

ISCTE

Beyond CUPID:

bolometric technological developments toward the exploration of the normal hierarchy



Andrea Giuliani

IJCLab, Orsay, France



Laboratoire de Physique des 2 Infinis

Outline

- Prospects in double beta decay search the bolometric way
- Background components in bolometric experiments: challenges and mitigations
- R&D and solutions aiming at background control
 - \circ CROSS β surface sensitivity
 - BINGO set of innovative methods
 - US CUPID-R&D
 - CALDER Ight detectors with sensors different from the CUPID baseline
 - AMoRE

Double beta decay: status and prospects

Current generation (concluded – running – on-commissioning projects)

Next generation (projects to be started in the next decade)



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Prospects of some promising projects





Prospects of CUPID

Unlike liquid-helium-bath cryostat ("wet"), current **pulsetube cryostats ("dry")** – like the CUORE system – have **no cryogenic intrinsic constrain** on the **cryostat diameter**

Very large cryostats are possible

New available technologies allow for control of vibrations induced by pulse-tubes

Record energy resolution achieved on a 45x45x45 mm³ Li₂MoO₄ CUPID-like crystal in the Canfranc laboratory in the CROSS **pulse-tube cryostat**







Full α rejection down to BI < 10⁻⁵ counts/keV/kg/y is assumed

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Background for ¹⁰⁰Mo: status





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Paths for improvement: radiopurity



Solutions for background reduction



However, improvements can be obtained in the following items:

- Replace part of the internal lead shield with high purity copper in the existing CUORE cryostat – Redesign the shields with radiopure materials in new cryostats
- Improve copper radiopurity using advanced techniques (electroforming)
- Improve purification processes of Li₂CO₃ and MoO₃ powders used to form the Li₂MoO₄ molecule to be crystallized (R&D ongoing)

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Solutions for background reduction

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 US CUPID R&D – AMORE – CALDER





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β surface radioactivity from holder/crystals

These processes become challenging at the surface \rightarrow it may happen that α escapes detection and β is (partially) absorbed





CROSS: tagging surface events

CROSS rationale \rightarrow **metallic film coating** of crystals works as a **pulse-shape modifier** for surface events

First tests with **superconductive films (AI)** to decrease the heat capacity

 \rightarrow Expectations: risetime longer for near-film events

J. Low Temp. Phys. 167 (2012) 1029

 $T_{c}=1.2 \text{ K}$

- Al film captures high-energy (30 K) athermal phonons
- Energy is stored in quasiparticles for milliseconds
- Quasiparticles recombine and lower-energy athermal phonons at 1.2 K are emitted
- → delayed component of the signal in near-film events

→ Experience: this works only for an athermal-phonon sensor (deposited film)

Metallic coating



J. High Energ. Phys. (2020) 2020: 18 Appl. Phys. Lett. 118, 184105 (2021



Metallic coating

Athermal phonon

hot spot

~ 1 mm

Bulk

event

Phonon energy

down-conversion

Surface

event

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→ Experience: this works only for an athermal-phonon sensor (deposited film)

\rightarrow Reality: risetime shorter for near-film events!

- Tests were done with CUORE/CUPID readout: thermal phonon sensors (NTDs)
- The film converts athermal to low-energy phonons faster than spontaneous phonon decay
- \rightarrow thermalization is accelerated by the film in near-film events

To validate this interpretation, we decided to perform a direct test with a **double-sensor device** Both NbSi film and NTD Ge thermistor on the same Al-coated (one face) TeO₂ crystal

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Test on prototypes: the Al way

erc

SS



Test on prototypes: the Al way

erc

SS



Test on prototypes: the Al way

erc

SS



Test on prototypes: the Pd way

Potential **drawback of Al**: competition between accelerated thermalization and slowing quasiparticle trapping Move to a potentially more efficient thermalizer \rightarrow a **normal metal** \rightarrow **Pd** chosen for possibility of thin continuous films **Pd 10 nm thick film** \rightarrow much thinner film in order to reduce the normal electron heat capacity



Test on prototypes: the Pd-Al way

Pd-Al 10 nm-100 nm bilayer to recover heat sensitivity \rightarrow superconductivity induced in Pd by **proximity effect**

Quantitative test of the $\boldsymbol{\beta}$ surface sensitivity

- 1. Use α peaks to model the source depth profile
- 2. Generate the expected $\boldsymbol{\beta}$ spectrum by MonteCarlo
- 3. Fit real data considering background
- 4. Evaluate number of β counts contained in spectrum \rightarrow 3526(81)
- 5. Estimate this number independently by α counts \rightarrow 3455(273)



Test on prototypes: the Pd-Al way



Move from continuous films to grids

- Further decrease of the heat capacity
- Possibility to extract light → redundancy
- Geometrical parameters to optimize surface sensitivity



Current objectives

- Transfer technology to fully-coated large crystals
- Test reproducibility in the CROSS demonstrator (Canfranc underground laboratory)



Bases of the BINGO technology

Isotopes

Q = 3034 keV > 2615 keV – AI: 9.7% Q = 2527 keV < 2615 keV – AI: 34%

Crystals



¹⁰⁰Mo

¹³⁰Te

Excellent bolometric properties High radiopurity ______ CUPID-Mo Extensively tested CUORE

heat-light Li₂MO₄: scintillator
 double readout TeO₂: only Cherenkov light ~

Feeble light yield:
Outstanding light detector required

NTD readout / Electronics —— Keep successful CUORE / CUPID-Mo solutions



arXiv:2204.14161

The BINGO technology is based on three innovations in the detector structure and conception

Crystal + light detector (Ge) $(Li_2MoO_4 \text{ or } TeO_2)$



The BINGO technology is based on **three innovations** in the detector structure and conception

① Innovative assembly of the detectors

The crystal will be surrounded only by active elements

 \rightarrow Mitigate surface radioactivity by anticoincidence



BINGO

erc

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² Full active shield inside the experimental space,

for the first time in bolometric technology. The shields will consist of BGO/ZnWO₄ radiopure scintillators with bolometric light read-out \rightarrow Mitigate external γ background





ements in **BINGO**

novations in the detector structure and conception



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DBD: the road to NH sensitivity - Lisboa, 7/6/2022



Study on active shield materials

Essential requirements to suppress the background from the ROI of TeO_2 :

If a 2615 keV γ deposit a small amount of energy in the surrounding material (~80 keV) and the rest in TeO₂ \rightarrow bkg in ROI (2527 keV)

• Very low energy threshold of the light detector \rightarrow low threshold of ZnWO₄ or BGO LD Neganov-Trofimov-Luke LD to increase signal to noise ratio \rightarrow low threshold Active shield Surrounding materials The target BINGO threshold is 50 keV in ZnWO₄ and BGO **BGO** results Light detector output gammas Energy spectrum NTD Ann+511 keV 352 keV 207Bi 570 keV+583 keV+609 keV Q ^{1.01} ► 0 600 bolometer Compton 207Bi 911 keV 400 968 keV scattering 0.99 muons 2615 keV 200 Th source 200 2000 2500 scintillator Energy (keV) Energy (keV)

50% efficiency threshold in BGO: ~ 20 keV without NTL effect (< 5 keV feasible) Light yield in $ZnWO_4$ about ½ of BGO



The BINGO technology is based on three innovations in the detector structure and conception

① Innovative assembly of the detectors
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 → Mitigate surface radioactivity by anticoincidence

② Full active shield inside the experimental space, for the first time in bolometric technology. The shields will consist of BGO/ZnWO₄ radiopure scintillators with bolometric light read-out \rightarrow Mitigate external γ background

③ Innovative light detector based on the Neganov-Trofimov-Luke (NTL) effect

Increase by a factor 10-20 current sensitivity reach < 10 eV rms baseline width – Multipurpose :

$\rightarrow \alpha/\beta$ discrimination in TeO_2 by Cherenkov light

- \rightarrow Read-out scintillator light of the active shield
- \rightarrow Increase S/N to perform pile-up rejection sensitivity Lisboa,



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erc "White" and "grey" zones in BINGO

Top view of the arrangement





BUNGO

Large scale implementation

BUNGO





MINI-BINGO demonstrator

MINI-BINGO is a small-scale demonstrator of the BINGO technology in a dedicated underground cryostat

[Modane laboratory]



Validation of all the BINGO elements

- 2 isotopes: ¹⁰⁰Mo and ¹³⁰Te
- 2 towers of 12 crystals each
- Crystals will see nothing else that is not active
- BGO or ZnWO₄ crystals as active shields
- Innovative light detectors

Data taking: 2024

Simulations in progress to evaluate the external background suppression factor in MINI-BINGO and in large-scale experiments (as a function of threshold, thickness, geometry)

TES-based light detectors: concepts

Current features of NTD-based light detectors (CUPID baseline) used in CUPID-Mo, CUPID-0, CUPID tests: \rightarrow Rise-times ~ 1-2 ms – Baseline width ~ 100 eV FWHM (improvable by a factor ~ 10 by NTL effect)

→ Limited by gluing, electron-phonon decoupling, high impedance readout

A new phonon-sensor technology can substabtially improve these figures mitigating the 2ν pile-up background



Time resolution with TES can be much better than with NTD

- Low impedance and SQUID readout Multiplexing
- Direct coupling to the wafer without glue
- Technical bolometric feature: negative electrothermal feedback

→ Objective: ≤ 150 µs rise-time

CUPID-R&D

ANL, UCB, LBL

is possible

TES-based light detectors: devices

- Control of the critical temperature T_c via proximity effect
- IrPt bilayer \rightarrow T_c in the 20-110 mK range
- Variable TES sizes: 300-500 μm x 300-500 μm
- Si wafer 5.08 cm Ø





Complex thermal network Important to **increase TES-wafer coupling**

B. Welliwer, 15th Pisa Meeting on Advanced Detectors 2022

ANL, UCB, LBL **TES-based light detectors: performance**



Gold pads to improve the TES **thermal conductance** to the Si



Test with optical fibers



- Allows for pulse-shape discrimination for pile-up events
- Baseline noise measured with photon statistics: ~ 120 eV (FWHM)

Implementation in CUPID

At least 1500 TESs are required for light detector readout Frequency domain multiplexing \rightarrow multiplexing factor of the order of 10

calder Content of the calder o

Cooper pairs in a superconductor act as an inductance

A KID is a **superconductive devices** that detect **photons** and **phonons** through the **change of its inductance** induced by a change of Cooper pair density

A KID is included in a **resonating circuit** biased with a microwave signal (GHz)

Signal from amplitude and phase shift



"Natural" multiplexing up to a factor 1000



Final results from CALDER

- Si wafer 5x5 cm²
 - Single AlTiAl three-layer resonator

- Baseline noise measured with photon statistics: ~ 210 (FWHM)
 → limitied by microphonic noise
- Pulse rise-time: 120 μs

EPJC 81, 636 (2021)

AMoRE approach: MMC readout

AMoRE is a bolometric experiment searching for neutrinoless double beta decay of ¹⁰⁰Mo with the **Magnetic Metallic Calorimeter (MMC)** technology

An **MMC** is a device that measures a given amount of heat deposited in it through a **magnetization change**



An MMC requires the application of a **magnetic field** and a **SQUID readout**

arXiv:1512.05957

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Light detector



- **Si wafer** 5.08 cm Ø
- Rise-time 200 μs, independent of operation temperature



Summary



Full α rejection down to BI < 10⁻⁵ counts/keV/kg/y is assumed

Active veto

Two types of scintillating crystals have been successfully tested at 10 mK with bolometric light detector readout: BGO and ZnWO₄ Prototype barrel under preparation [BINGO]

Surface sensitive detectors

Surface sensitivity to β particles by pulse shape discrimination assisted by film coating proved in small-scale prototypes Trasnsfer to final crystals in progress [CROSS]

Redesign the holder structure

Almost full active surrounding for each crystal Proof of concept achieved [BINGO]

Increase light detector signal-to-noise ratio Exploit Neganov-Trofimov-Luke effect + NTD readout

Proof of concept achieved

[BINGO]

Increase light detector speed

Several possible solutions:

TES – tested in prototype [US CUPID-R&D]

KIDs – tested in prototypes [CALDER]

MMC – adopted in [AMoRE]