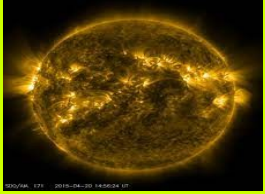


neutrino theory

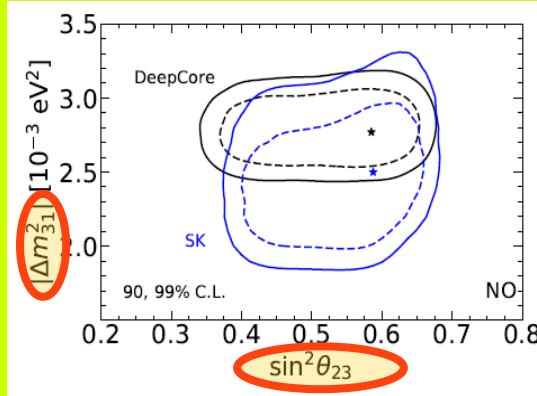
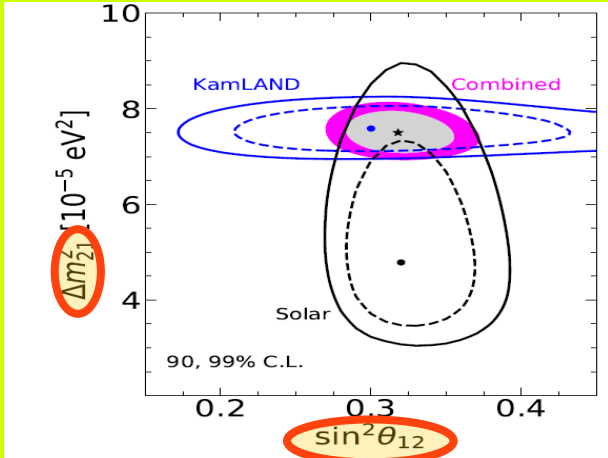
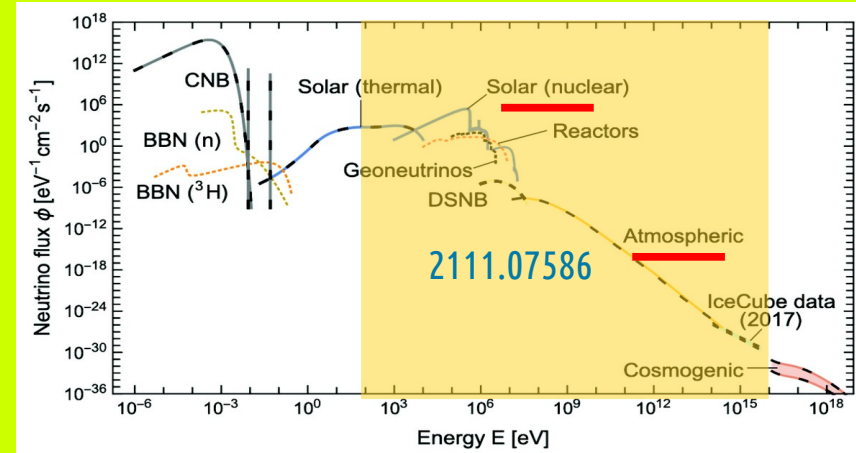
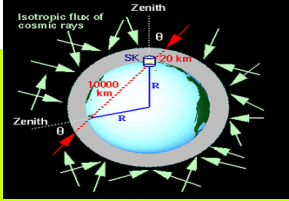
JOSÉ W F VALLE

Double Beta Decay: the road
to normal hierarchy sensitivity

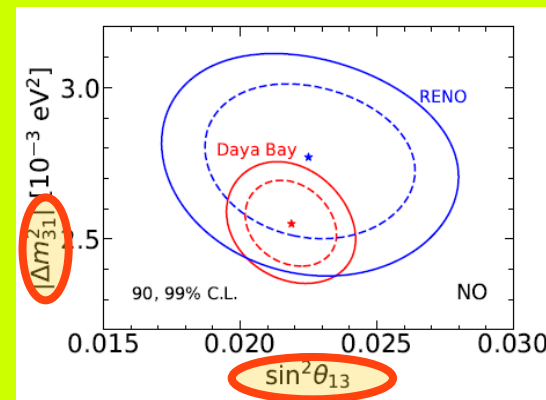
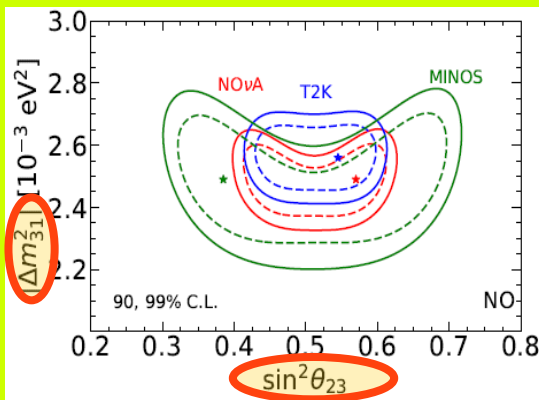
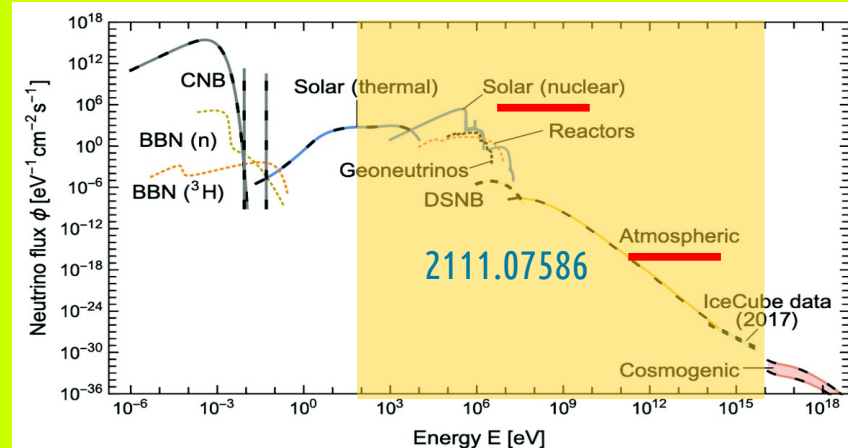
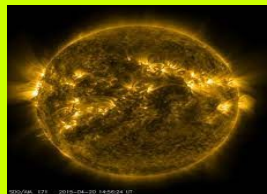
ASTROPARTICLES
Astroparticles and High Energy Physics Group



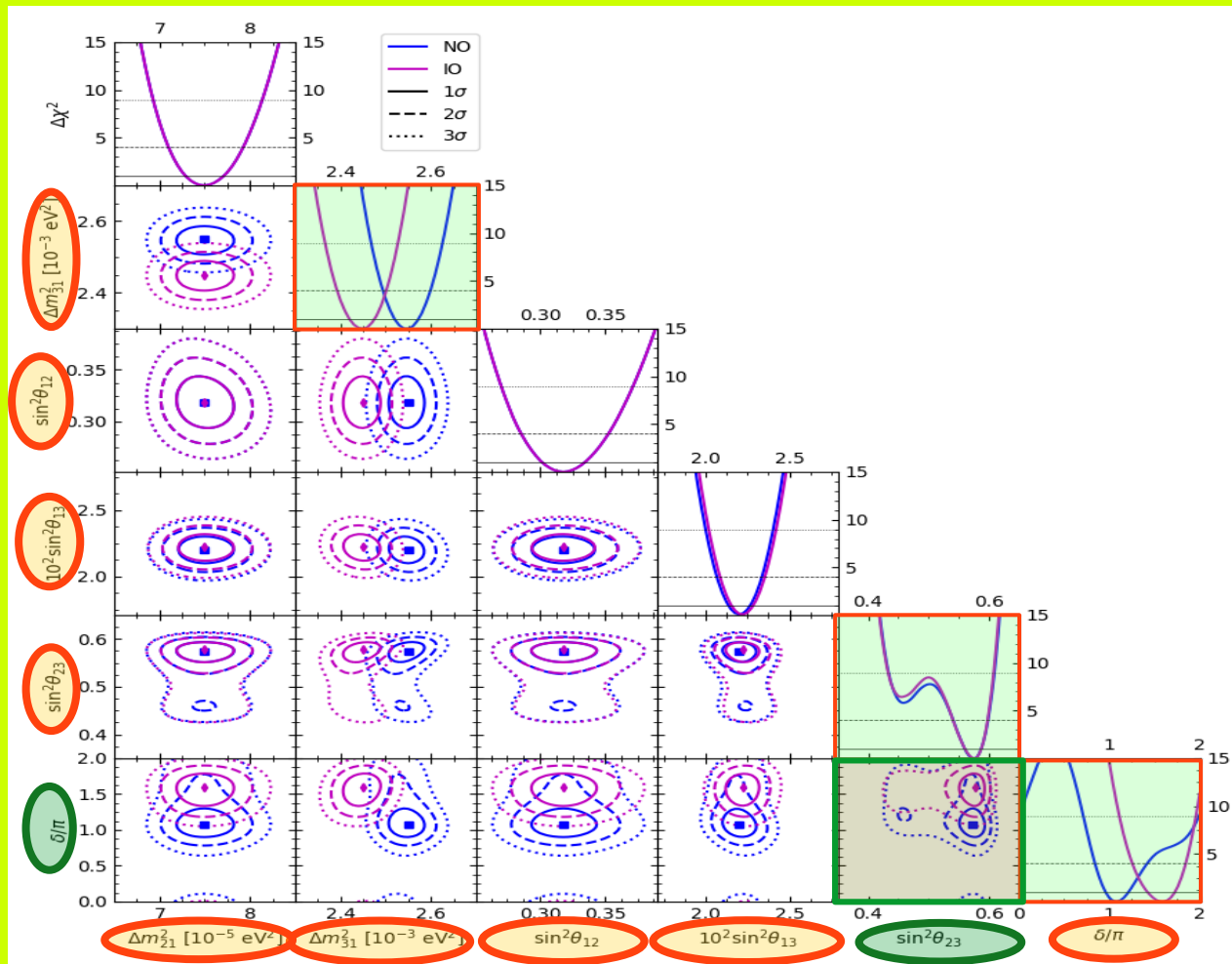
oscillations



P.F. de Salas et al JHEP 02 (2021) 071



neutrino oscillation maps

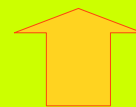


P.F. de Salas et al JHEP 02 (2021) 071

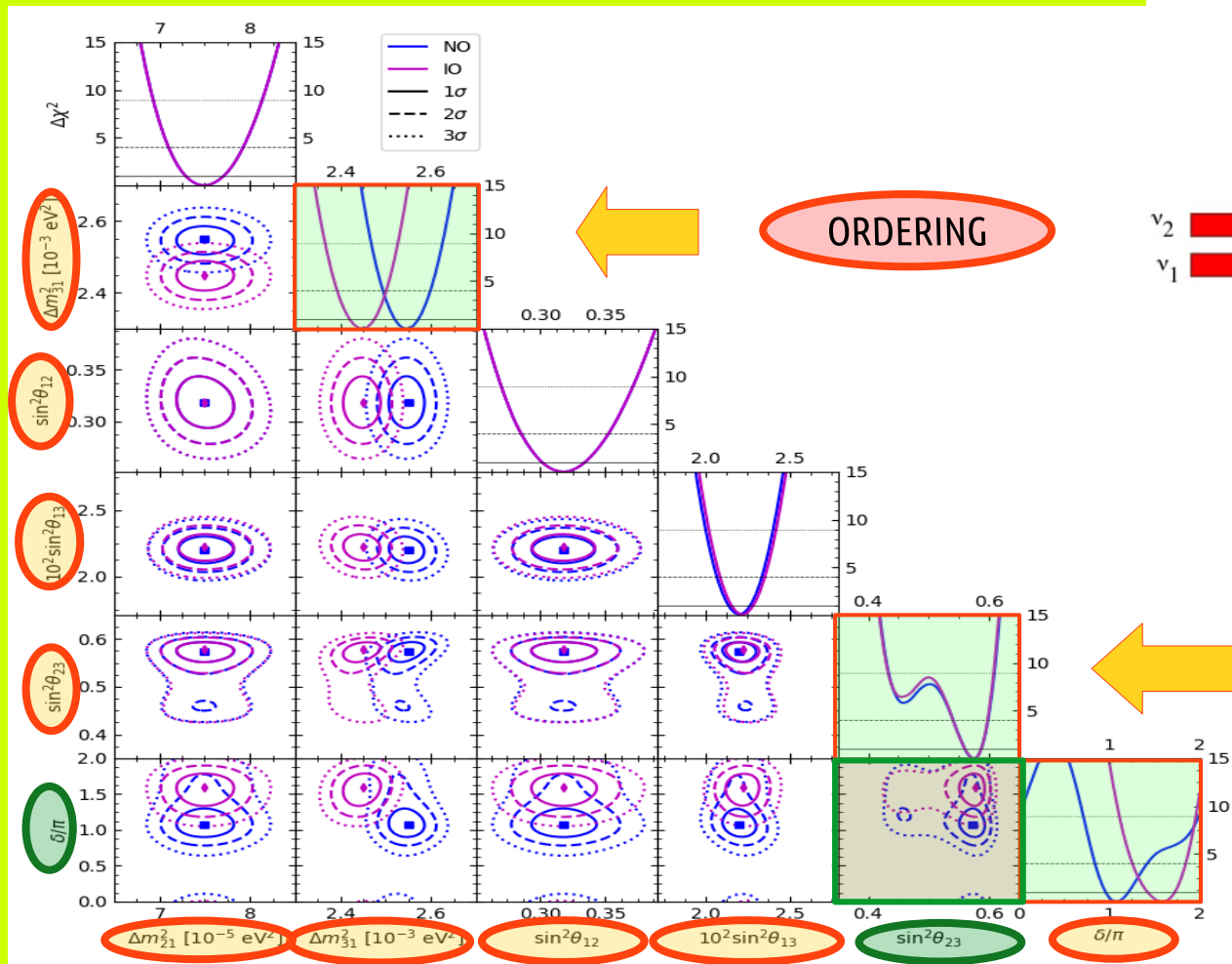
Chi2 maps downloadable from
<https://globalfit.astroparticles.es/>

@jwvalle 3

pairwise correlations



neutrino oscillation maps



NORMAL

INVERTED



ν_e

ν_μ

ν_τ

NO

IO



OCTANT

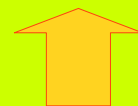
CP

P.F. de Salas et al JHEP 02 (2021) 071

Chi2 maps downloadable from
<https://globalfit.astroparticles.es/>

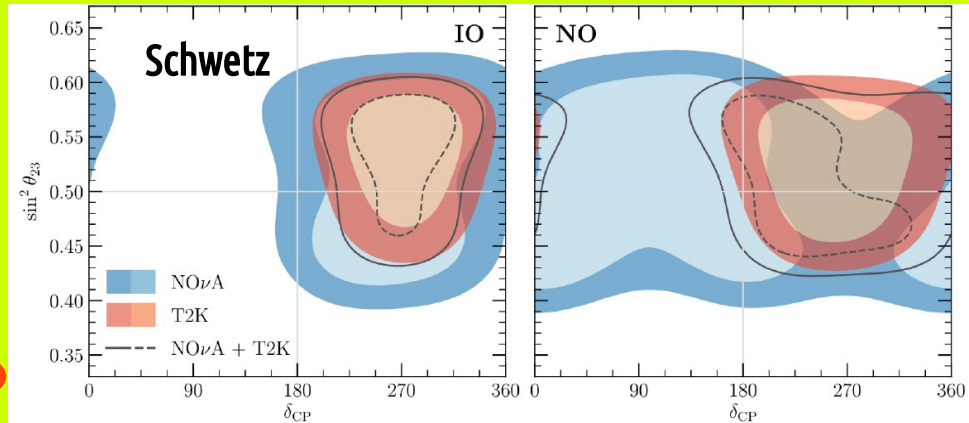
@jwvalle 3

pairwise correlations



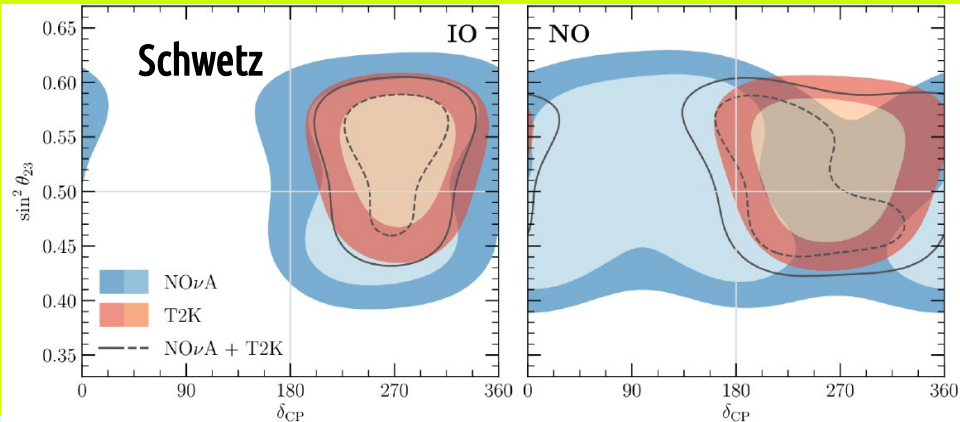
CP

ORDERING



CP

ORDERING



Sanford Underground Research Facility

EXISTING LABS
UNDERGROUND PARTICLE DETECTOR

800 miles (1300 kilometers)

PARTICLE DETECTOR

NEUTRINO PRODUCTION

PROTON ACCELERATOR

Fermilab

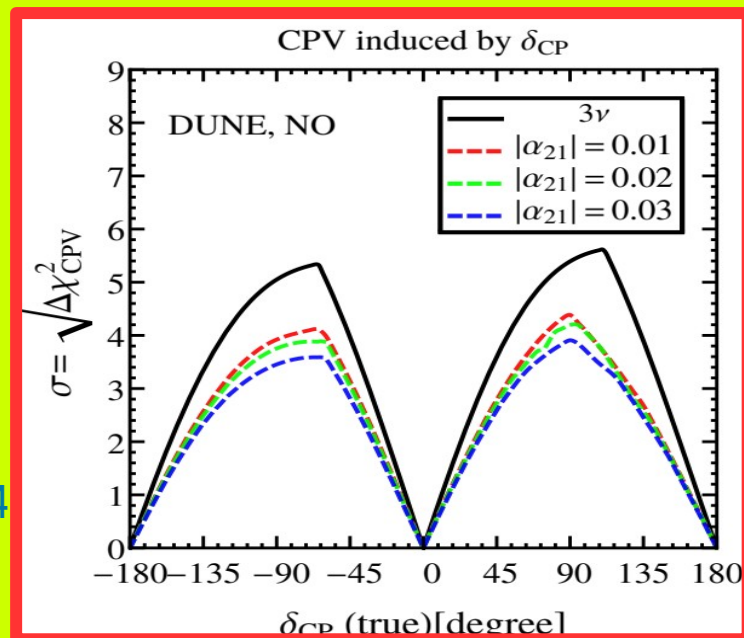
DUNE

Expected Sensitivity:
3nu vs Unitarity
violation

Phys.Rev.Lett.117 (2016) 061804

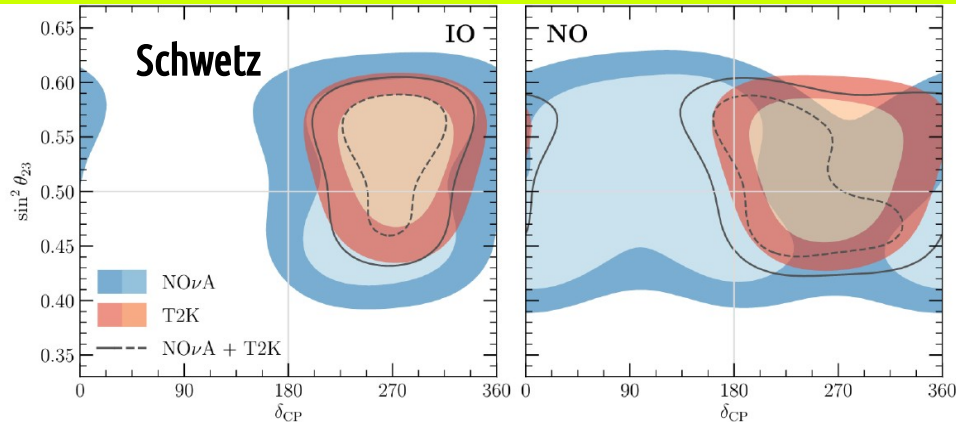
New J.Phys. 19 (2017) 9, 093005

PhysRevD97 (2018) 095026



CP

ORDERING



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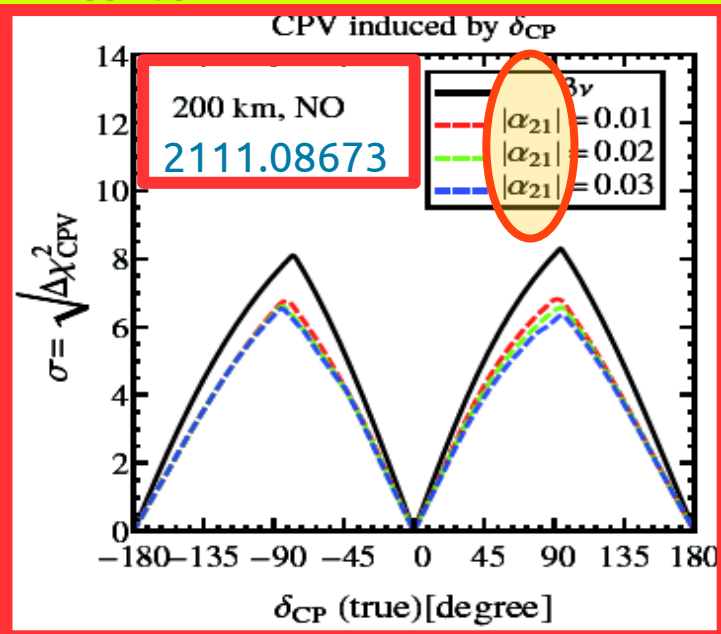
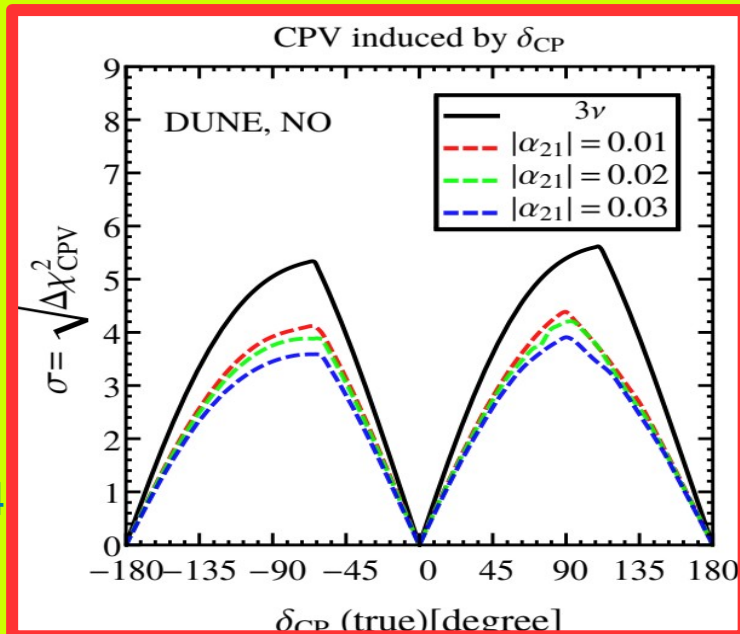
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Phys.Rev.Lett.117 (2016) 061804

New J.Phys. 19 (2017) 9, 093005

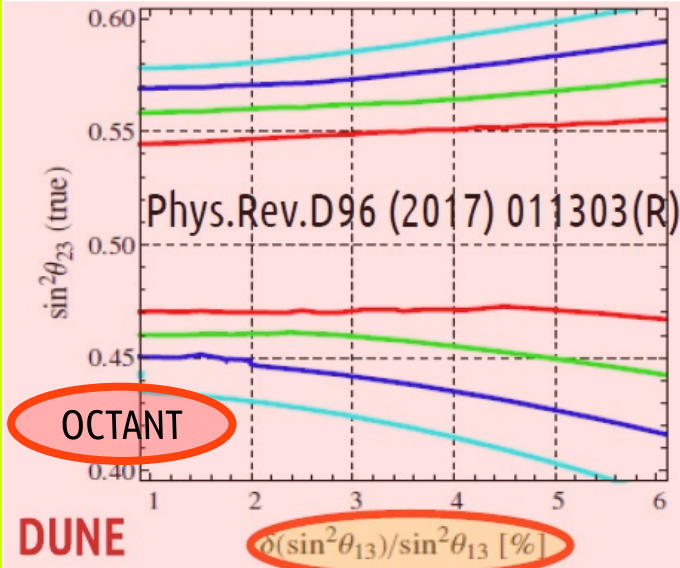
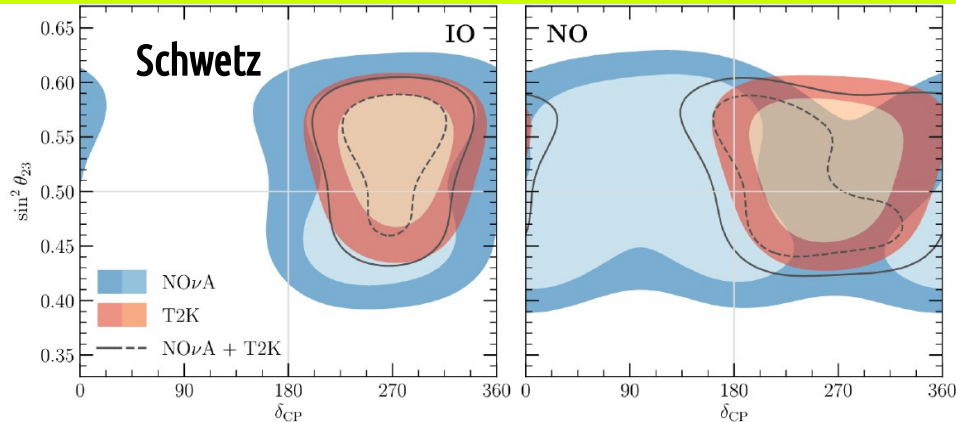
PhysRevD97 (2018) 095026

ESSnuSB



CP

ORDERING



ESSnuSB

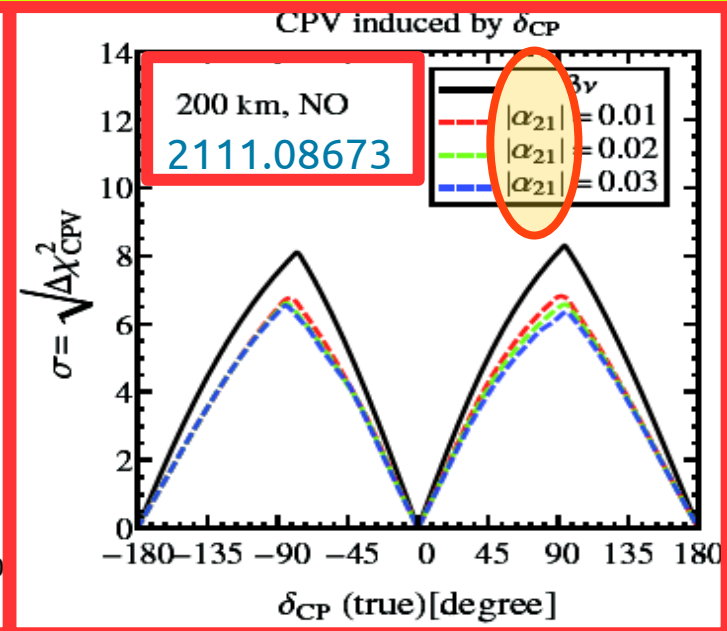
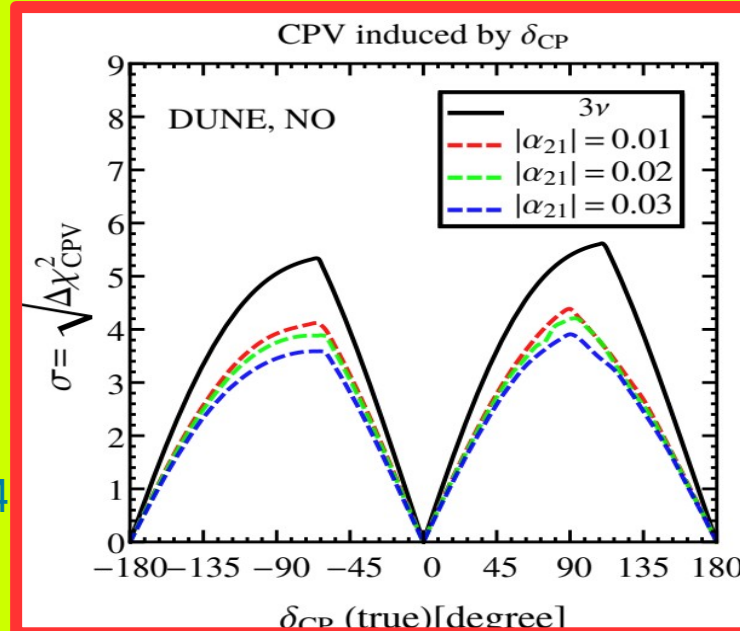
DUNE

Expected Sensitivity:
3nu vs Unitarity
violation

Phys.Rev.Lett.117 (2016) 061804

New J.Phys. 19 (2017) 9, 093005

PhysRevD97 (2018) 095026



@jwvalle4

neutrino properties from colliders

current oscillation data
can reconstruct m_{ν}
so that it can be tested
at high-energies

Miranda et al Phys.Rev.D105 (2022) 095020

seesaw mediator produced in
@ e^+e^- / pp collisions



Miranda et al PLB 829 (2022) 137110

neutrino properties from colliders

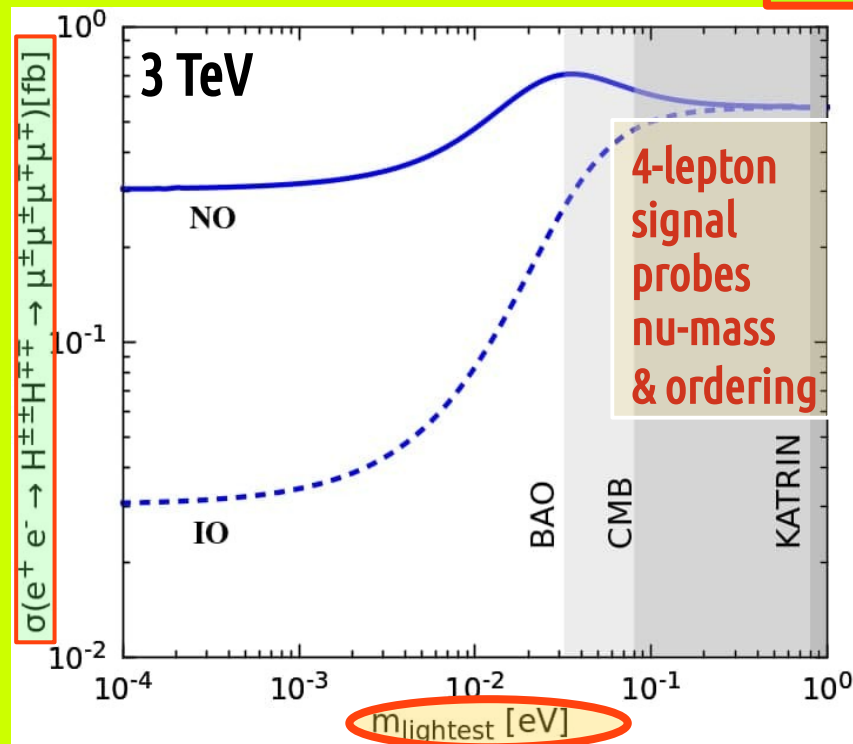
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Miranda et al Phys.Rev.D105 (2022) 095020

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Miranda et al PLB 829 (2022) 137110



neutrino properties from colliders

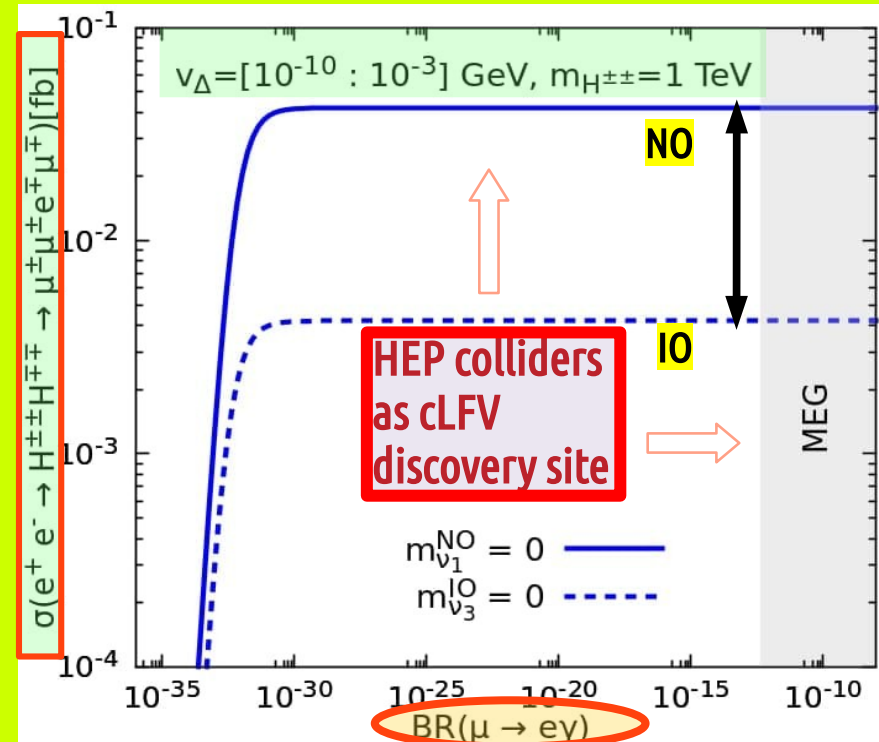
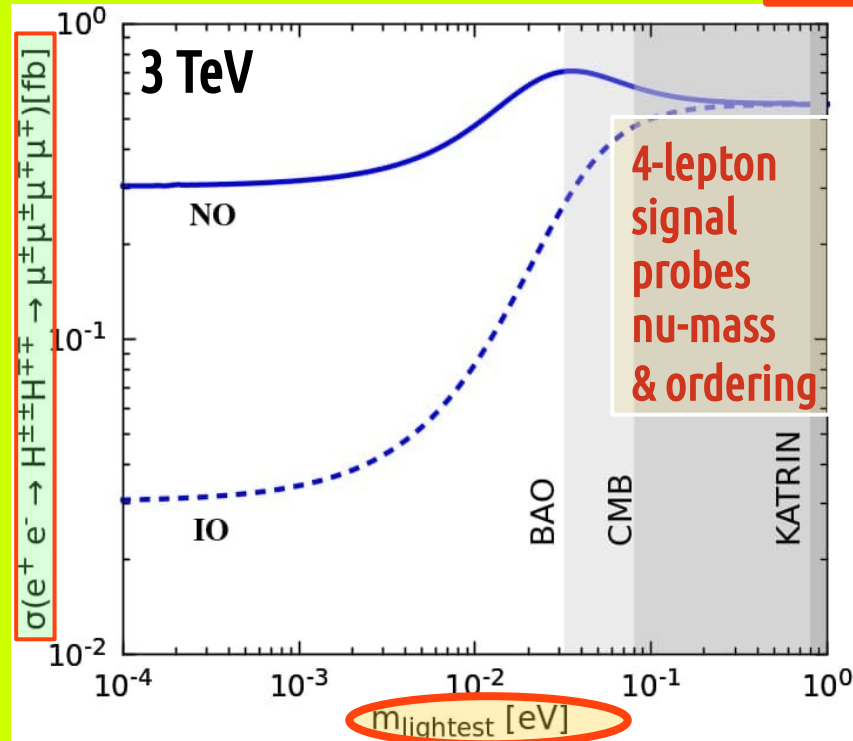
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Miranda et al Phys.Rev.D105 (2022) 095020

seesaw mediator produced in
@ e+e- / pp collisions



Miranda et al PLB 829 (2022) 137110



TBM interpretation

Harrison,
Scott
& Perkins
2002

$$\begin{bmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{bmatrix}$$

θ_{13}



CP

systematic revamping

Chen et al

Phys.Lett. B753 (2016) 644

Phys.Rev. D94 (2016) 033002

JHEP 1807 (2018) 077

Phys.Lett. B792 (2019) 461

Phys.Rev. D99 (2019) 075005

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θ_{13}



CP

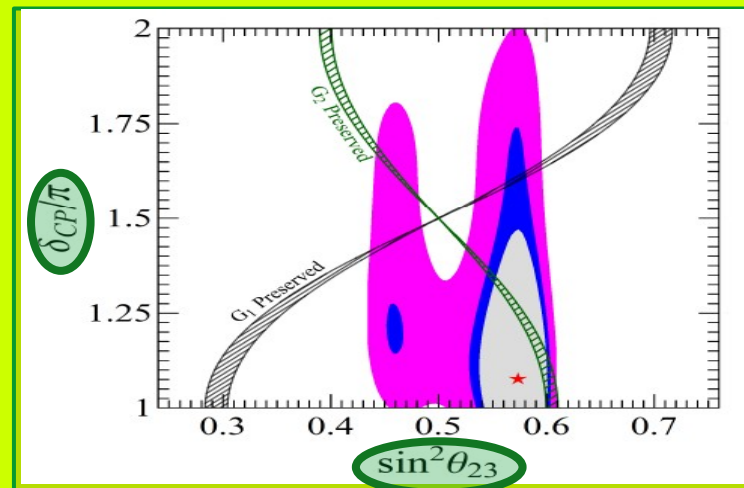
Phys.Rev.D98(2018)055019

$$\sin^2 \theta_{12} \cos^2 \theta_{13} = \frac{1}{3}$$

$$\tan 2\theta_{23} \cos \delta_{CP} = \frac{\cos 2\theta_{13}}{\sin \theta_{13} \sqrt{2 - 3 \sin^2 \theta_{13}}}$$

systematic revamping

Chen et al
Phys.Lett. B753 (2016) 644
Phys.Rev. D94 (2016) 033002
JHEP 1807 (2018) 077
Phys.Lett. B792 (2019) 461
Phys.Rev. D99 (2019) 075005



many examples

Bi-Large lepton mixing pattern

$$\begin{bmatrix} 1 - \frac{1}{2}\lambda^2 & -\lambda e^{i\phi} & A\lambda^3 e^{i\phi} \\ \lambda e^{-i\phi} & 1 - \frac{1}{2}\lambda^2 & -A\lambda^2 \\ 0 & A\lambda^2 & 1 \end{bmatrix} \quad \begin{bmatrix} 1 - \frac{5\lambda^2}{2} & 2\lambda & -\lambda \\ -2\lambda + 3\lambda^2 & 1 - \frac{13\lambda^2}{2} & 3\lambda \\ \lambda + 6\lambda^2 & -3\lambda + 2\lambda^2 & 1 - 5\lambda^2 \end{bmatrix}$$

Largest Q-mixing similar to smallest L-mixing
Cabibbo angle as universal seed for flavor
mixing

Phys.Rev. D86 (2012) 051301

Phys.Rev.D87 (2013) 053013

Phys.Lett. B748 (2015) 1-4

$\sin \theta_{12}^{\text{CKM}} = \lambda$ and $\sin \theta_{23}^{\text{CKM}} = A\lambda^2$, where $\lambda = 0.22453 \pm 0.00044$, $A = 0.836 \pm 0.015$

predicting solar & atm

Bi-Large lepton mixing pattern

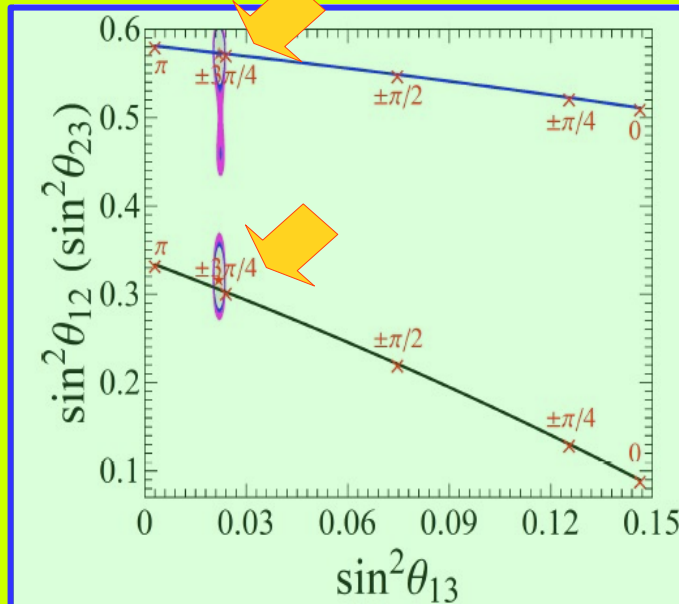
$$\begin{bmatrix} 1 - \frac{1}{2}\lambda^2 & -\lambda e^{i\phi} & A\lambda^3 e^{i\phi} \\ \lambda e^{-i\phi} & 1 - \frac{1}{2}\lambda^2 & -A\lambda^2 \\ 0 & A\lambda^2 & 1 \end{bmatrix} \quad \begin{bmatrix} 1 - \frac{5\lambda^2}{2} & 2\lambda & -\lambda \\ -2\lambda + 3\lambda^2 & 1 - \frac{13\lambda^2}{2} & 3\lambda \\ \lambda + 6\lambda^2 & -3\lambda + 2\lambda^2 & 1 - 5\lambda^2 \end{bmatrix}$$

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predicting solar & atm



many probes of neutrino mixing hypotheses
e.g. at DUNE

e.g. Phys.Rev.D97(2018)095025

From
Phys.Lett. B792 (2019) 461

Less tight realization
Phys.Lett.B 796 (2019) 162

@jwvalle 7

neutrinoless double beta decay

$$\left| \sum_j U_{ej}^2 m_j \right| = \left| c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{2i\phi_{12}} + s_{13}^2 m_3 e^{2i\phi_{13}} \right|$$

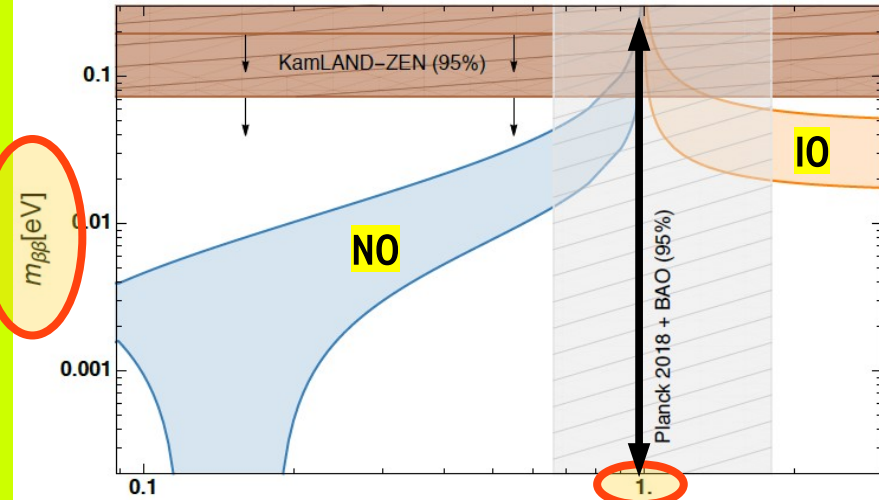
Schechter & JV PRD22 (1980) 2227

Rodejohann, JV Phys.Rev. D84 (2011) 073011

➤ Quasi-degenerate strongly-disfavored

Lattanzi, Gerbino, Freese, Kane, Valle
JHEP 10 (2020) 213

$m_{\beta\beta}[\text{eV}]$



η

degeneracy parameter

neutrinoless double beta decay

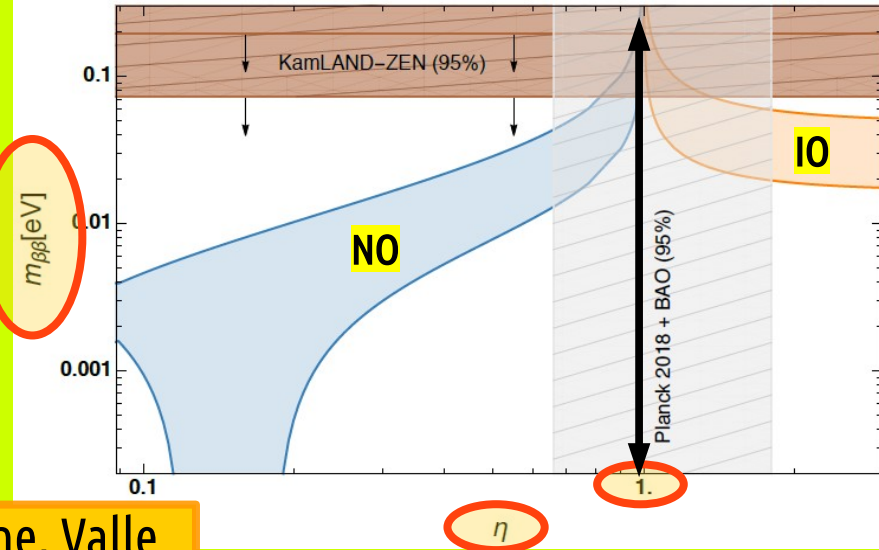
$$\left| \sum_j U_{ej}^2 m_j \right| = \left| \underbrace{c_{12}^2 c_{13}^2 m_1}_{\text{NO}} + \underbrace{s_{12}^2 c_{13}^2 m_2 e^{2i\phi_{12}}}_{\text{IO}} + \underbrace{s_{13}^2 m_3 e^{2i\phi_{13}}}_{\text{IO}} \right|$$

Schechter & JV PRD22 (1980) 2227

Rodejohann, JV Phys.Rev. D84 (2011) 073011

➤ Quasi-degenerate strongly-disfavored

Lattanzi, Gerbino, Freese, Kane, Valle
JHEP 10 (2020) 213



degeneracy parameter

➤ One-massless neutrino

Reig et al Phys.Lett. B790 (2019)303

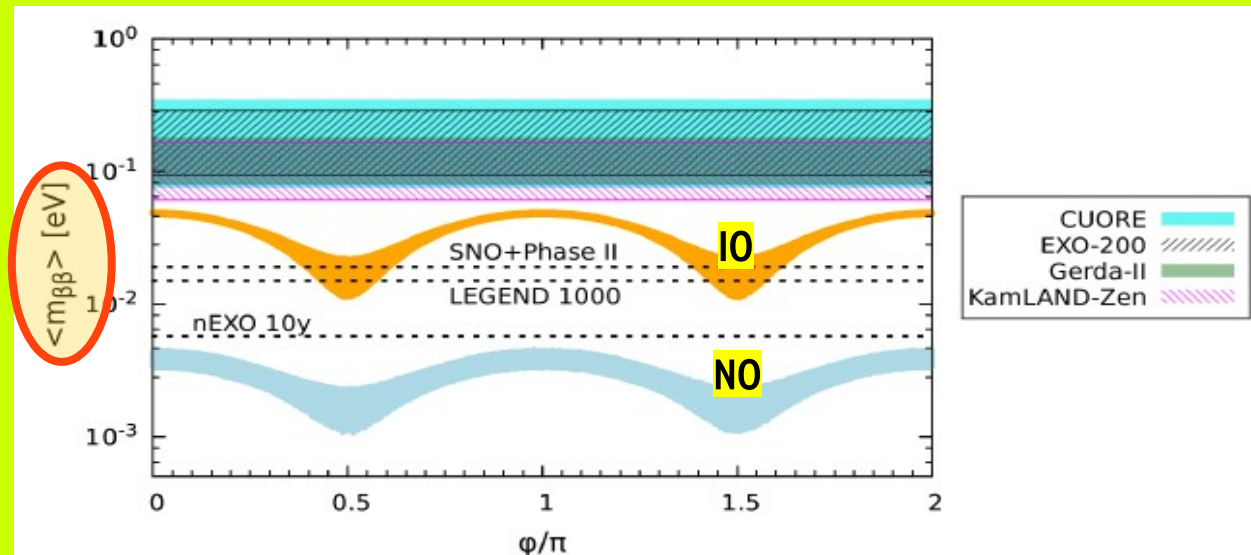
Barreiros, Felipe & Joaquim JHEP (2019) 223

Mandal et al PLB789 (2019) 132

Avila et al Eur.Phys.J.C 80 (2020) 10, 908

Agostini et al. Science 365 (2019) 1445

Final Gerda ... 2009.06079



majorana phase

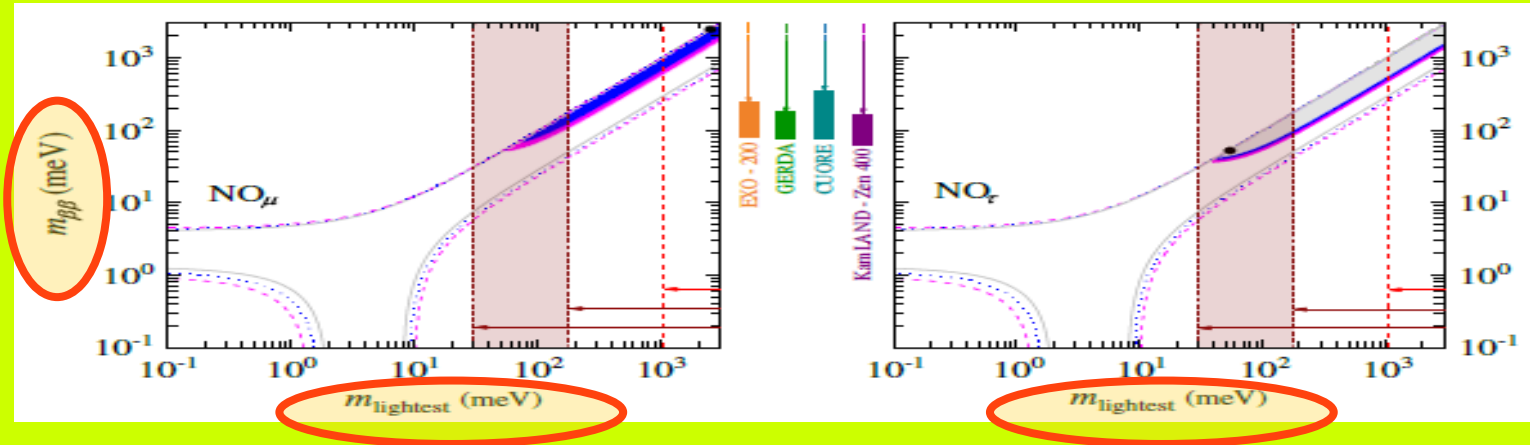
neutrinoless double beta decay

Lower bounds from family symmetries

Dorame et al PhysRevD86(2012)056001

Dorame et al Nucl.Phys.B 861 (2012) 259-270

King et al Phys.Lett. B 724 (2013) 68-72 etc



From Barreiros et al JHEP04(2021)249

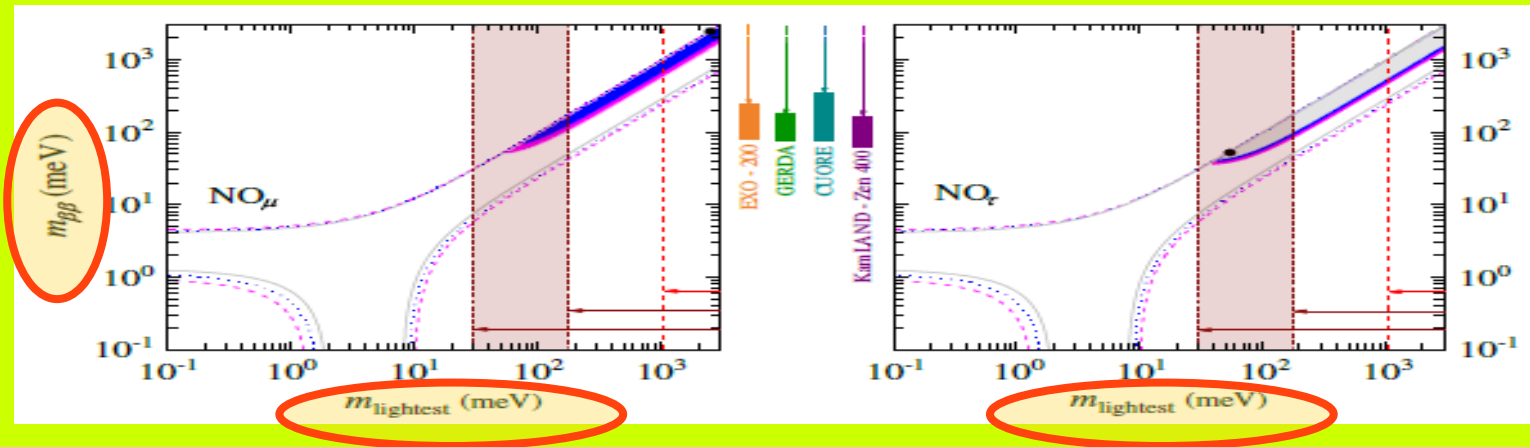
neutrinoless double beta decay

Lower bounds from family symmetries

Dorame et al PhysRevD86(2012)056001

Dorame et al Nucl.Phys.B 861 (2012) 259-270

King et al Phys.Lett. B 724 (2013) 68-72 etc

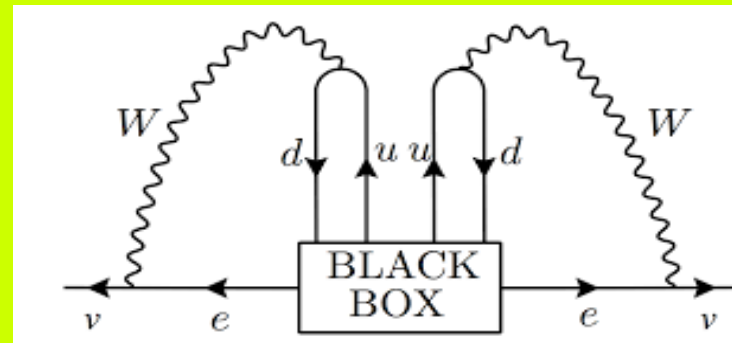


From Barreiros et al JHEP04(2021)249

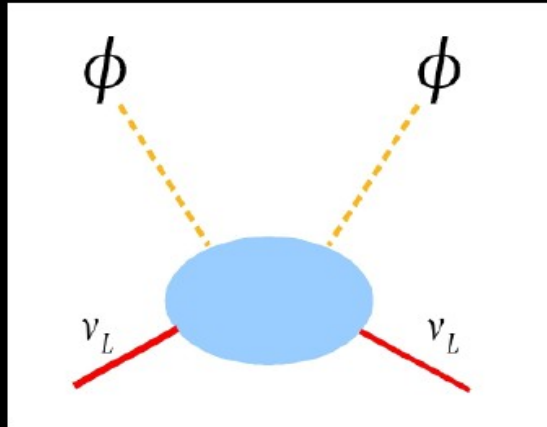
Schechter, Valle PhysRev D25 (1982) 2951

Duerr, Lindner, Merle JHEP06(2011)091

B.J.P. Jones 2108.09364 (TASI 2020)



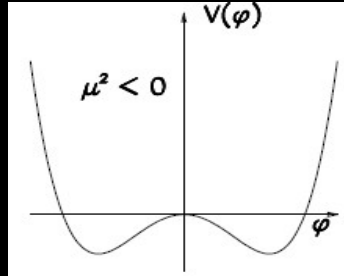
Origin of neutrino mass



Origin of neutrino mass

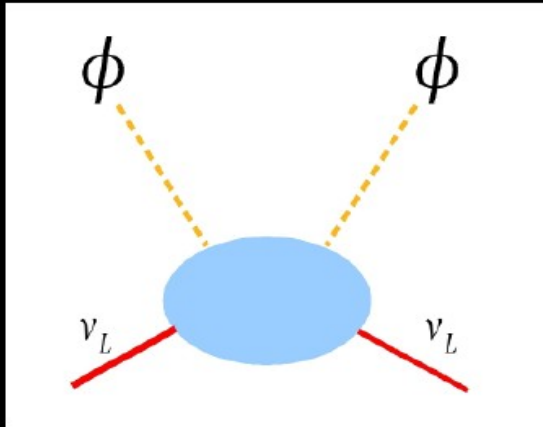
SEESAW dynamics

$$v_3 v_1 \sim v_2^2$$



Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



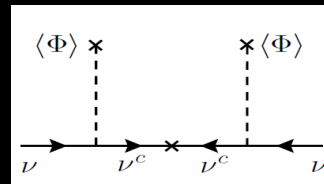
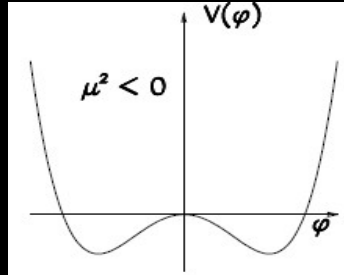
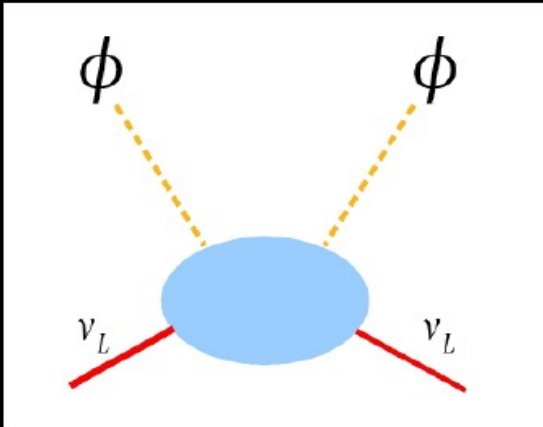
Origin of neutrino mass

SEESAW dynamics

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Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



TYPE I

Minkowski 77

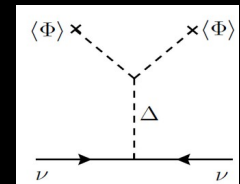
Gellman Ramond Slansky 80

Glashow, Yanagida 79

Mohapatra Senjanovic 80

Lazarides Shafi Wetterich 81

Schechter-Valle 80 & 82



TYPE II

Schechter-Valle 80 & 82

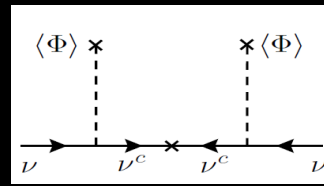
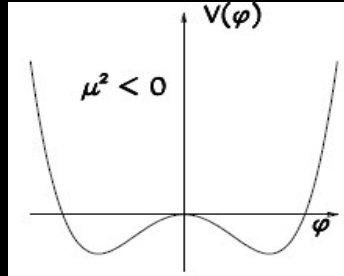
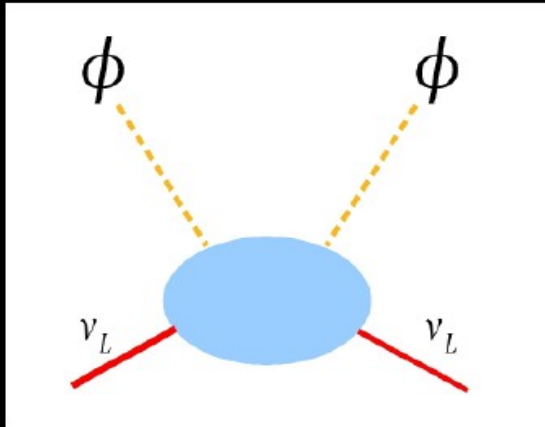
Origin of neutrino mass

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Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



TYPE I

Minkowski 77

Gellman Ramond Slansky 80

Glashow, Yanagida 79

Mohapatra Senjanovic 80

Lazarides Shafi Wetterich 81

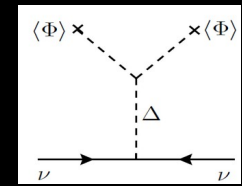
Schechter-Valle 80 & 82

L-R Type1 seesaw

SM Type1 seesaw

of Rs = # Ls

of singlets arbitrary



TYPE II

Schechter-Valle 80 & 82

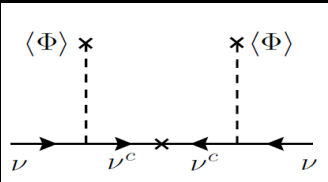
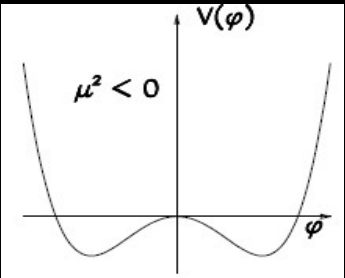
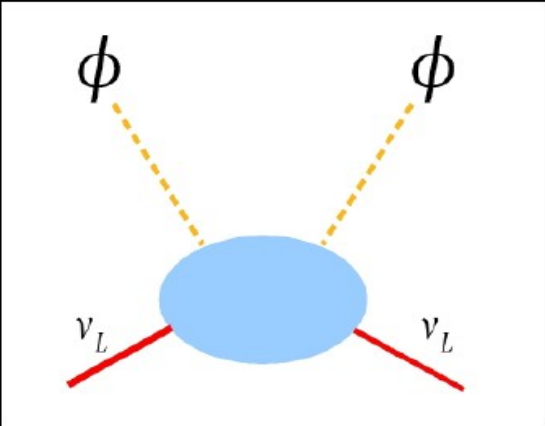
Origin of neutrino mass

SEESAW dynamics

$$v_3 v_1 \sim v_2^2$$

Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



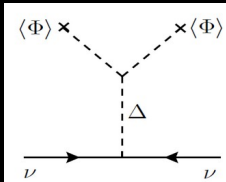
TYPE I

Minkowski 77
 Gellman Ramond Slansky 80
 Glashow, Yanagida 79
 Mohapatra Senjanovic 80
 Lazarides Shafi Weterrich 81
 Schechter-Valle 80 & 82

L-R Type1 seesaw

SM Type1 seesaw

■ MISSING PARTNER SEESAW



TYPE II

Schechter-Valle 80 & 82

of Rs = # Ls

of singlets arbitrary

(3,2) min viable type1 seesaw
 (3,1) scoto-seesaw template

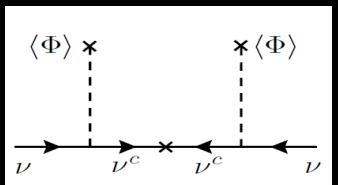
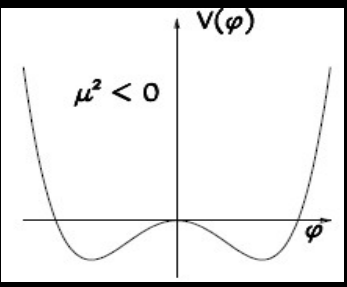
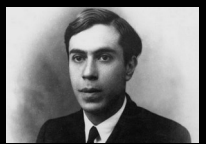
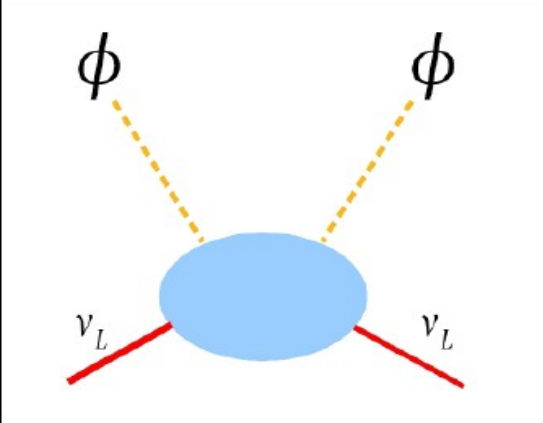
Origin of neutrino mass

SEESAW dynamics

$$v_3 v_1 \sim v_2^2$$

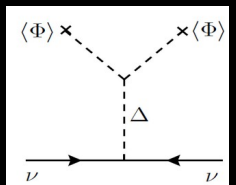
Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



TYPE I

Minkowski 77
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TYPE II

Schechter-Valle 80 & 82

L-R Type1 seesaw

of Rs = # Ls

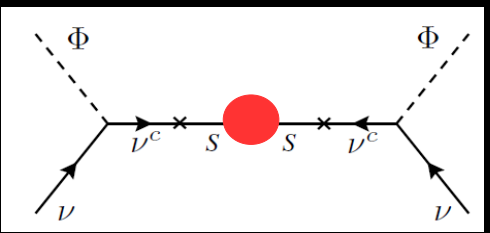
SM Type1 seesaw

of singlets arbitrary

MISSING PARTNER SEESAW

(3,2) min viable type1 seesaw
(3,1) scoto-seesaw template

LOW-SCALE Type1 SEESAW (3,6) ISS & LSS

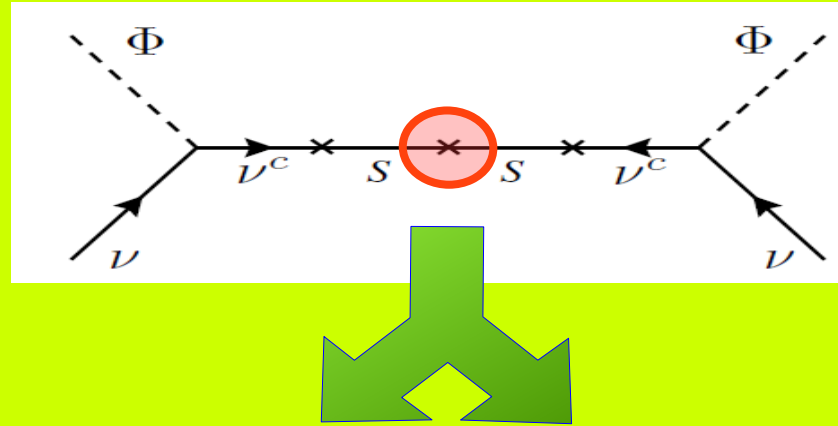


Mohapatra,Valle 86

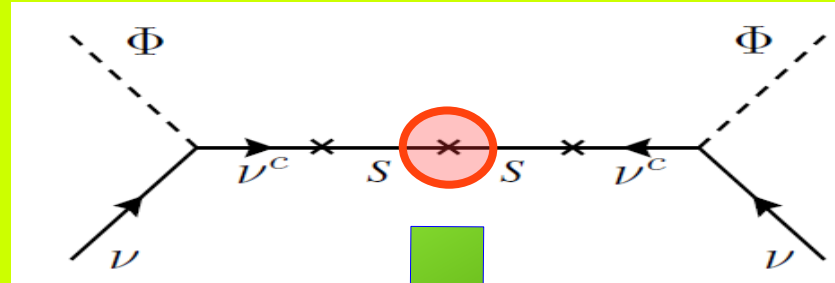
Akhmedov et al Phys.Rev.D53 (1996) 2752
PhysLettB368 (1996) 270

Malinsky et al PhysRevLett95(2005)161801

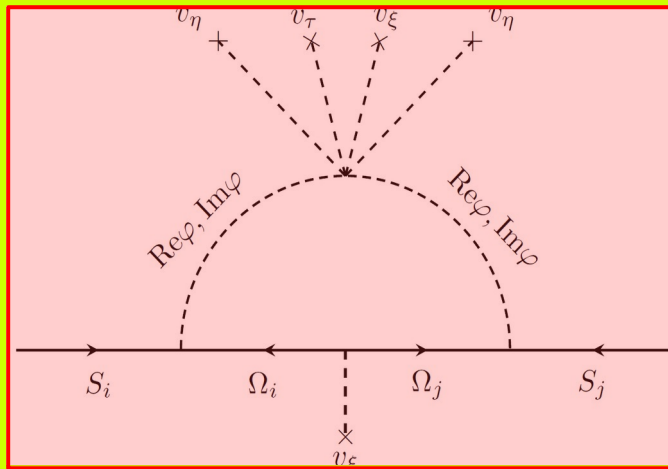
doubly protected inverse seesaw



doubly protected inverse seesaw

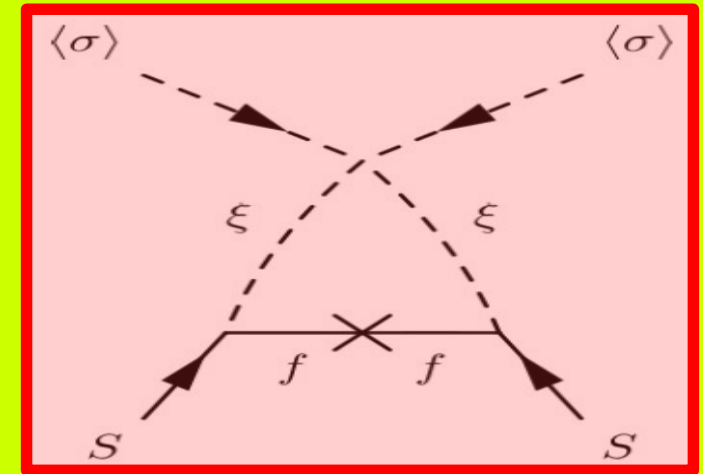


radiative
inverse seesaw



L-R scheme

Cárcamo Hernández et al JHEP 1902 (2019) 065



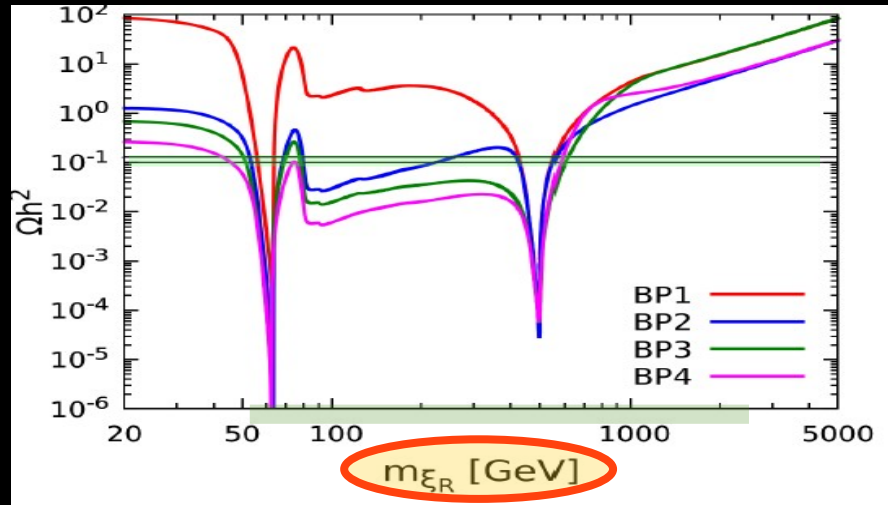
dark inverse seesaw

Mandal et al Phys.Lett.B821 (2021) 136609

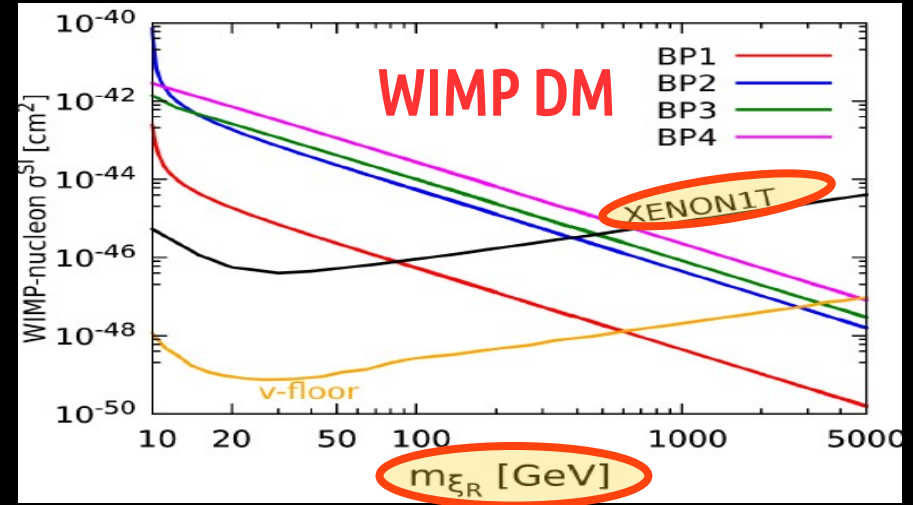


dark inverse typeI seesaw mechanism

Phys.Lett.B 821 (2021) 136609

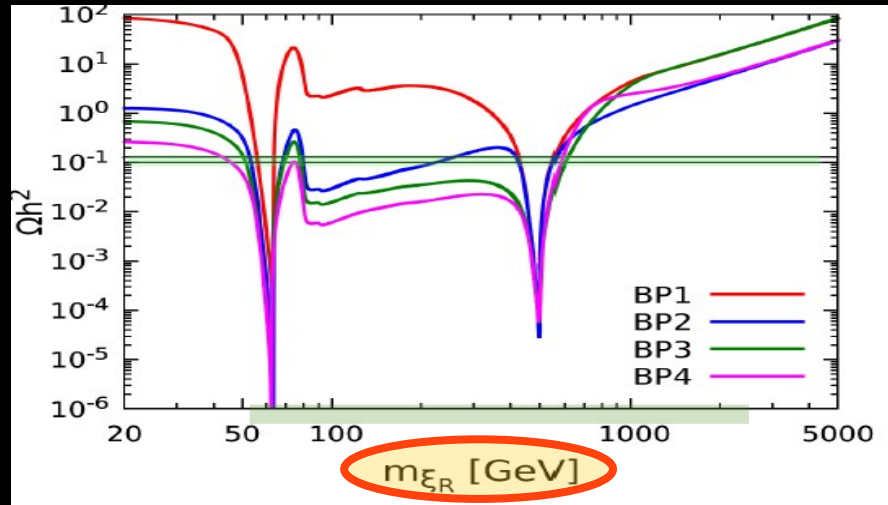


Xenon1T PhysRevLett.121.111302

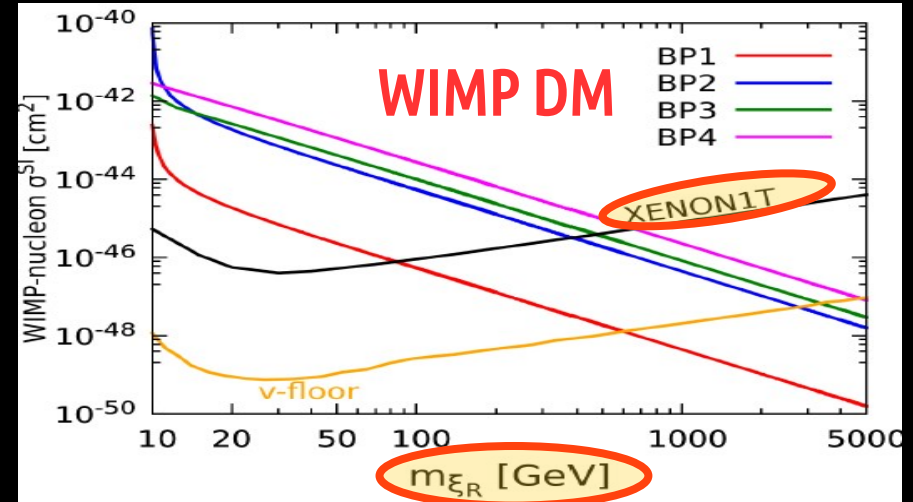


dark inverse type I seesaw mechanism

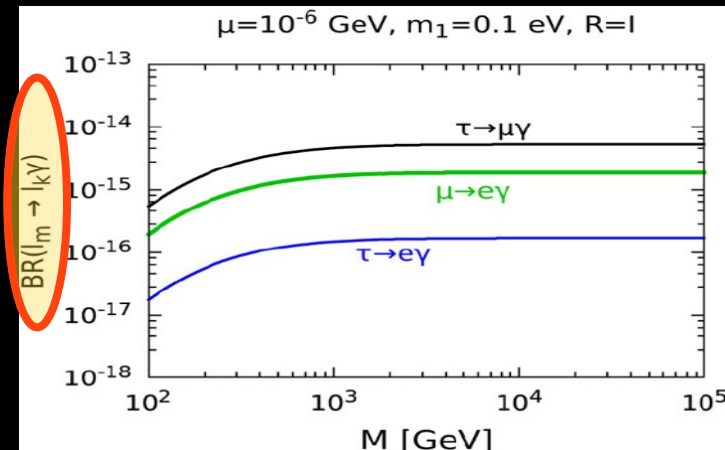
Phys.Lett.B 821 (2021) 136609



Xenon1T PhysRevLett.121.111302

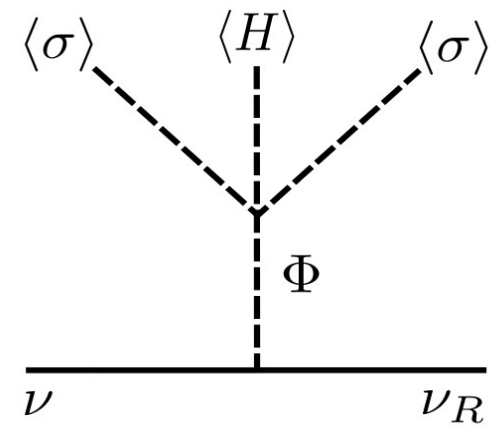
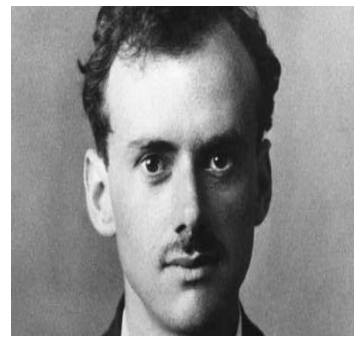
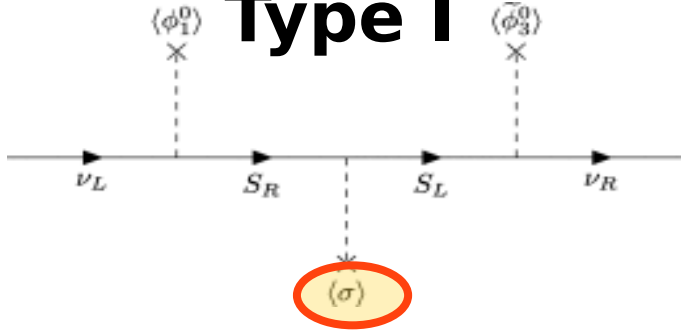


large cLFV from inverse
type I seesaw
Mandal et al
Phys.Lett.B 821 (2021) 136609



Seesawing a la

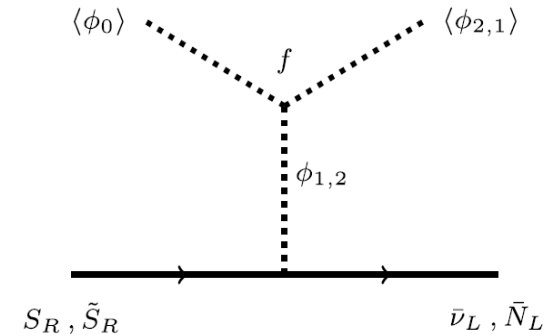
Type I



Type II

Phys.Lett. B762 (2016) 162-165

Phys.Rev. D94 (2016) 033012



Phys.Lett. B761 (2016) 431-436

Phys.Lett. B767 (2017) 209-213

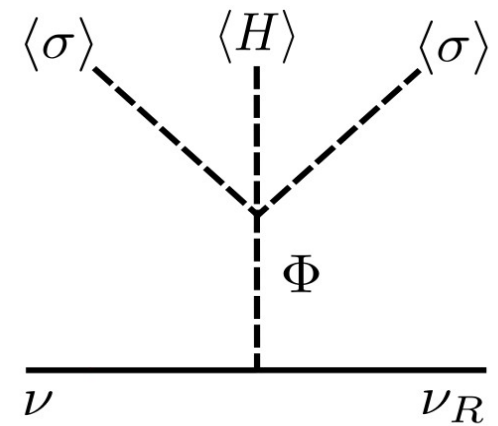
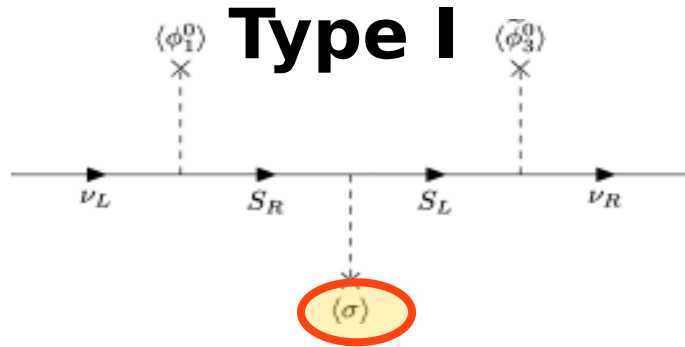
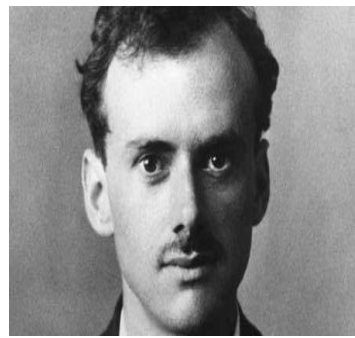
Phys.Rev. D98 (2018) 035009

Phys.Lett. B781 (2018) 122-128

Addazi et al Phys.Lett. B759 (2016) 471-478

Phys.Lett. B755 (2016) 363-366

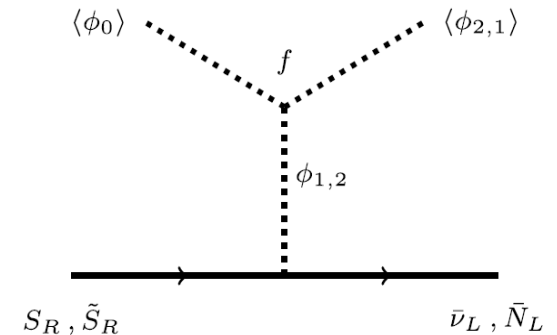
Seesawing a la



Type II

Phys.Lett. B762 (2016) 162-165

Phys.Rev. D94 (2016) 033012



symmetry protecting small neutrino mass
+ Diracness

Peccei-Quinn symmetry [Phys.Lett.B 810 \(2020\) 135829](#)

$$m_\nu^D \simeq \frac{y^{\nu_1} (y^S)^{-1} (y^{\nu_2})^T}{\sqrt{2}} \frac{v}{v_\sigma} \begin{matrix} \leftarrow \text{SU3L} \\ \leftarrow \text{PQ} \end{matrix}$$

Phys.Lett. B761 (2016) 431-436

Phys.Lett. B767 (2017) 209-213

[Phys.Rev. D98 \(2018\) 035009](#)

[Phys.Lett. B781 \(2018\) 122-128](#)

Addazi et al Phys.Lett. B759 (2016) 471-478

Phys.Lett. B755 (2016) 363-366



SM lacks an
organizing principle
to understand Flavor

Quarks



Leptons



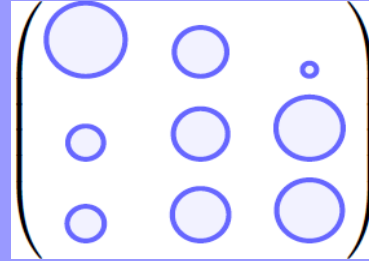
Force Carriers



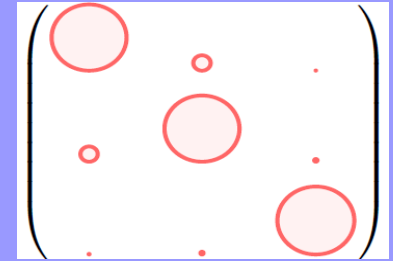
H
Higgs boson

flavour legacy of oscillations

Q/L mixing pattern

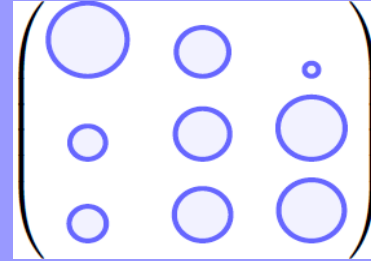


versus

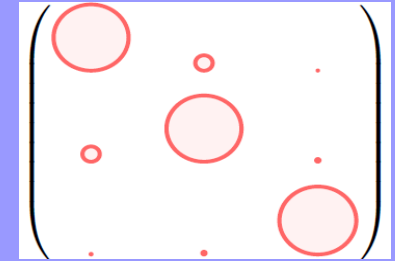


flavour legacy of oscillations

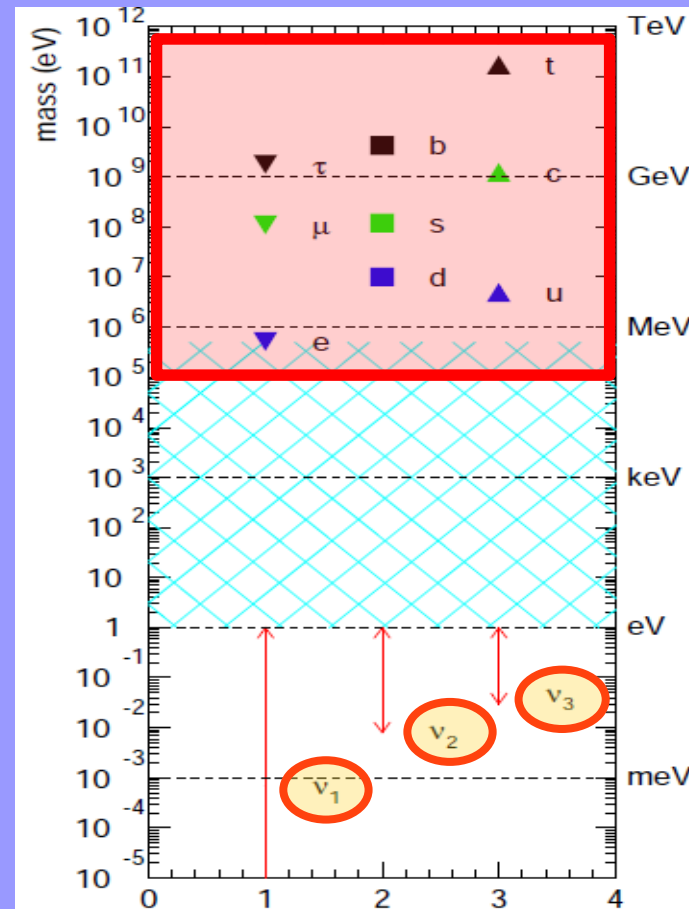
Q/L mixing pattern



versus

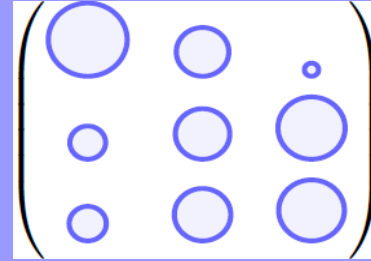


Q/L mass hierarchies

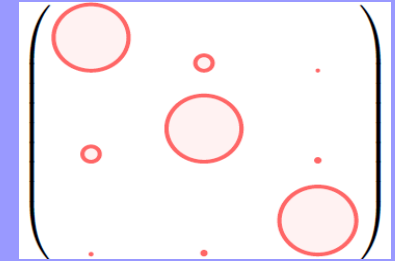


flavour legacy of oscillations

Q/L mixing pattern



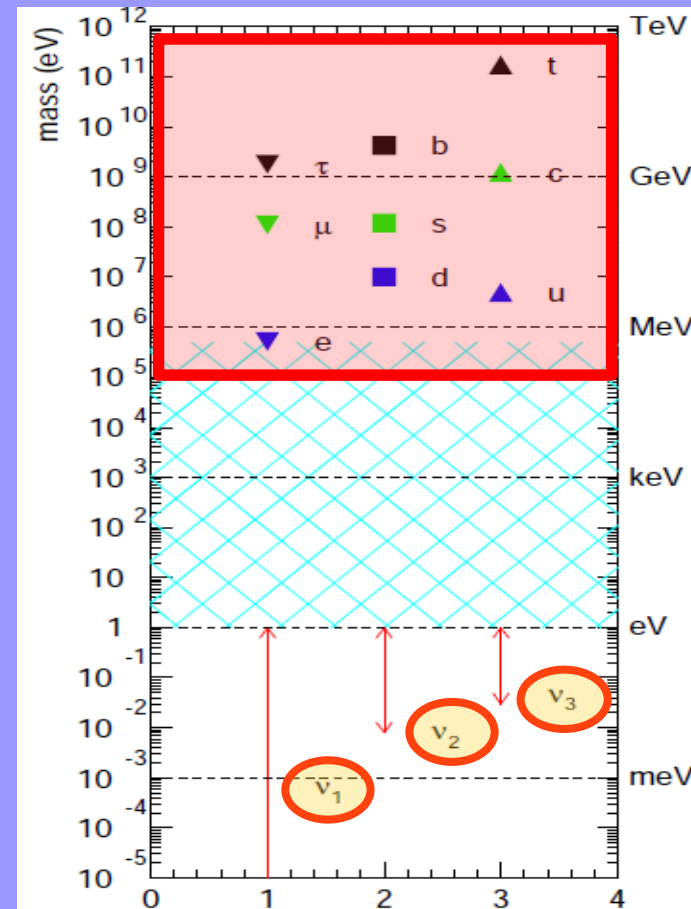
versus



Q/L mass hierarchies

Golden
Q-L mass
relation

$$\frac{m_\tau}{\sqrt{m_e m_\mu}} \approx \frac{m_b}{\sqrt{m_d m_s}}$$



Morisi et al	Phys.Rev. D84 (2011) 036003
King et al	Phys. Lett. B 724 (2013) 68
Morisi et al	Phys.Rev. D88 (2013) 036001
Bonilla et al	Phys.Lett. B742 (2015) 99
Reig, JV, Wilczek	Phys.Rev. D98 (2018) 095008

Warped flavour dynamics

Randall-Sundrum Phys.Rev.Lett. 83 (1999) 3370

- **mass hierarchies from geometry**

Arkani-Hamed & Schmaltz hep-ph/9903417

- **mixing angles from family symmetry**

Warped flavour dynamics

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TM mixing pattern predicted from T'

$$\cos^2 \theta_{12} \cos^2 \theta_{13} = \frac{2}{3} \quad \text{TM1}$$

$$\cos \delta_{CP} = \frac{(3 \cos 2\theta_{12} - 2) \cos 2\theta_{23}}{3 \sin 2\theta_{23} \sin 2\theta_{12} \sin \theta_{13}}$$

Chen et al

Phys. Rev. D 102, 095014 (2020)

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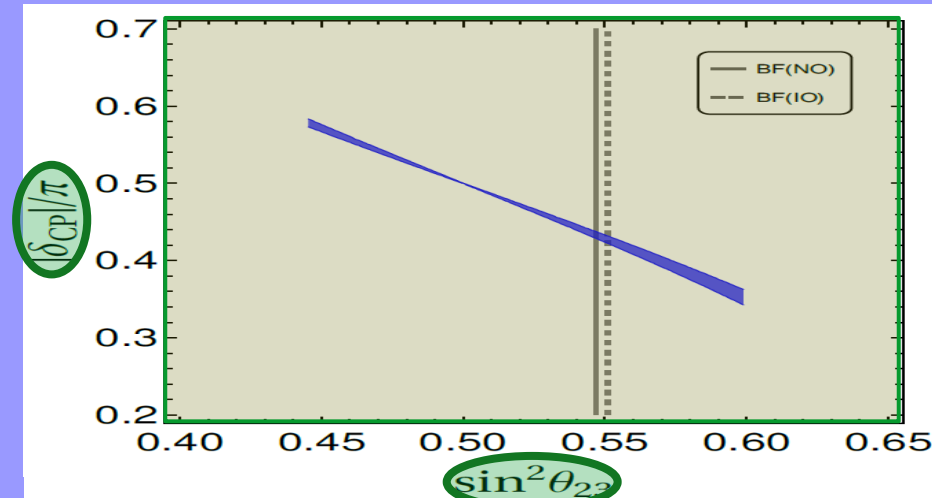
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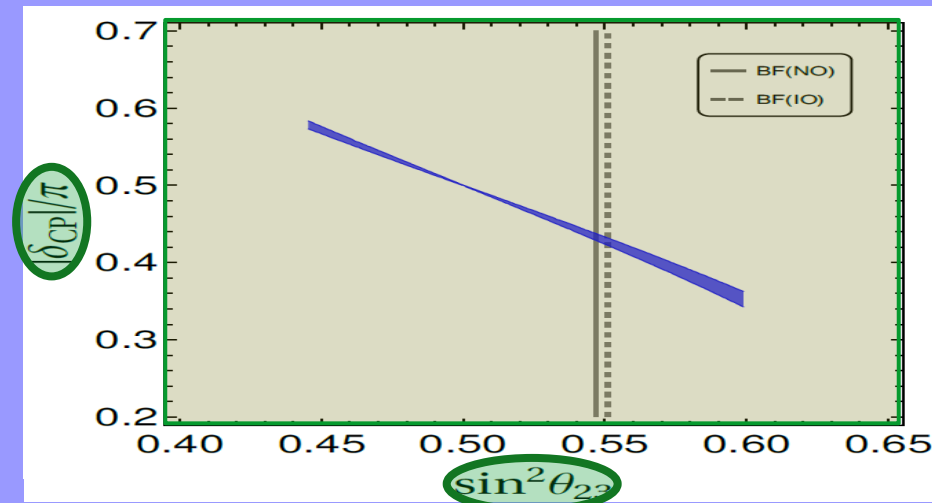
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TM1

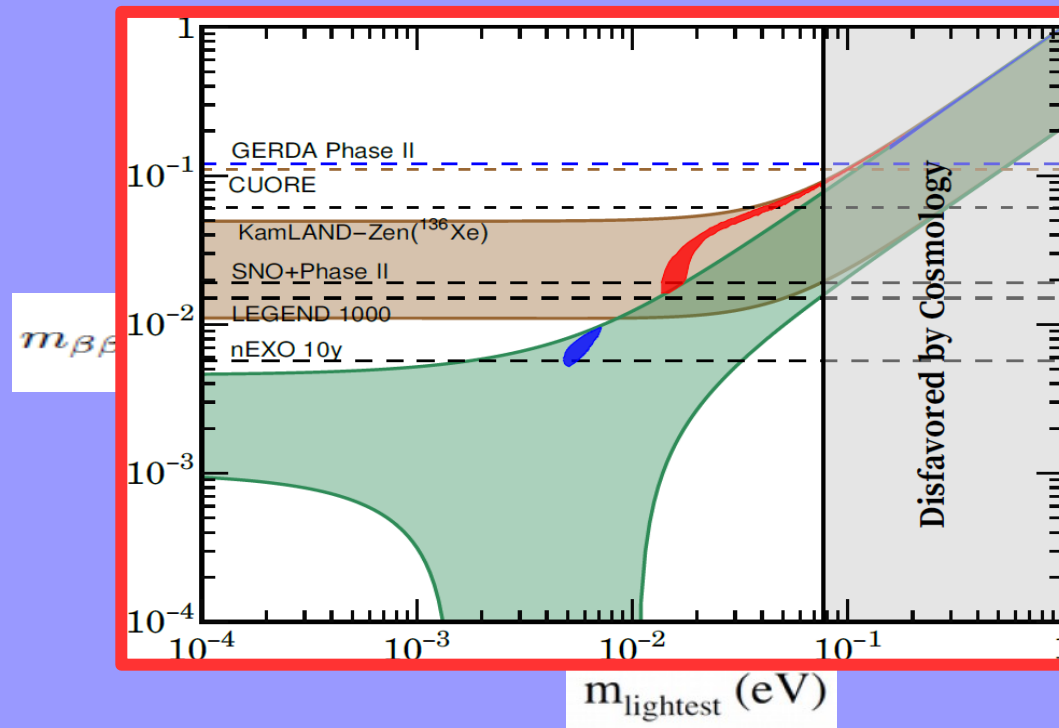
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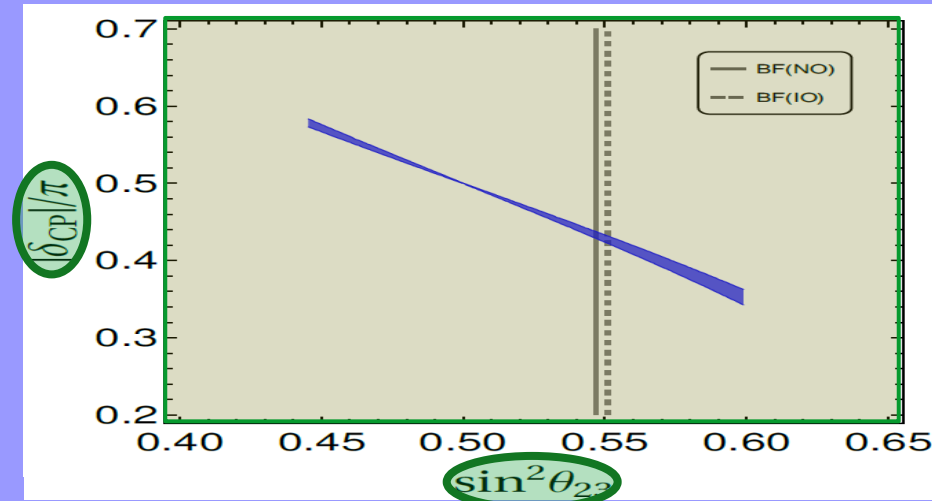
Chen et al

Phys. Rev. D 102, 095014 (2020)

For an alternative see also Chen et al
JHEP01(2016)007

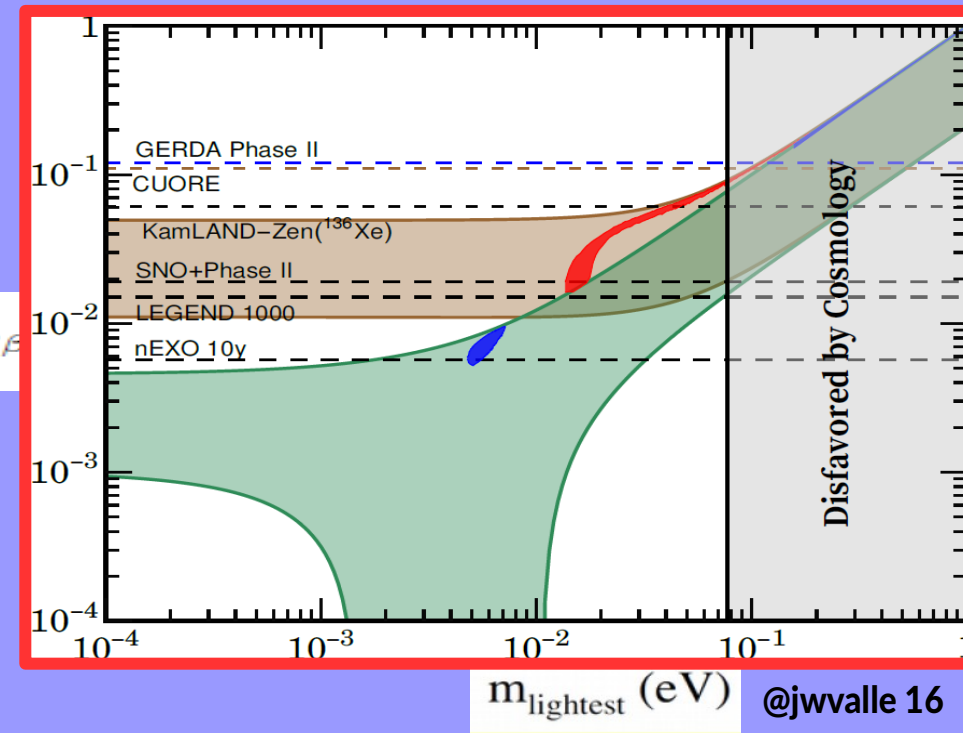
TM2

Phys. Rev. D95 (2017) 095030
Phys.Lett. B771 (2017) 524



TM1

$m_{\beta\beta}$



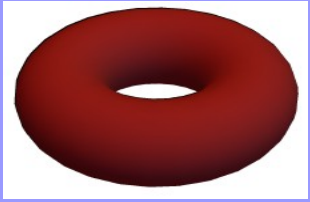
@jwvvalle 16

family symmetry from orbifolding

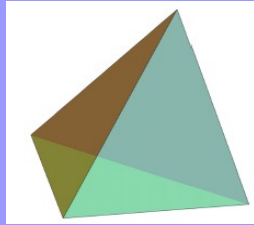
$$\mathcal{M} = \mathbb{M}^4 \times (\mathbb{T}^2 / \mathbb{Z}_2)$$

Phys.Lett.B 801 (2020) 135195

Phys.Rev.D 105 (2022) 055030



**A4 family symmetry
as emergent**

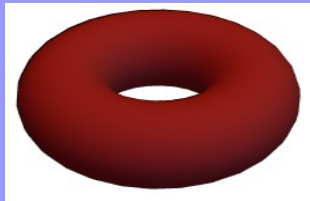


family symmetry from orbifolding

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Phys.Lett.B 801 (2020) 135195

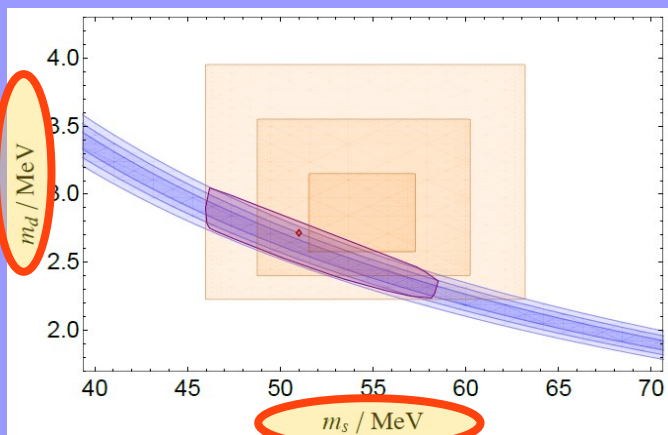
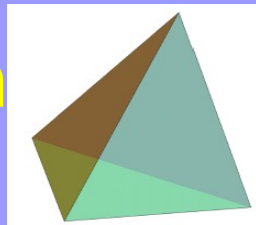
Phys.Rev.D 105 (2022) 055030



**A4 family symmetry
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Golden Q-L relation

$$\frac{m_\tau}{\sqrt{m_\mu m_e}} \approx \frac{m_b}{\sqrt{m_s m_d}}$$



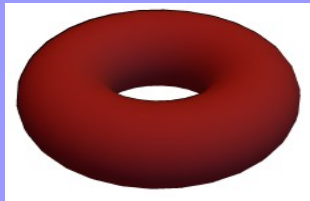
Good global fit of flavor observables

family symmetry from orbifolding

$$\mathcal{M} = \mathbb{M}^4 \times (\mathbb{T}^2 / \mathbb{Z}_2)$$

Phys.Lett.B 801 (2020) 135195

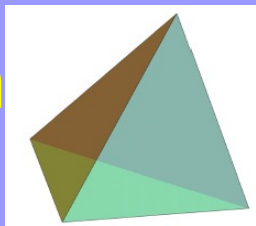
Phys.Rev.D 105 (2022) 055030



A4 family symmetry
as emergent

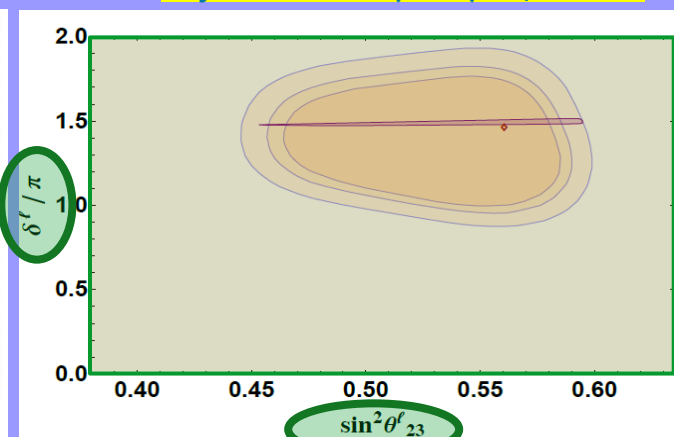
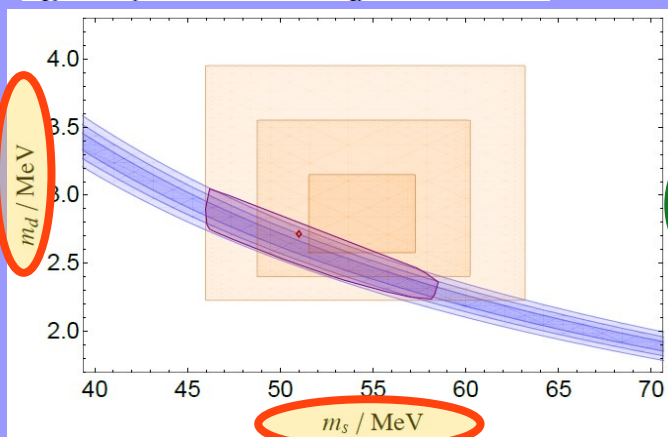
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Sign of CP phase!

Phys.Rev.D 101 (2020) 11, 116012



Good global fit of flavor observables

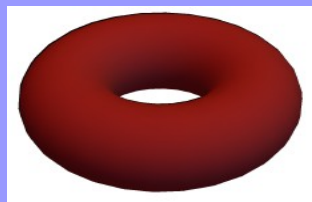
Dark Matter incorporated

Anda, Antoniadis et al

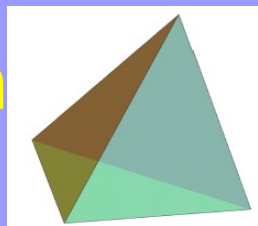
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Phys.Lett.B 801 (2020) 135195
Phys.Rev.D 105 (2022) 055030



A4 family symmetry
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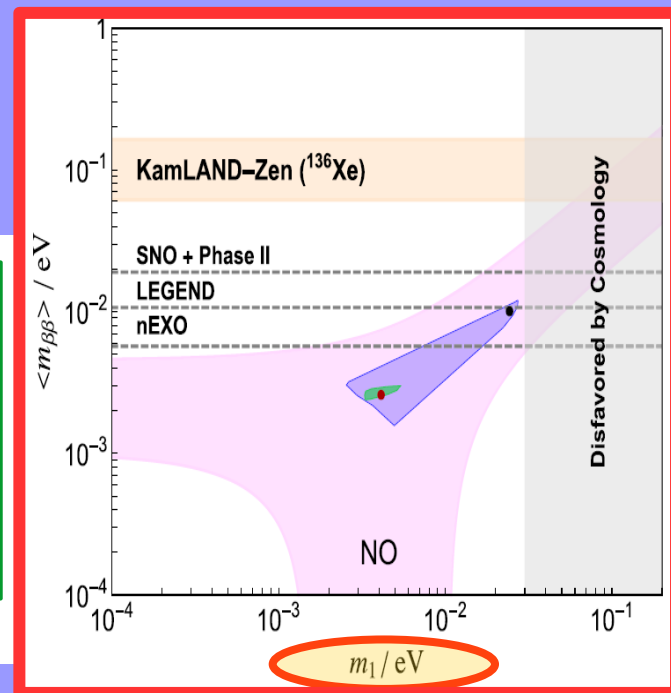
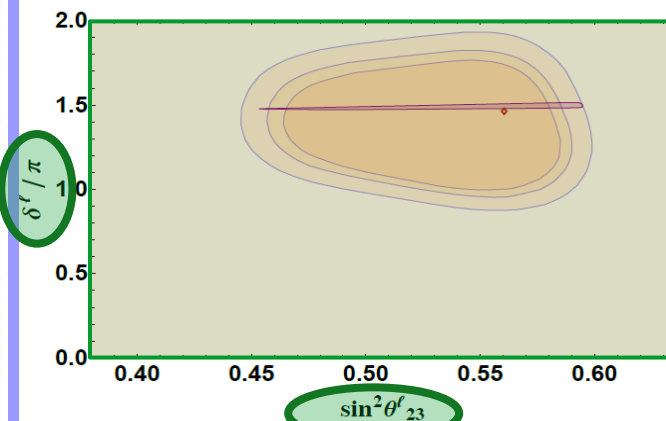
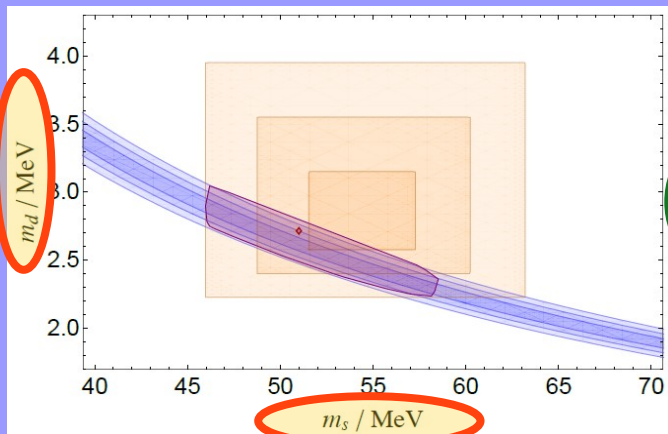


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Phys.Rev.D 101 (2020) 11, 116012

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Dark Matter incorporated

Anda, Antoniadis et al JHEP10(2020)190

Good global fit of flavor observables

@jwvvalle 17

**neutrinos at
the spotlight**

oscillation program, CP, octant, ordering, unitarity/NSI **CEvNS**
0 ν DBD
other searches at HE
& **cLFV**

THANK YOU



neutrinos at
the spotlight

oscillation program, CP, octant, ordering, unitarity/NSI **CEvNS**

0nuDBD

other searches at HE
& **cLFV**

EW vacuum
flavor puzzle & unification
strong CP problem



THANK YOU

WIMP DM as neutrino-mass mediator, **DM stability** from Diracness ...

baryon asymmetry from neutrinos, GWs from lepton # violation.. @jwvalle 18

