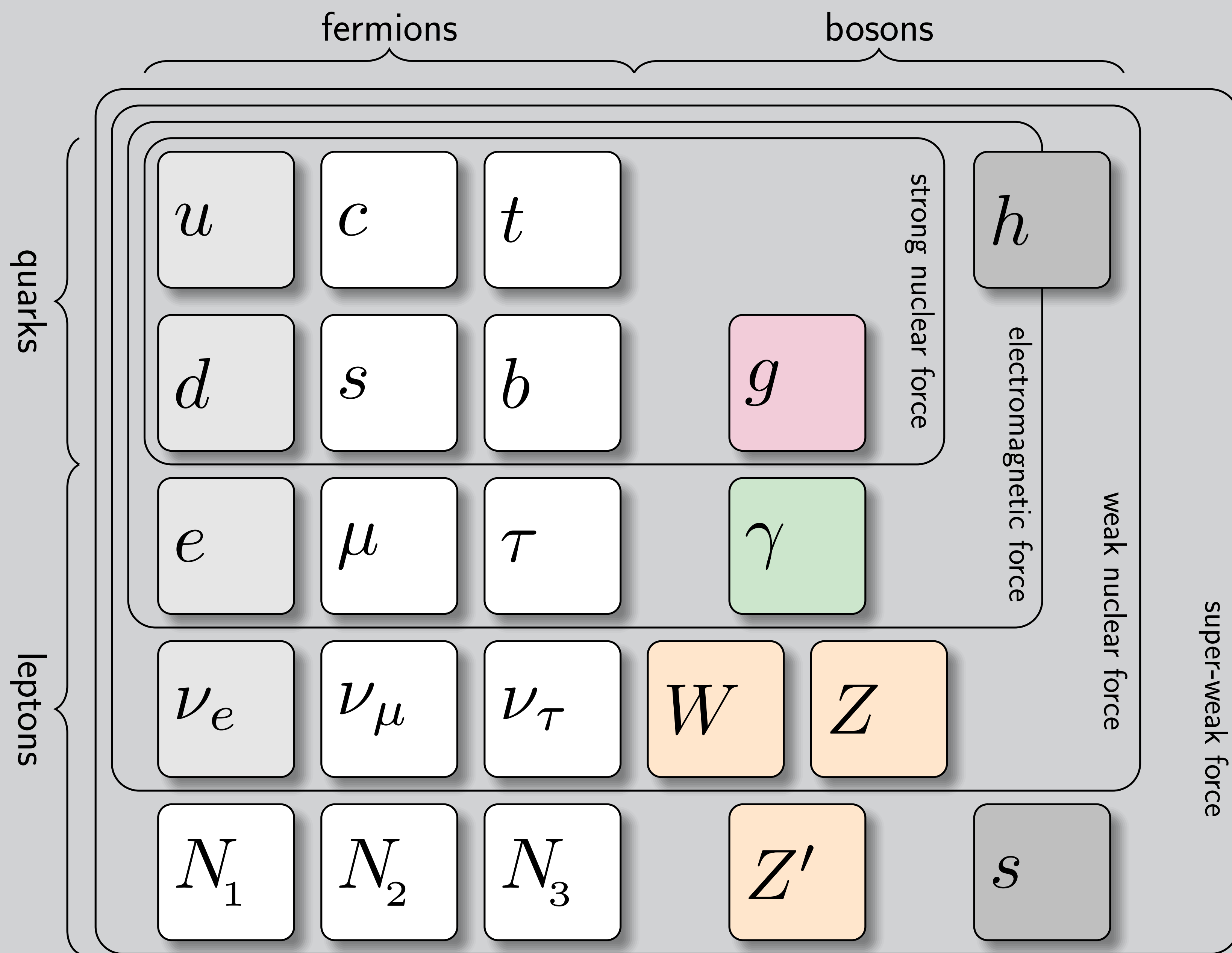


# Neutrino mass and phenomenology from a super-weak $U(1)_Z$ symmetry

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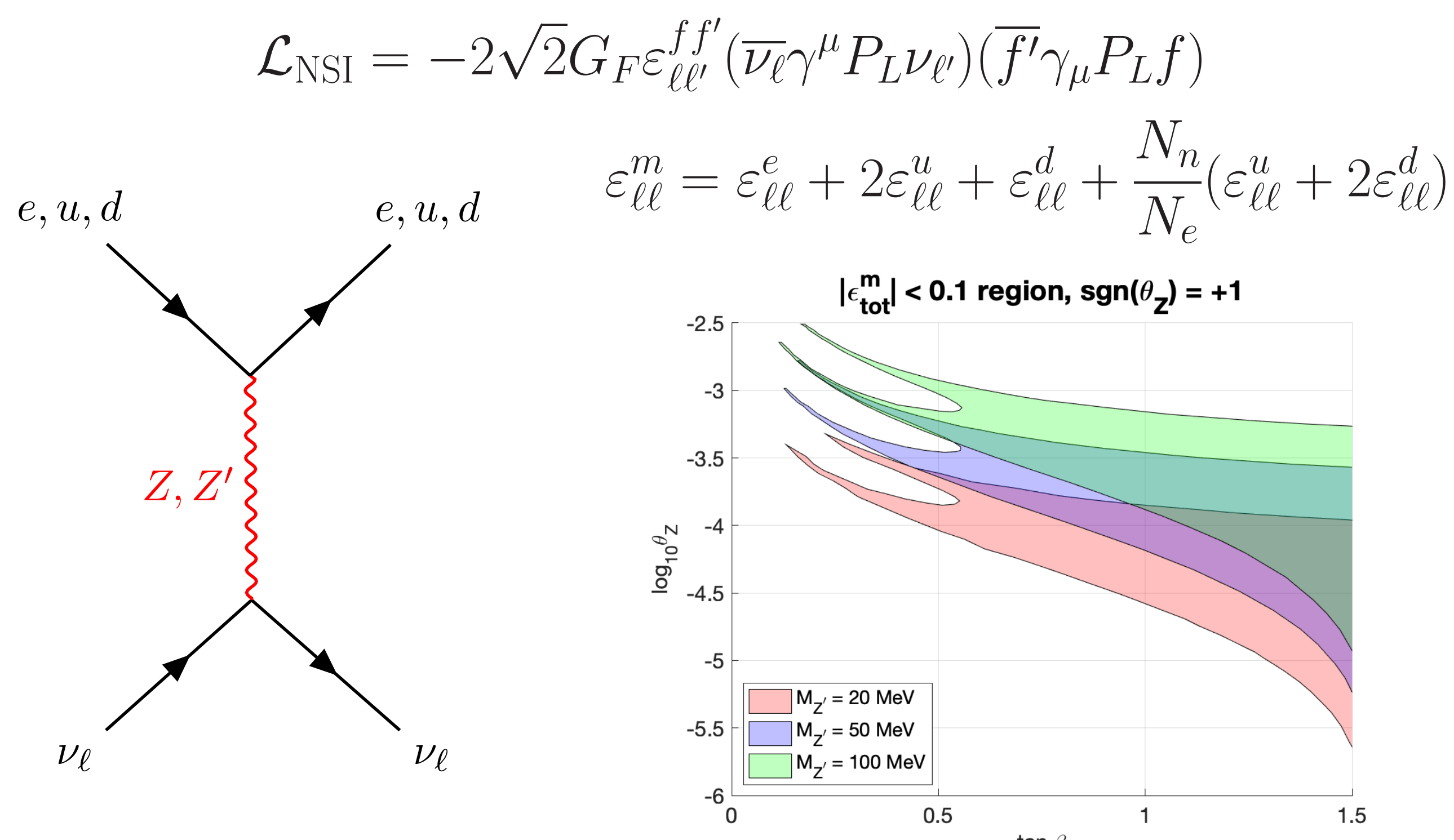


	$SU(2)_L$	$U(1)_Y$	$U(1)_Z$
$Q_L$	<b>2</b>	$\frac{1}{6}$	$\frac{1}{6}$
$u_R$	<b>1</b>	$\frac{2}{3}$	$\frac{7}{6}$
$d_R$	<b>1</b>	$-\frac{1}{3}$	$-\frac{5}{6}$
$L_L$	<b>2</b>	$-\frac{1}{2}$	$-\frac{1}{2}$
$\ell_R$	<b>1</b>	$-1$	$-\frac{3}{2}$
$\phi$	<b>2</b>	$\frac{1}{2}$	<b>1</b>
$N_R$	<b>1</b>	<b>0</b>	$\frac{1}{2}$
$\chi$	<b>1</b>	<b>0</b>	$-1$

## Introduction

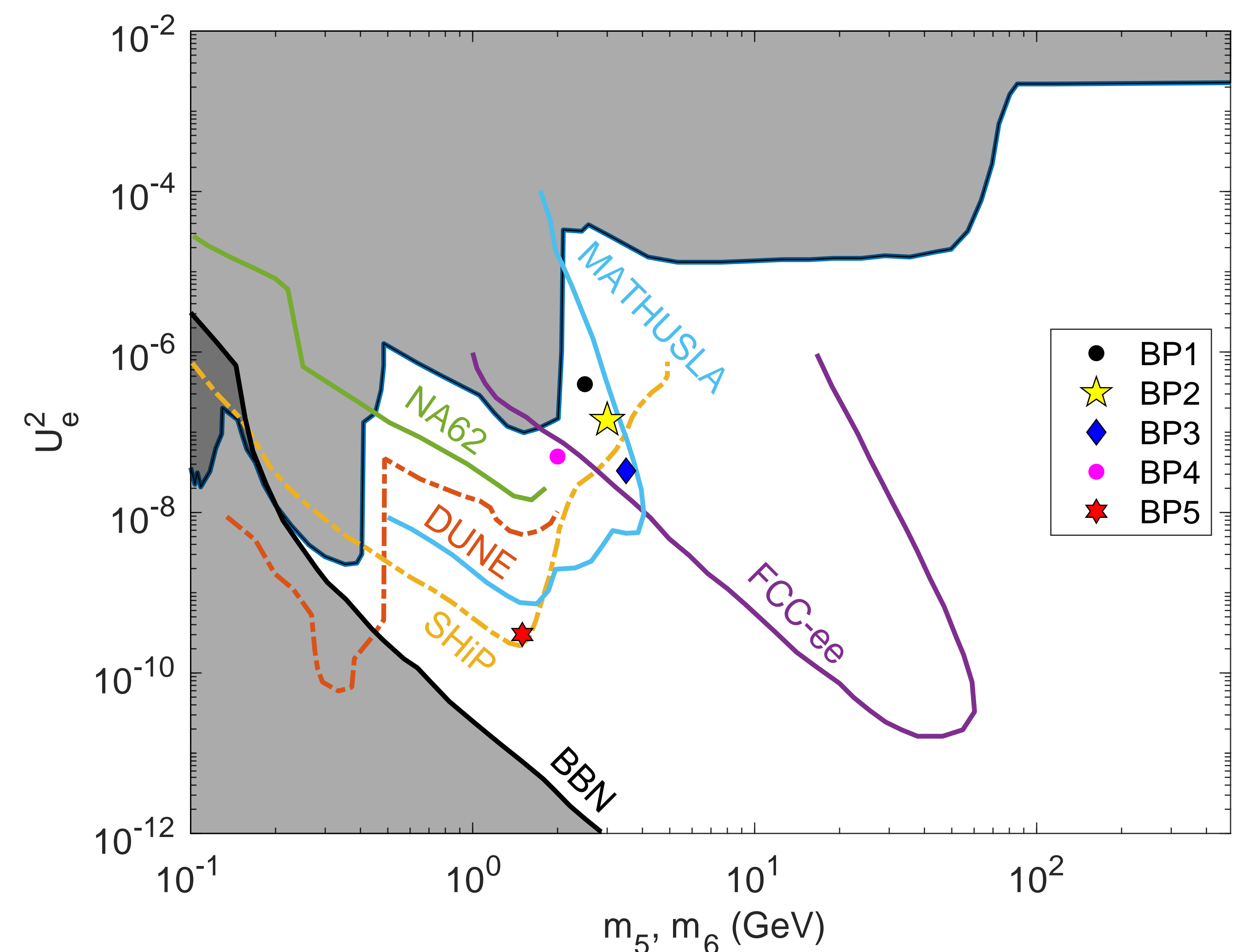
The super-weak model [1] is an economical  $U(1)$ -extension of the SM, aiming to explain the origin of neutrino mass and oscillations [2,3], dark matter [4], matter-antimatter asymmetry of the universe, cosmic inflation and to stabilize the vacuum [5]. New physics is manifested by nonstandard interactions and active-sterile neutrino mixing. The benchmark points provided fall within the sensitivity region of SHiP and MATHUSLA experiments.

## Nonstandard interactions (NSI)



Particle spectrum contains an extra neutral gauge boson,  $Z'$ , which mixes with  $Z$ . The mixing angle  $\theta_Z$  is constrained by NSI. We fix the  $Z'$  boson mass on MeV scale.

## Experiments are sensitive to GeV neutrinos



Benchmark point	BP1	BP2	BP3	BP4	BP5
$m_1$ (meV)	10	1	0	0.1	5
$m_4$ (keV)	30	7.1	40	50	25000
$m_{5,6}$ (GeV)	2.5	3.0	3.5	2.0	1.5
$w$ (GeV)	100	750	250	500	175

## One-loop correction to light neutrino masses

$$\delta M_L = \frac{1}{16\pi^2} \sum_{k=1,2} \left[ 3(Z_G)_{k1}^2 \frac{M_{V_k}^2}{v^2} \mathbf{F}(M_{V_k}^2) + (Z_S)_{k1}^2 \frac{M_{S_k}^2}{v^2} \mathbf{F}(M_{S_k}^2) \right]$$

We calculated the 1-loop  $m_\nu$  correction and found it to be  $\mathcal{O}(0.1)$  %.

$$\mathbf{F}_{ij}(M^2) = \sum_{a=1}^6 (\mathbf{U}_L^\dagger)_{ia} (\mathbf{U}_L)_{aj} \frac{m_a^3 \ln \frac{m_a^2}{M^2}}{M^2 m_a^2 - 1}$$

## References

- [1] Z. Trócsányi, Symmetry **12**, 107 (2020), arXiv:1812.11189.
- [2] S. Iwamoto, T. J. Kärkkäinen, Z. Péli and Z. Trócsányi, arXiv:2104.14571
- [3] T. J. Kärkkäinen and Z. Trócsányi, arXiv: 2105.13360
- [4] S. Iwamoto, K. Sella and Z. Trócsányi, arXiv: 2104.11248
- [5] Z. Péli, I. Nándori and Z. Trócsányi, Phys. Rev. D **101** (2020), arXiv: 1911.07082

