

Signatures of Primordial Black Holes in Theories of Large Extra Dimensions

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Overview

- Primordial Black Holes
- Large Extra Dimensions
- Ways To Observe Primordial Black Holes
- Conclusions

Primordial Black Holes

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- PBHs **evaporate** via **Hawking radiation**
- If PBHs survive they act as **Cold Dark Matter**
- PBHs have been studied extensively in 4D
 - But in **extra-dimensions** black holes behave differently!

Large Extra-Dimensions

- ADD model proposes M_* is true scale of Quantum Gravity (Arkani-Hamed et al. arxiv:hep-ph/9803315)

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- When $r \ll R$

$$V(r) = \frac{m_1 m_2}{M_*^{2+N_{\text{ED}}} r^{1+N_{\text{ED}}}}$$

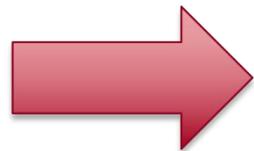
Large Extra-Dimensions

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- N_{ED} extra dimensions are **compactified** to size R
- When $r \ll R$

$$V(r) = \frac{m_1 m_2}{M_*^{2+N_{\text{ED}}} r^{1+N_{\text{ED}}}}$$

- When $r \gg R$

$$V(r) \approx \frac{m_1 m_2}{M_*^{2+N_{\text{ED}}} R^{N_{\text{ED}}} r}$$



$$M_*^{2+N_{\text{ED}}} R^{N_{\text{ED}}} = M_{\text{pl}}^2$$

Extra Dimensional Black Holes

- Black holes with $r_s \ll R$ have **modified size and temperature**

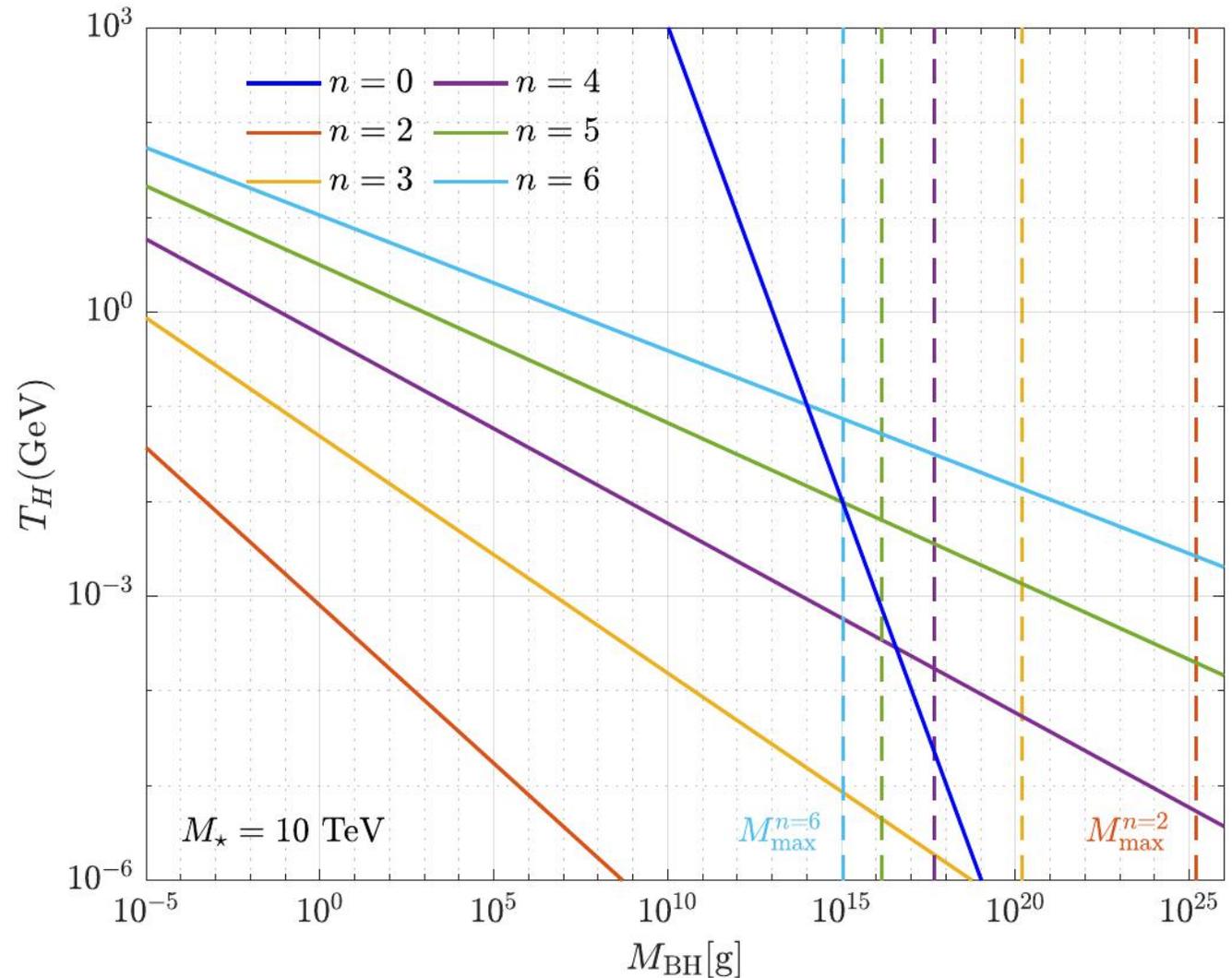
$$r_s = \frac{a(N_{\text{ED}})}{M_*} \left(\frac{M}{M_*} \right)^{\frac{1}{N_{\text{ED}}+1}}$$

$$T_{BH} = \frac{N_{\text{ED}} + 1}{4\pi r_s}$$

See Conley and Wizanksy
arxiv:hep-ph/0611091

Extra Dimensional Black Holes

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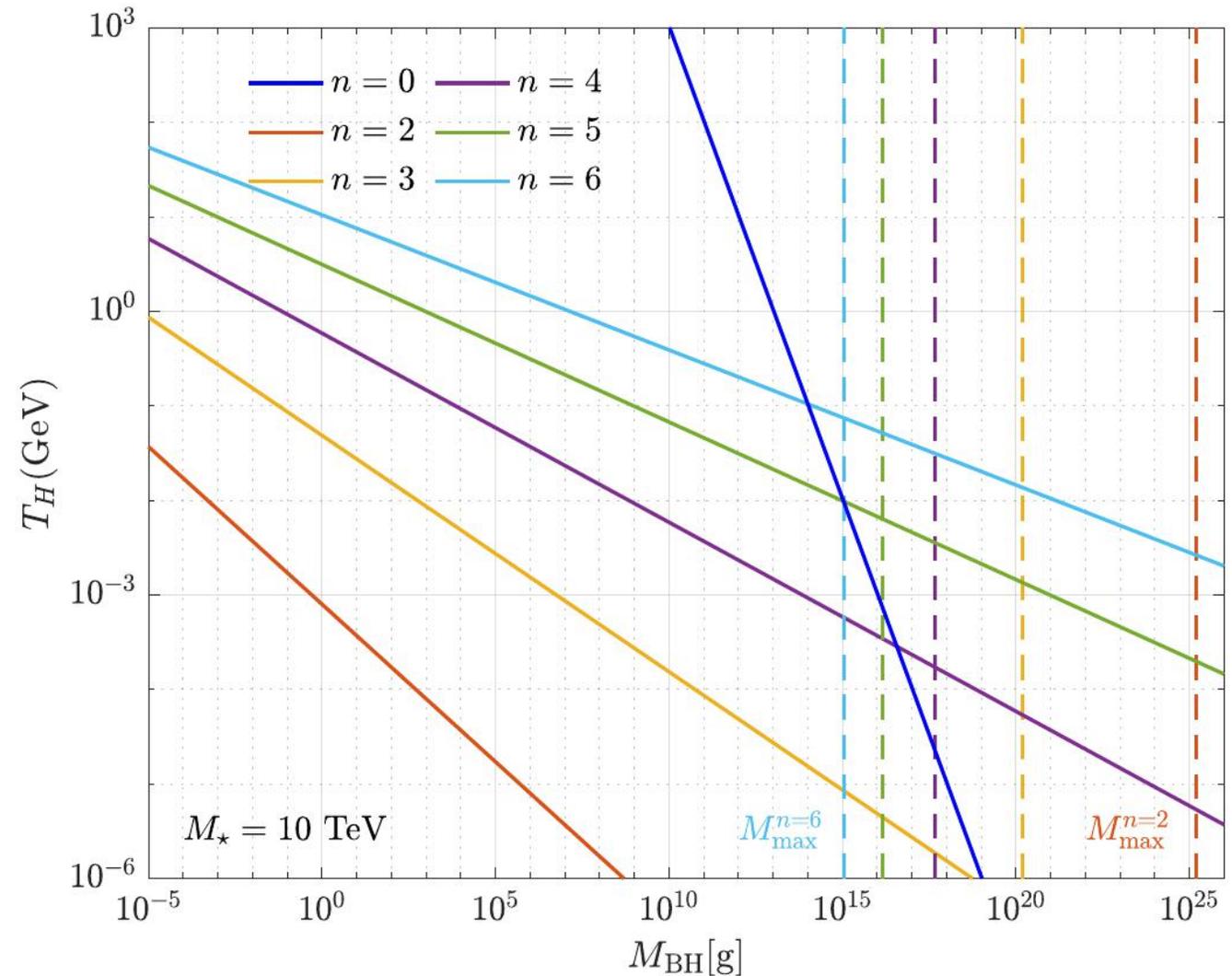


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Extra Dimensional Black Holes

- Black holes with $r_s \ll R$ have **modified size and temperature**
- Particle collisions at $E > M_*$ can **produce microscopic black holes**

See Conley and Wizanksy
arxiv:hep-ph/0611091



Black Hole Evolution

- Black hole mass evolves due to **accretion** and **evaporation**

$$\frac{dM}{dt} = \left(-\alpha + \beta \frac{T^4}{T_H^4} \right) T_H^2$$

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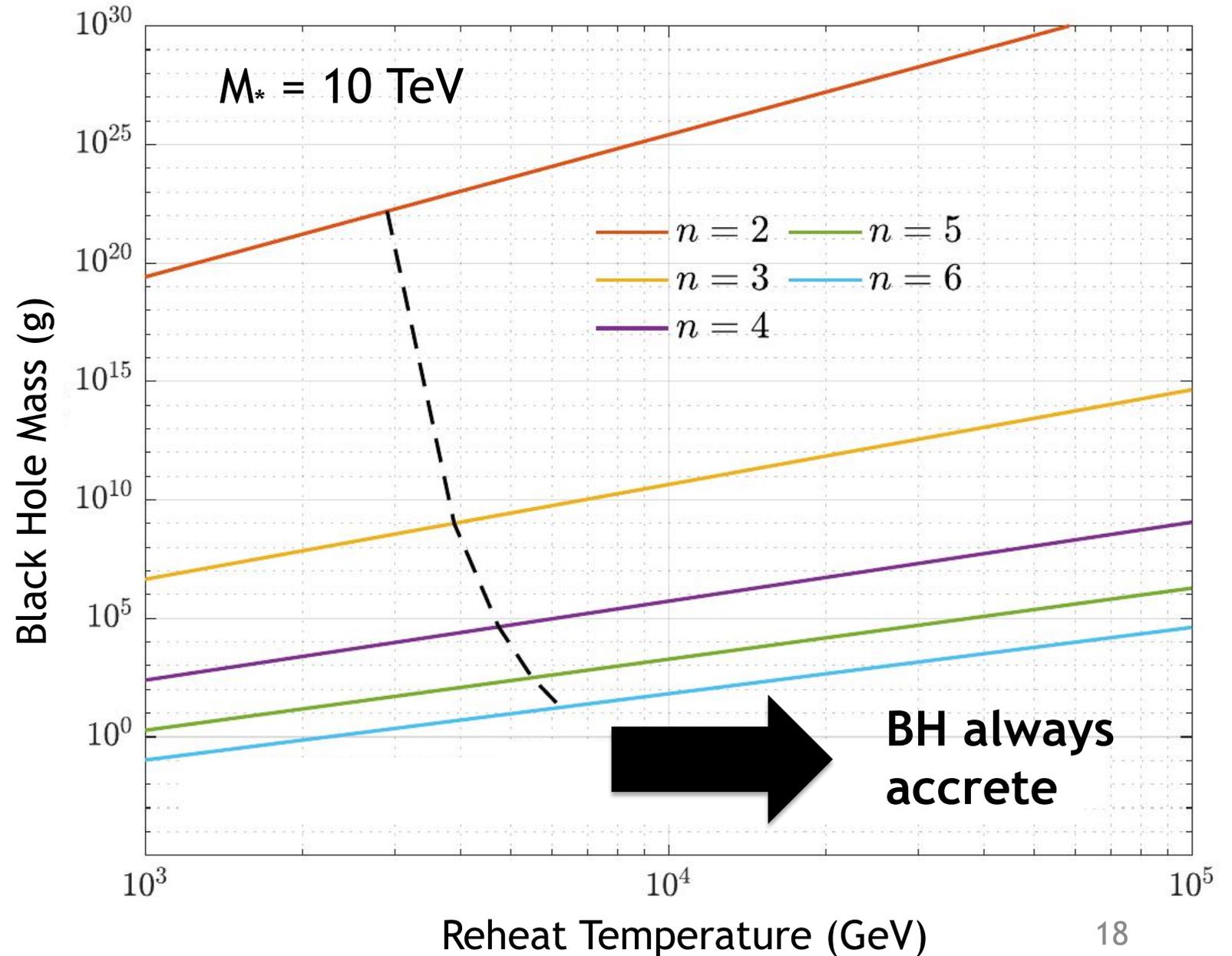
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$$\frac{dM}{dt} = \left(\underbrace{-\alpha}_{\text{Evaporation}} + \underbrace{\beta \frac{T^4}{T_H^4}}_{\text{Accretion}} \right) T_H^2$$

- Typically either accretion or evaporation dominate

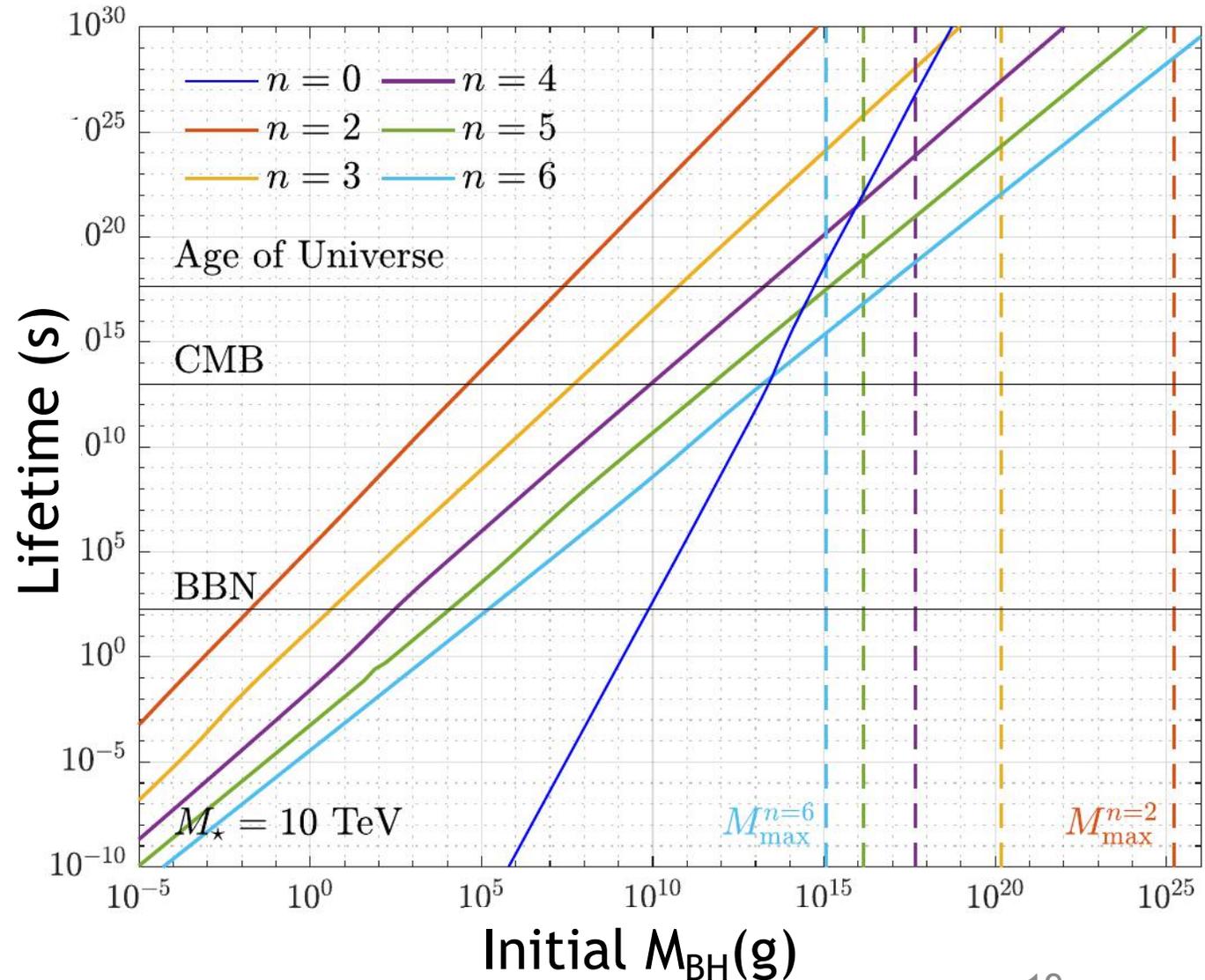
Black Hole Growth

Black holes in extra-dimensions initially **grow via accretion** depending if they are created when the universe is sufficiently hot



Black Hole Evaporation

- After accretion stops due to the universe cooling, black holes evaporate via Hawking radiation
- Lifetime depends on initial mass and number of dimensions



Evaporation Spectrum

- Black holes evaporate to **all sufficiently light particles** with a grey-body spectrum

$$dP = \sigma_i \frac{\omega}{\exp(\omega/T_H) \mp 1} \frac{d^3 p}{(2\pi)^3}$$

Evaporation Spectrum

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- **Secondary photons** and **electrons** produced from **cascades** of unstable evaporation products
 - PPC4DMID used for secondary spectrum

Cirelli et al. arxiv:1012.4515

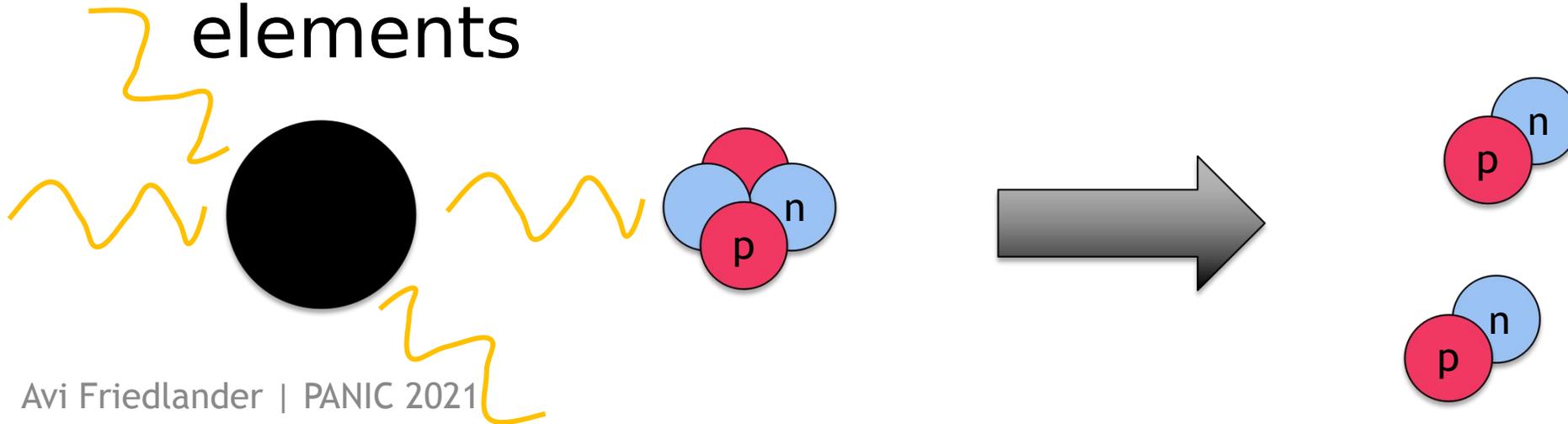
Constraining Large Extra-Dimensions

- Astrophysical constraints on PBHs come from a variety of sources:

- Big Bang Nucleosynthesis (**BBN**)
- **Galactic centre** photon flux
- **Isotropic x-ray** and **gamma-ray** flux
- **CMB** angular power spectrum
- **21cm** Cosmology

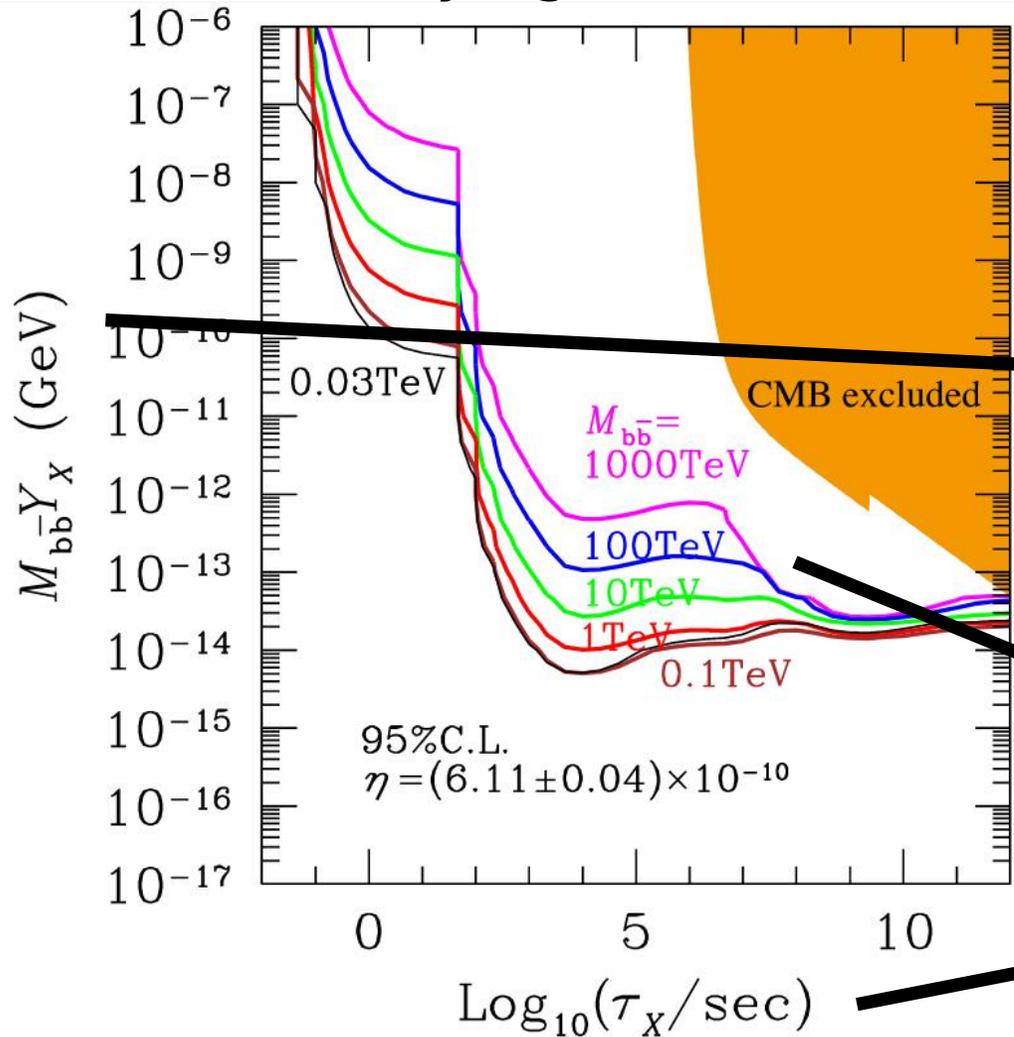
PBH and Big Bang Nucleosynthesis

- When the universe cools to ~ 1 MeV light **nuclei form**
- PBHs can change the **expansion rate** during BBN
- Evaporation products can **dissociate** primordial elements

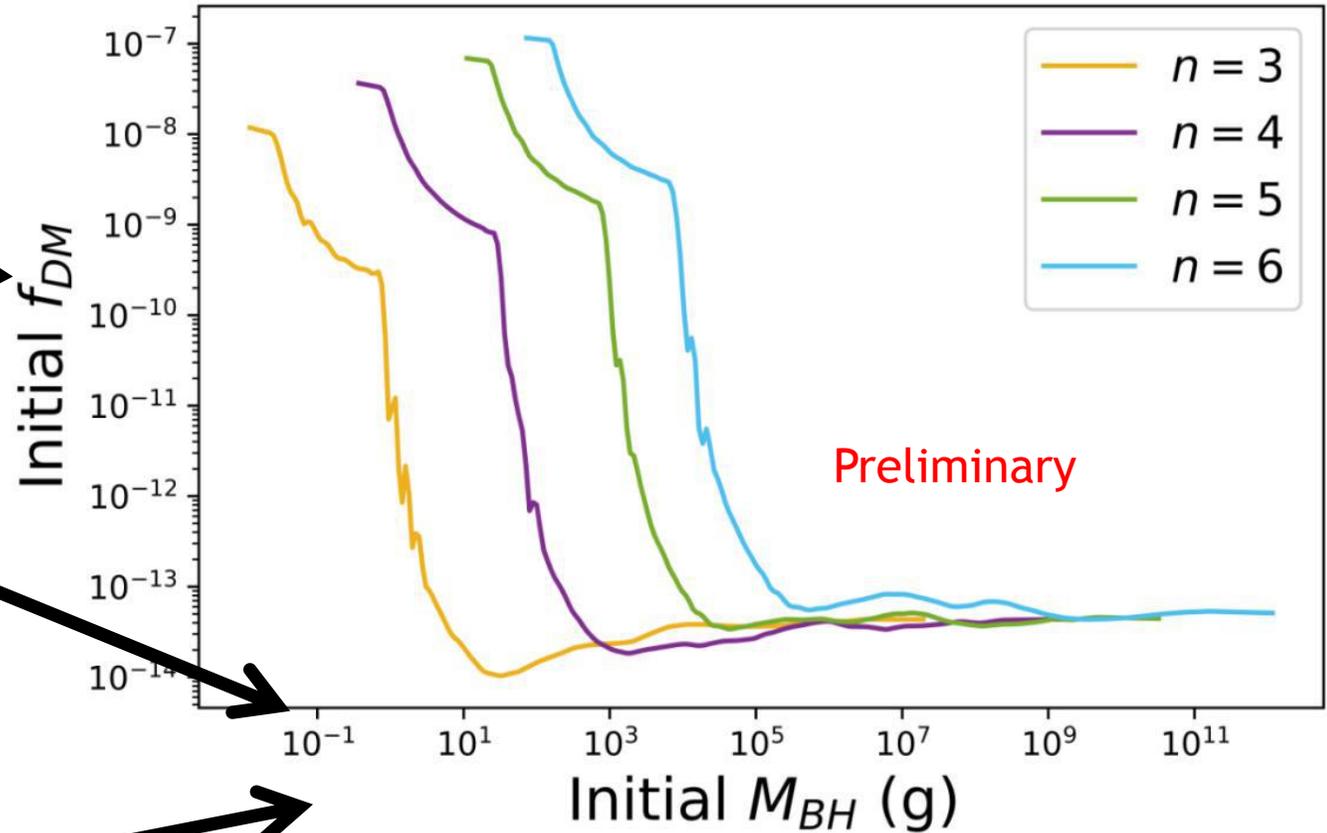


Preliminary BBN Results

Decaying Dark Matter



Primordial Black Holes



Kawasaki et al. arxiv:1709.01211

See Kieth et al. arxiv:2006.03608
for recasting details

Galactic Centre Photons

$$\frac{d\Phi_\gamma}{dEd\Omega} = \frac{f_{PBH}}{4\pi M_{BH}} \frac{dN}{dEdt} \frac{\mathcal{D}(\Omega)}{\Delta\Omega}$$

$$\mathcal{D}(\Omega) \equiv \int_{l.o.s. \Delta\Omega} \rho_{DM}(\vec{x}) d\Omega dx$$

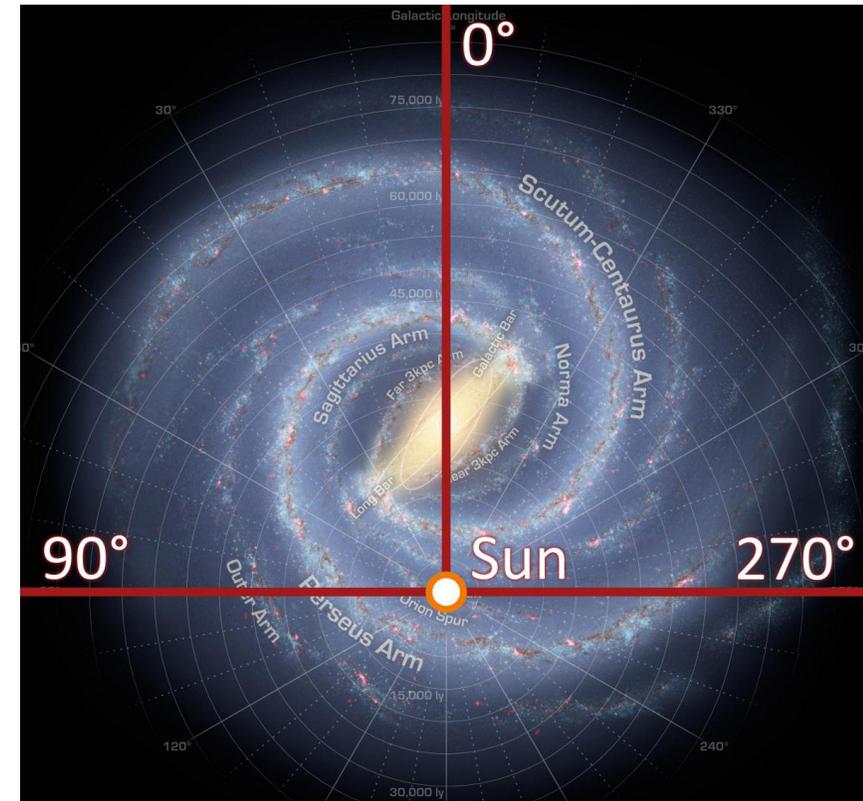


Image from NASA

- Compared to previously processed 6-years of INTEGRAL/SPI data (Bouchet et al. arxiv:1107.0200)

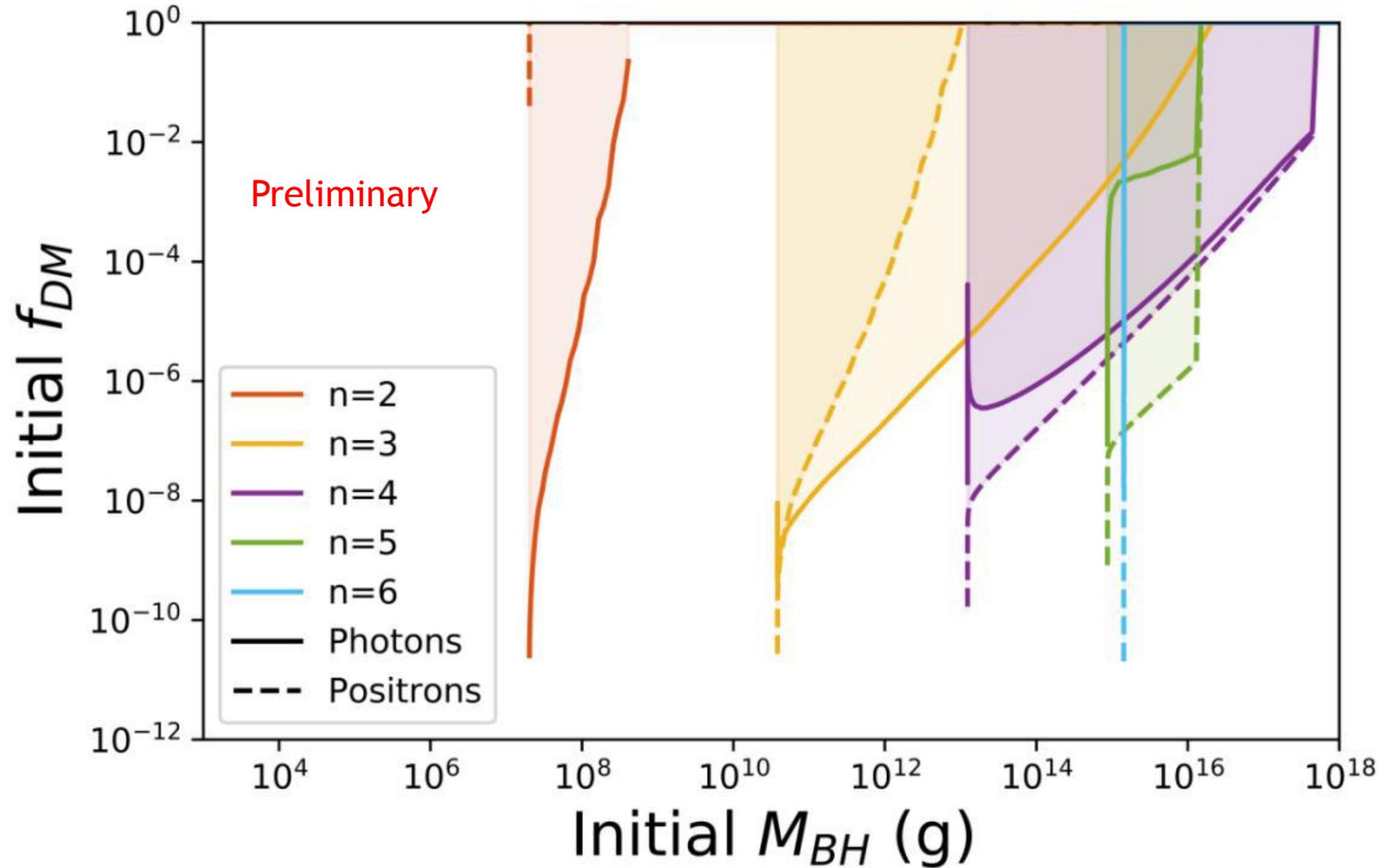
Galactic Centre Positrons

- If PBHs evaporate to positrons, they will annihilate to produce an additional 511 keV signal

$$\frac{d\Phi_{511}}{d\Omega} = 2(1 - 0.75f_P) \frac{dN_{e^+}}{dt} \frac{1}{4\pi} \frac{1}{M_{BH}} \frac{1}{\Delta\Omega} \mathcal{D}(\Omega)$$

- Compared to INTEGRAL/SPI 511 keV line data (Siegert et al. arxiv:1906.00498)

Preliminary Galactic Centre Results



Extragalactic Background Light

- Photon spectrum tracked from recombination to today

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- Each redshift step updated using:

$$\frac{d\Phi_{\gamma,\text{EBL}}}{dE}(E, z_i) = \frac{dE'}{dE} \frac{d\Phi_{\gamma,\text{EBL}}}{dE'}(E', z_{i-1}) e^{-\tau} + n_{BH}(z_i) \frac{d^2 N_\gamma}{dE dt} \frac{z_{i-1} - z_i}{H(z_i)(1 + z_i)}$$

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Energy Loss

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Extragalactic Background Light

- Photon spectrum tracked from recombination to today
- Each redshift step updated using:

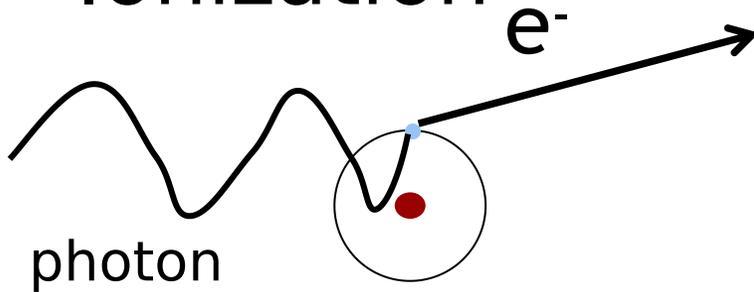
$$\frac{d\Phi_{\gamma,\text{EBL}}}{dE}(E, z_i) = \underbrace{\frac{dE'}{dE} \frac{d\Phi_{\gamma,\text{EBL}}}{dE'}(E', z_{i-1}) e^{-\tau}}_{\text{Energy Loss}} + \underbrace{n_{\text{BH}}(z_i) \frac{d^2 N_{\gamma}}{dE dt} \frac{z_{i-1} - z_i}{H(z_i)(1 + z_i)}}_{\text{Added Photons}}$$

Photon Energy Loss

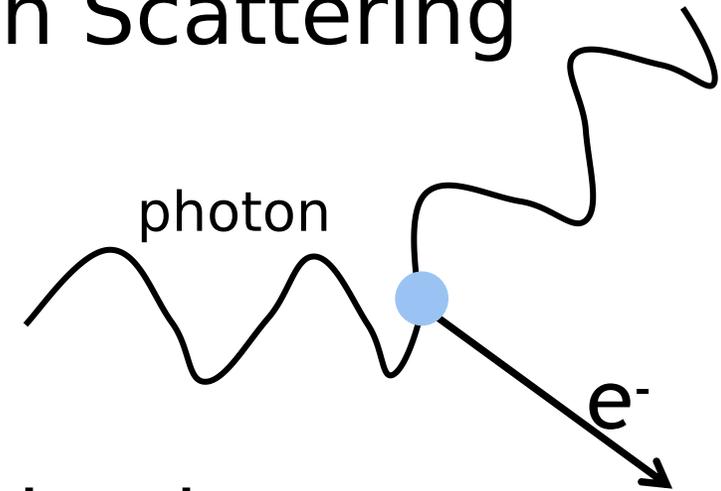
Redshifting



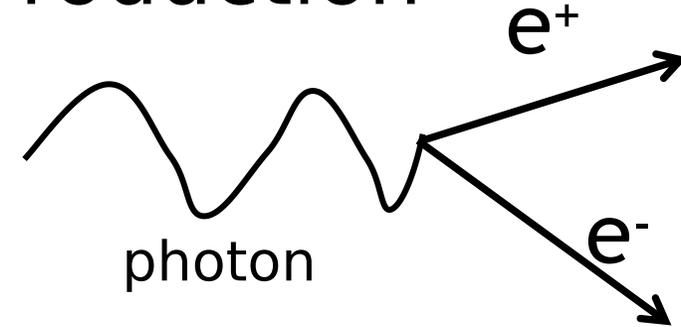
Ionization



Compton Scattering



Pair Production

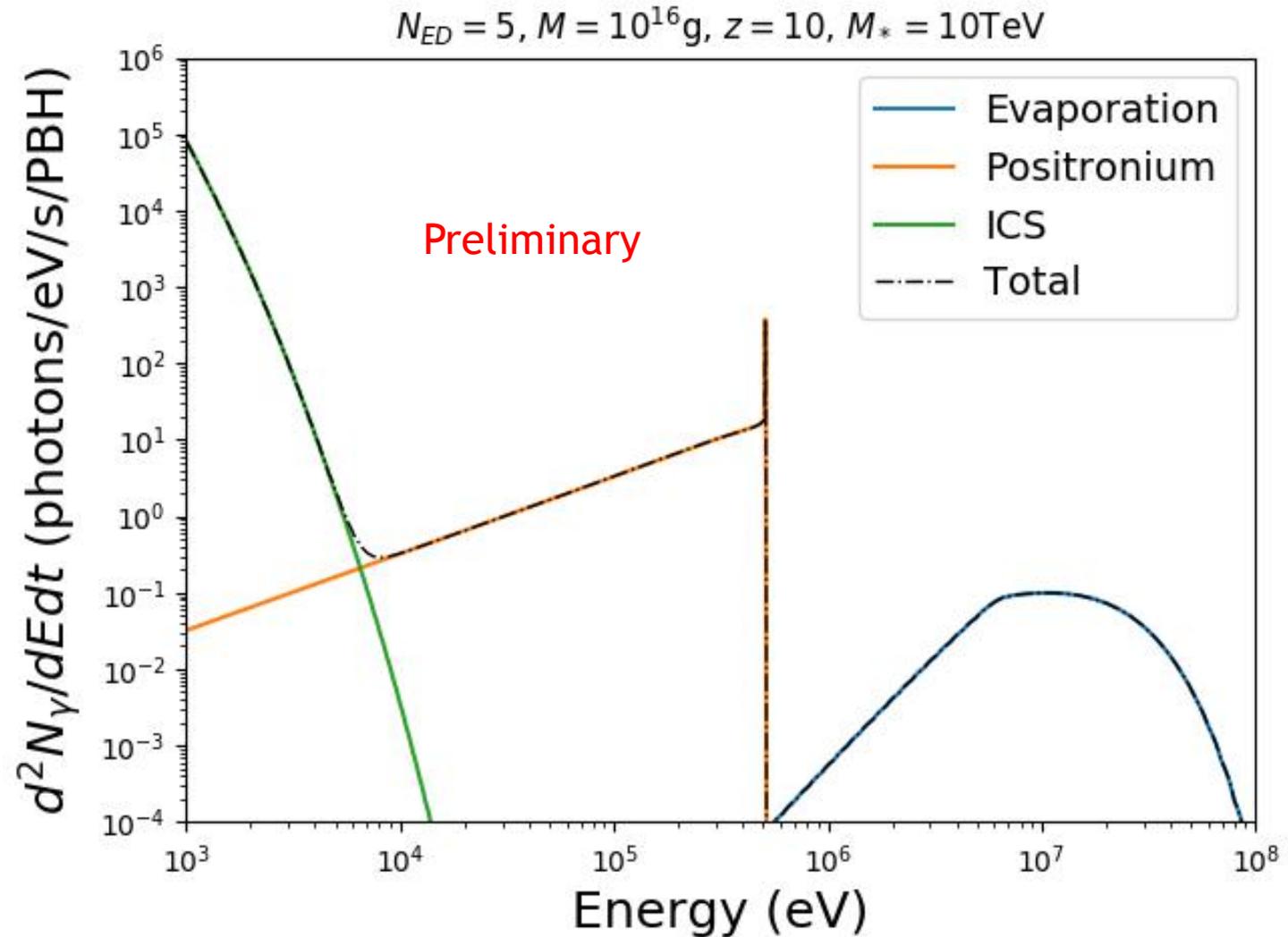


For the observed energy range (1 keV - 100 MeV) attenuation is important for high redshift ($z > 100$)

Extragalactic Photon Production

Extragalactic Photons produced via 3 mechanisms:

- Evaporation
- Positronium annihilation
- Inverse Compton scattering (ICS)



Galactic Isotropic Signal

- Galactic photon flux is highly anisotropic but non-zero in all directions
- Isotropic component can be determined with

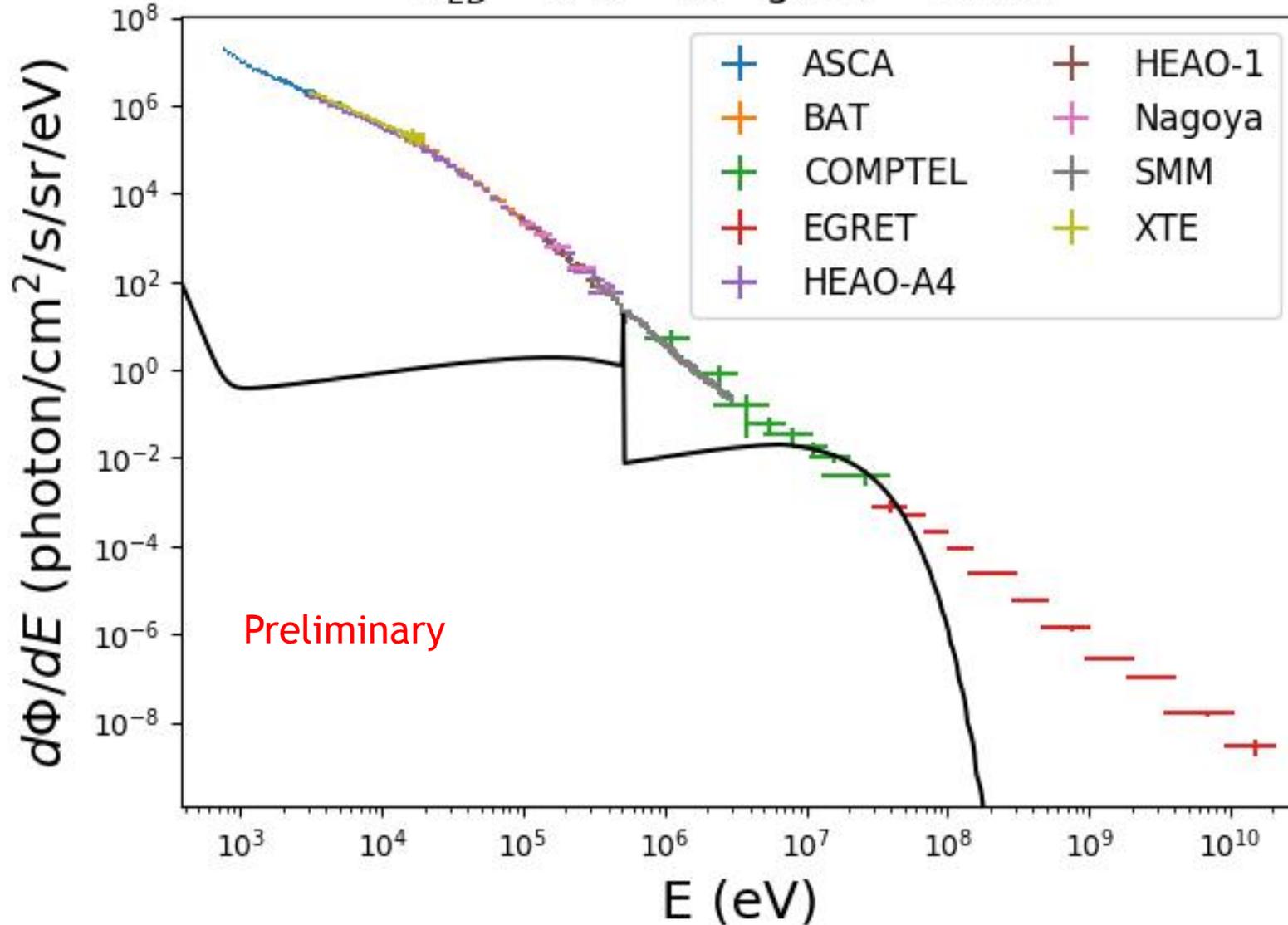
$$\frac{d\Phi_{\gamma,\text{gal}}}{dE} = \frac{f_{PBH}}{4\pi M_{BH}} \frac{d^2 N_{\gamma}}{dE dt}(E, M_{BH}) \mathcal{D}_{\text{min}}$$

$$\mathcal{D}_{\text{min}} = \int_{r_{\text{earth}}}^{\infty} dr \rho_{NFW}(r)$$

See: Iguaz et al.
arxiv:2104.03145

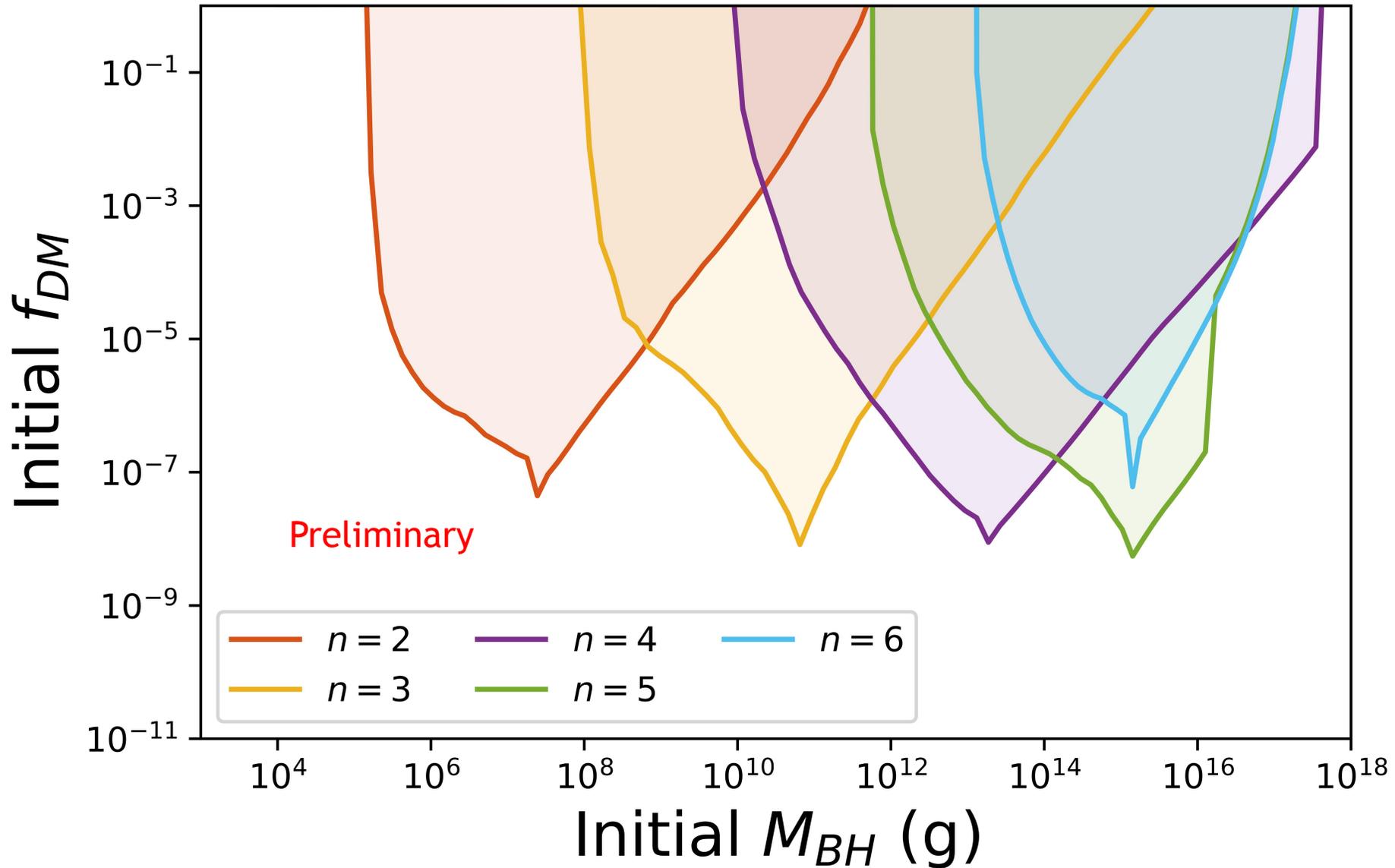
Observed Flux

$$N_{ED} = 5, M = 10^{16}g, M_* = 10\text{TeV}$$



Calculated flux is compared to various **x-ray and gamma-ray telescopes** as compiled by Ajello et al. (arxiv:0808.3377)

Preliminary Isotropic Light Constraints



Conclusions

- Theories of Large Extra Dimensions predict the existence of Primordial Black Holes
- Astrophysical observables of LED PBHs is very different from regular PBHs
- Constraints from BBN, galactic centre, and isotropic background light change significantly with additional large dimensions

Questions?

Extra Slides

The Hierarchy Problem

- Two scales exist in the Standard Model
 - Electroweak scale ($\sim 10^3$ GeV)
 - Planck/Quantum Gravity scale ($\sim 10^{18}$ GeV)
- Higgs boson mass is set by the electroweak scale but without fine tuning, quantum corrections would be expected to increase Higgs mass to the Planck

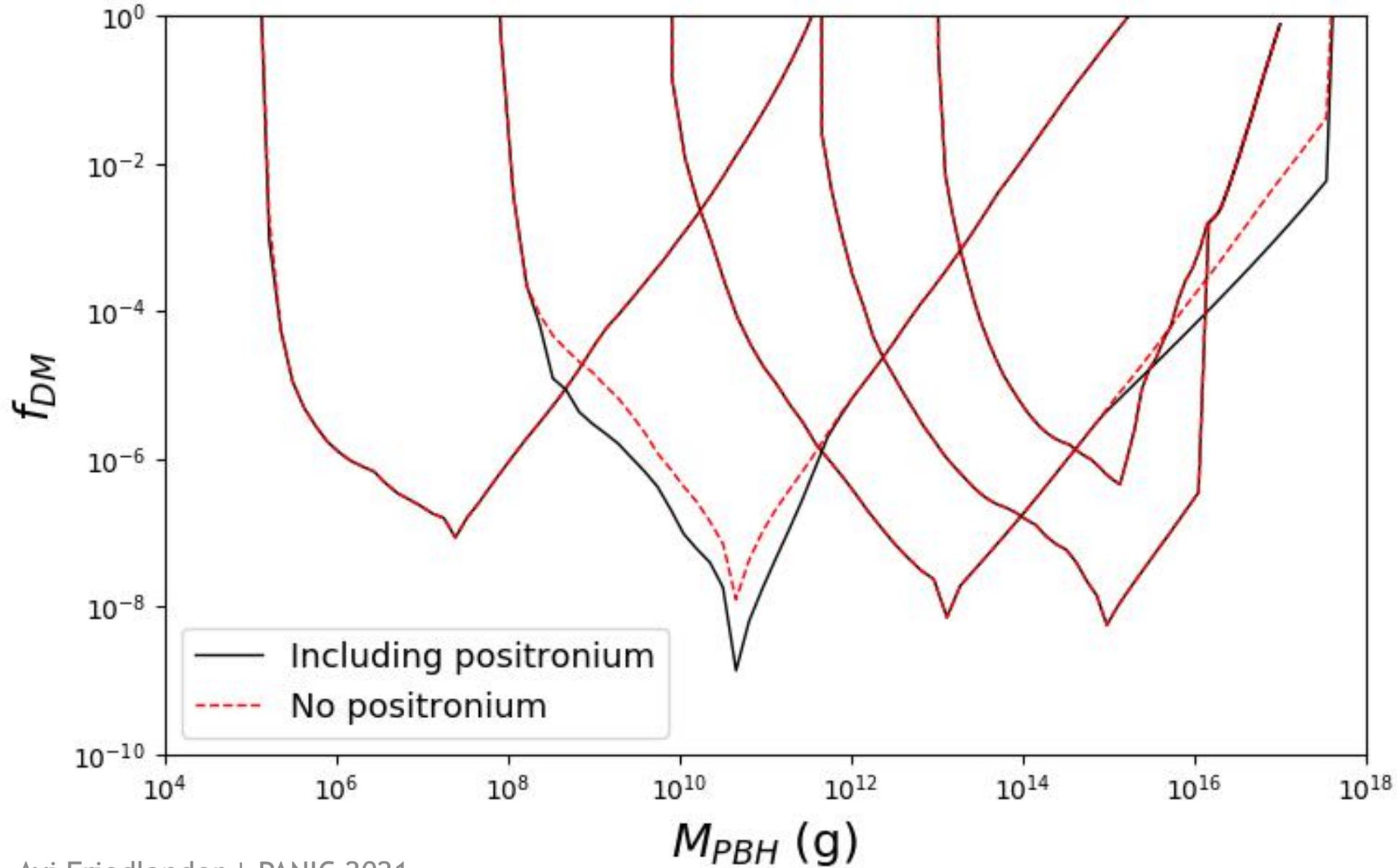
Detailed BBN Effects

- There are four main mechanisms for PBHs to affect BBN
 - Increased universe expansion rate causing earlier neutron freeze-out
 - Hadrons and mesons converting protons to neutrons after freeze-out
 - Energetic mesons dissociating Helium nuclei
 - Energetic photons dissociating Helium nuclei

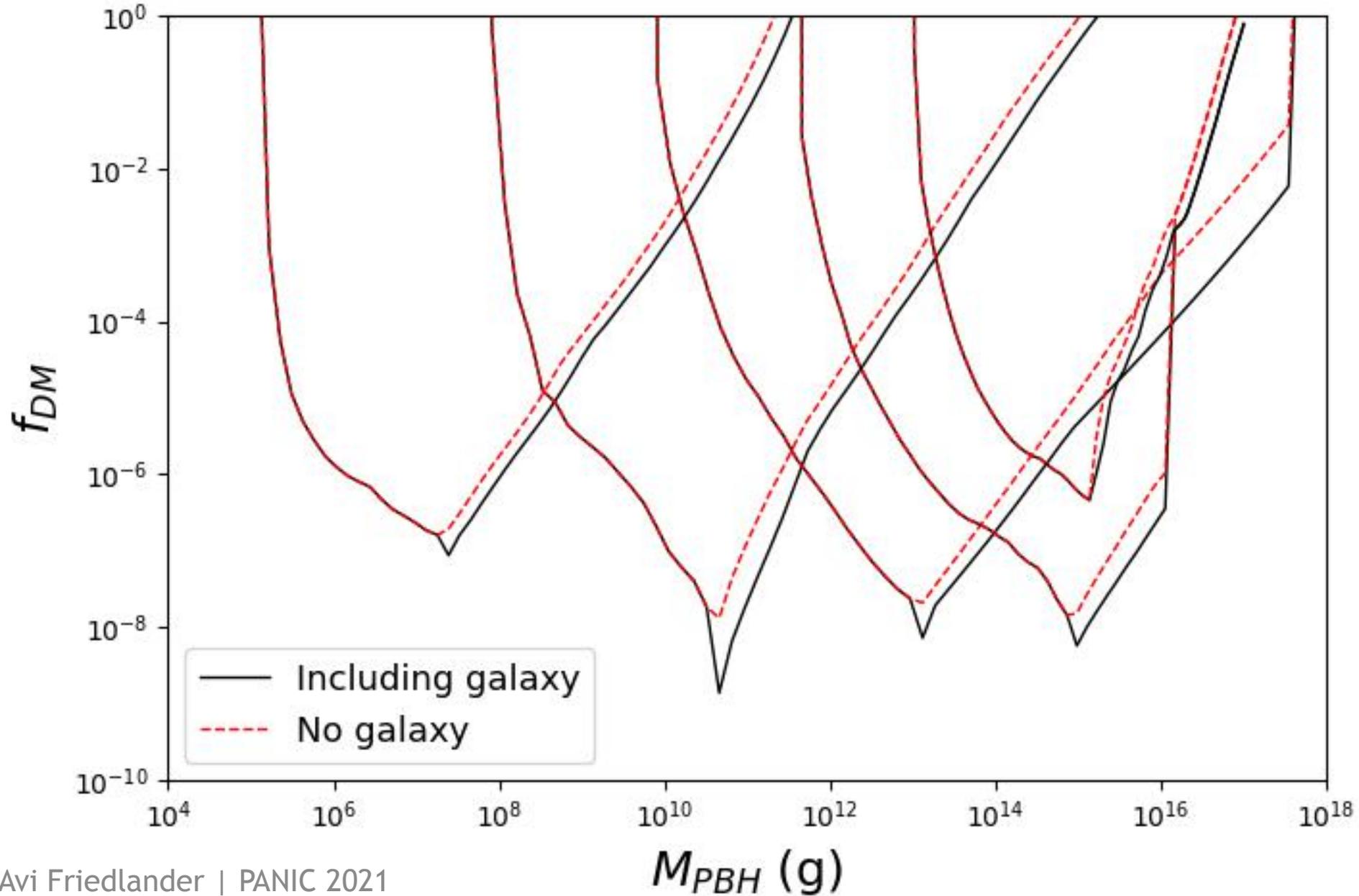
Calculating BBN Constraints

- The effect of decaying dark matter on BBN has been studied in detail (Kawasaki et al. arxiv:1709.01211)
- PBHs can be mapped onto decaying dark matter models ensuring key properties match
 - Dark matter/PBH density
 - Average injected fermion energy
 - Average time of injected energy (Kieth et al. arxiv:2006.03608)

Impact of Positronium



Impact of Milky Way



Inverse Compton Scattering

- High energy electrons and positrons upscatter CMB photons

$$\frac{d^2 N_{\gamma, \text{ics}}}{dE dt}(E, M_{BH}) = 2 \int_0^\infty dE_e \frac{d^2 N_{e^-}}{dE dt}(E_e, M_{BH}) \frac{d\tilde{N}_{\gamma, \text{ics}}}{dE}(E, E_e, T_{CMB})$$

Photons
produced by ICS

Electron
spectrum from
evaporating PBH

ICS photon spectrum
per electron
calculated by
DarkHistory
(arxiv:1904.09296)

Positronium

- Positrons annihilate into photons via formation of positronium
 - Each annihilation produces spectrum

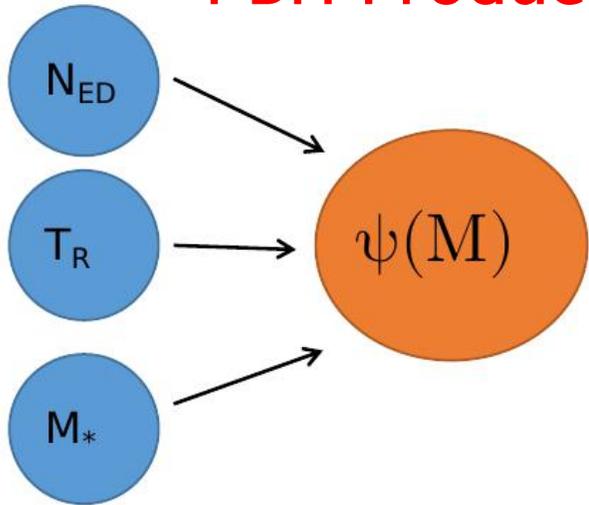
$$\frac{d\tilde{N}_\gamma^{\text{ann}}}{dE}(E) = \frac{1}{2}\delta(E - m_e) + \frac{3}{4} \frac{dN_\gamma^{\text{ann}}}{dE} \Big|_{\text{triplet}}$$

- All positrons assumed to immediately annihilate

See: Iguaz et al.
arxiv: 2104.03145

Next Steps

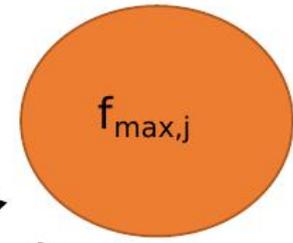
PBH Production



M_* - T_R exclusion curves

$$\sum_{j=1}^N \left(\int dM \frac{\psi(M)}{f_{\max,j}(M)} \right)^2 \leq 1,$$

Combine PBH constraints and production to constrain LEDs



Set constraints with other observables

