

The Southern Wide-Field Gamma-ray Observatory

Ruben Conceição



Gamma-rays

- ✧ Photons **extremely energetic**
 - ✧ Up to 100 TeV
 - ✧ Point to production source
 - ✧ Travel long distances
- ✧ Probes for the **most violent processes** known in the Universe
 - ✧ Supernovae; Active Galaxy Nuclei;
- ✧ Allows to perform strong **tests to fundamental physics**



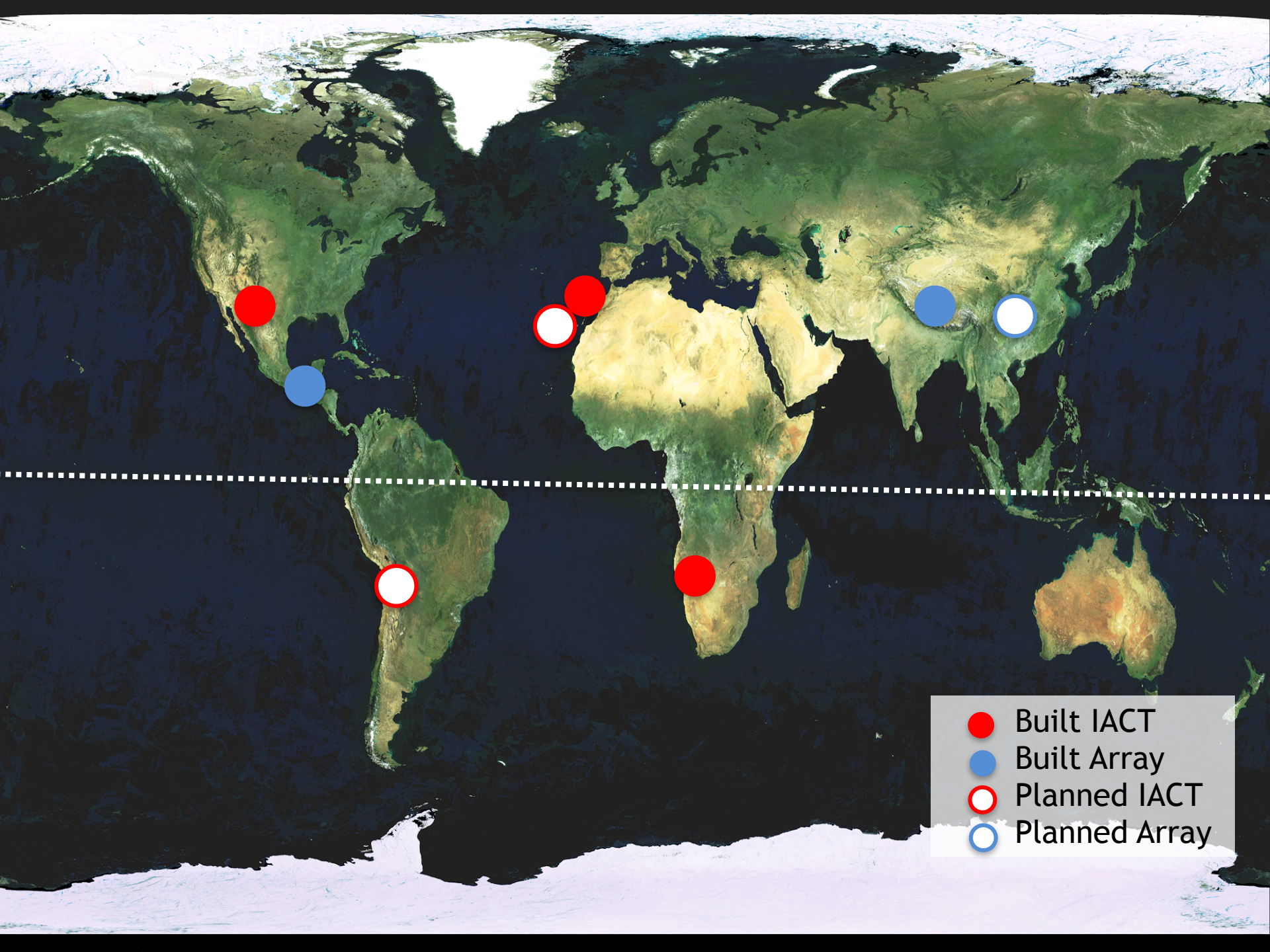
CTA

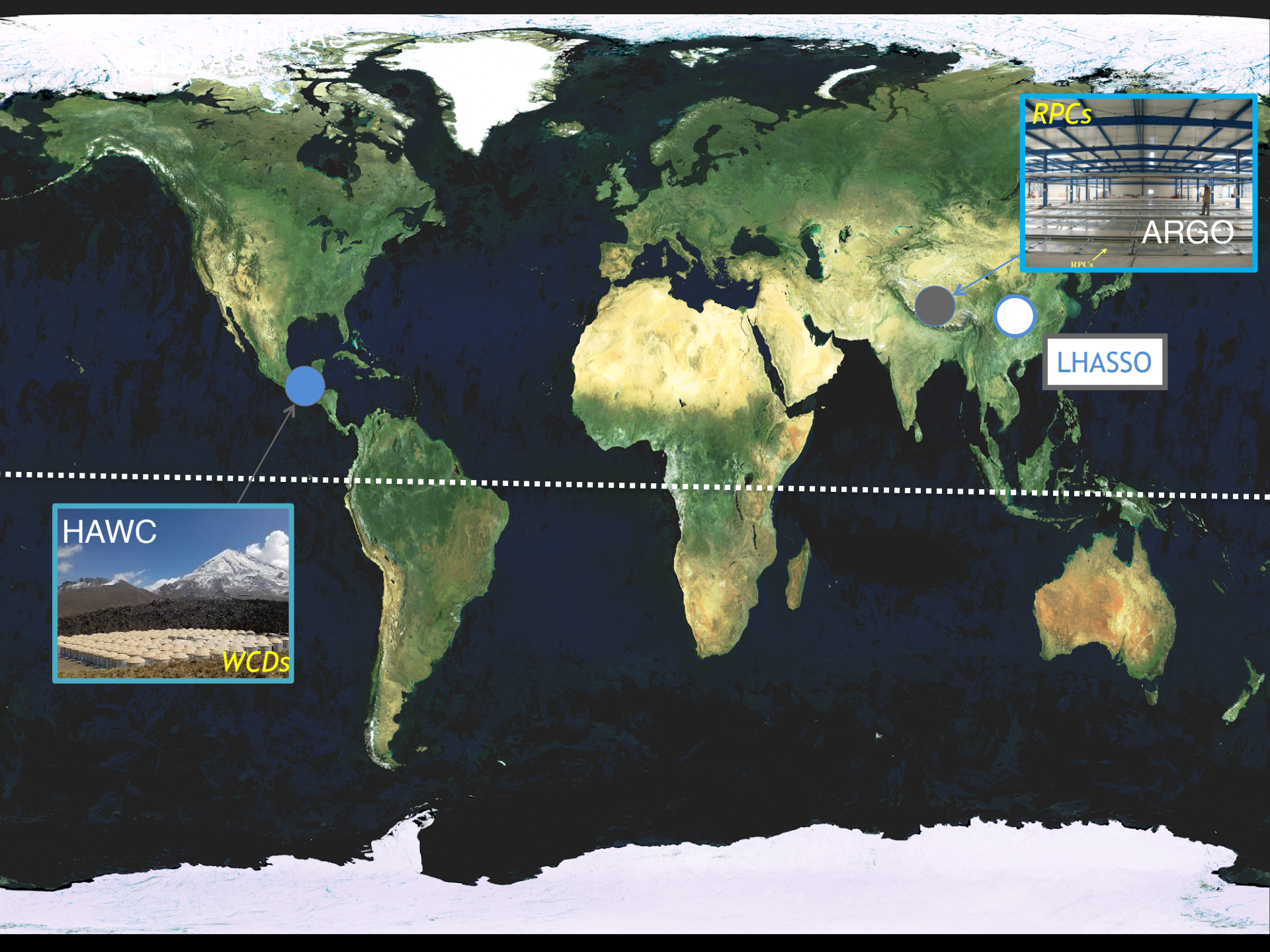


CTA



- Built IACT
- Built Array
- Planned IACT
- Planned Array





HAWC

WCDs

RPCs

ARGO

LHASO



RPCs

ARGO

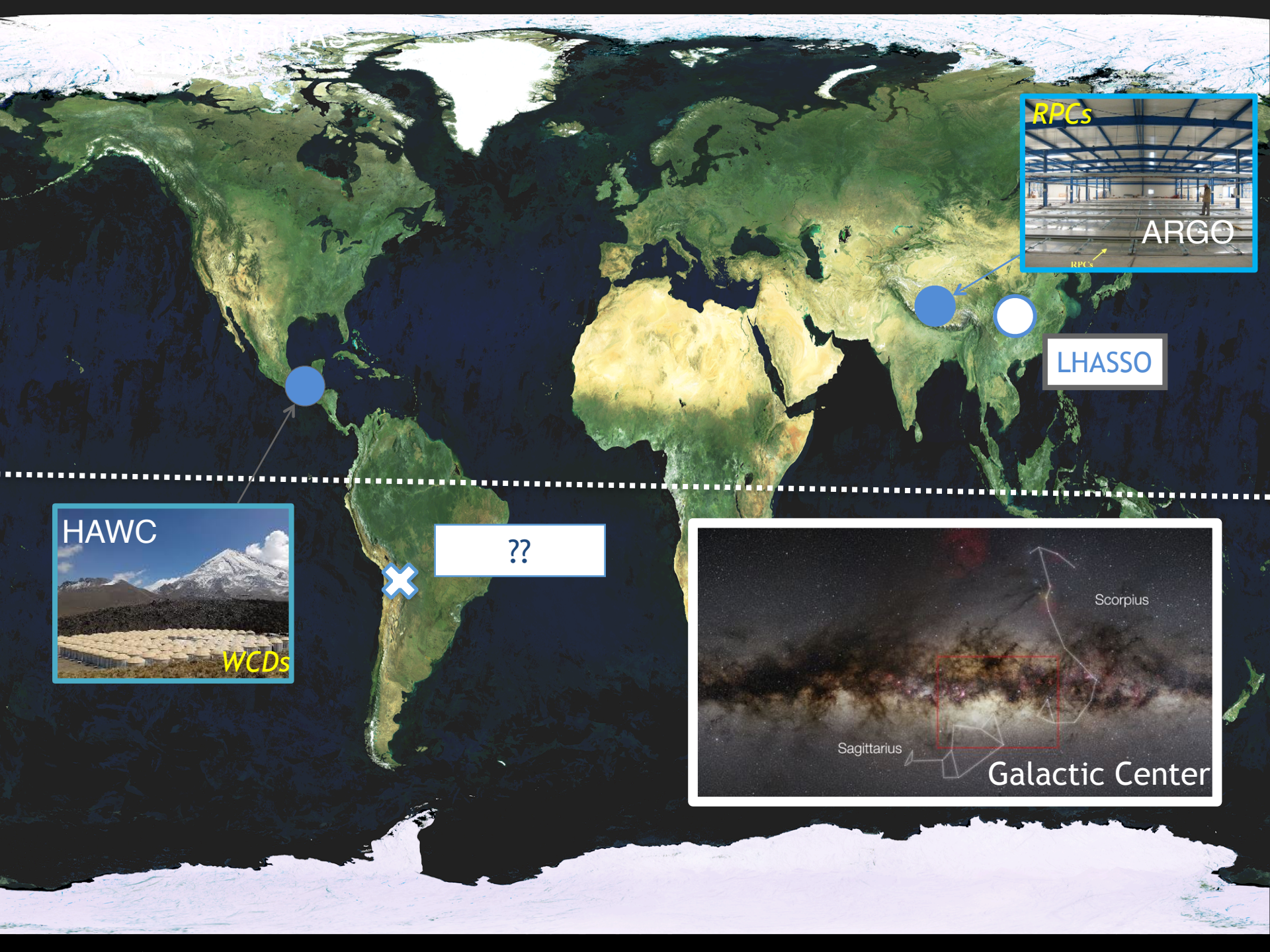
LHASO

HAWC

??

WCDs

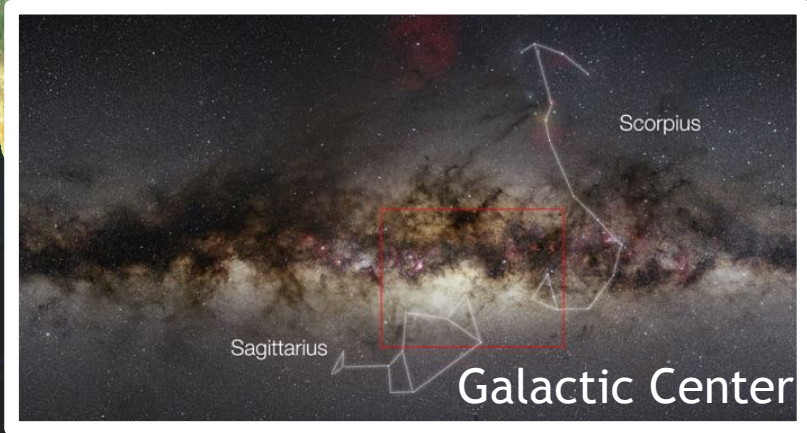
RPCs

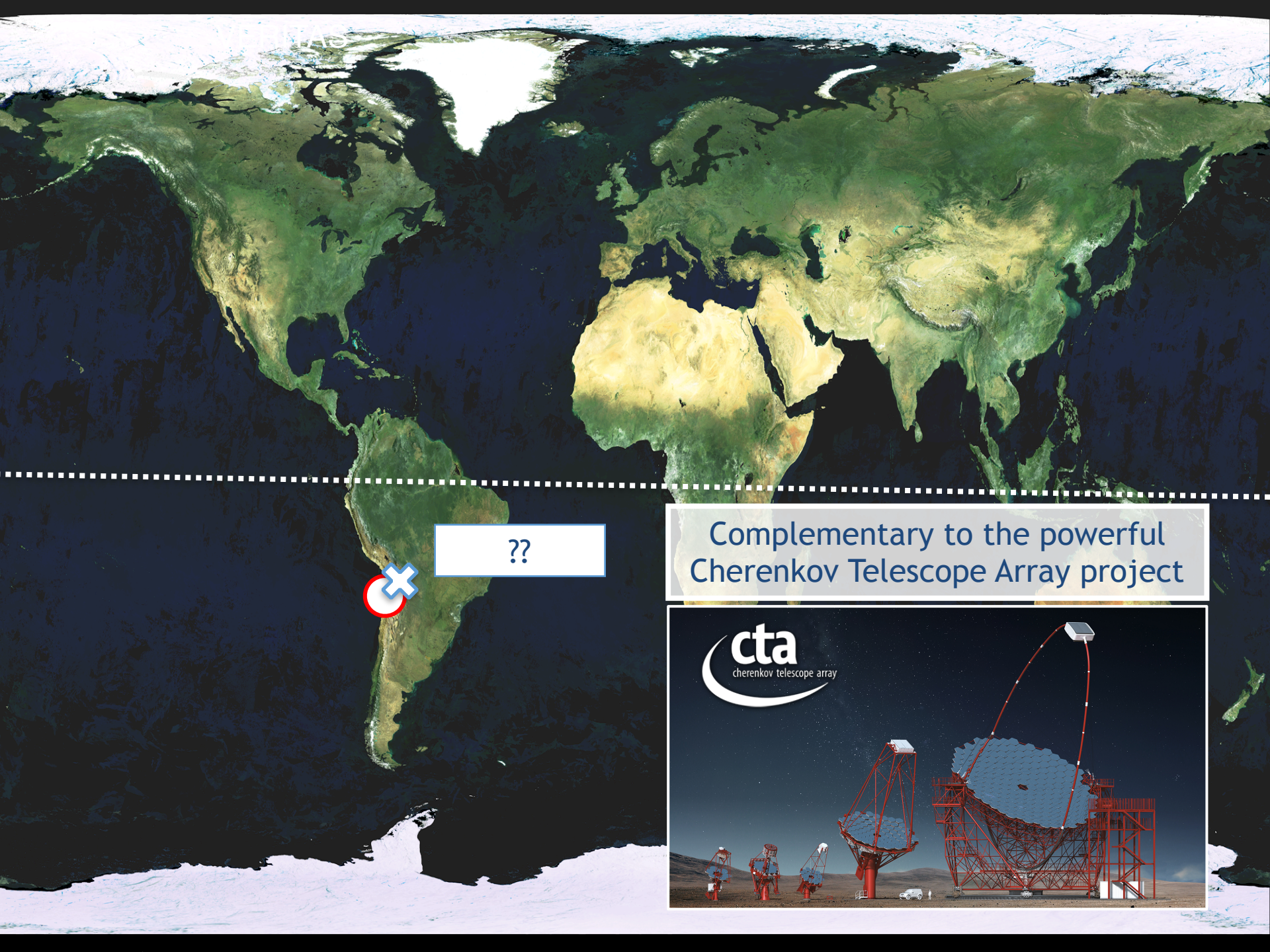


LHASO



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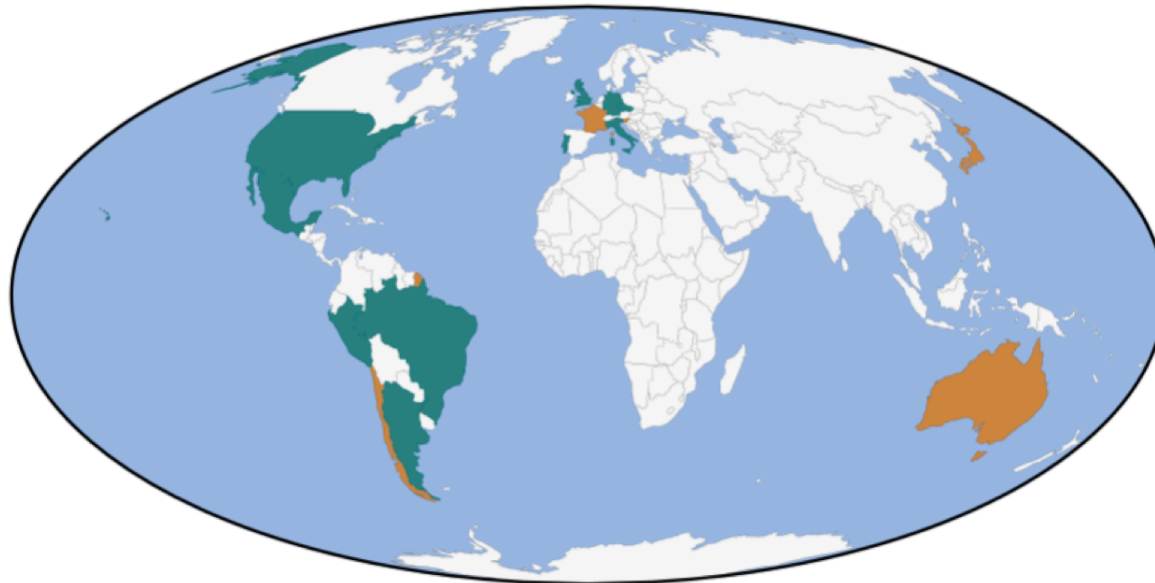
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Complementary to the powerful Cherenkov Telescope Array project



SWGGO collaboration

Next generation wide field-of-view gamma-ray observatory



Countries in SWGGO

Institutes

Argentina*, Brazil,
Czech Republic,
Germany*, Italy,
Mexico, Peru, Portugal,
United Kingdom,
United States*

Supporting scientists

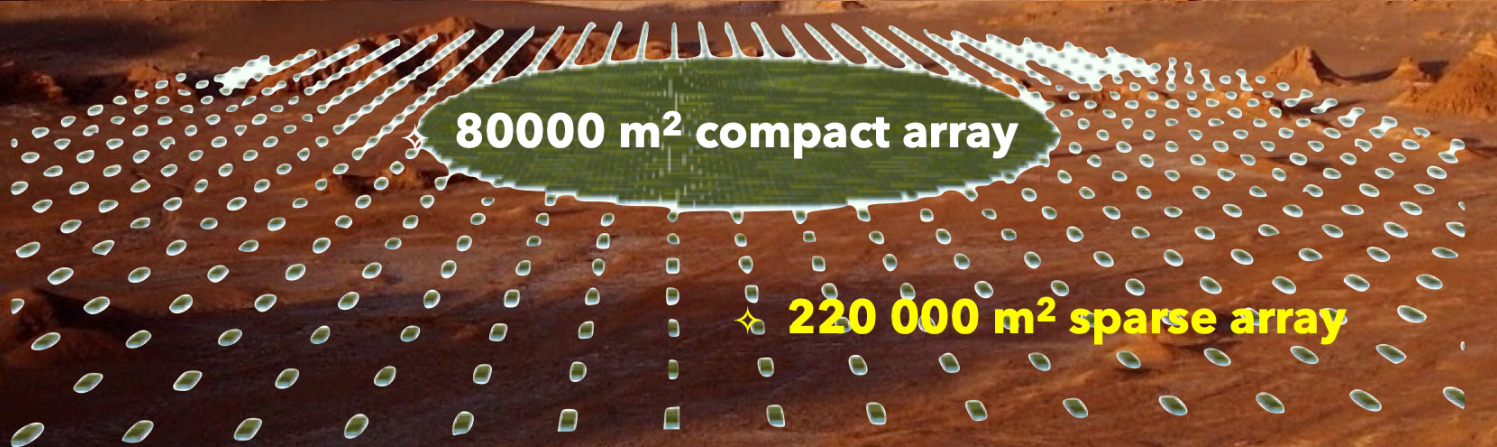
Australia, Chile,
France, Japan,
Slovenia

**also supporting
scientists*

3-year R&D project to design and plan the best experiment

Final workshop on Autumn 2022

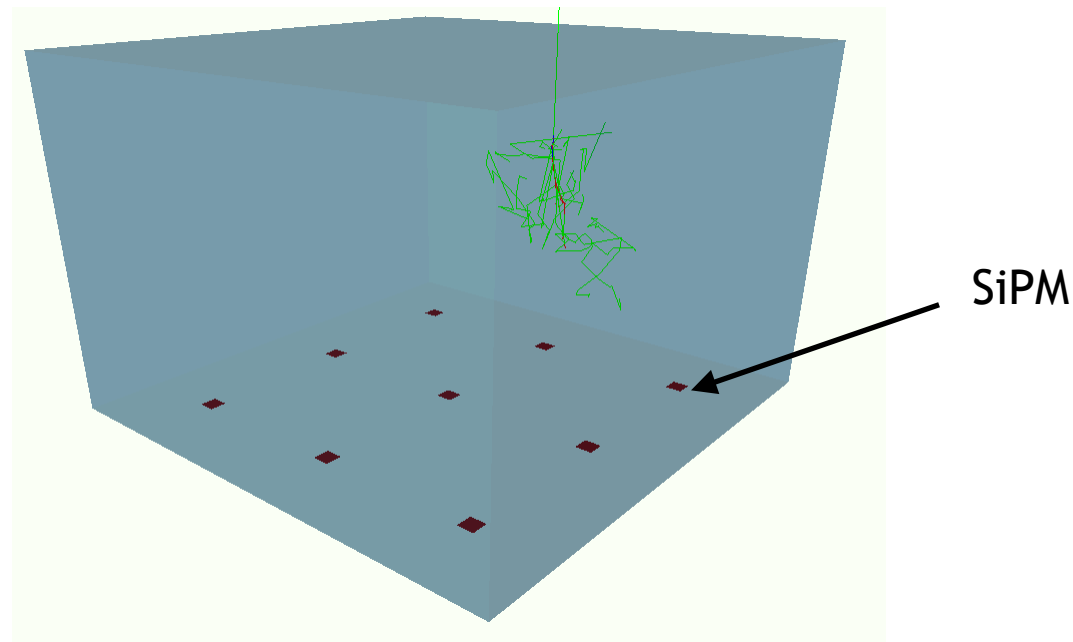
The challenge



Build a huge array at an altitude of **5000 m** based on the **Water Cherenkov detection** technology

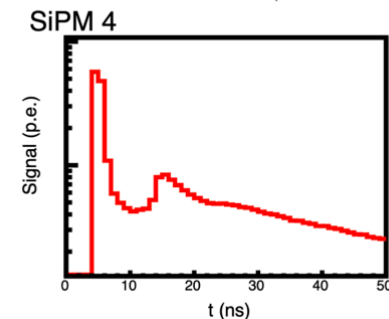
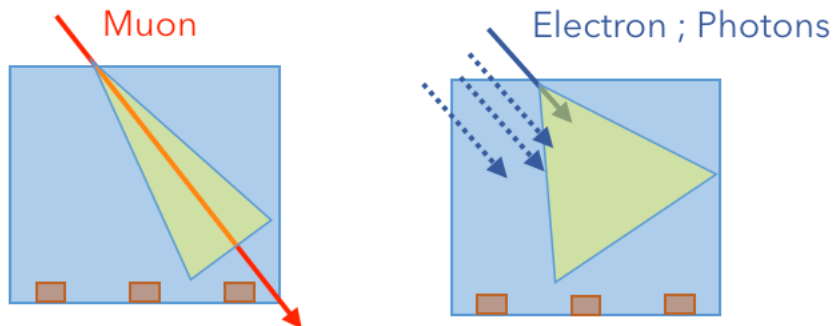
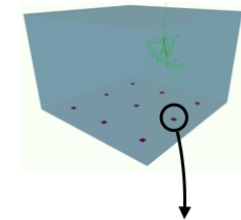
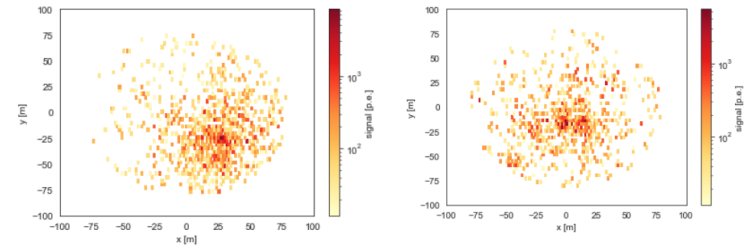
Activities @ LIP

- ✧ Design of the detector unit station
 - ✧ Can the water volume be decreased having a smarter light collection?
 - ✧ New inventive readout solutions based on SiPM



Activities @ LIP

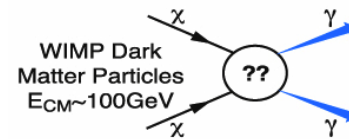
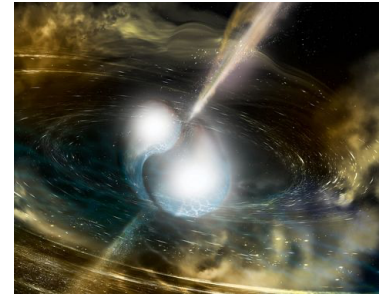
- ✧ Improve reconstruction analysis
 - ✧ Gamma/hadron discrimination
 - ✧ Analyse shower patterns at the ground with machine learning techniques
 - ✧ Explore the WCD signal time trace to identify the presence of muons



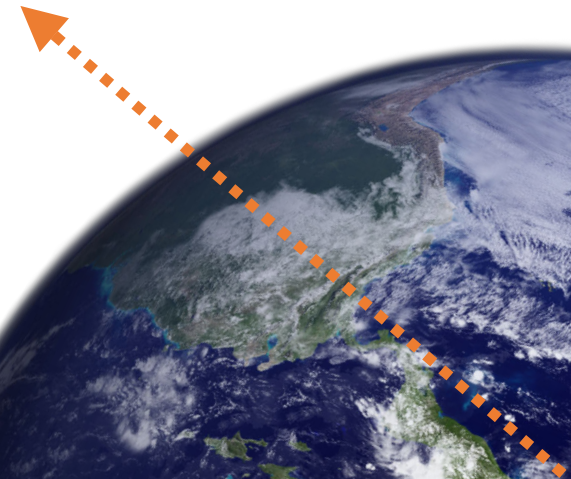
Activities @ LIP

- ❖ Science capabilities
 - ❖ Studies on the sensitivity to:
 - ❖ Transient astrophysical phenomena
 - ❖ Explore hadronic interactions the forward region
 - ❖ Dark matter annihilation
 - ❖ Detect Neutrino physics
 - ❖ BSM physics

Neutron Stars Merger



tau-neutrino



SWGGO

- ✧ Exciting project with an interesting timescale for a Master thesis
 - ✧ Work may be supported by a grant
- ✧ The thesis can have a significant impact on the design of the experiment
- ✧ More information:
 - ✧ ruben@lip.pt, pimenta@lip.pt

Summary

- ✧ Astroparticle physics (Multi-Messengers)
 - ✧ Use **astrophysical messengers** and known **particle physics** to gain a deeper understanding of the dynamics of our Universe
 - ✧ Rapidly evolving field
 - ✧ **Lots of ambitious projects**
 - ✧ Will soon provide important tests to our knowledge over fundamental physics

Backup slides



Newly formed international collaboration to make this experiment possible!



ÚLTIMA HORA Operação Antídoto. PJ detém médicos em buscas a consultórios

RTP NOTÍCIAS

2. Jul. 2019 | 10:48

DESPORTO PAÍS MUNDO POLÍTICA ECONOMIA

f PARTILHE NO FACEBOOK t PARTILHE NO TWITTER in PARTILHE NO LINKEDIN

Portugal junta-se a oito países para construir observatório de raios gama nos Andes



Terça-feira, 2 Julho 2019

Flutuação 28° 28° 19°

tvi24

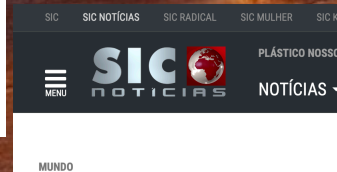
INÍCIO NOTÍCIAS VÍDEOS FOTOS DIRETO GUIA TV DOSSIERS EU VI

Últimas Opinião Sociedade Política Economia Internacional Desporto Motores Tecnologia Música

Investigações Alexandra Borges Investigações Ana Leal Deus e o Diabo Gente Que

Portugal junta-se a oito países para construir observatório de raios gama nos Andes

Será o primeiro observatório de raios gama no hemisfério sul. Já existe um do género, mas no hemisfério norte, no México



SIC NOTÍCIAS SIC RADICAL SIC MULHER SIC K

PLÁSTICO NOSSO D

NOTÍCIAS

MUNDO

Portugal junta-se a oito países para construir observatório de raios gama nos Andes



JN DN TSP Dinheiro Vivo V Digital Plataforma O Jogo Motor 24 Men's Health Women's Health Evidas Volta ao Mundo NM N-TV Delas

MENU 19

Diário de Notícias

INÍCIO / LUSA

Portugal junta-se a oito países para construir observatório de raios gama nos Andes

Lisboa, 01 jul 2019 (Lusa) - Portugal e mais oito países juntam-se a partir de hoje numa colaboração internacional para construir um observatório de raios gama na região dos Andes, para procurar sinais de matéria escura no centro da Via Láctea, foi hoje anunciado.



CIÊNCIA > ESPAÇO MEDICINA ECOSFERA

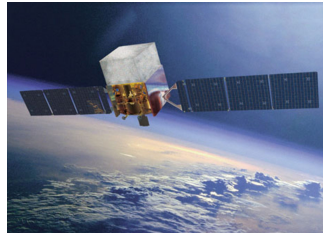
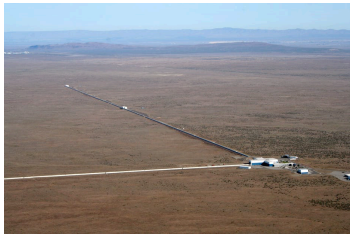
UNIVERSO

Portugal junta-se a oito países para construir observatório de raios gama nos Andes

Empreendimento marcará a edificação do primeiro laboratório de observação de raios gama no hemisfério sul. Custo da construção estimado em, pelo menos, 50 milhões de euros.

Lusa - 1 de Julho de 2019, 21:00

57 PARTILHAS f t in p



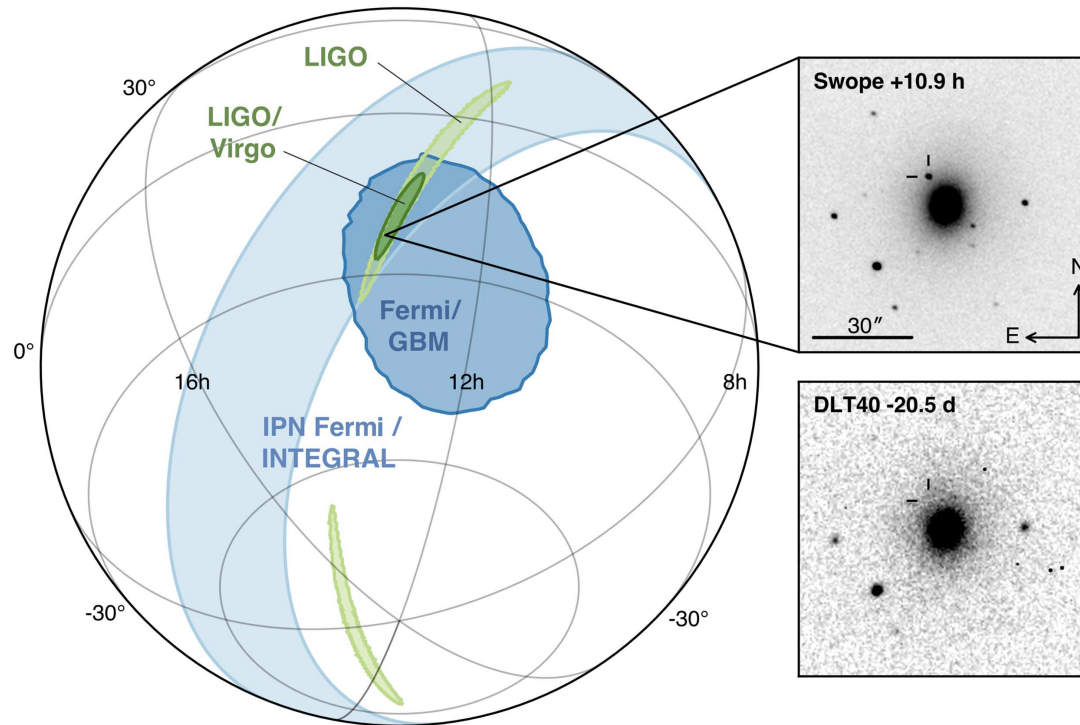
Multi-messengers

The opening of a new era...

Multi-messenger observation of a Binary Neutron Star Merger



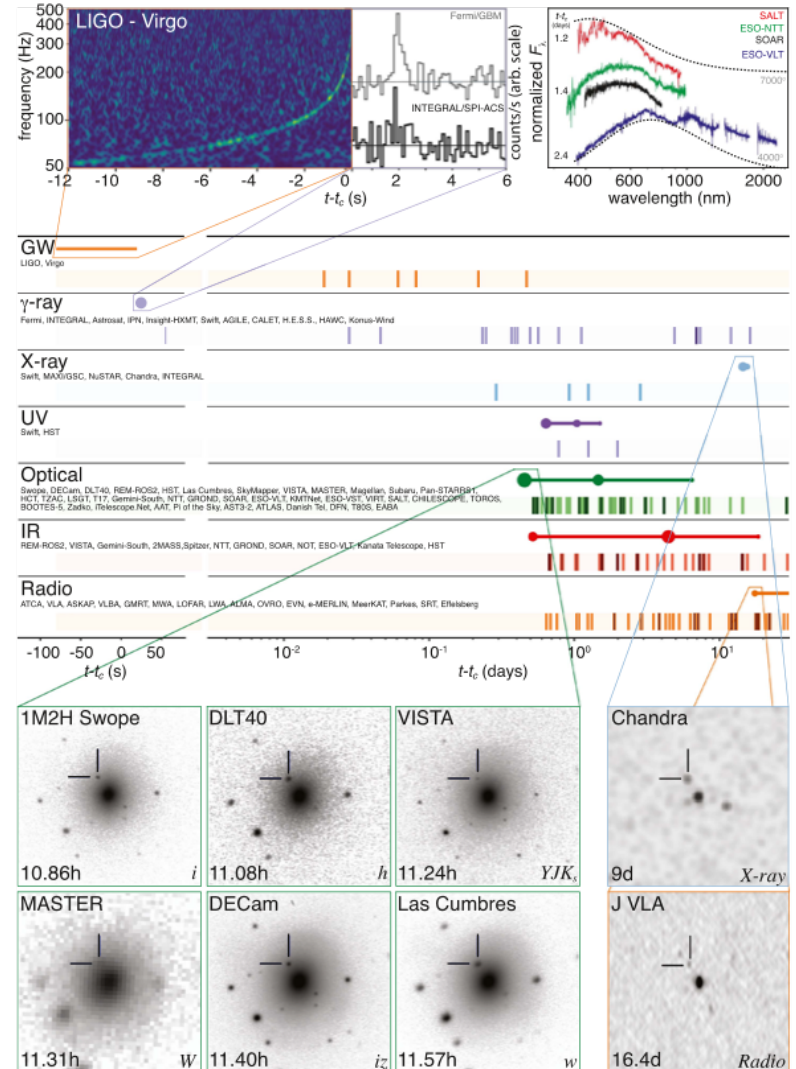
Joint publication of LIGO, VIRGO, INTEGRAL, Fermi, IceCube, Pierre Auger ...



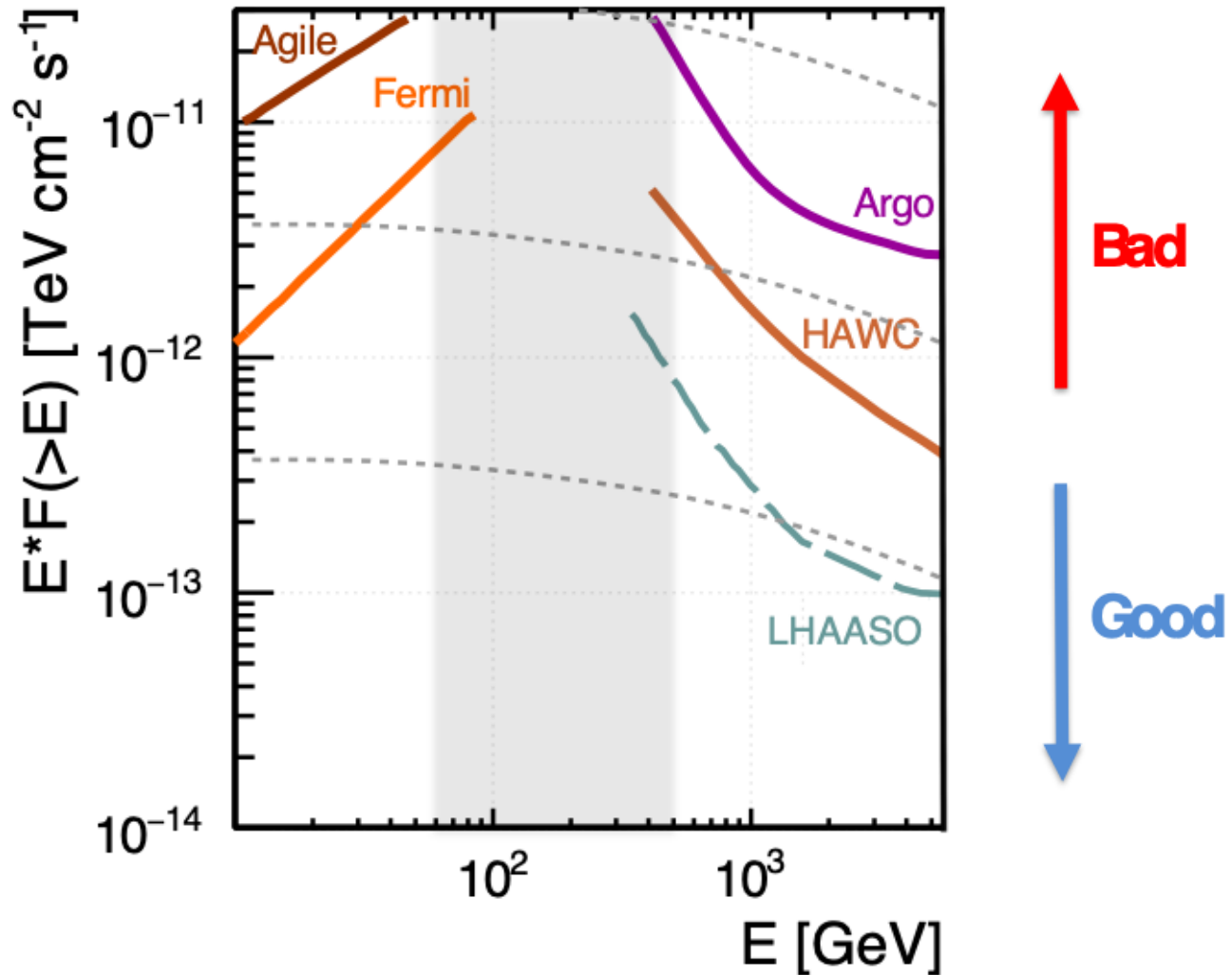
- ❖ Simultaneous observation of a **Gravitational Wave + electromagnetic** counter parts
- ❖ Allows to test the dynamics of our surrounding Universe
- ❖ Study of **transient phenomena in all energy** regions is one of the main ingredients

Multi-messenger observation of a Binary Neutron Star Merger

- ✧ Observe the same phenomenon with **different instruments**
- ✧ Follow the **evolution in time**
- ✧ Different wavelengths \Rightarrow different kind of interactions \Rightarrow different phenomena

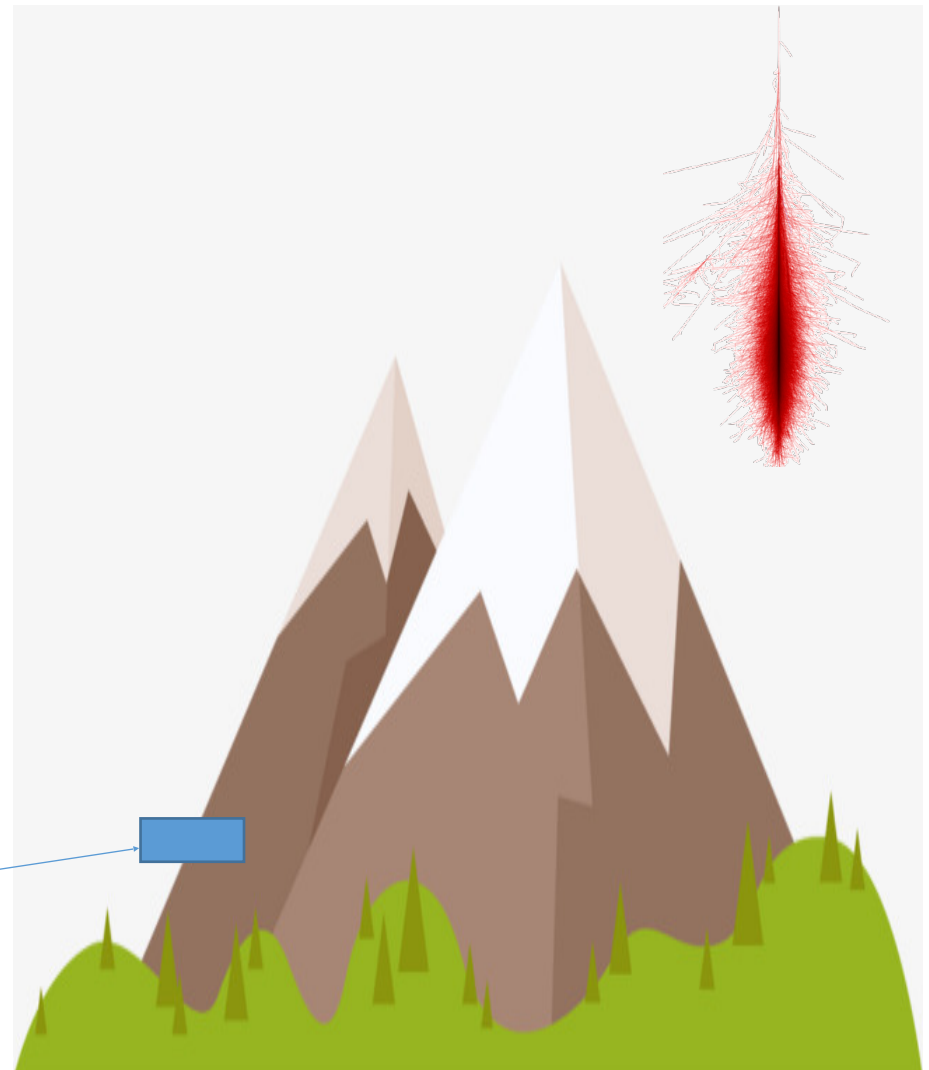


Current Wide Field-of-View Gamma-Ray Observatories



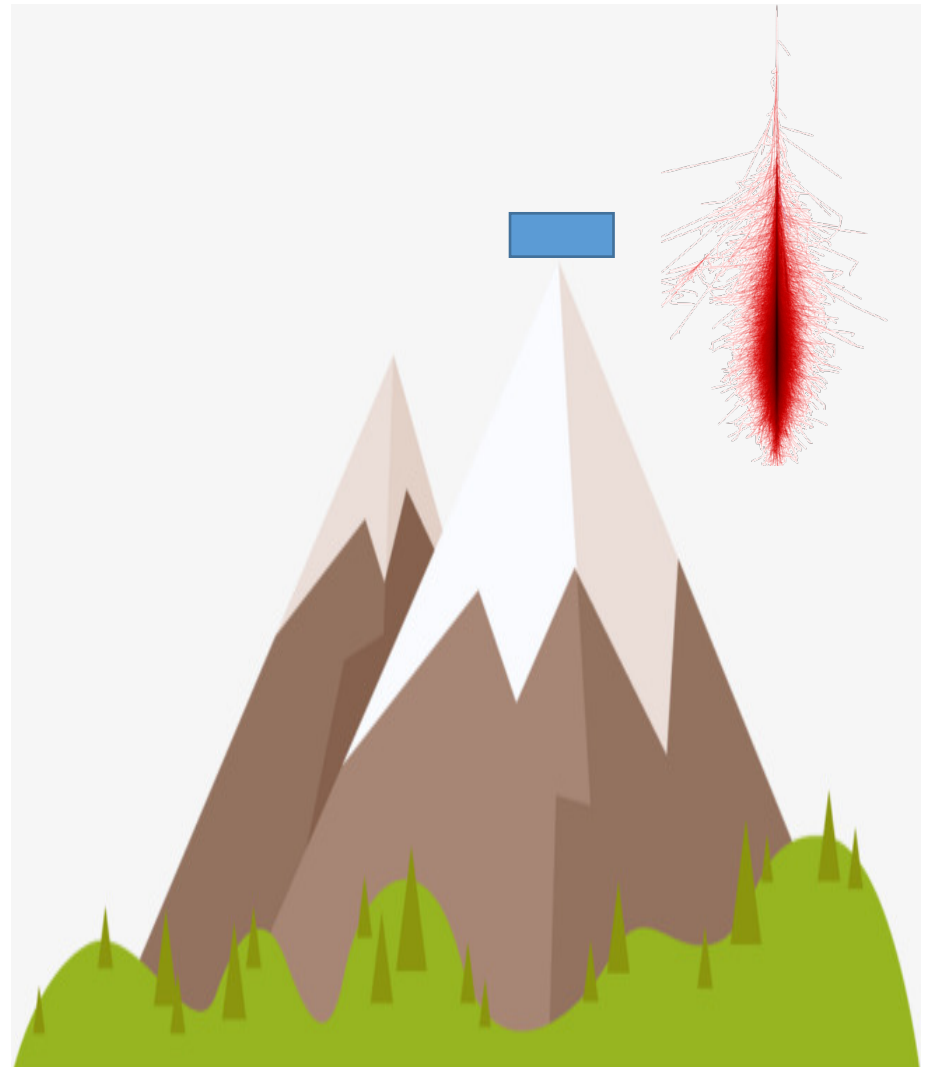
How to lower the energy threshold?

- ✧ Put the experiment at higher altitude
- ✧ Gamma-ray EAS arrays have typically 20 000 m²



How to lower the energy threshold?

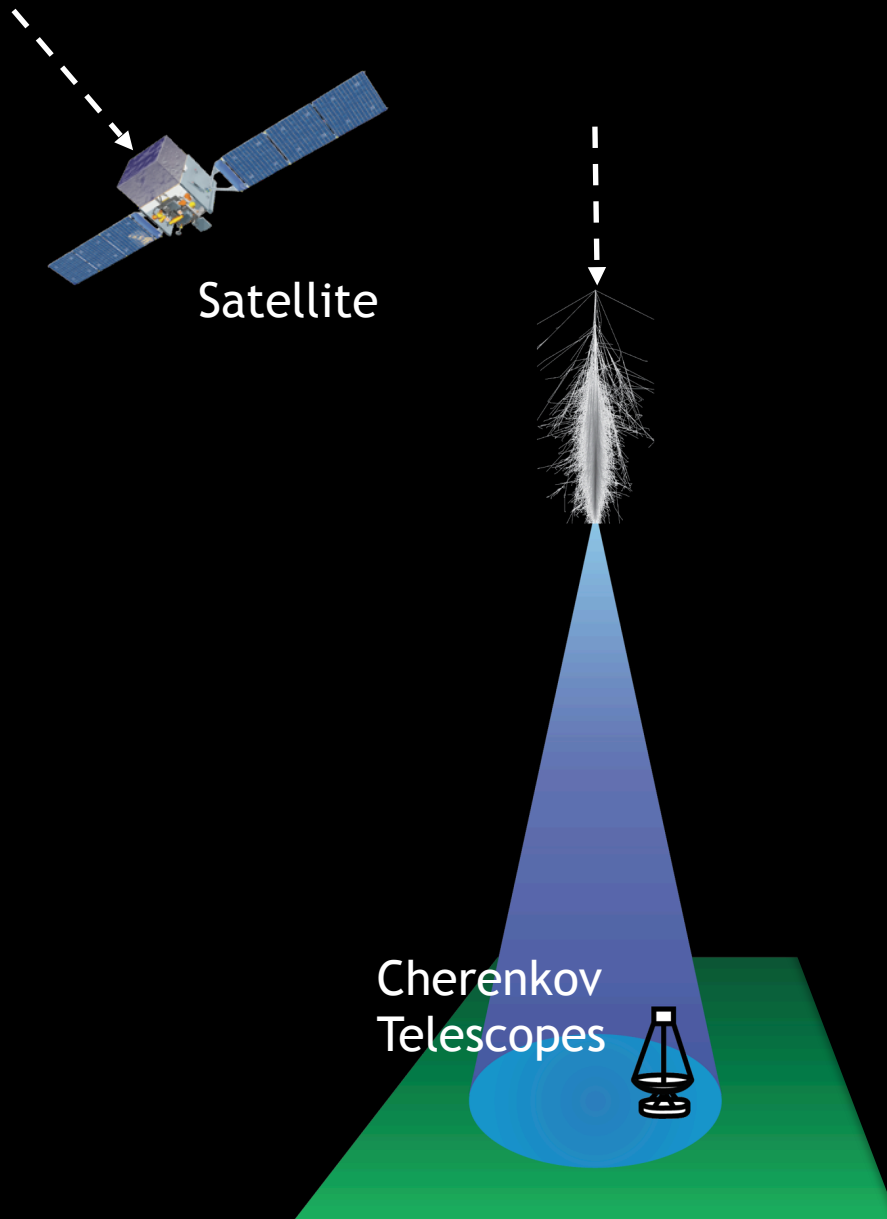
- ✧ Put the experiment at higher altitude
- ✧ Gamma-ray EAS arrays have typically 20 000 m²
- ✧ It is possible to find sites with **≈5000 m of altitude**
 - ✧ Atacama desert, Northern Chile
- ✧ *Can the detector concept be improved?*



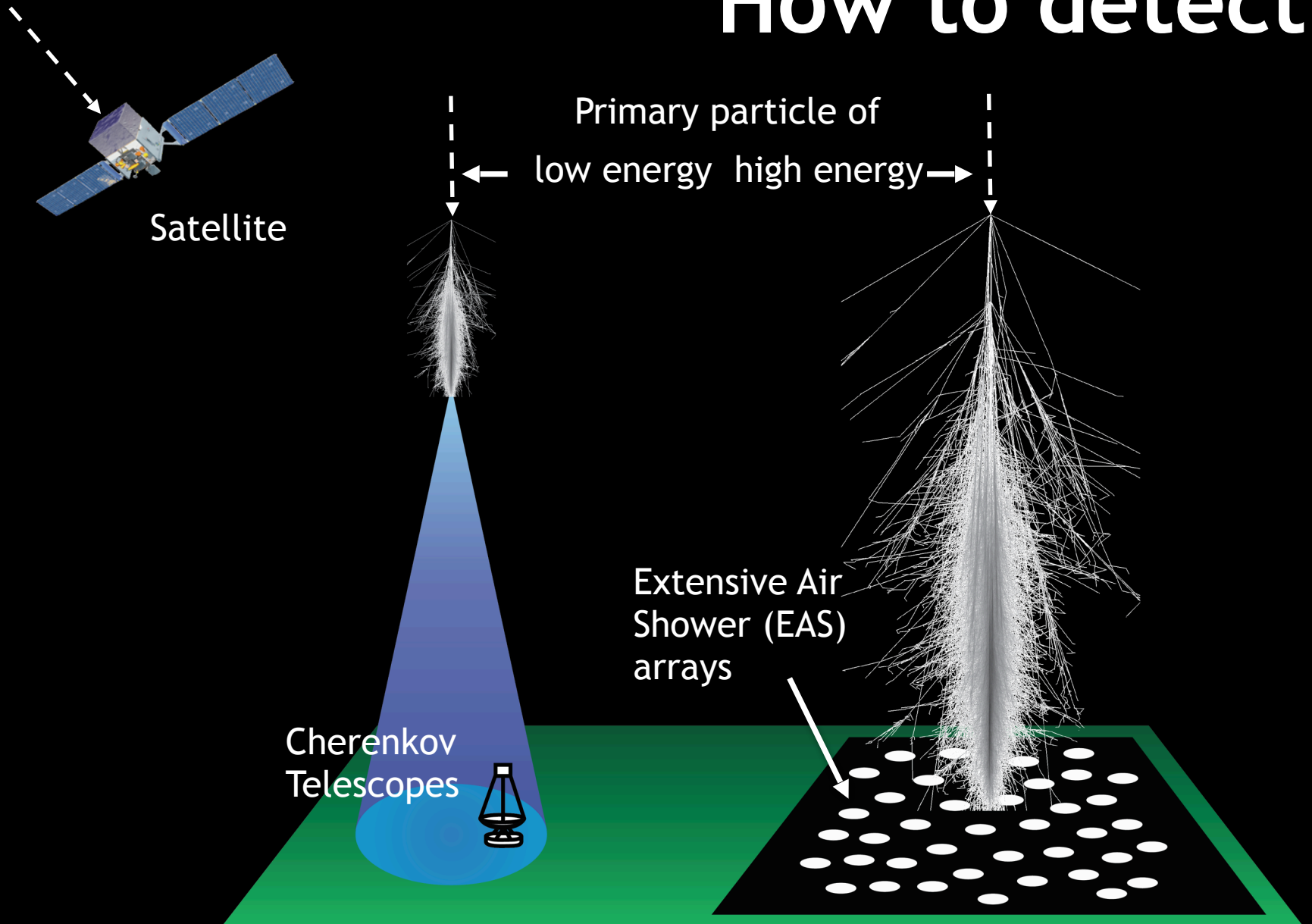
How to detect?



How to detect?



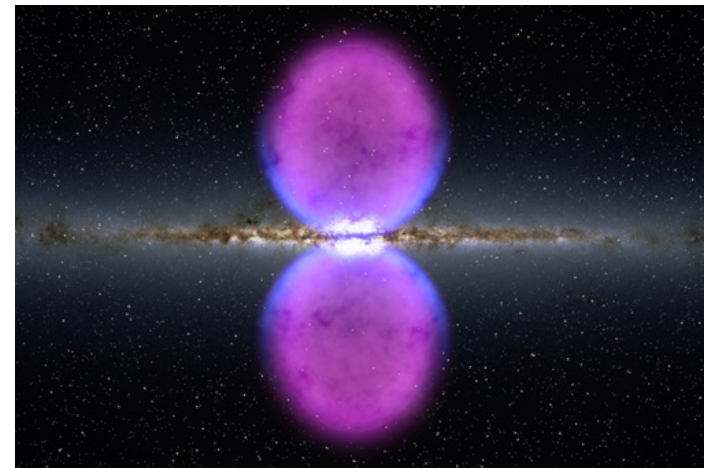
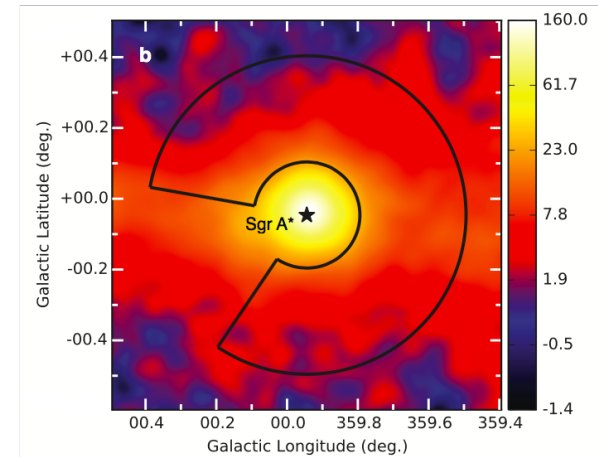
How to detect?



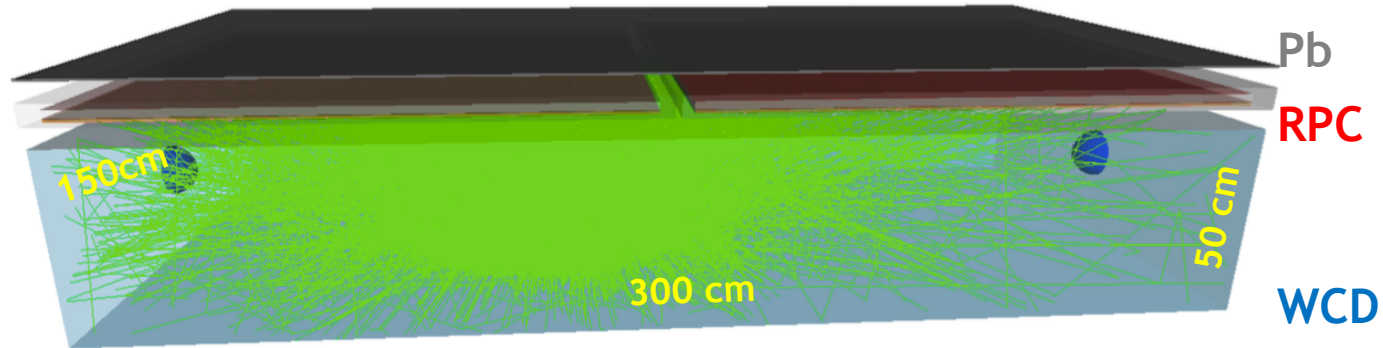
Arrays at high-altitude = large field of view + large duty cycle + low energy

What we know so far...

- ❖ Protons are known to be accelerated in the galaxy up to PeV energies ($E = 10^{15}$ eV)
- ❖ All current **acceleration models** encounter non-trivial **difficulties** at these energies
- ❖ HESS data suggests that there might be a **PeVatron source in the galactic center**
- ❖ **Fermi bubbles** - gamma ray emission in outbursts from our galaxy

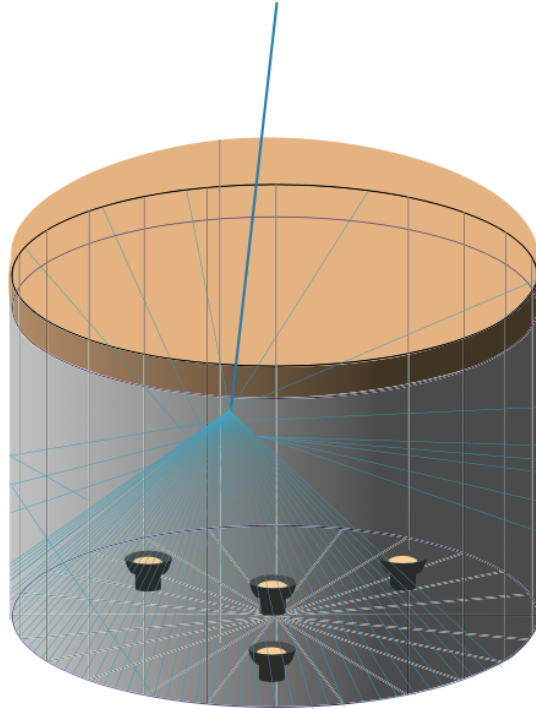


Improve detector concept!

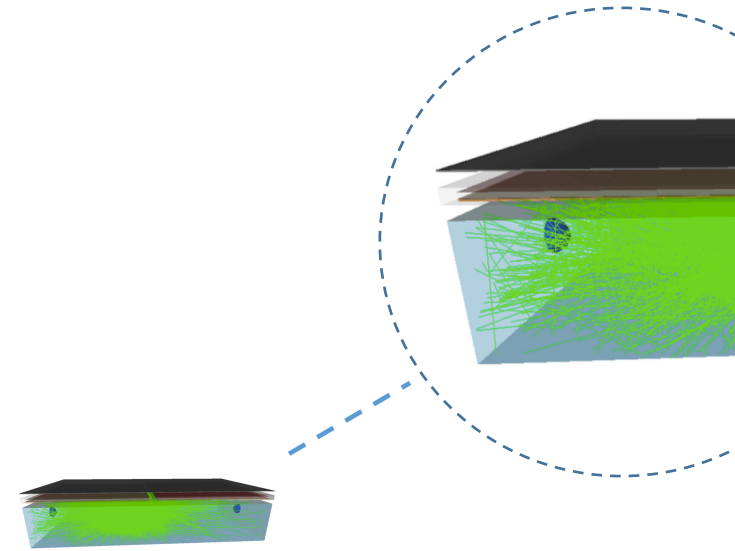


- ❖ **Thin lead converter plate (Pb)**
 - ❖ Improve shower geometry reconstruction
- ❖ **Resistive Plate Chamber (RPC)**
 - ❖ Measure charged particles with high spatial and time resolution
- ❖ **Water Cherenkov Detector (WCD)**
 - ❖ Collect shower secondary photons/electrons to improve trigger at low energy

The station



HAWC
(present detector)



...
(next generation)

**caveat: R&D phase, which means that the detector concept continues to evolve...*