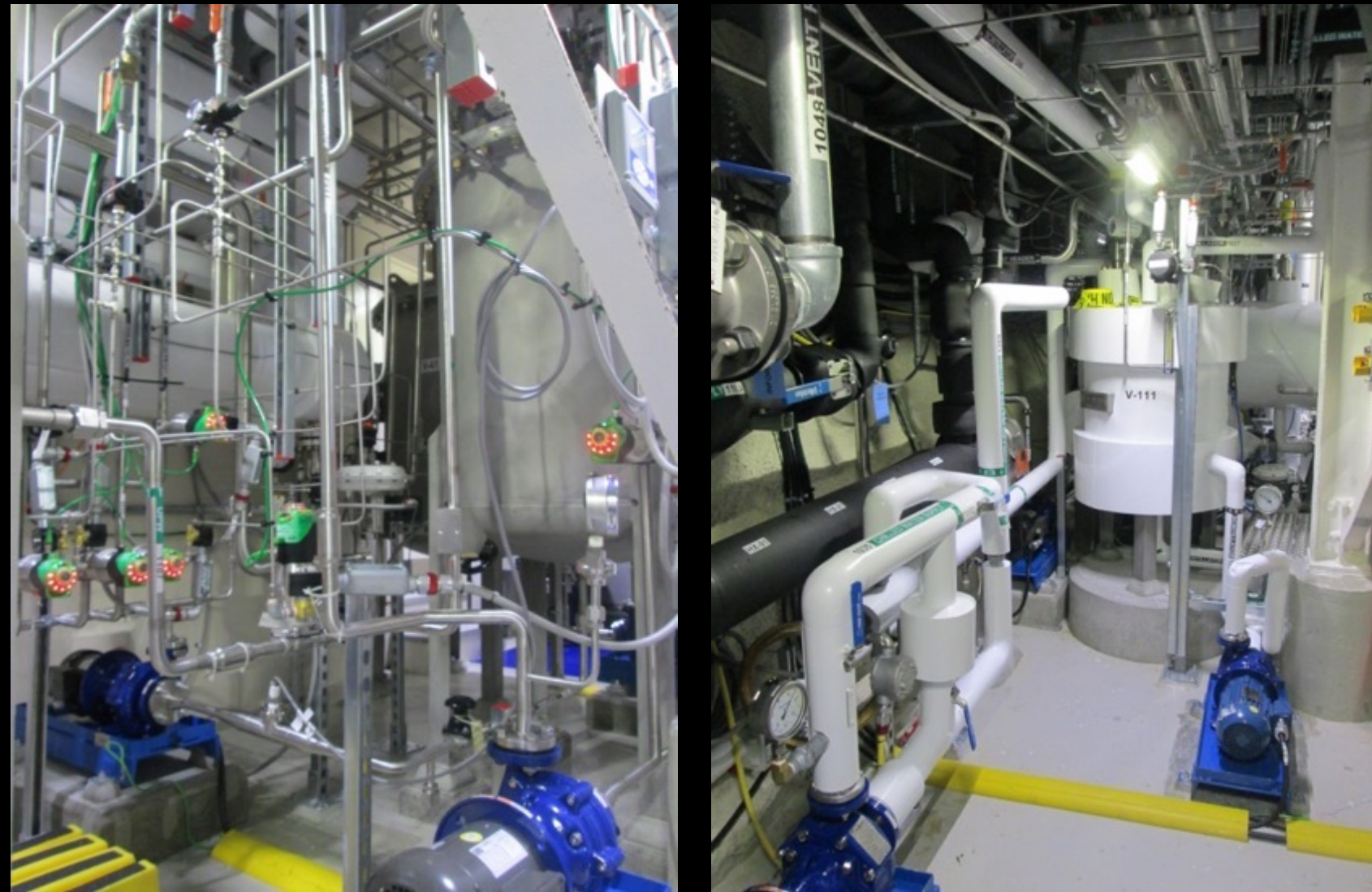
this result does not challenge our understanding of Δm_{21}^2 ; but does draw attention to upcoming SNO+ measurements with full LS that will

SNO+ Scintillator Purification Plant

- reinforced mezzanine steel
- installed columns, vessels, heat exchangers, tank, pumps, valves, high-grade sanitary piping (orbital- welded, electropolished 2012 stainless steel tubing)
- utility plumbing (cooling water, compressed air, vent, boil-off nitrogen)
- process control, wiring, instrumentation, electrical
- firewalls, fire detection and suppression



Scintillator Filling



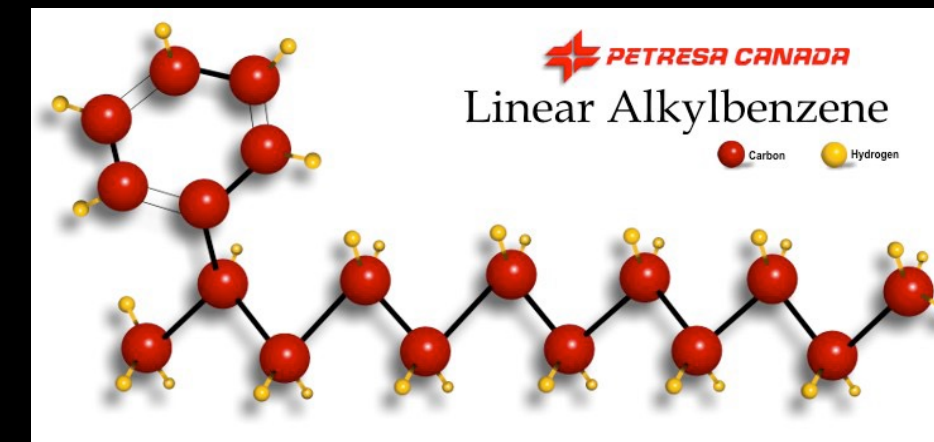
Purification and filling systems underground

LAB, Master Solution, and final scintillator assessed for quality hourly during purification plant operation and detector filling

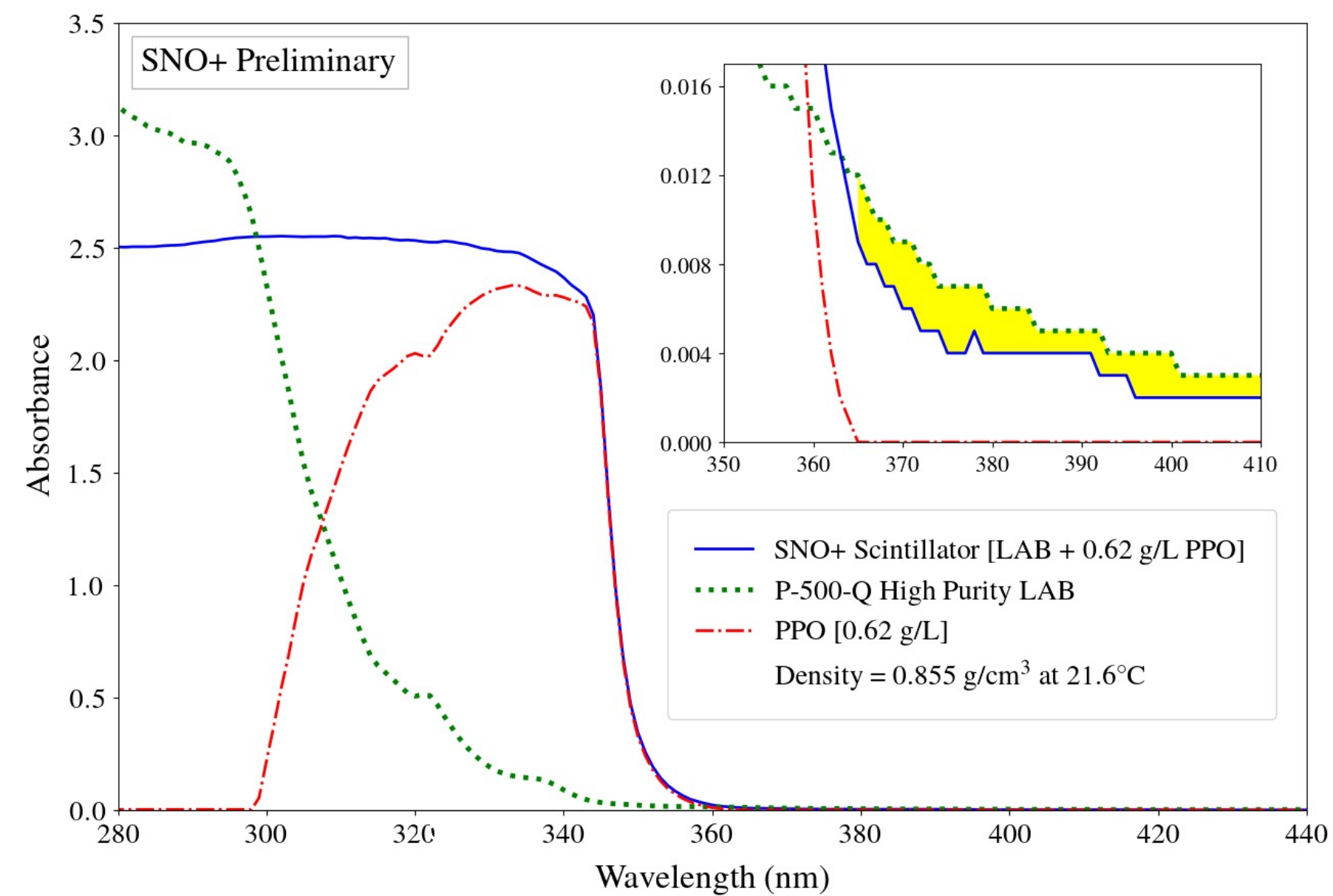
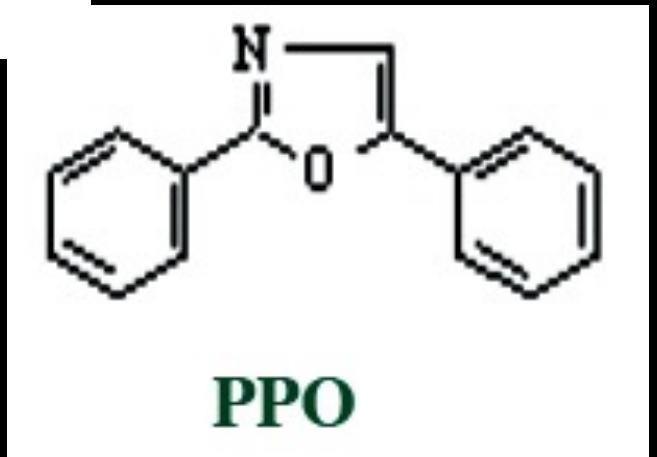
- Observe excellent clarity above PPO absorption (UV-Vis spectroscopy)
- Light yield in excess of calibration standards



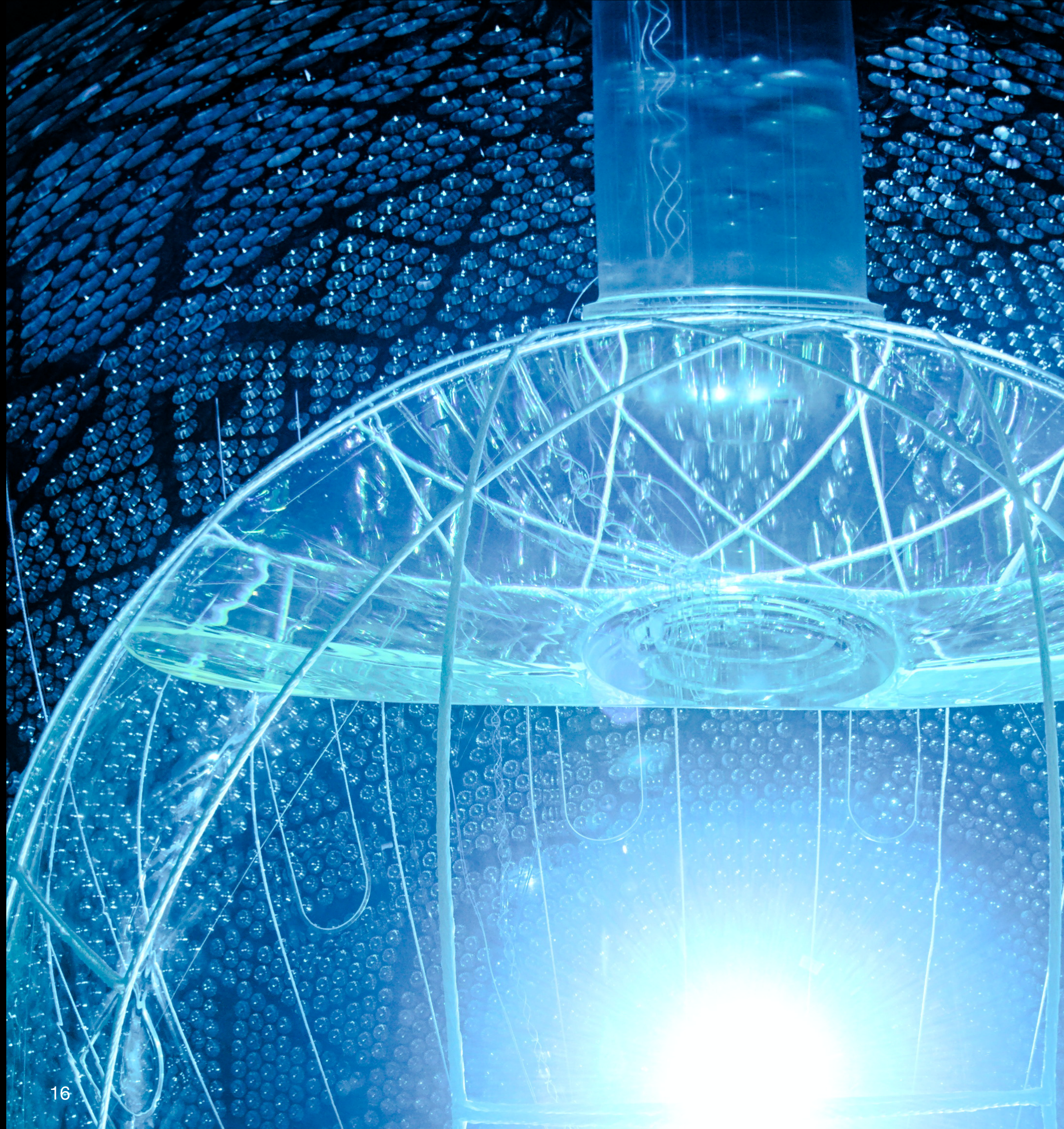
Transfer via railcar from surface to underground



LAB + 2g/L PPO

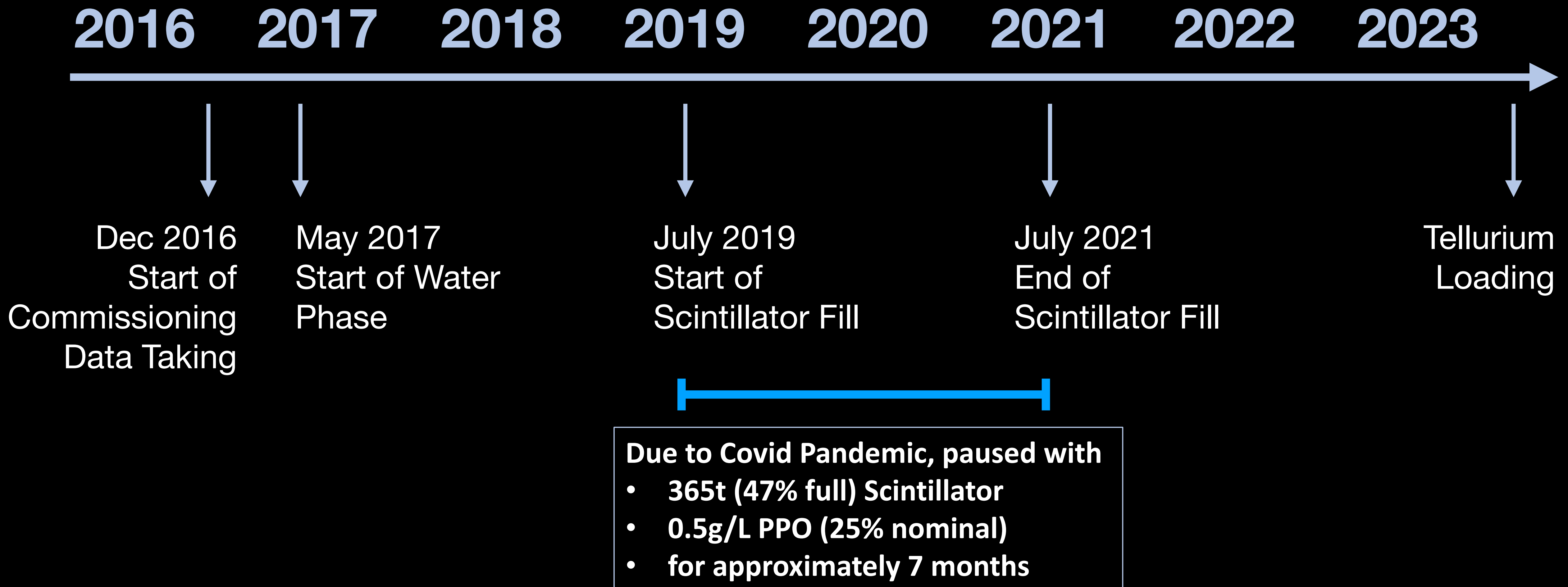


SNO+ Partial Scintillator-Fill Phase



SNO+ Timeline

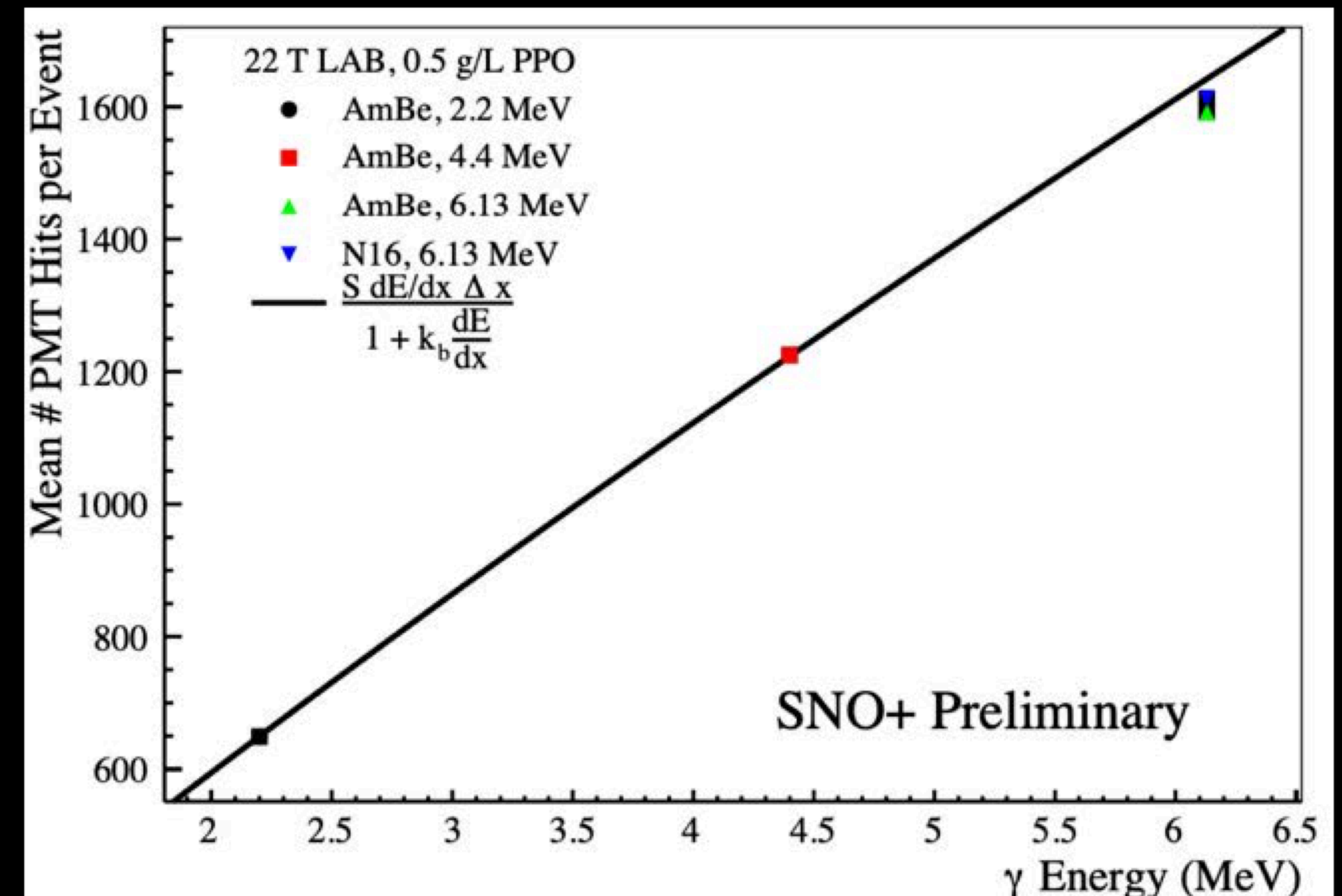
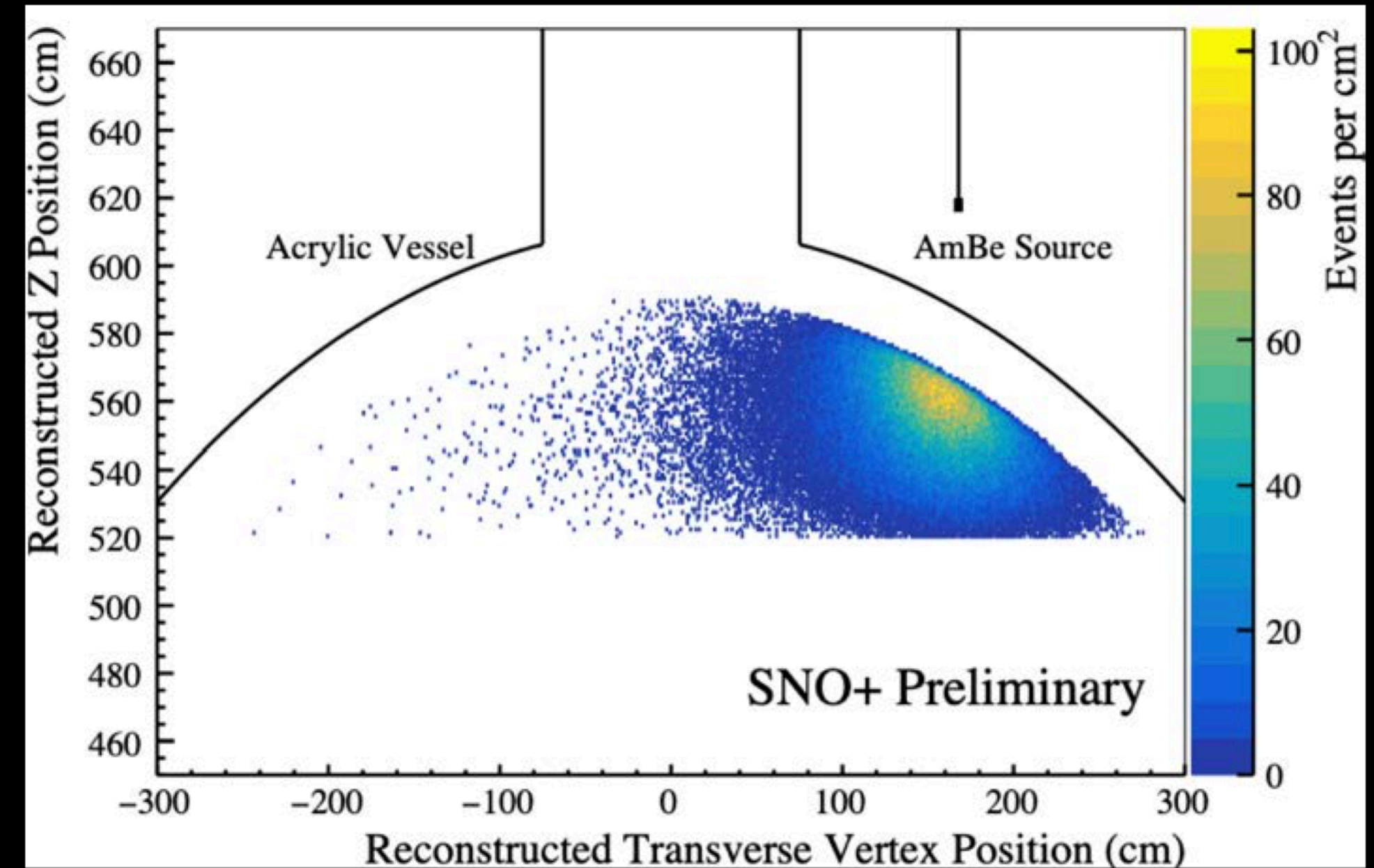
Partial Fill Phase



SNO+Partial Fill

Scintillator Light Yield

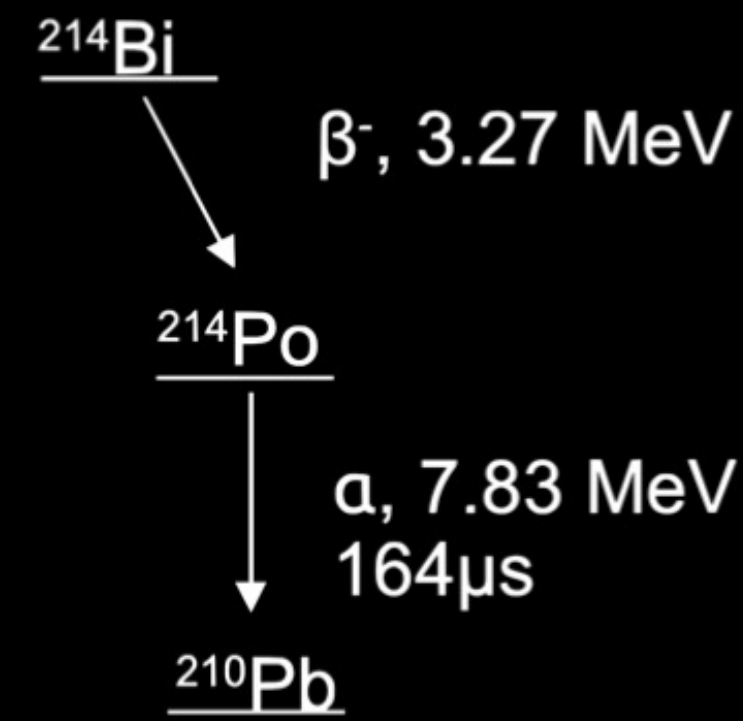
- Calibration sources (^{16}N , AmBe) deployed through guide tubes into the external water region
- With a PPO concentration of 0.5 g/L (25% of nominal concentration) we saw a LY equivalent of ~ 300 p.e. / MeV



SNO+Partial Fill

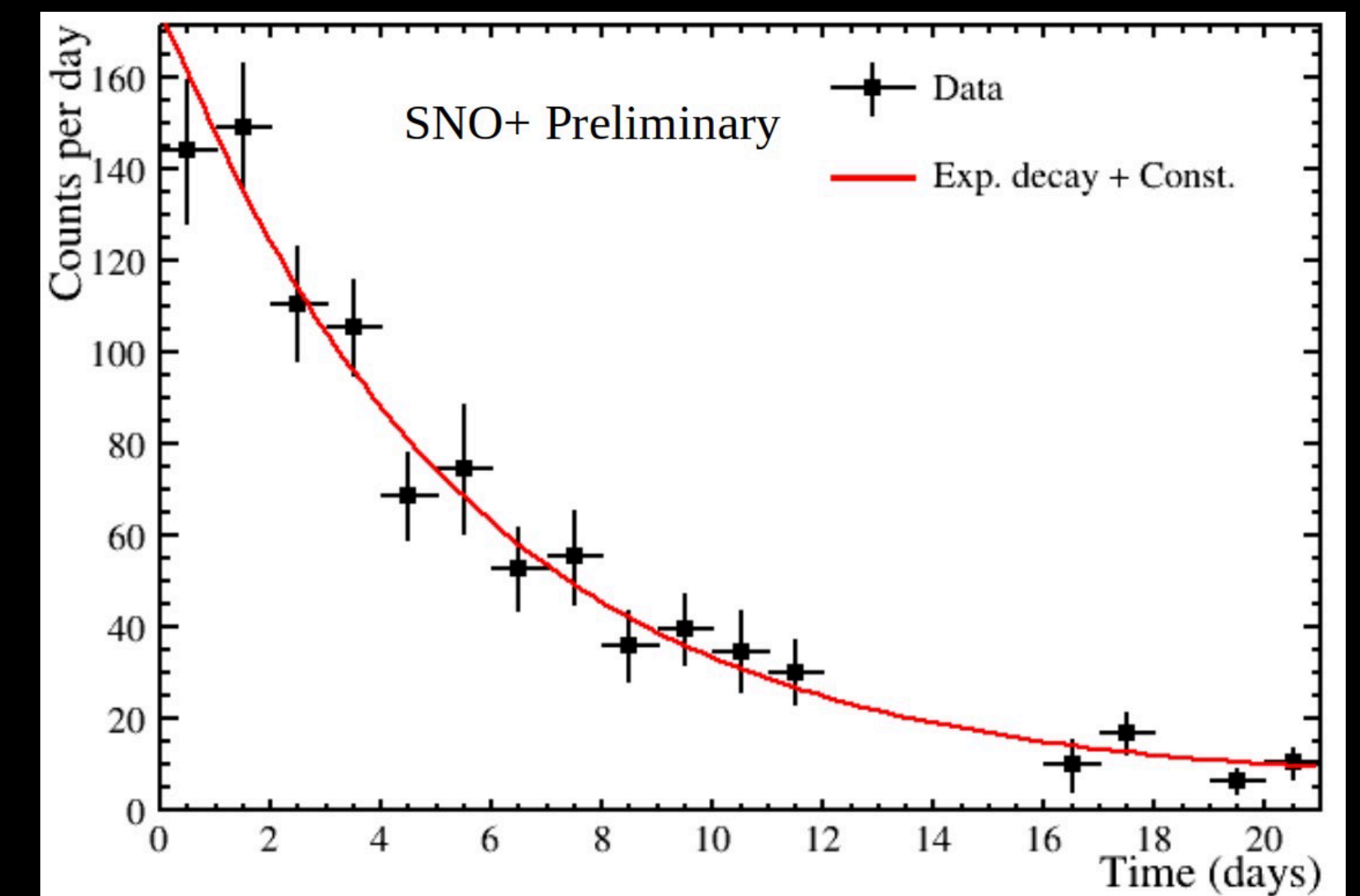
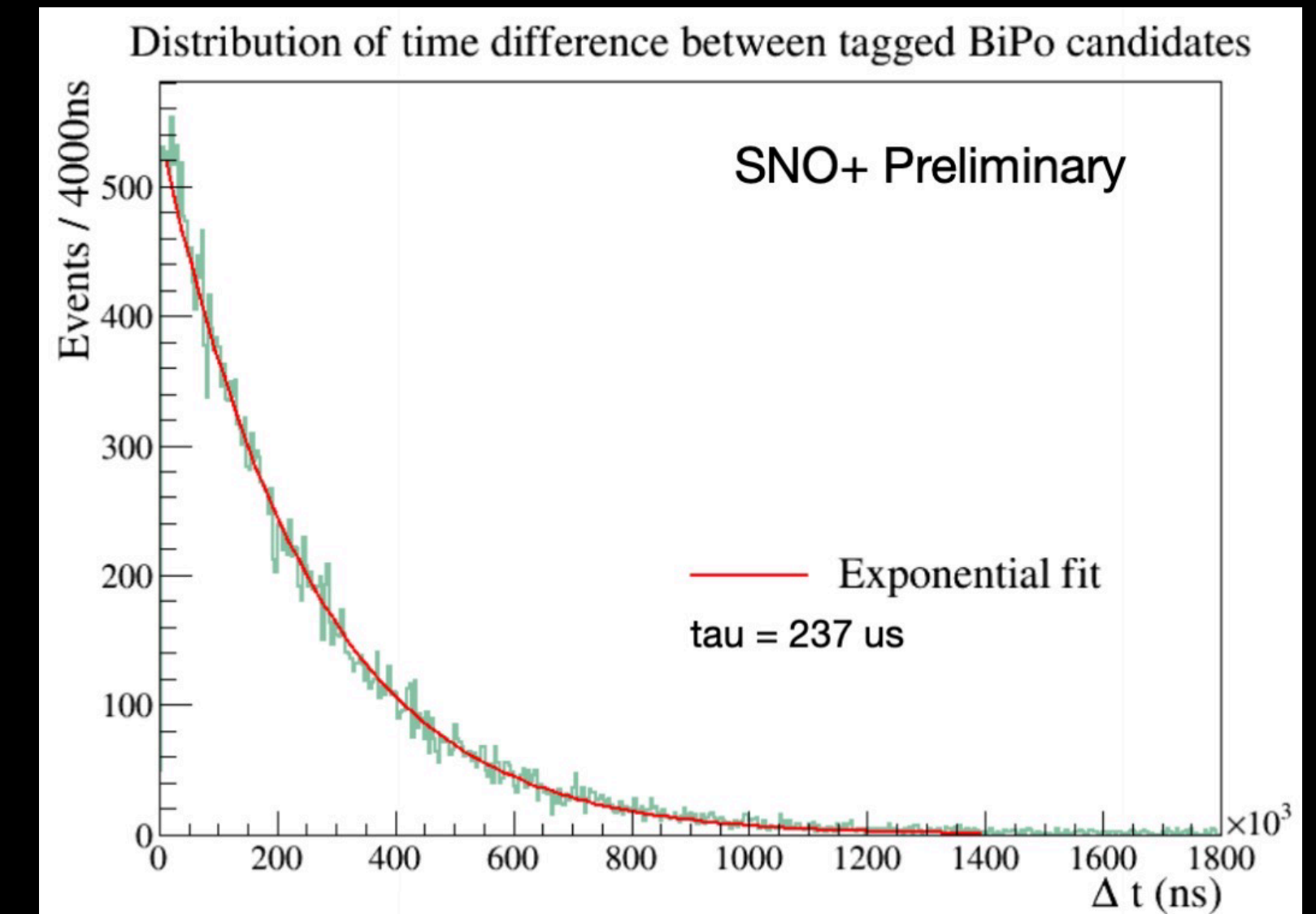
Scintillator Background Measurement

- LS backgrounds measured at
 - $^{214}\text{BiPo}$ delayed coincidences for U chain ($T_{1/2} = 164 \mu\text{s}$)
 - $(4.7 \pm 1.2) \times 10^{-17} g_{\text{U}}/g_{\text{LAB}}$
 - $^{212}\text{BiPo}$ delayed coincidences for Th chain
 - $(5.3 \pm 1.5) \times 10^{-17} g_{\text{Th}}/g_{\text{LAB}}$



Base rate below requirements for $0\nu\beta\beta$

- Opportunity to perform a “target-out” $\beta\beta$ measurement
 - Test analysis methods
 - Check background model

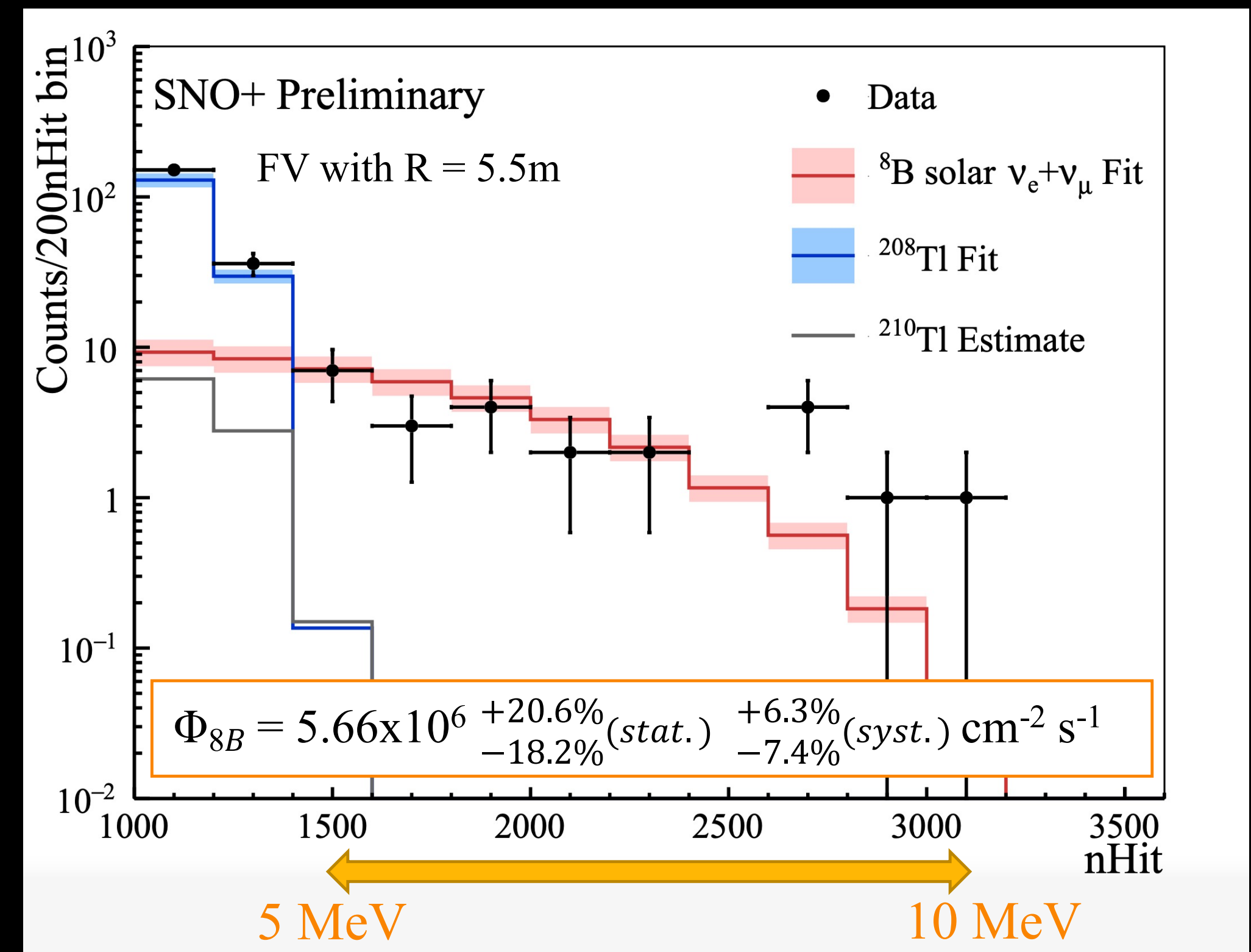


Some Radon (^{222}Rn , $T_{1/2} = 3.82$ days) introduced with filling that decays away

SNO+Partial Fill

Scintillator Physics

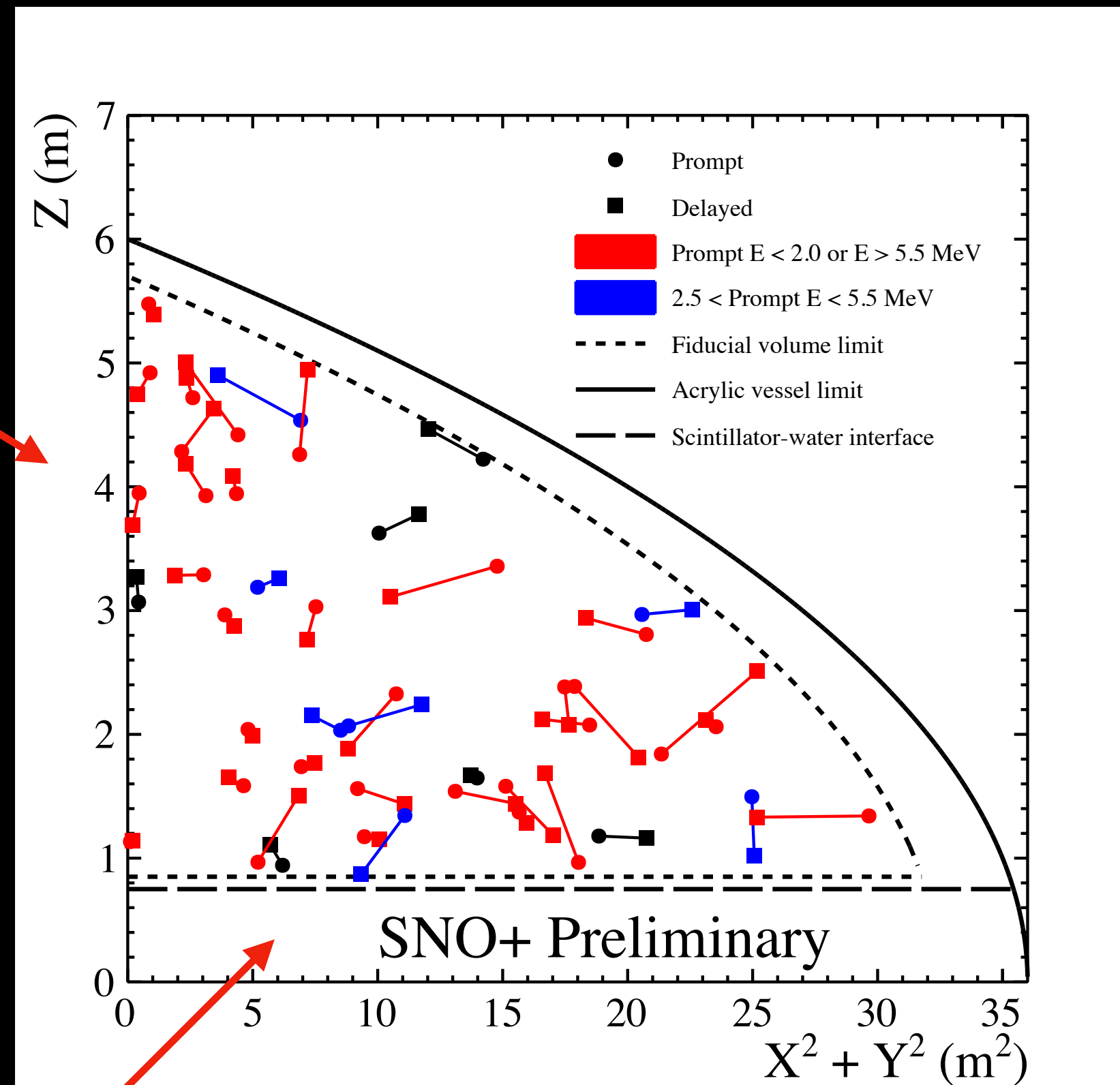
- ^8B solar ν + background fit to partial fill data
 - 5.5 m fiducial radius
 - Includes preliminary systematics
 - Flux fitted result compatible with other measurements
- Live for supernova
 - Burst monitoring part of SNEWS-1
 - Pre-supernova monitor enabled (IBD)
- Anti-neutrino measurement
 - Using time coincidence of IBD



SNO+ reactor antineutrinos

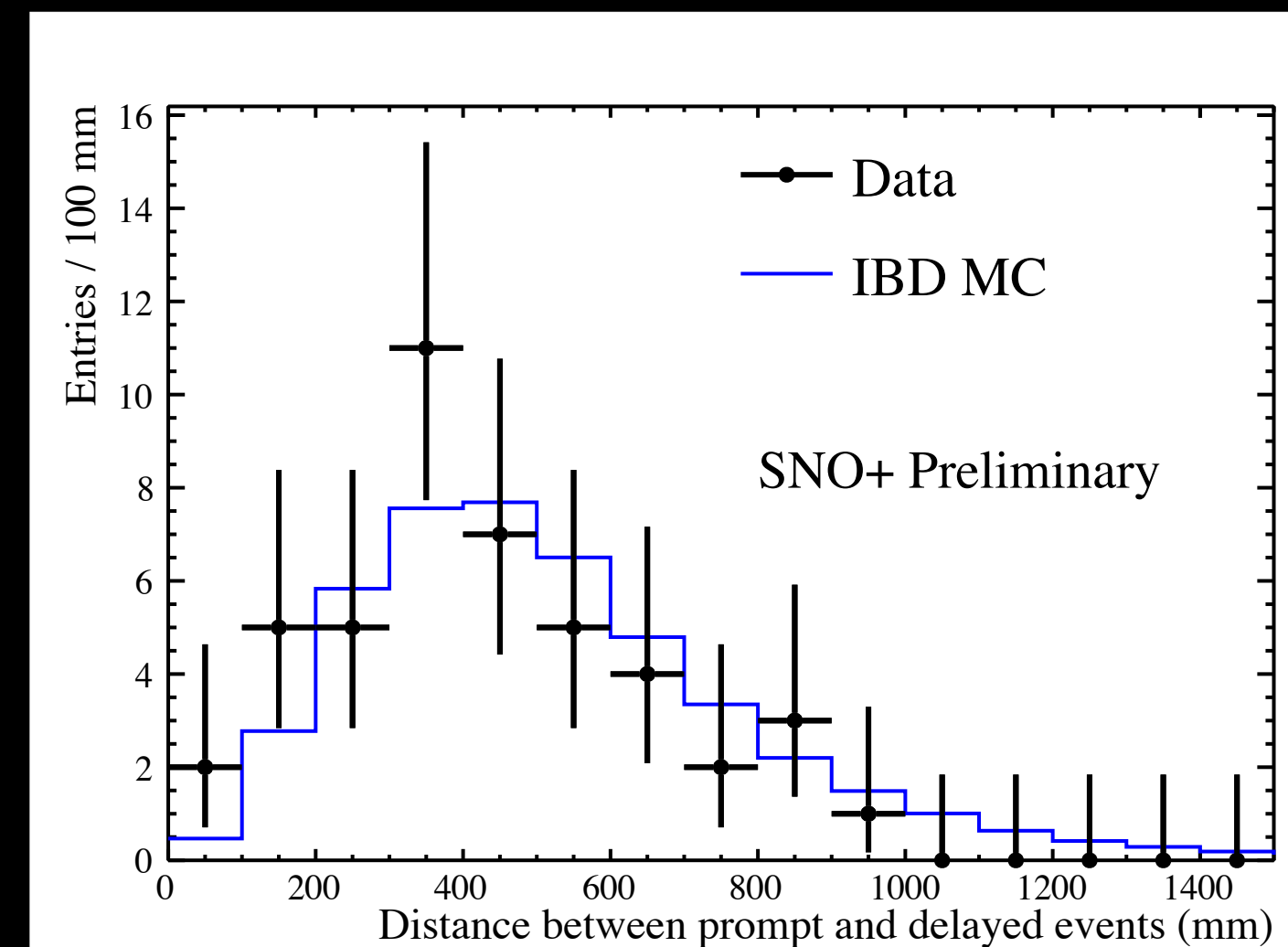
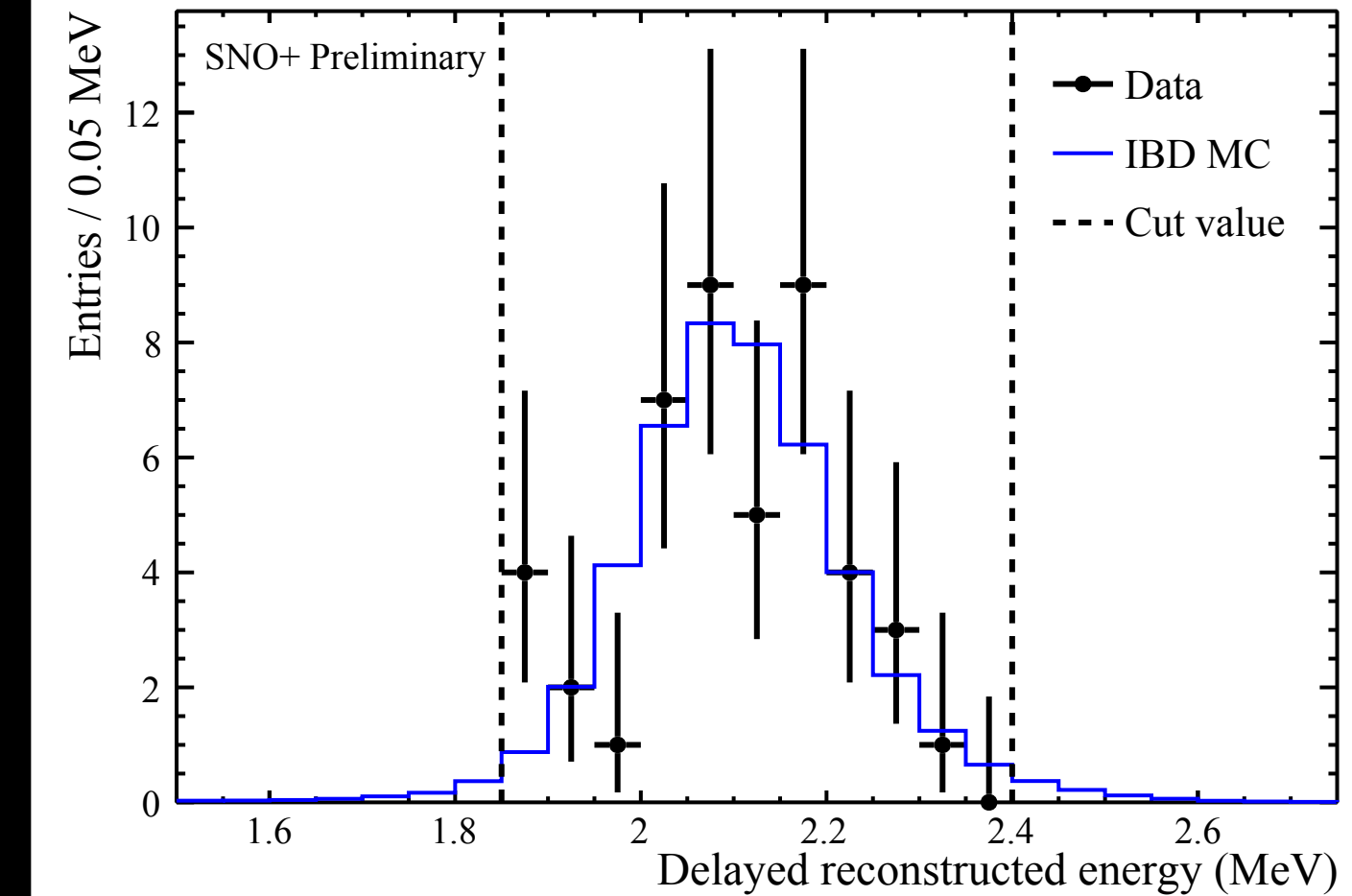
...in partial-fill

Events uniformly distributed in the detector



Fiducialisation to reduce background events from acrylic

Publication in preparation

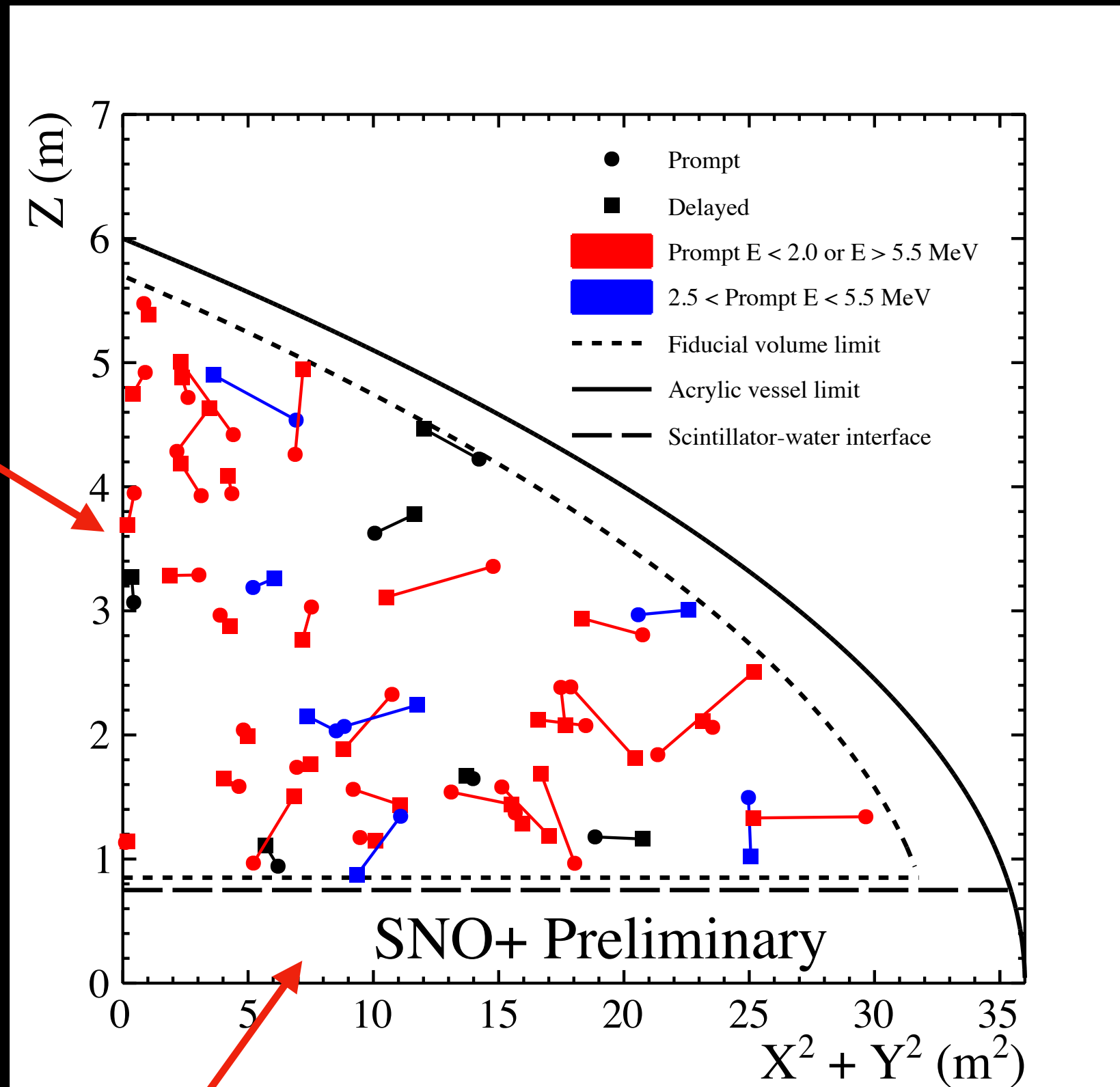


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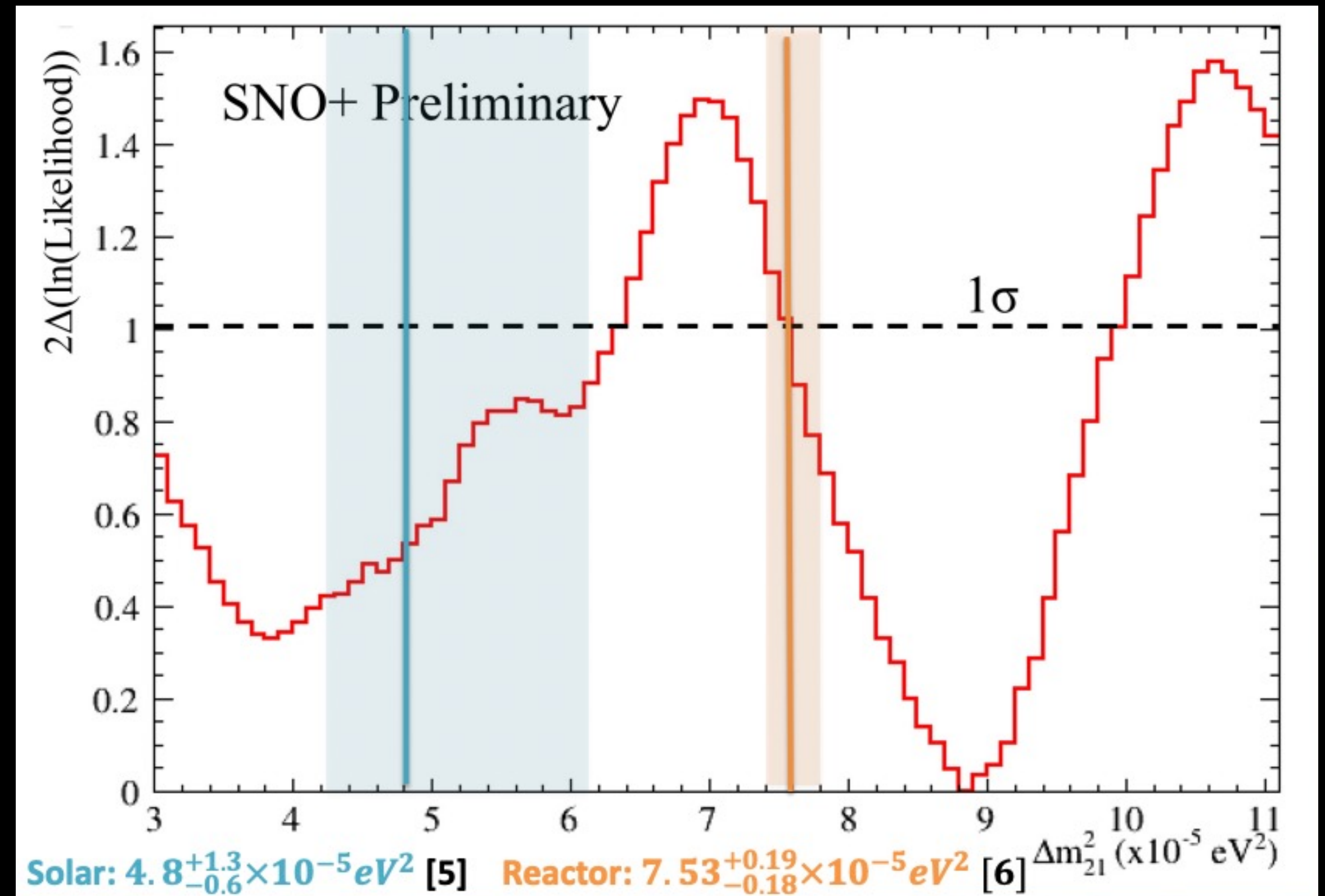
...in partial-fill

Publication in preparation

Events uniformly distributed in the detector



Fiducialisation to reduce background events from acrylic



Δm_{21}^2 values corresponding to **solar** and **reactor** neutrino measurements allowed within 1σ

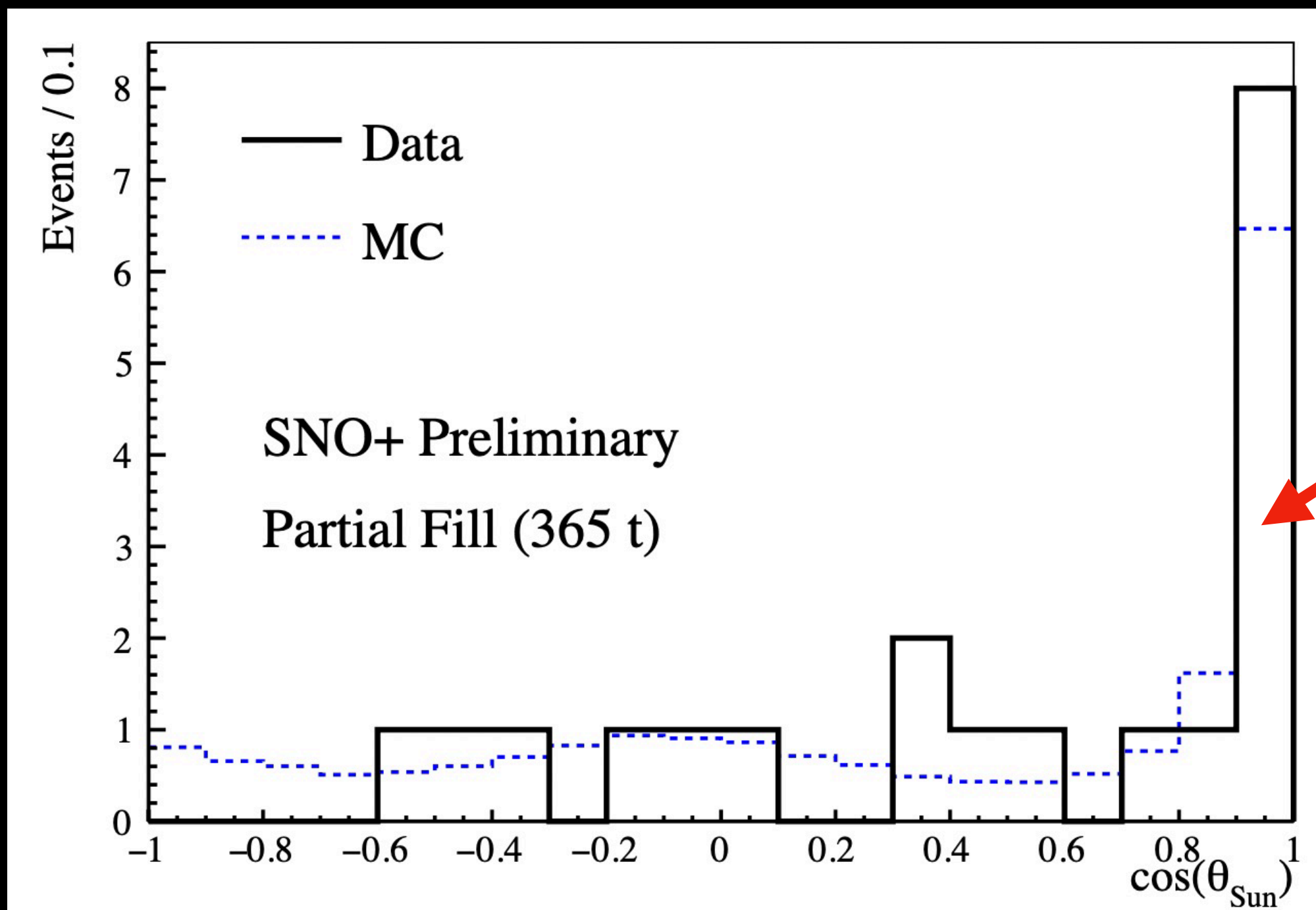
SNO+ Timeline

PPO top-up phase (from 0.5 g/L to 2.2 g/L)



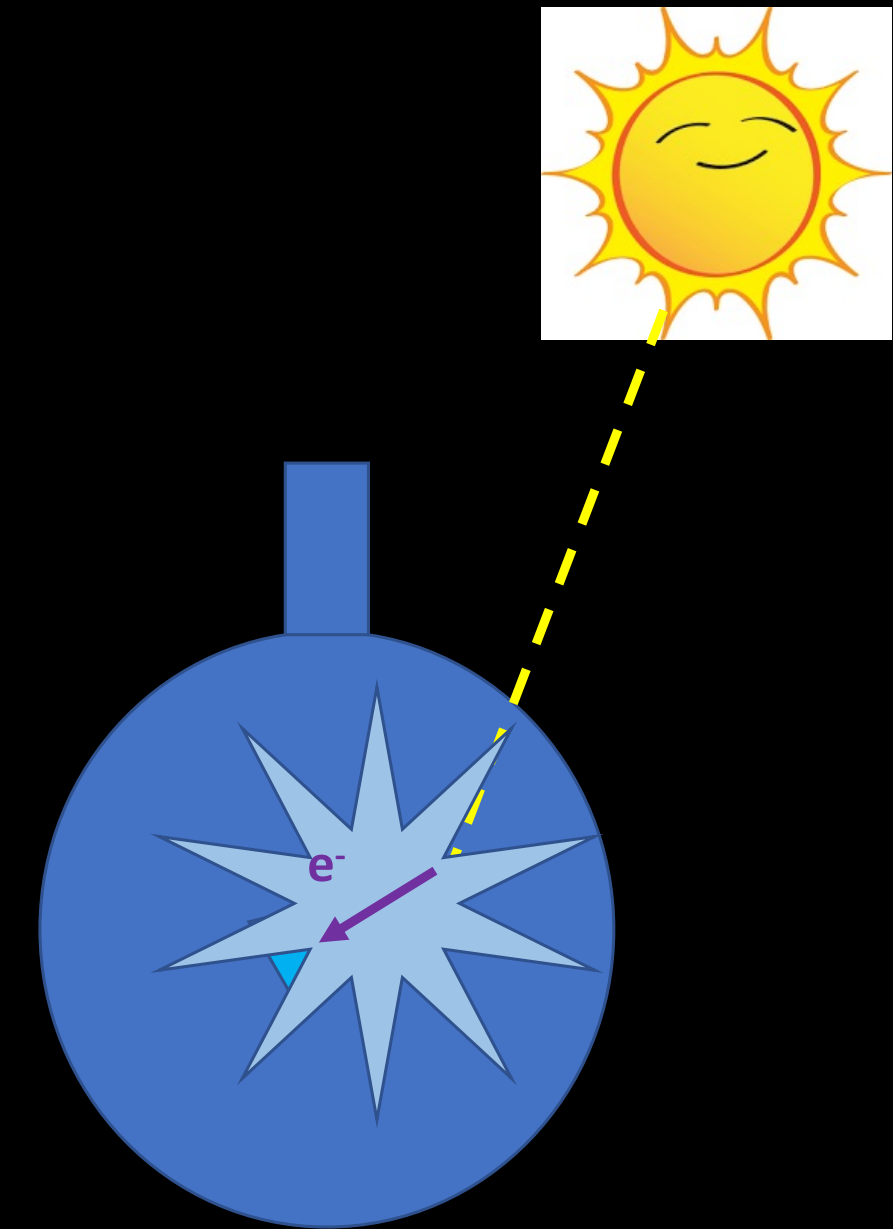
SNO+ direction in scintillation

- Data collected with lower PPO concentration (~ 0.5 g/L) than nominal (2.2 g/L)
- Lower LY and slower rise time permits to distinguish early (directional) Cherenkov light



Direction reconstruction
from simultaneous
Cherenkov + Scintillation
fit

Event-by-event direction
reconstruction



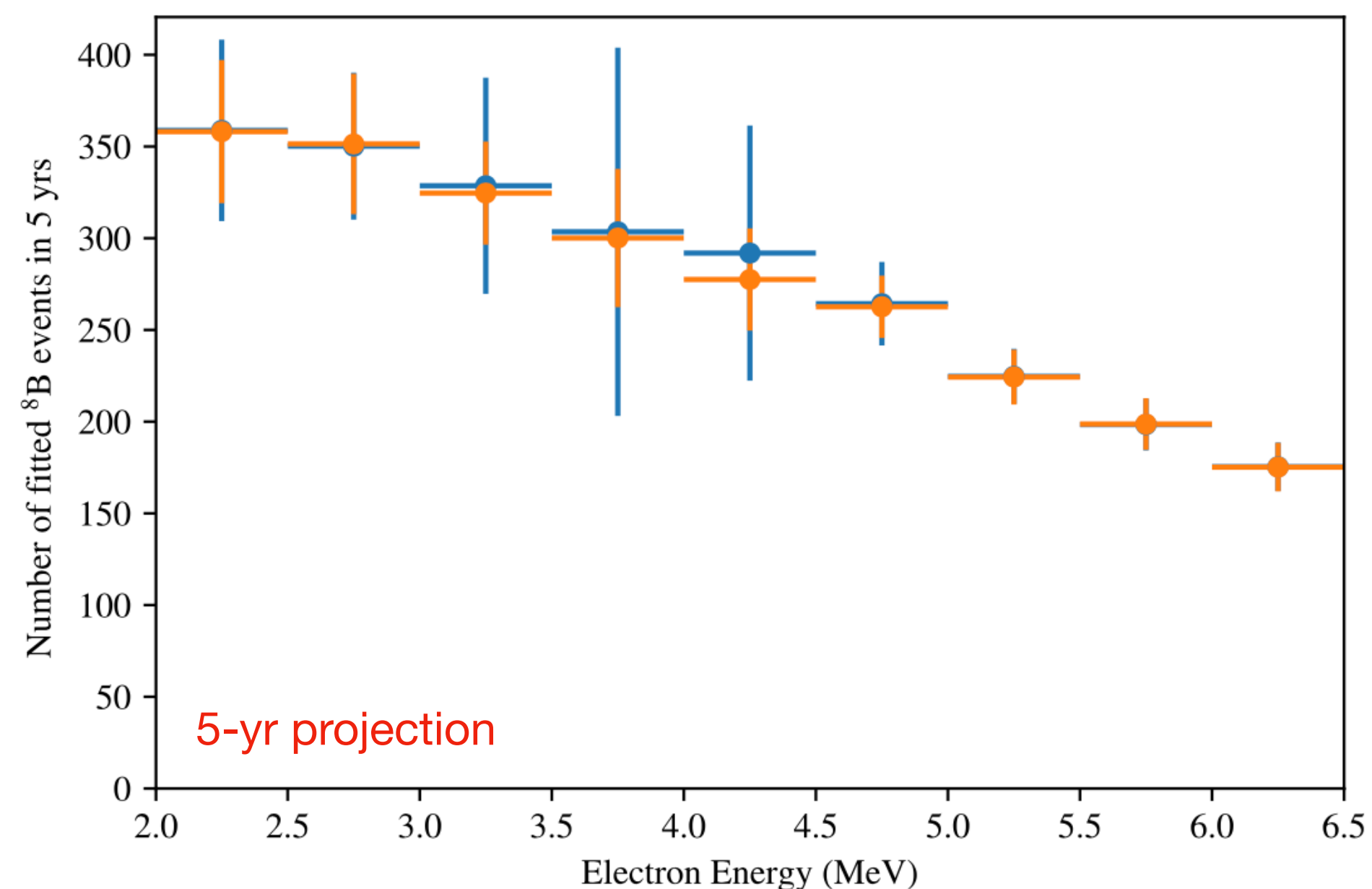
SNO+ Scintillator Phase



Objectives for SNO+ Scintillator Phase

^8B Solar ν

- Attempt to measure ^8B below 3 MeV
 - Probe into the transition region
 - larger fiducial volume than Borexino
 - cosmogenic backgrounds lower than KamLAND (e.g., no ^{10}C , ^{11}C)



Blue : $5 \times 10^{-17} \text{ g}^{238}\text{U}/\text{g}_{\text{LAB}}, 5 \times 10^{-17} \text{ g}^{232}\text{Th}/\text{g}_{\text{LAB}}$
Orange : $5 \times 10^{-18} \text{ g}^{238}\text{U}/\text{g}_{\text{LAB}}, 5 \times 10^{-18} \text{ g}^{232}\text{Th}/\text{g}_{\text{LAB}}$

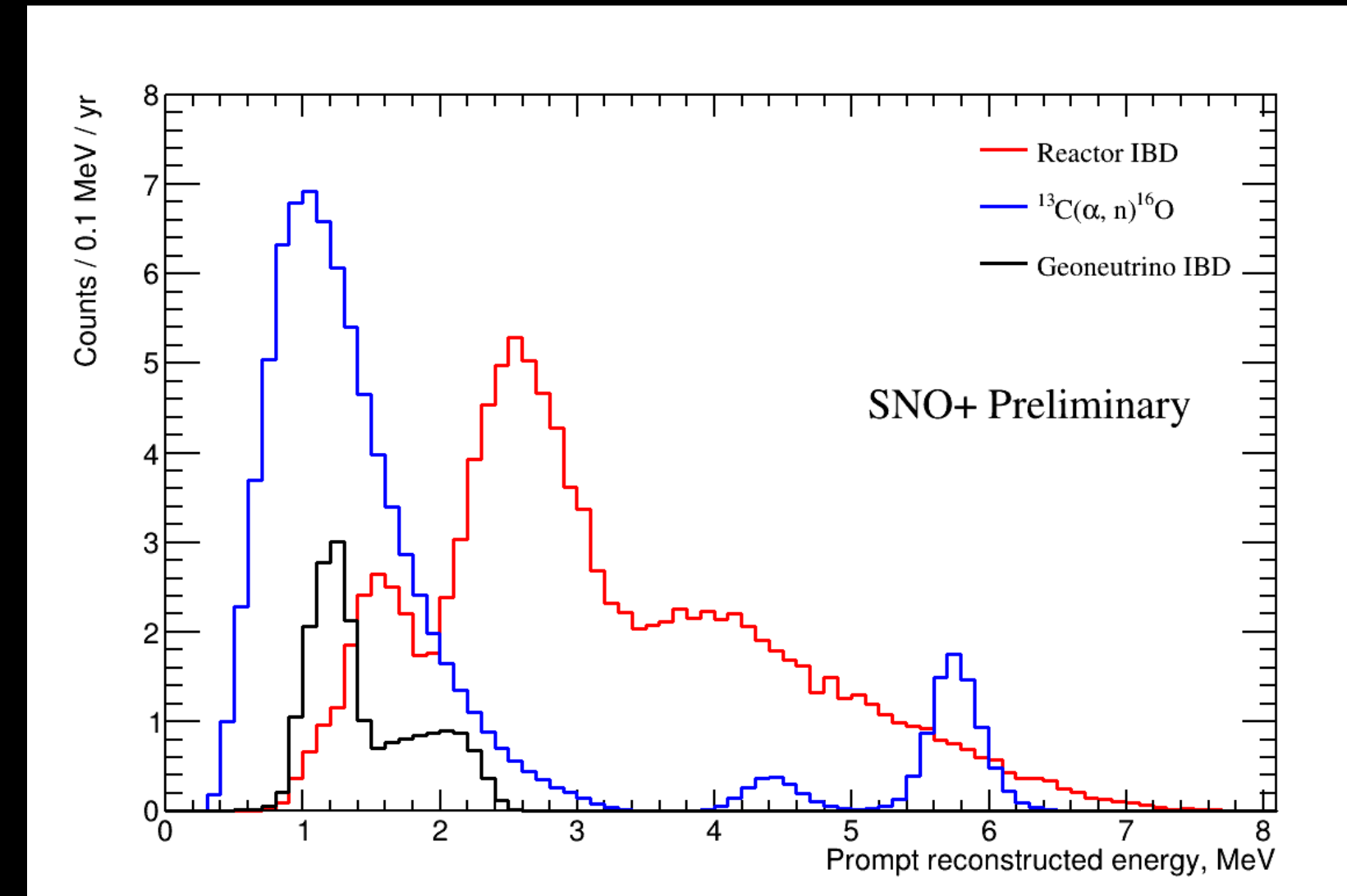
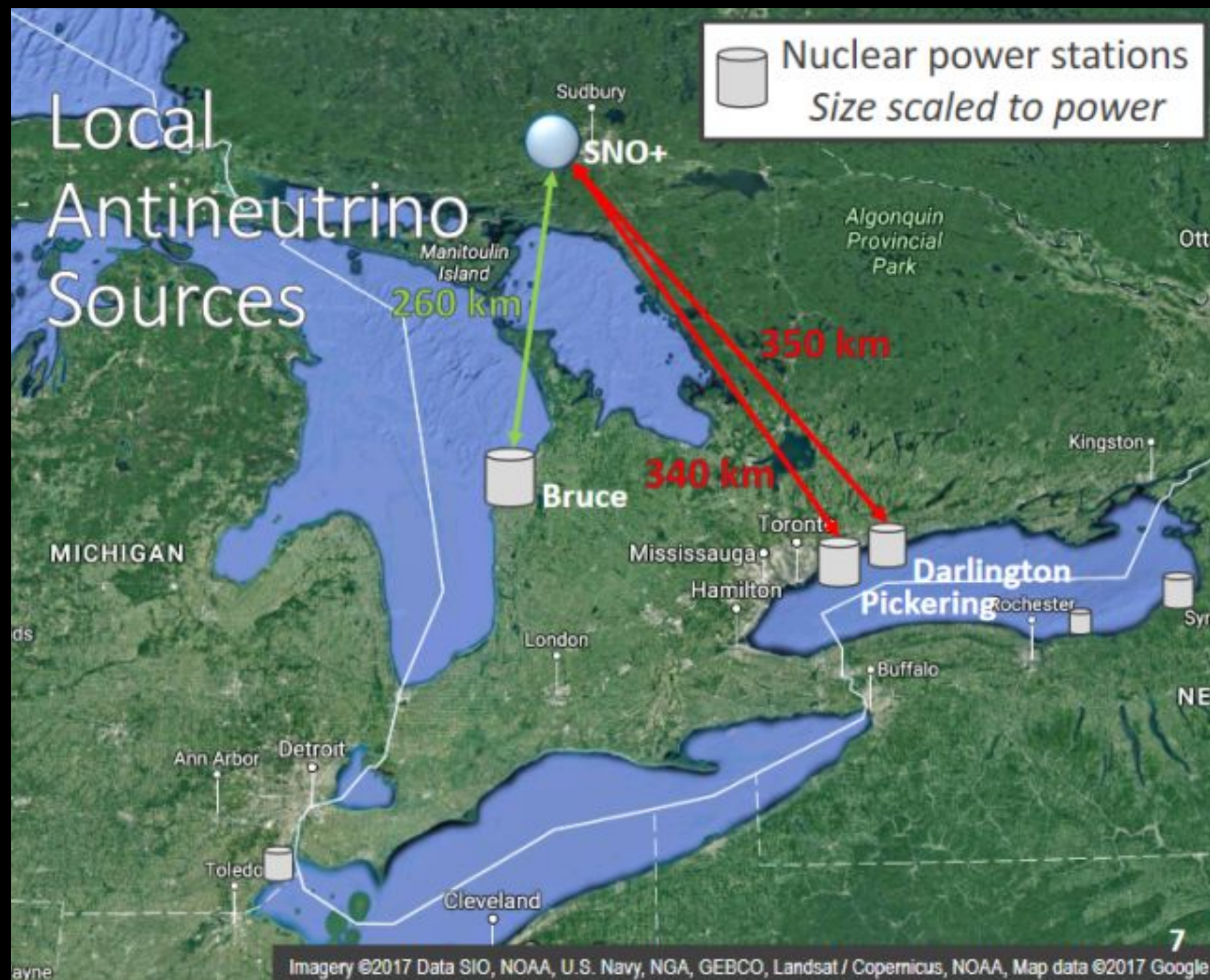
Blue : U and Th at partial fill level

Orange : U and Th below 10^{-17} g/g

Objectives for SNO+ Scintillator Phase

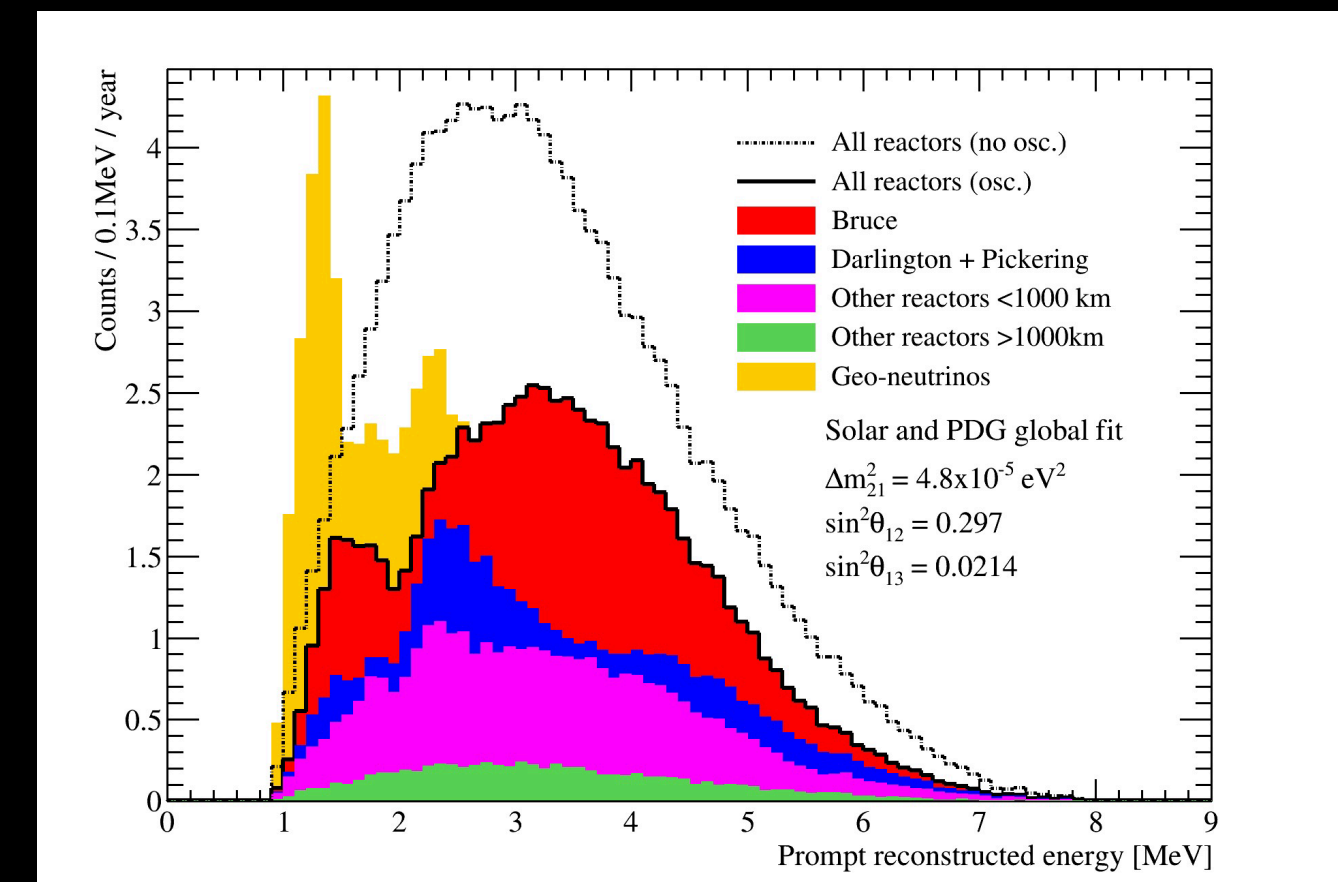
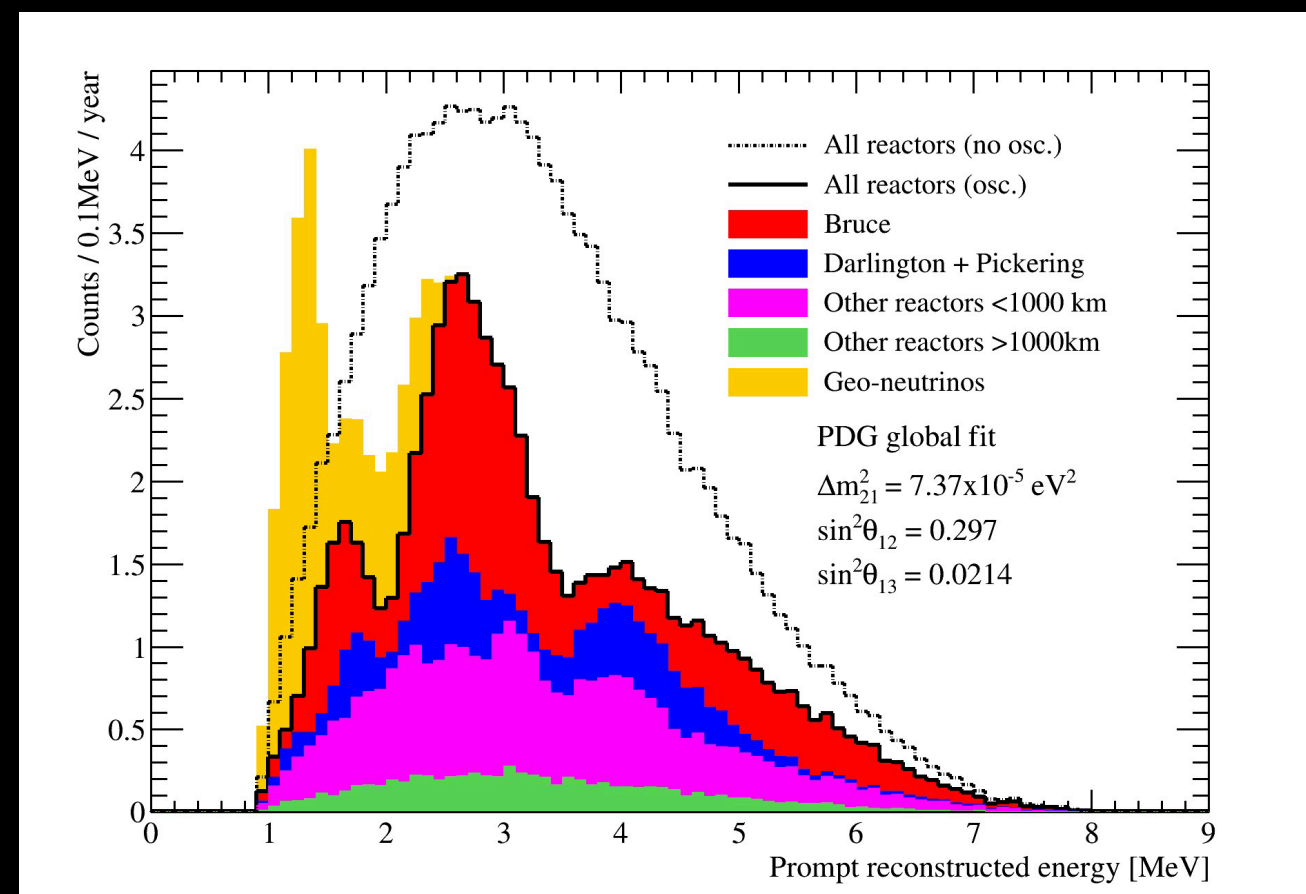
Reactor antineutrinos (Δm_{21}^2)

- Sensitivity to probe the tension already after 1 year
 - Shape is more important than rate
- SNO+ location yields only two distinct baselines



KamLAND ($\sim 7 \times 10^{-5} \text{ eV}^2$)

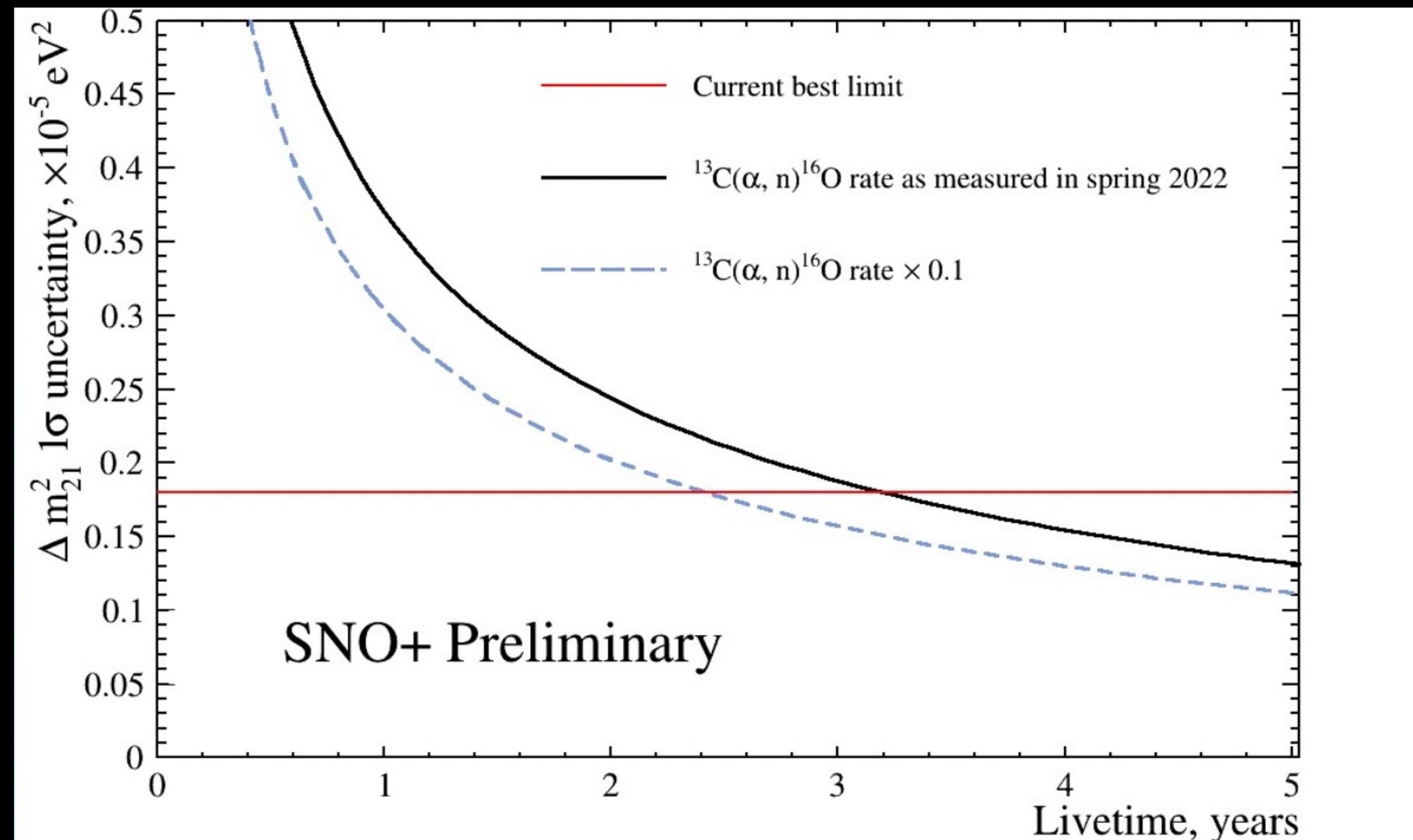
Global ($\sim 5 \times 10^{-5} \text{ eV}^2$)



Objectives for SNO+ Scintillator Phase

Reactor antineutrinos (Δm_{21}^2)

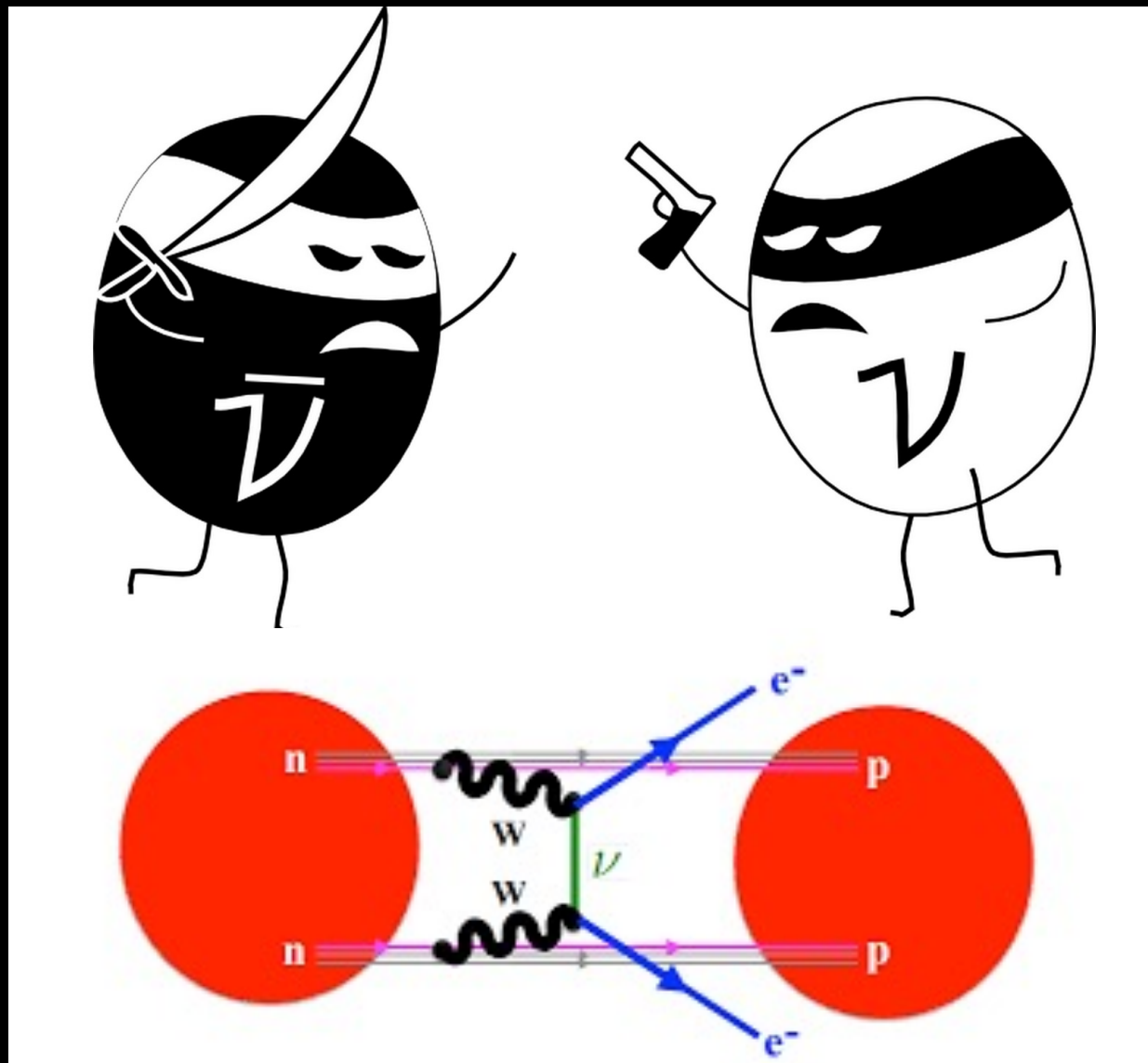
- Sensitivity to probe the tension already after 1 year
 - Shape is more important than rate
- SNO+ location yields only two distinct baselines



Why is this relevant for SNO+ $0\nu\beta\beta$?

- The advantages of a well-understood detector with low backgrounds are being demonstrated!
- SNO+ has a diverse program of neutrino (and other) physics that is being pursued.
 - The physics themselves are relevant and interesting
 - ...better understanding of the detector means better background model
- With the detector performing well; with all background components being measured and constrained (most coming in at or below target levels), it looks promising for the final phase of SNO+...

SNO+ Tellurium Double Beta Phase



Neutrinoless Double Beta ($0\nu\beta\beta$) Decay in SNO+

Strategy of Tellurium-Loaded Liquid Scintillator

- **Massive Detector:**

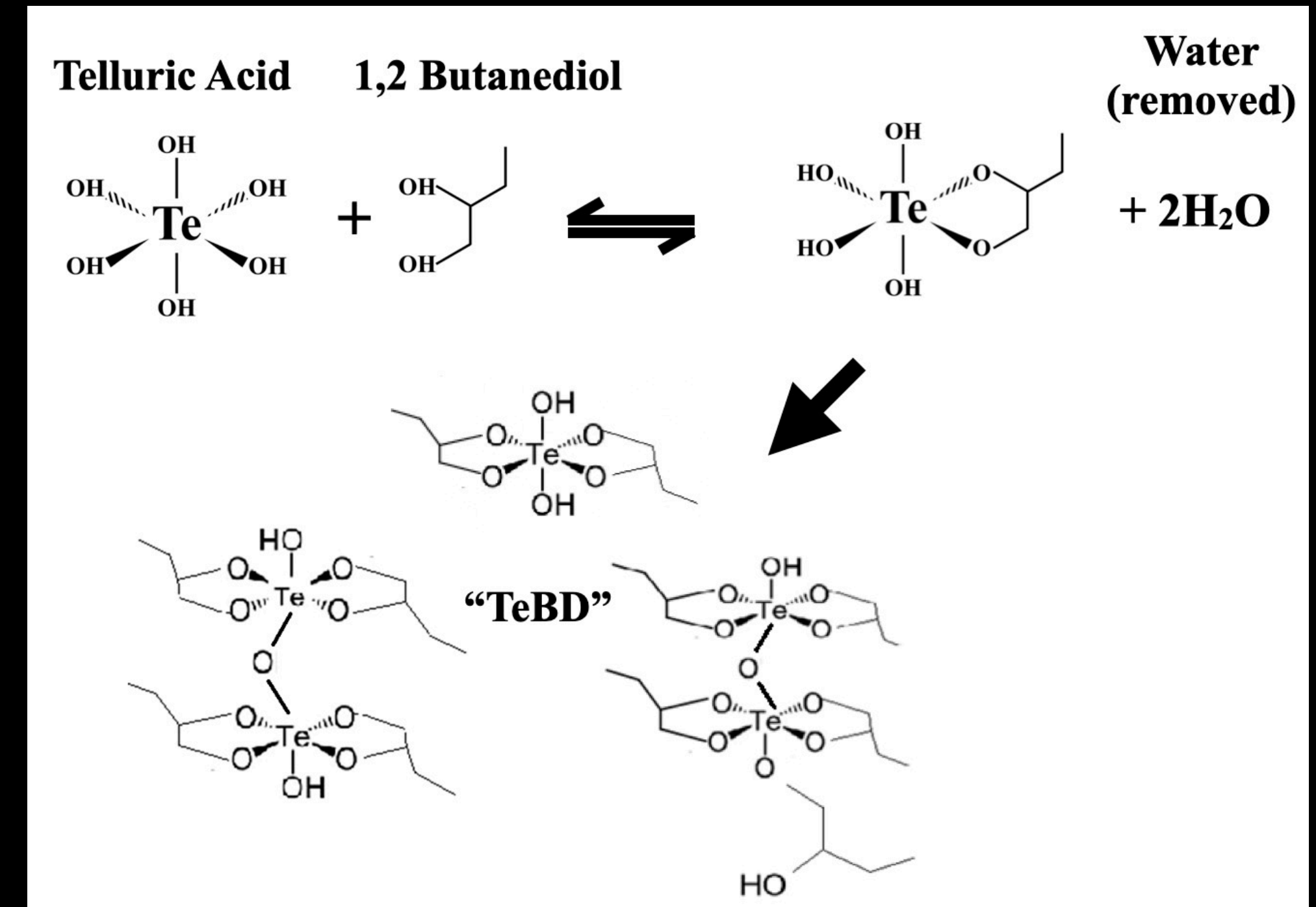
- High statistics
- Shielding through fiducialisation

- **Liquid Scintillator**

- Purification methods
- Scalable loading
- Homogeneous loading of isotope in detection medium

- **Tellurium-130**

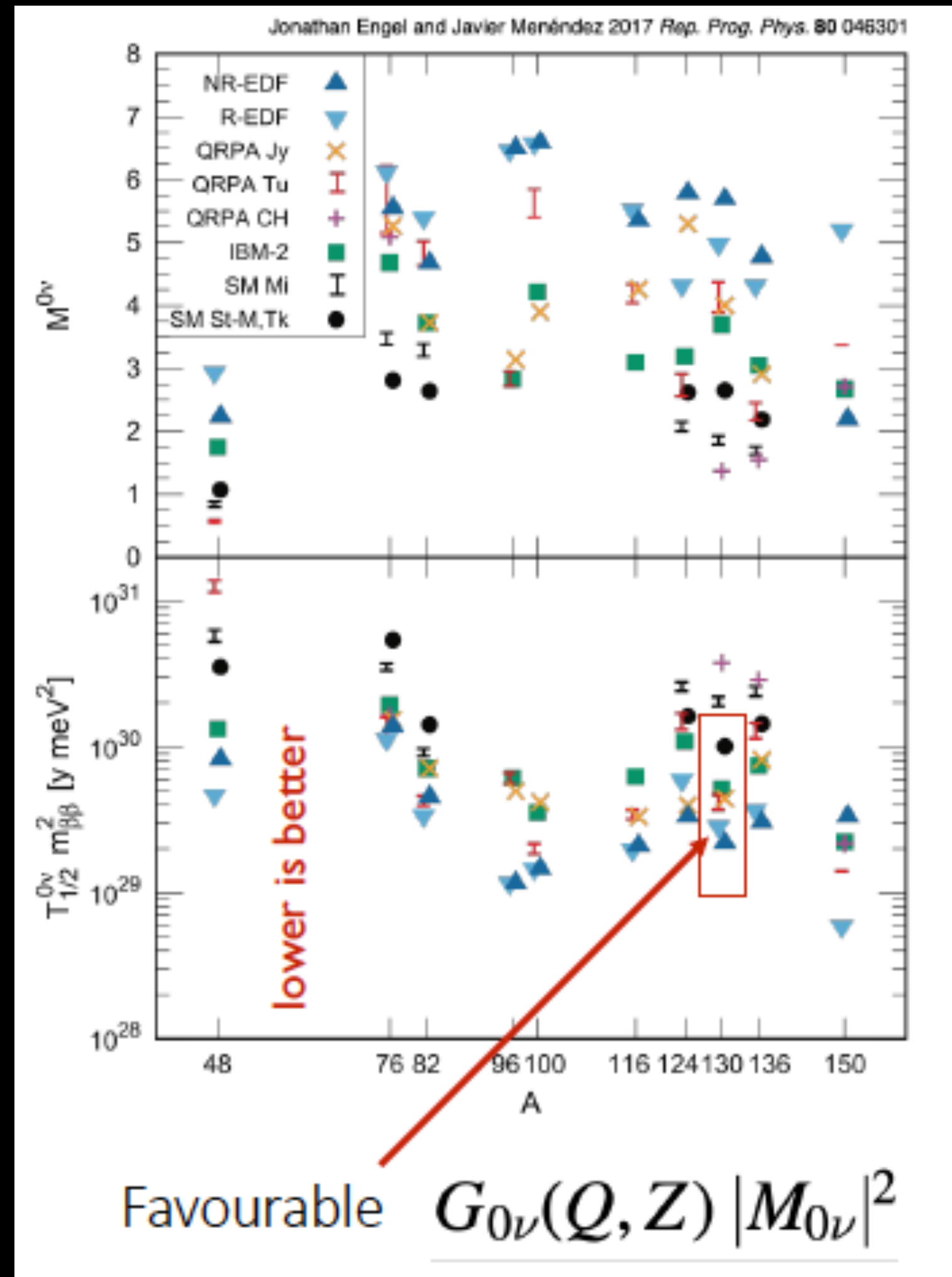
- Highest natural abundance, no enrichment required, low cost



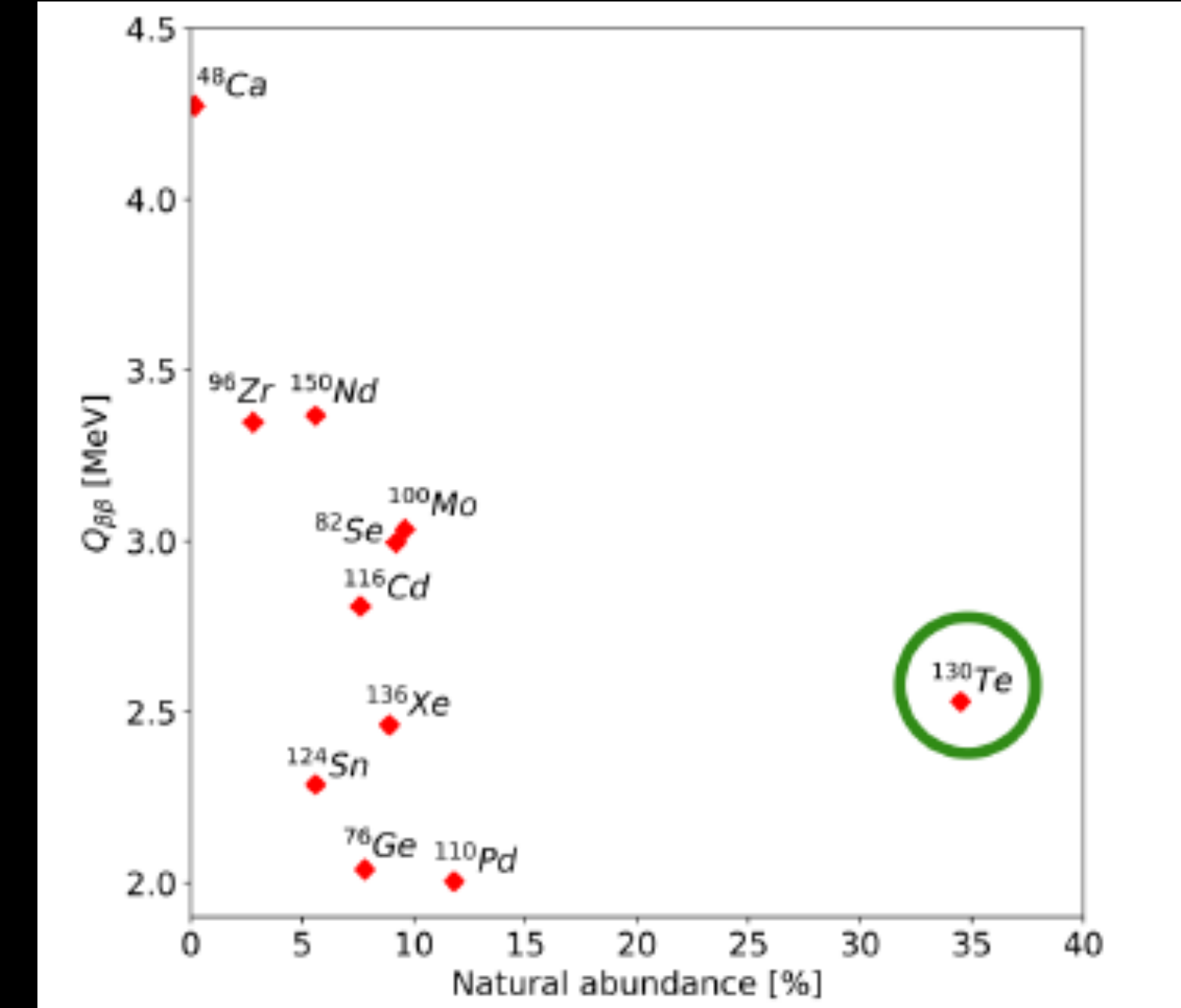
Tellurium as a DBD candidate

Why Te?

Favourable Phase Space



Very high abundance



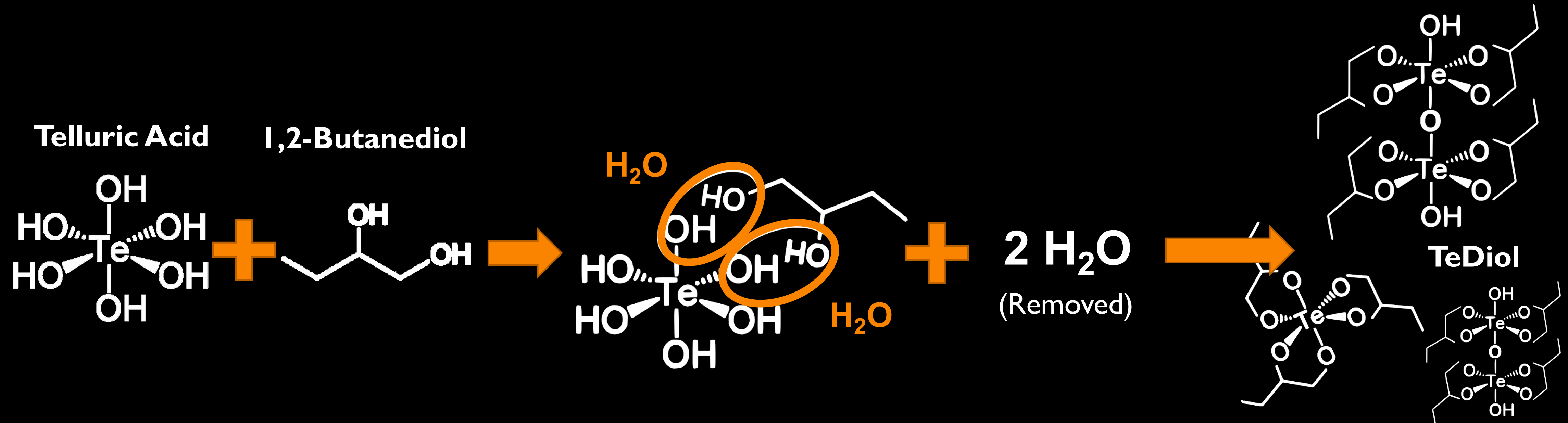
Tellurium loaded LS

Why liquid scintillator?

- Very low backgrounds: 5×10^{-7} counts/keV/kg_{fiducial detector}/yr
- Homogeneous detector volume – reliable background model
- “target out” – ability to measure/constrain backgrounds before isotope added
- “sideband analysis” – not just counts in a bin but distributions in position and energy verify detector response and background model
- liquid detector permits: assays, chemistry; liquid medium can be modified in situ (e.g., adding more Te, more fluor)

The dependence of a putative signal with amount of isotope would be a strong confirmation!

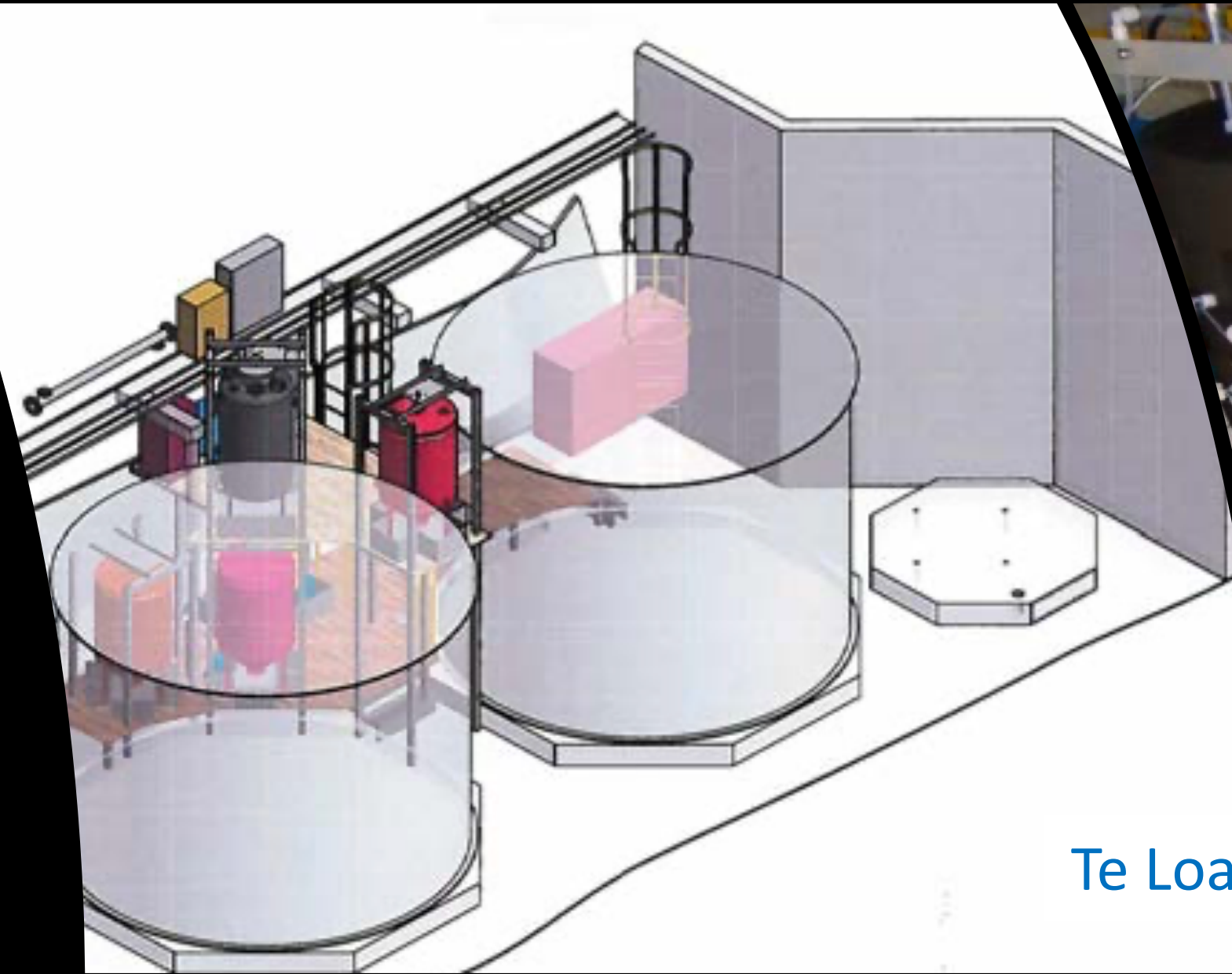
Tellurium Loading



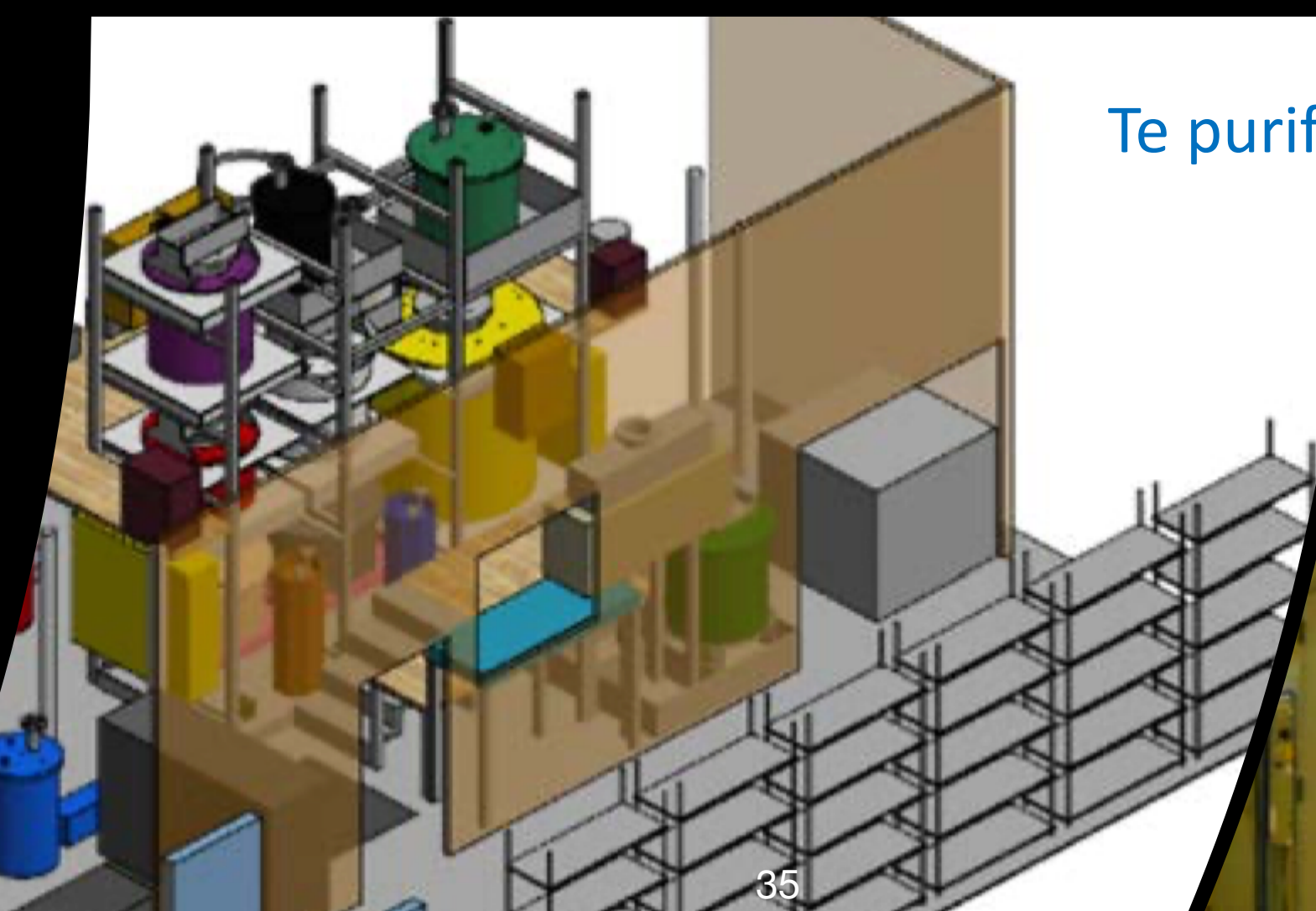
- TeDiol is mixed directly into the LAB+PPO with 15 mg/L bis-MSB and a stabilizer called Dimethyldodecylamine (DDA).
- Final Te-loaded LS cocktail expected to produce ~460 p.e. / MeV in SNO+ for 0.5% Te loading.

Tellurium Plants

- ~8 tons of telluric acid (TeA) has been “cooling” underground for several years.
- Ton-scale underground purification of TeA for further background reduction.
- Target purification for Te cocktail:
 - $\sim 10^{-15}$ g/g U
 - $\sim 10^{-16}$ g/g Th



Te Loading plant



Te purification plant



SNO+ Te-phase background projections

Major contributions and mitigation strategies

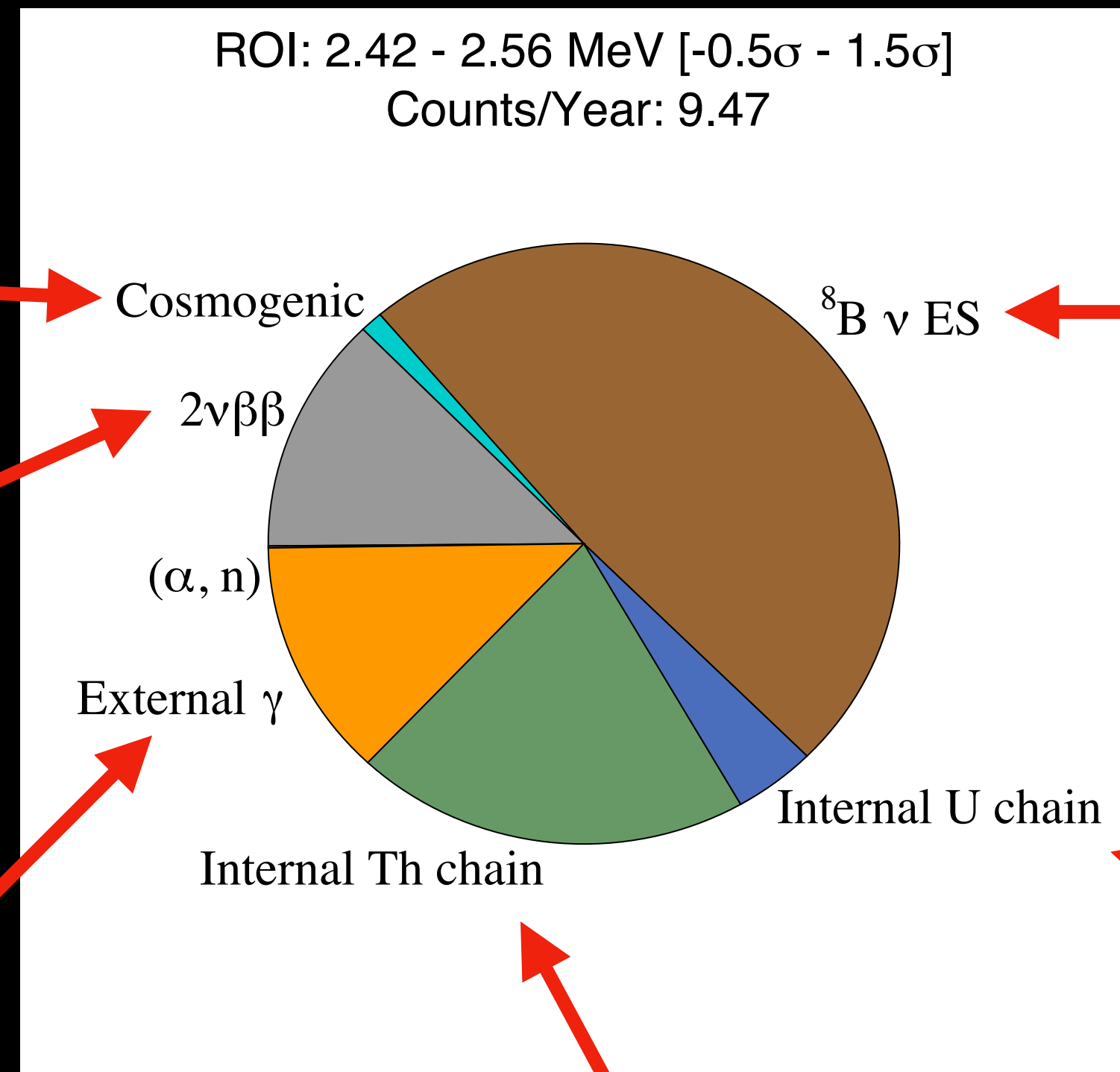
Staged Te-loading allows us to assess Te-backgrounds

Telluric acid has been “cooling”
underground for ~8 years

+ Te purification
+ multi-site rejection

Suppressed by asymmetric
ROI: (assuming ~460 PMT hits/
MeV)

Measured in water data
(reduced by 23%)



Well measured from previous
experiments

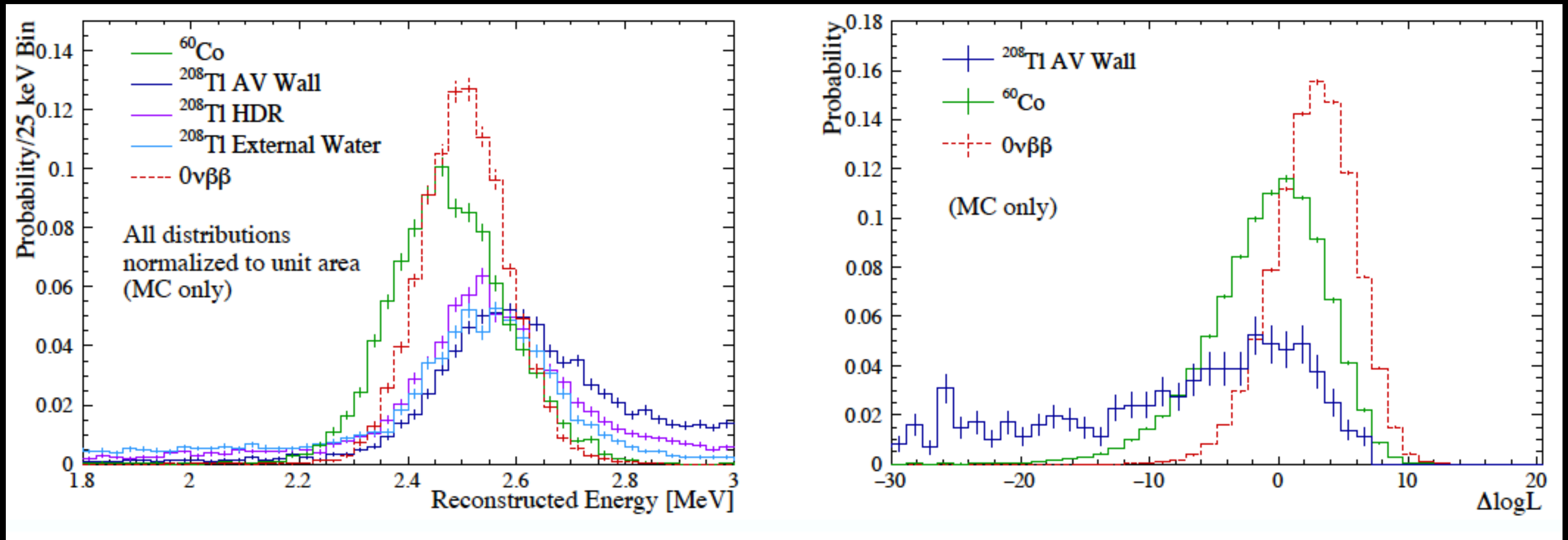
U levels already measured
below requirements in partial
fill ($<10^{-16}$ g/g)

Th levels already measured
below requirements in partial
fill ($<10^{-16}$ g/g)

SNO+ background suppression

Multi-site background constraint

- Analyses ongoing with promising results

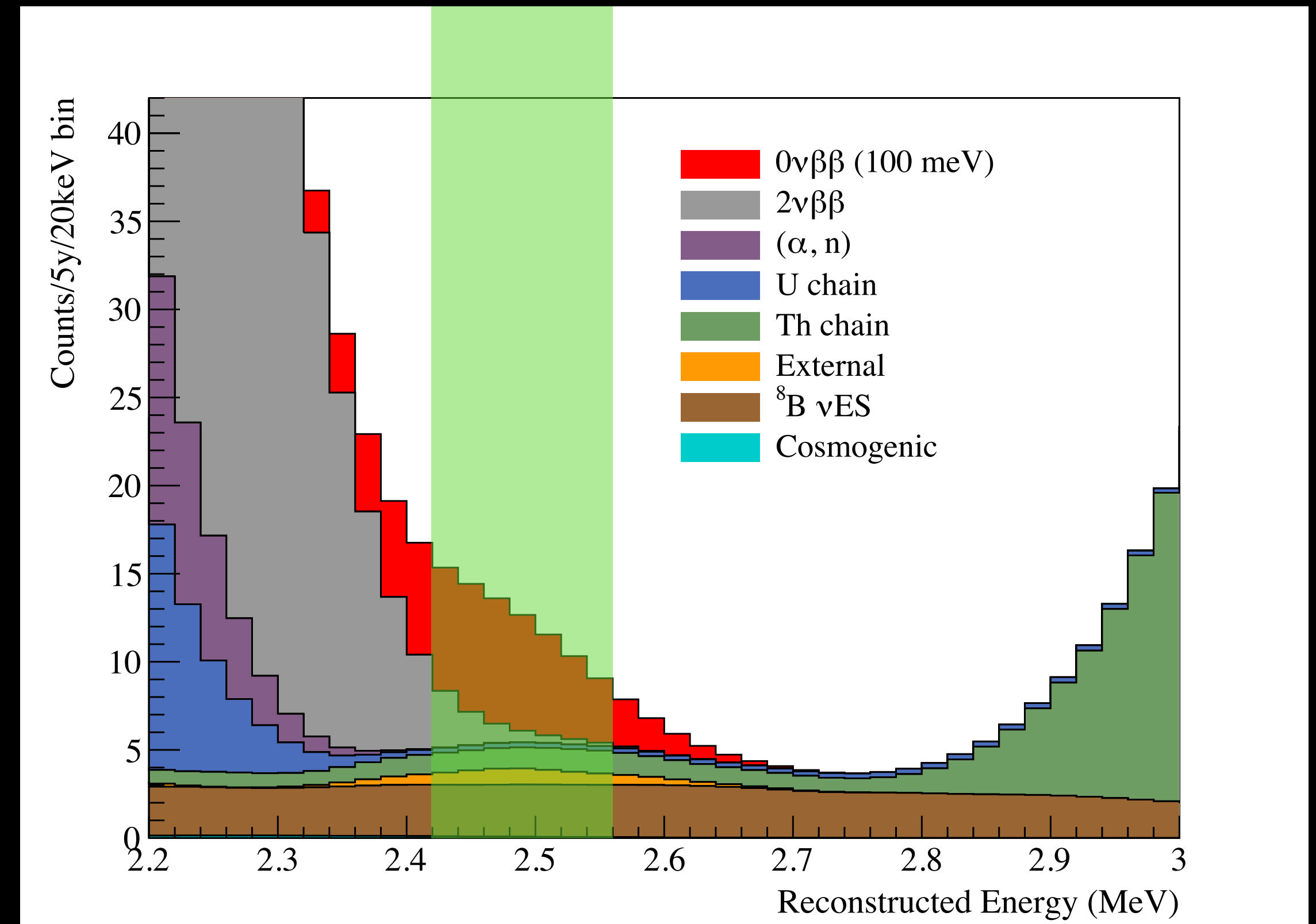


Projected spectrum

Assuming 0.5% loading

- Half-life sensitivity:
 - $T_{1/2} > 2 \times 10^{26}$ years
 - Sensitive to $m_{\beta\beta} \sim 37\text{-}89$ meV
- Full likelihood analysis using energy, radius and multi-site discriminators achieves this sensitivity with 3 years of data

ROI : 2.42 - 2.56 MeV

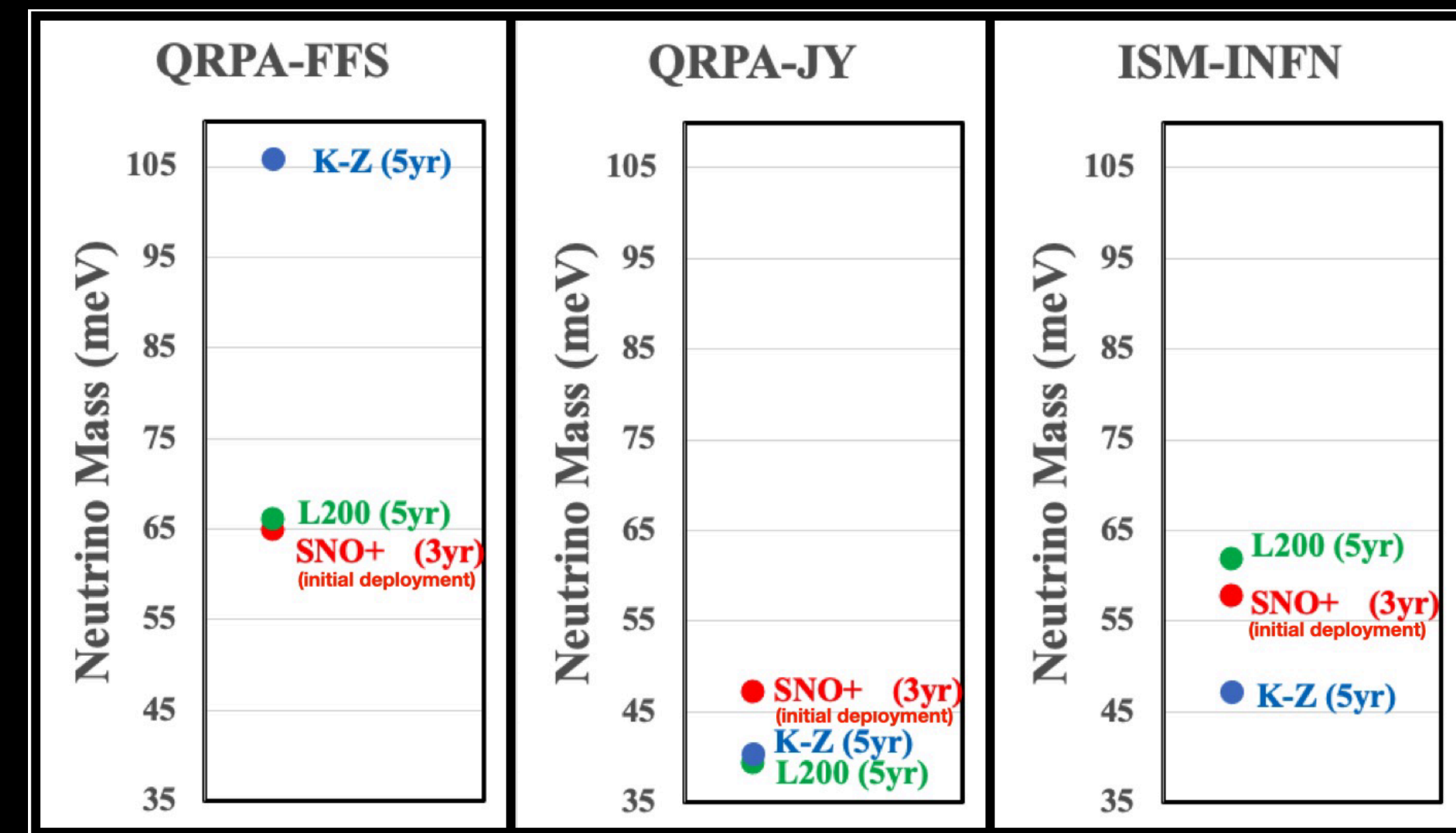


FV of 3.3 m (not optimised)

SNO+ Te-DBD

Additional Considerations

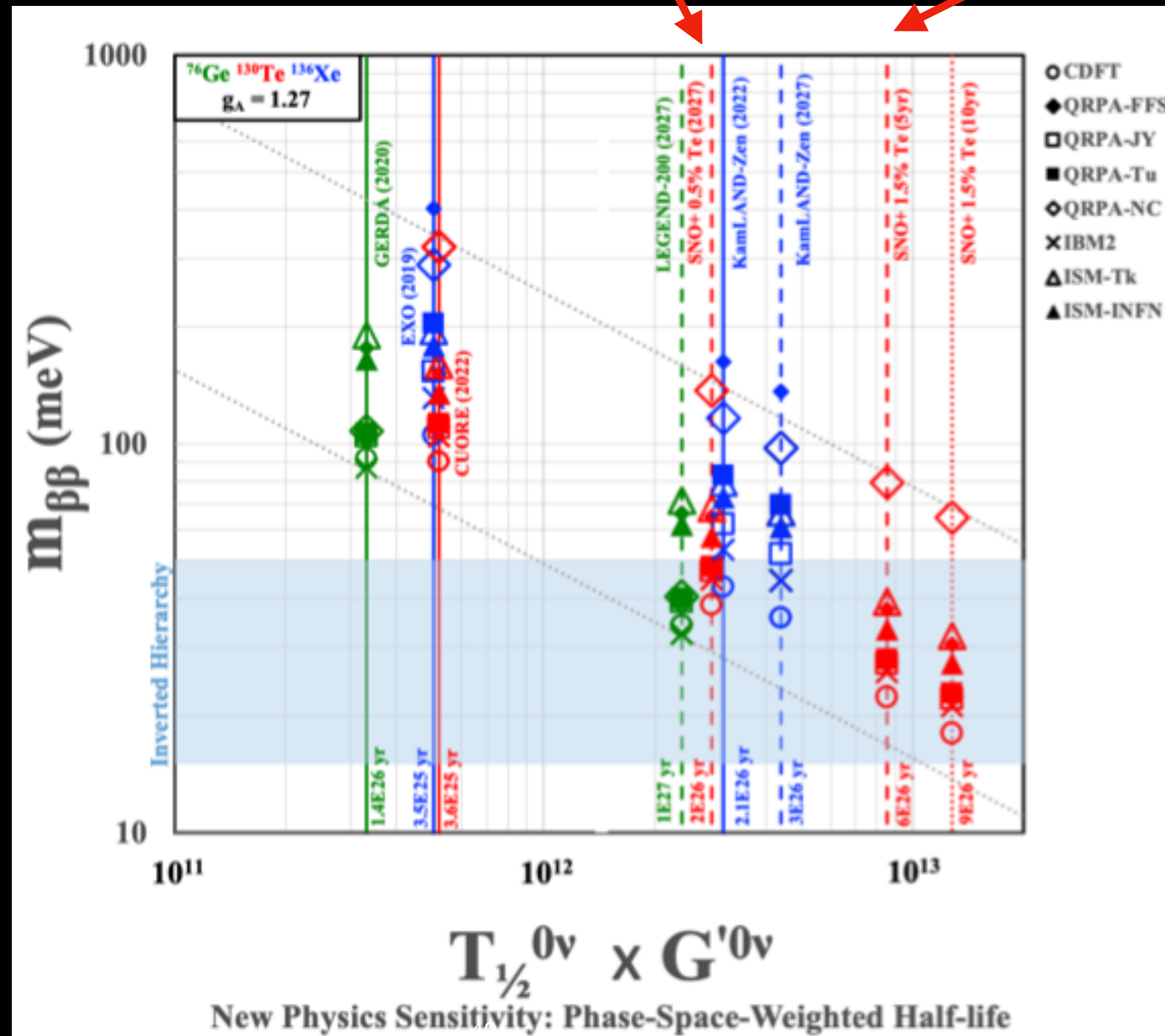
- ^{130}Te DBD is scalable, cost effective, unimpacted by geopolitical events that currently severely affect the availability of isotopic enrichment
- KL-Z 800 has world-leading sensitivity (upper limit 36-156 meV) and highlights the strength of the loaded LS DBD approach
- The competition does not make SNO+ Te DBD less relevant because of complementarity of isotope; NME model dependencies
- higher loading further extends SNO+ sensitivity and “fills the gap”, before larger experiments like nEXO come online
- initial deployment of Te would already be competitive and ready to test any hints of a positive signal
- purification of Te underground is novel technology
- “target out” analysis is a strong and unique feature; all non-Te backgrounds constrained prior to adding any Te
- SNO+ also has single-site/multi-site background constraining power



$0\nu\beta\beta$ sensitivity

2×10^{26} yr (0.5% Te)

6×10^{26} yr (1.5% Te)



Status of SNO+ Te-DBD

- Tellurium systems are built and ready for operation
 - SNO+ is ready to pursue **full-scale test batches in 2022 and 2023** to purify telluric acid and synthesize tellurium-diol, with SNOLAB assistance in this effort to help retire risks
 - Following demonstration of operations and subsequent approvals, aim to **begin loading Te in the detector in 2024** for the start of the double beta decay phase
- Meanwhile, the SNO+ project, with endorsement from SNOLAB, is completing R&D to **establish the viability and execution plans of Te loading at the 1.5%-2.0%** concentration, enabling reaching our goal of DBD sensitivity in the Inverted Mass Ordering region of parameter space



Telluric acid
purification



Te-Diol
synthesis

Summary

- SNO+ is an operating liquid scintillator neutrino detector filled with LAB + 2.2 g/L PPO and taking data
- Well understood detector from water phase data: backgrounds are low
 - Solar, nucleon decay, neutron capture, anti-neutrino physics
- Early scintillator data (partial fill and low PPO loading) analysed
 - scintillator performance and backgrounds, first scintillator physics measurements
- Diverse Scintillator Physics Program (some already underway)
 - $0\nu\beta\beta$, Solar, reactor- ν , geo- ν , Supernova ...
- Already-built underground tellurium plants represent novel technology in the field of low-radioactivity techniques
 - Operating the plants and demonstrating their capabilities is the next step towards preparing to load SNO+ with Te for the $0\nu\beta\beta$ phase
- Scalable approach of SNO+ has huge potential for extending $0\nu\beta\beta$ sensitivity