



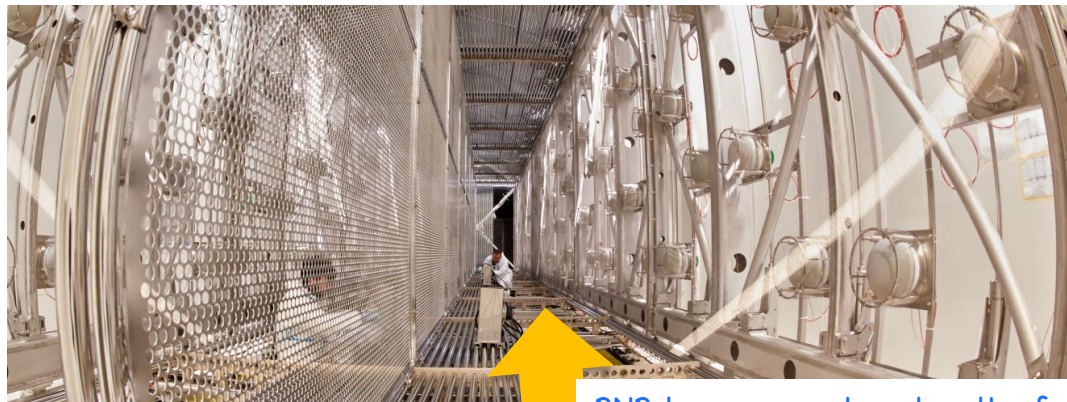
## Short-Baseline neutrino oscillation searches with the ICARUS detector

Bruce Howard – Fermilab  
On behalf of the ICARUS Collaboration

PANIC 2021 – 5 September 2021

# ICARUS at a glance

- Liquid Argon (LAr) time-projection chamber (TPC) detector
- First large LAr TPC: still one of the largest in operation
  - 2 modules, each  $19.6 \times 3.6 \times 3.9 \text{ m}^3$
  - Total 760 t LAr, 476 t active.
  - 2 TPCs per module with active drift distance  $\sim 1.5 \text{ m}$
- Originally deployed at LNGS, refurbished/moved to FNAL for Short Baseline Neutrino program (sterile  $\nu$  search)



BNB beam runs along length of detector

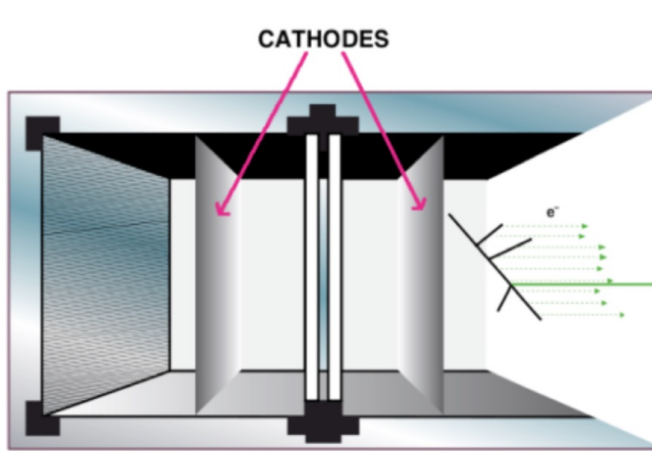


Above: photo of ICARUS being deployed at FNAL (FNAL VMS 18-0150-12). Left: inside the ICARUS detector.



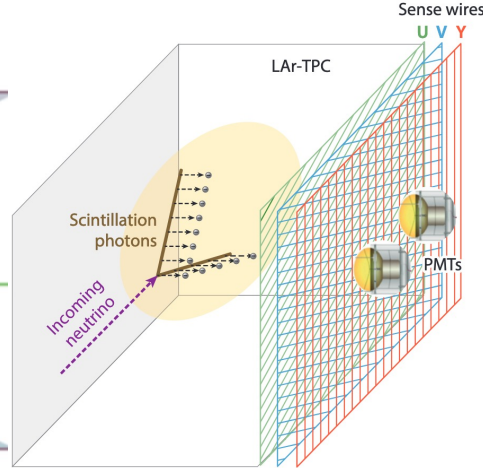
# Liquid argon TPCs

- LAr is powerful detection medium: charged particles produce ionization and scintillation
- E-field drifts ionization to wires to measure tracks/showers. Light provides prompt signal: in ICARUS, PMTs provide light detection
- Same detection principle as the future (multi-kt) DUNE experiment

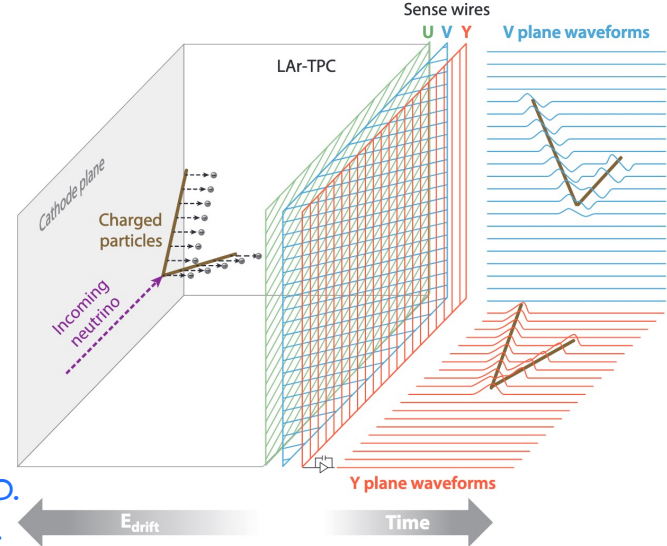


FRONT VIEW OF THE DETECTOR

(C. Rubbia et al. arXiv: 1106.0975)

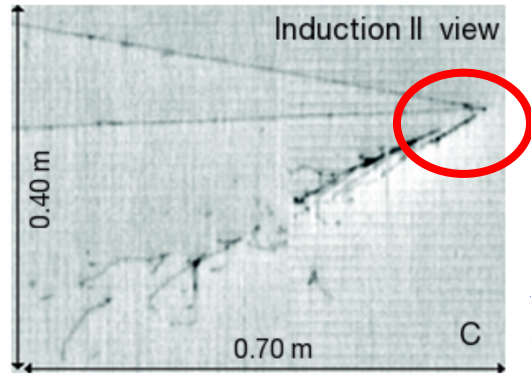


P. Machado, O. Palamara, and D. Schmitz. Annu. Rev. Nucl. Part. Sci. (2019). doi: 10.1146

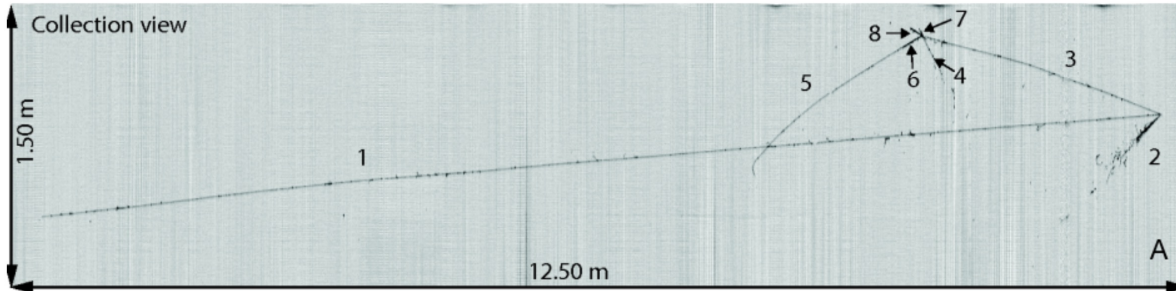


# Liquid argon TPCs

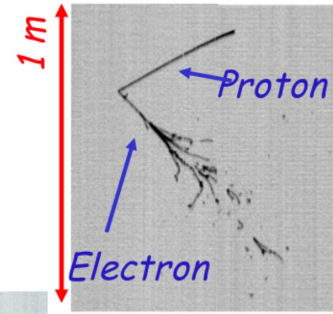
- LAr TPC detectors offer high resolution tracking (multiple planes of wires spaced mm apart) and calorimetry (proportional to amount of signal)
  - Determine properties of tracks (e.g. from muons) and showers (e.g.  $e/\gamma$  separation)



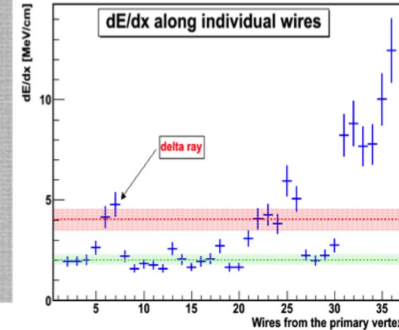
$\nu_\mu$  CC candidate from CNGS beam  
(C. Rubbia et al. arXiv: 1106.0975)



Atmospheric  $\nu_e$  CC event  
collected @LNGS



80 cm



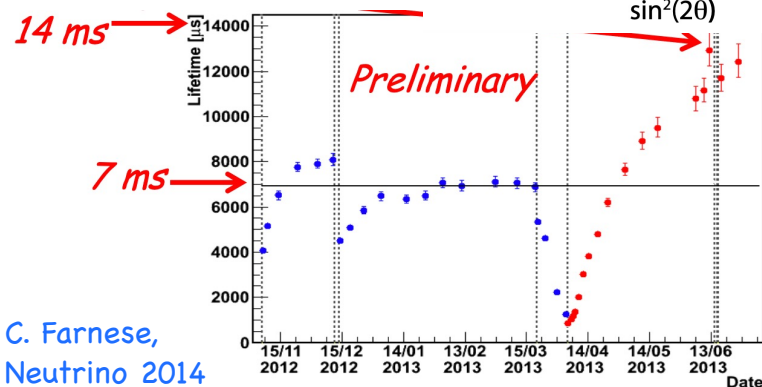
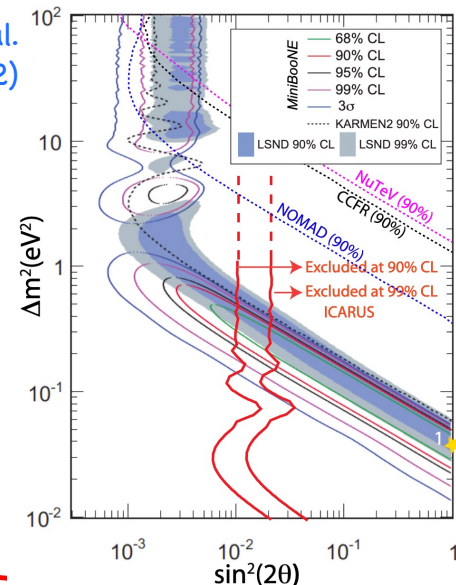
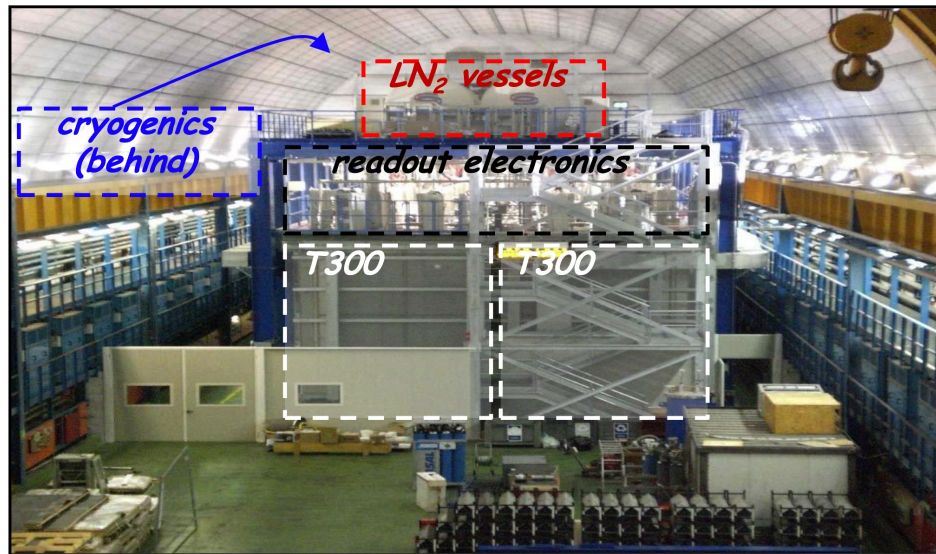
Profile of an electron candidate shower  
(Jan 2020 FNAL PAC meeting, D. Gibin  
and O. Palamara)



# ICARUS at LNGS

(M. Antonello et al.  
arXiv:1209.0122)

- ICARUS T600 run in LNGS spanned several years, with neutrino beam (CNGS from CERN) between 2010-2012
- Tests/demonstrations of many aspects of LAr detector development and analysis techniques
- Analyses e.g. sterile  $\nu$  search with CNGS ( $\sim 10\text{-}30$  GeV)

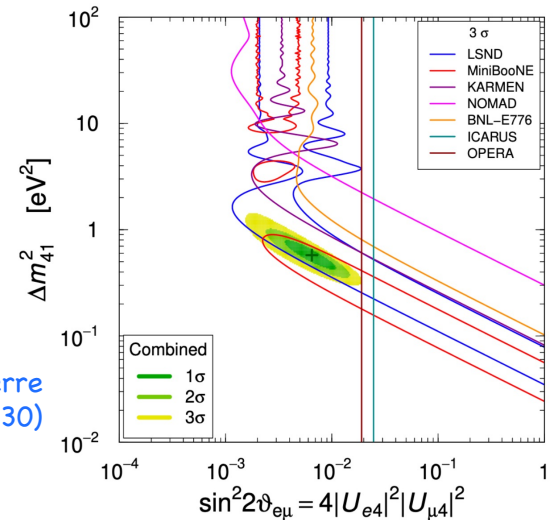


C. Farnese,  
Neutrino 2014

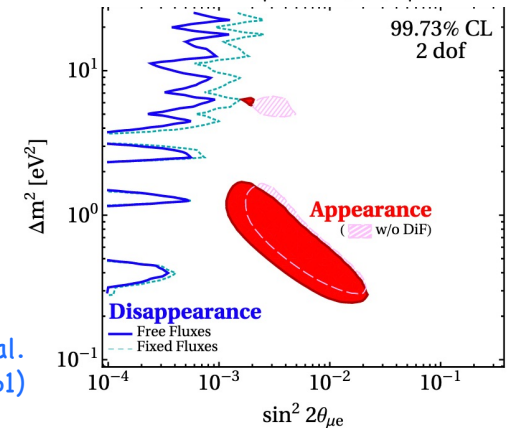
# Sterile neutrinos & SBL anomaly

- LSND and MiniBooNE experiments have found anomalous excess of low energy  $\nu_e$  candidates at short baseline (SBL), which is outside the scope of 3-flavor oscillation explanation.
  - An interpretation of this signature is oscillation of  $\nu_\mu$  to  $\nu_e$  via presence of  $\geq 1$  sterile  $\nu$
- The experimental landscape of other appearance-type experiments have placed limits on allowed sterile osc parameters (including previous iteration of ICARUS) and strong tension w/  $\nu_\mu$  disappearance channel
- More recently, the Neutrino-4 collaboration has interestingly reported a hint of oscillation signature at higher mass splitting ( $7.3 \text{ eV}^2$  and  $\sin^2 2\theta \sim 0.36$ ). See A. P. Serebrov et al. Phys. Rev. D **104**, 032003 (2021)
  - More on this in a bit

(C. Giunti, T. Lasserre  
arXiv:1901.08330)

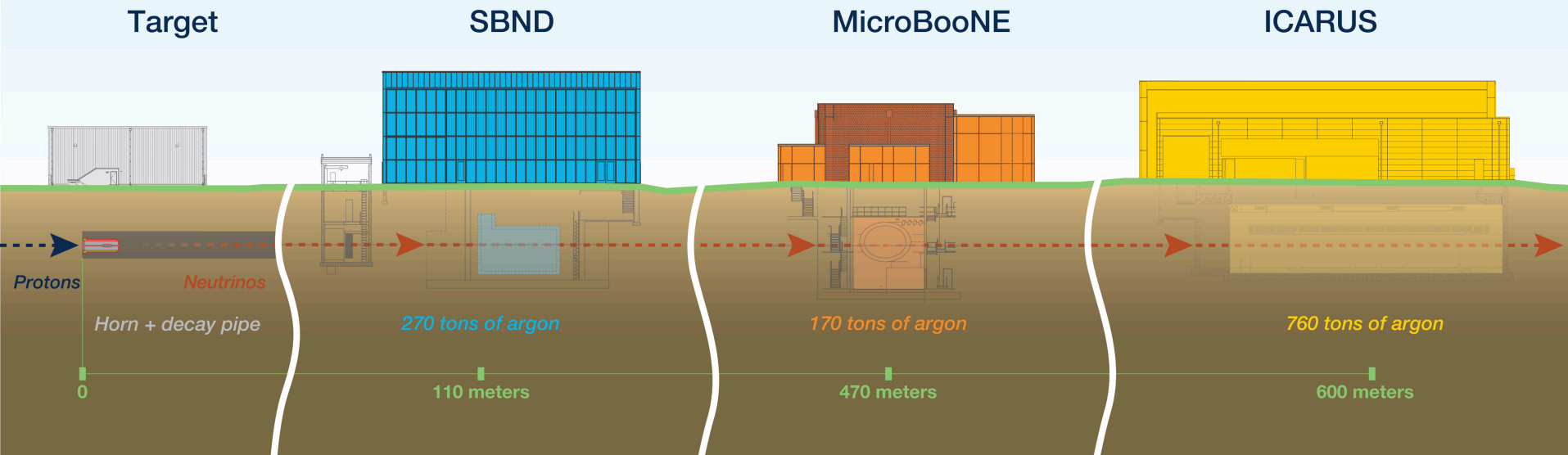


(M. Dentler et al.  
arXiv:1803.10661)



# SBN Program at FNAL

## Short-Baseline Neutrino Program at Fermilab

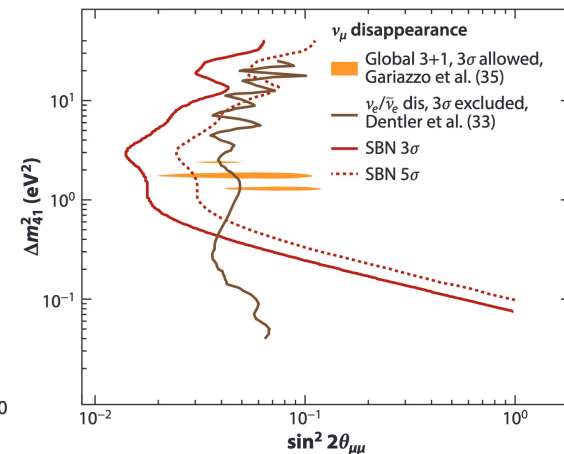
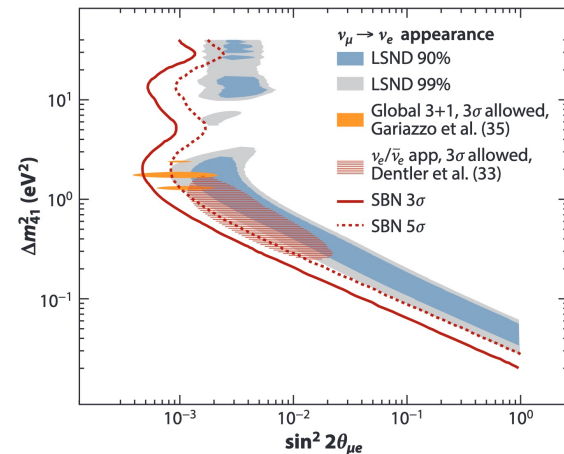
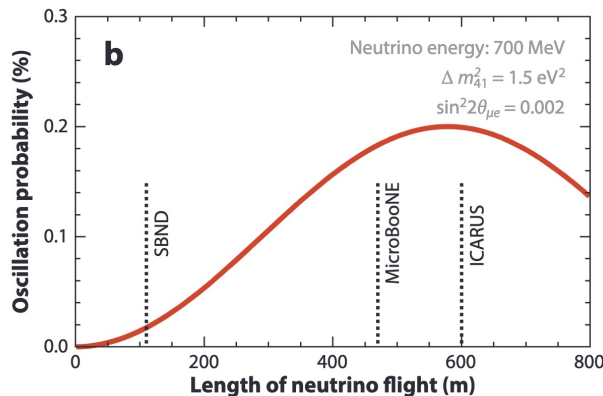
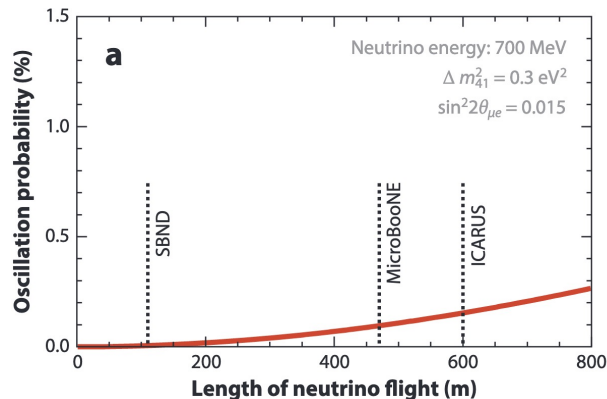


Fermilab VMS, 19-0107-02



# SBN Program at FNAL

- ICARUS largest of 3 LAr TPC experiments using BNB beamline and serves as a far detector in the osc study
- SBND will be the closest & serve as near detector
- Studies  $\nu_e$  appearance and  $\nu_\mu$  disappearance along beam
  - Aim to provide definitive answer on nature of MiniBooNE & LSND excess, and the sterile neutrino question



P. Machado, O. Palamara, and D. Schmitz. Annu. Rev. Nucl. Part. Sci. (2019). doi: 10.1146

# ICARUS at FNAL

- ICARUS T600 moved from Europe to the US and deployed at Fermilab
- Serves as the far detector in the Short Baseline Neutrino (SBN) program searching for definitive answer on this sterile neutrino question
- Able to detect NuMI (120 GeV p on target) at  $\sim 6$  degrees off-axis – input to another oscillation study, valuable cross-section measurements, light dark matter searches, etc.

BNB beam runs along  
length of detector

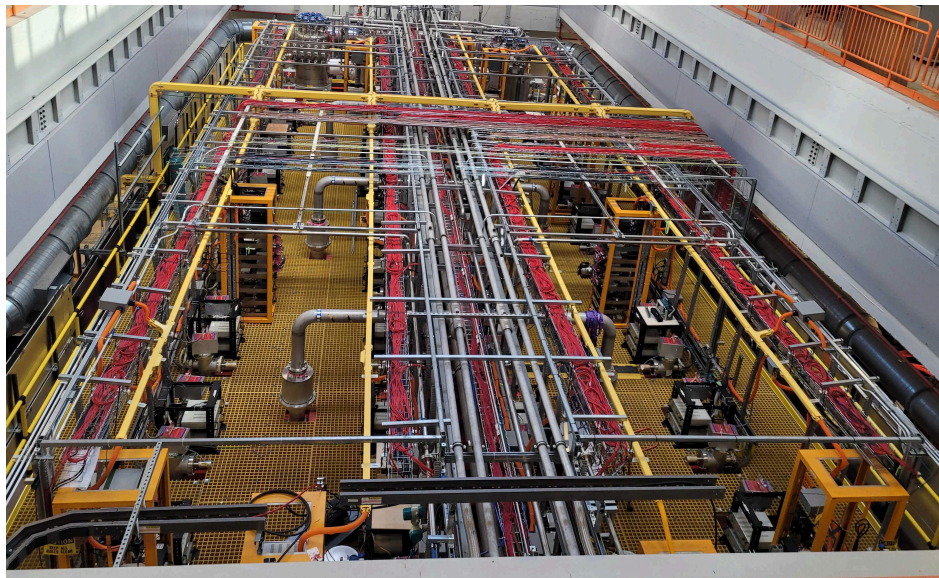
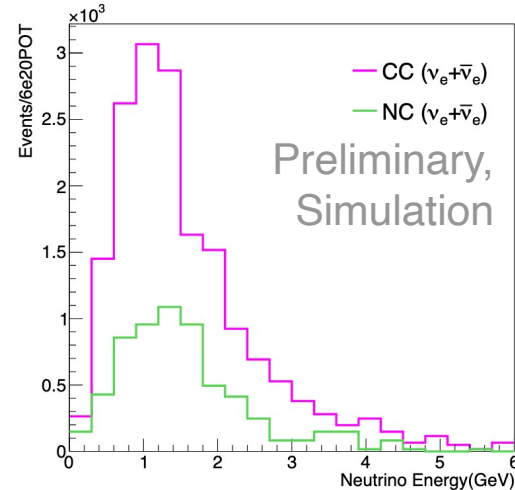
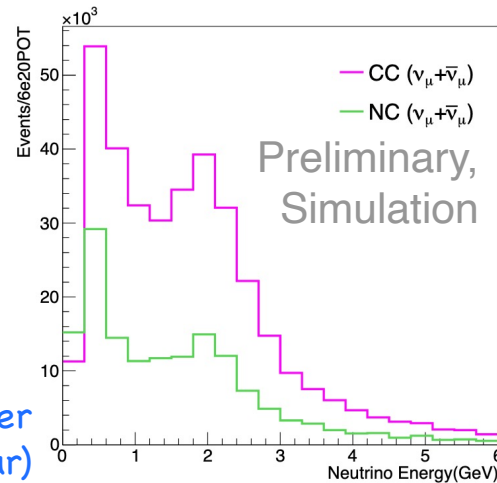


Photo from this spring  
showing electronics, piping,  
cables, etc. of the installed  
ICARUS detector at FNAL

# Cross sections with NuMI

- Understanding neutrino interactions with Ar like cross-sections, nuclear effects, final states are important to oscillation measurements and constraining systematics
  - Effects e.g. both the rates/signal efficiency understanding and energy resolution
- The NuMI beam provides ICARUS a wealth of  $\nu_\mu$  and  $\nu_e$  interactions, from order of a few hundred MeV to multi-GeV (higher energy  $\nu_e$  from K decays & are more abundant in NuMI)
- Provides both an interesting sample in comparison to BNB oscillation study and also to provide input to DUNE (covers 2<sup>nd</sup> oscillation maximum and extends well into 1<sup>st</sup>)
- NuMI off-axis measurements will also include light dark matter searches

Anticipated NuMI neutrino interactions per 6e20 protons on target (roughly 1 year)

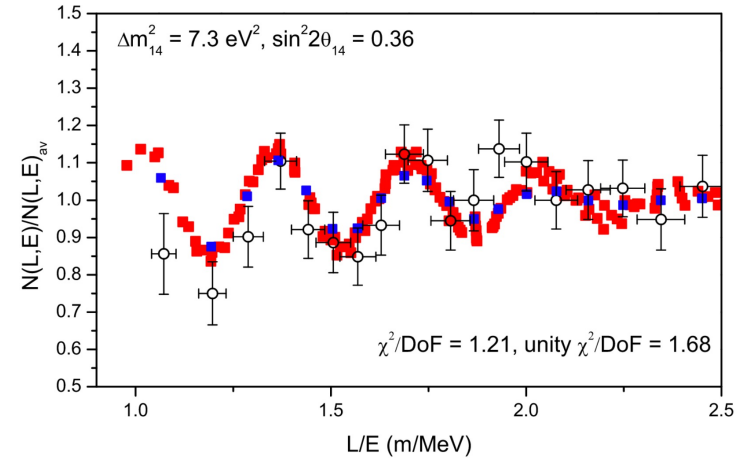




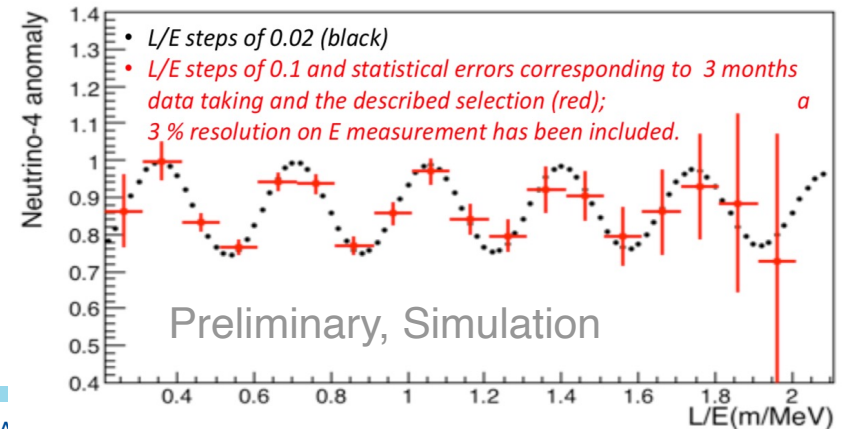
# Neutrino-4 study with ICARUS

- Neutrino-4 collaboration reported finding a reactor neutrino disappearance signal in the L/E range  $\sim 1$ -3
- ICARUS sensitive to similar L/E. Two (separate) samples and beam-off sample can test these findings:
  - $\nu_\mu$  disappearance in BNB beamline: focus on  $\nu_\mu$  CC QE sample w/ contained  $\mu \geq 50$  cm
  - $\nu_e$  disappearance in NuMI beamline: focus on  $\nu_e$  CC QE sample w/ contained electron shower

A. P. Serebrov et al. Phys. Rev. D **104**, 032003 (2021)



Example  $\nu_\mu$  survival probability at  $\Delta m^2 \sim 7.25$ ,  $\sin^2 2\theta \sim 0.26$  in ICARUS and corresponding anticipated measurement assuming 3 months data taking and 3% energy scale resolution.

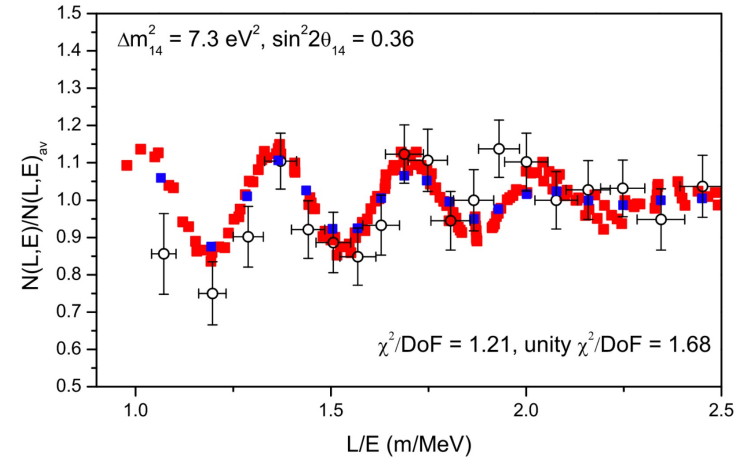


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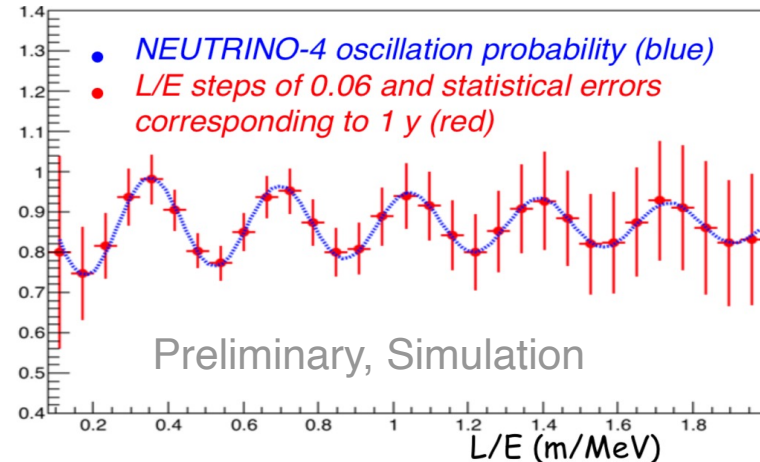
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Example  $\nu_e$  survival probability at  $\Delta m^2 \sim 7.25$ ,  $\sin^2 2\theta \sim 0.26$  in ICARUS and corresponding anticipated measurement assuming 1 year data taking (statistical uncertainties only).

A. P. Serebrov et al. Phys. Rev. D **104**, 032003 (2021)

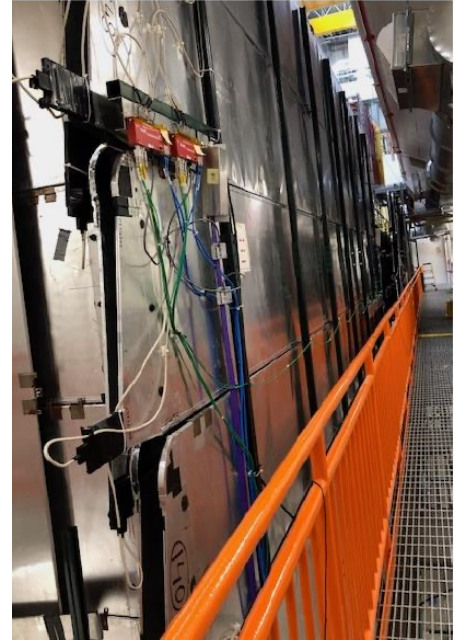


Neutrino-4 anomaly



# ICARUS at FNAL

- Because on surface and size, significant cosmic bkg expected ( $\sim 3$  (1-2) cosmic trigger events per 1 from  $\nu$  in LAr in BNB (NuMI))
- Multi-pronged mitigation strategy:
  - Some cosmic activity can be “clear” (e.g. appear too far outside expected detector bounds given “T=0” beam time)
  - Matching of charge signal from the TPC with the light signals in the PMTs can help determine if depositions are in-time
  - In addition to TPC/PMT systems noted earlier, cosmic ray tagger (CRT) is being installed to help identify cosmic tracks entering detector, especially helpful in-time activity
    - Possibility of CRT-PMT match to distinguish entering/exiting tracks under study
  - An overburden is to be placed above the detector upon the completion of installation

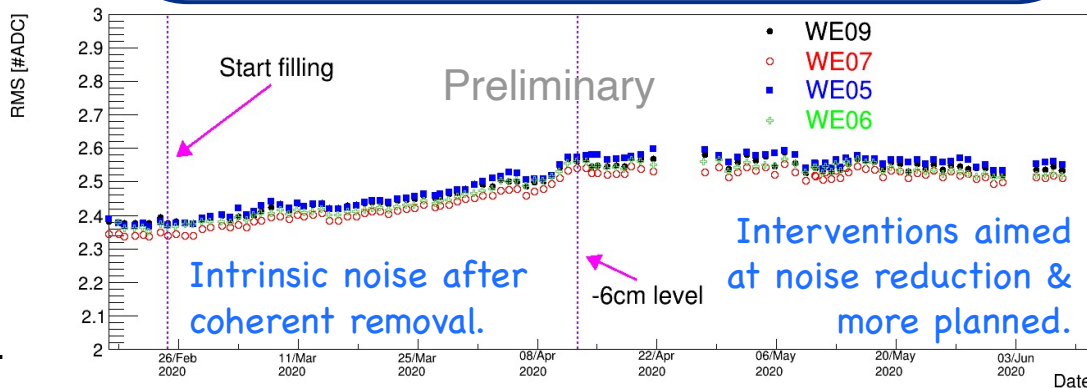
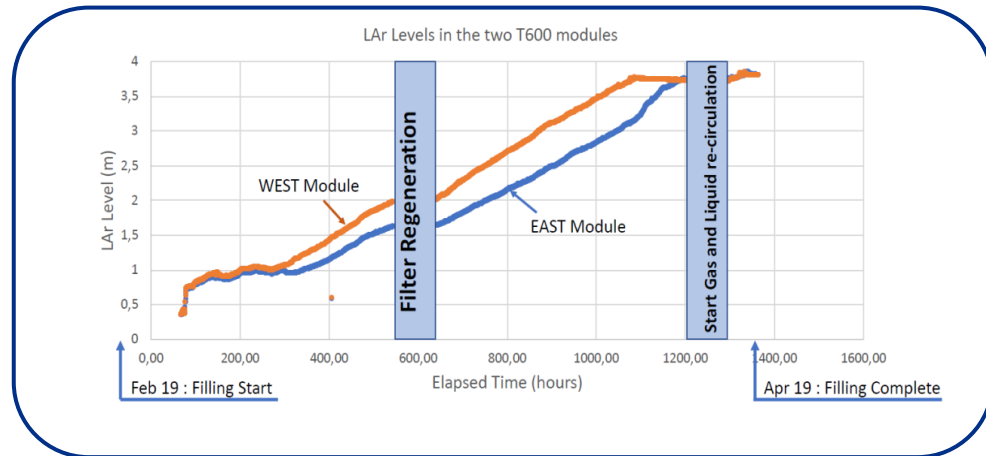


ICARUS Side CRT



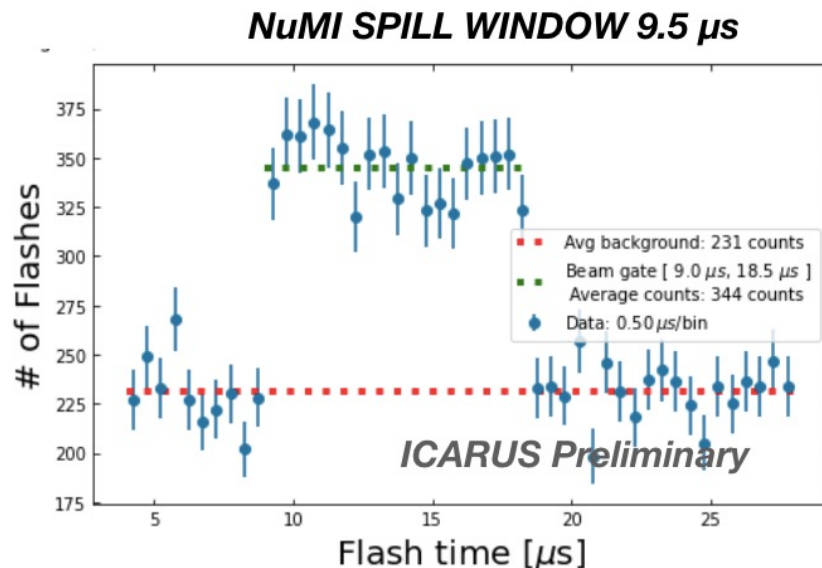
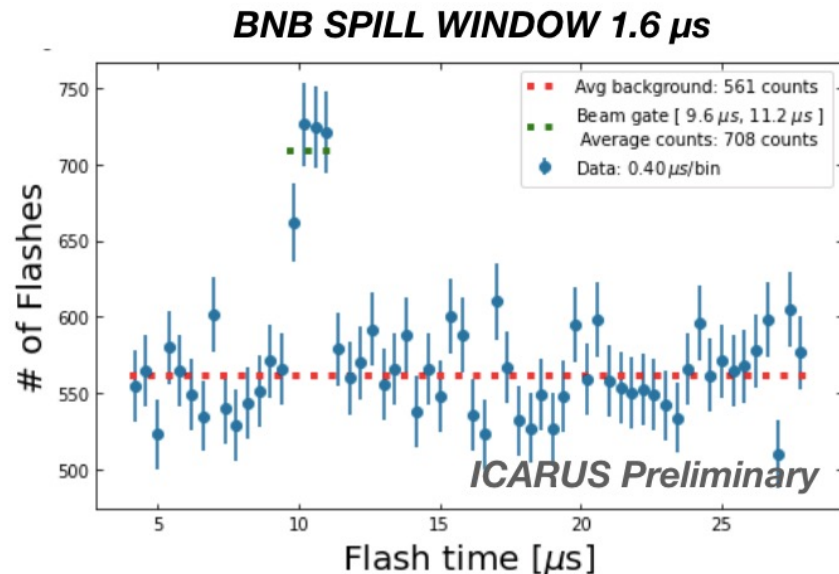
# ICARUS at FNAL: Activation & Commissioning

- After installation, ICARUS began cooldown and filling Feb 2020
- Taking detector monitoring shifts since then: briefly in person, switching to remote w/ pandemic
- By Fall 2020 detector was activated to the full electric drift field of -75kV (500 V/cm) was reached. PMTs were activated and gain calibration was underway, etc.
  - A few walls of CRT were installed prior to pandemic, other parts of side CRT installed during. Top CRT remains – is upcoming work.



# ICARUS at FNAL: Activation & Commissioning

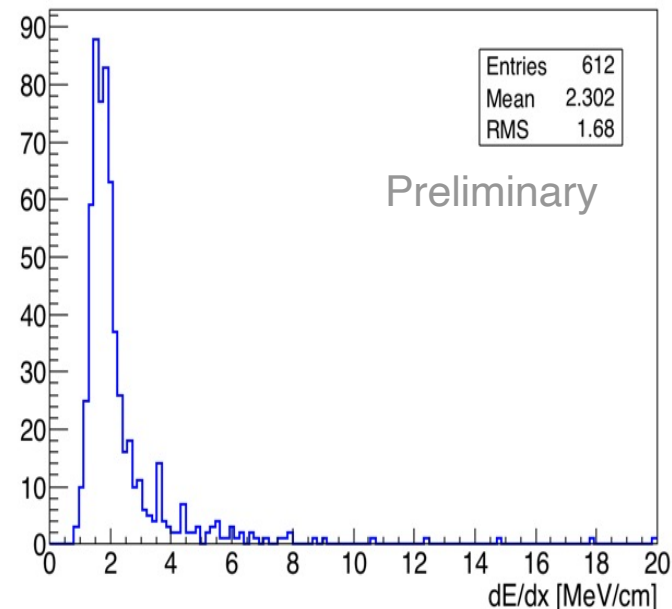
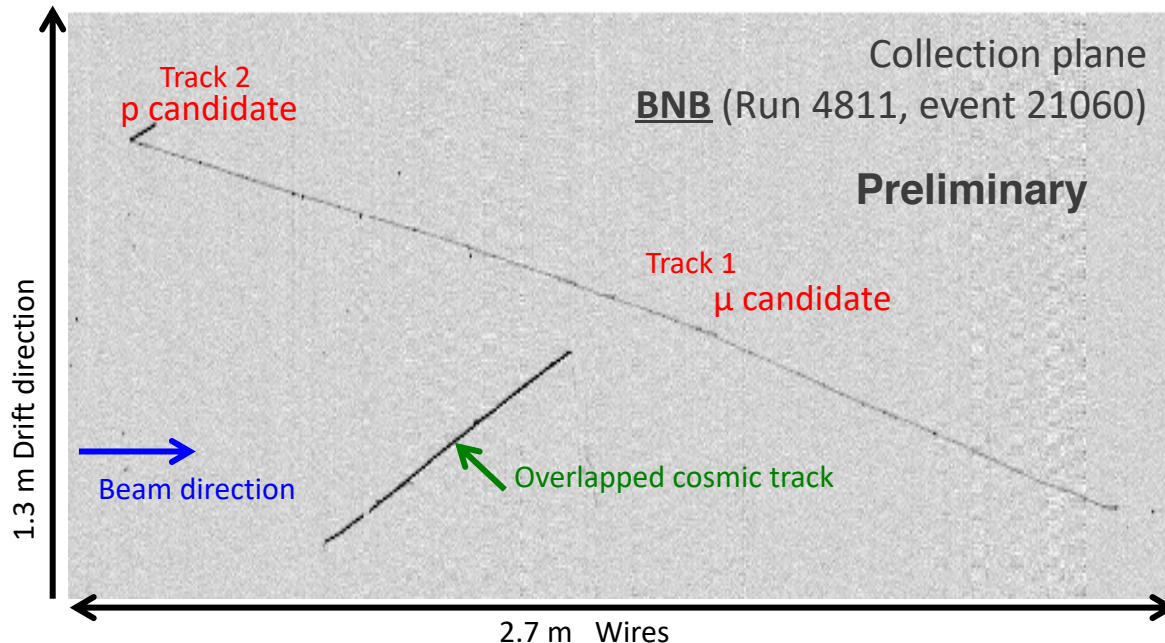
- Trigger commissioning included confirmation of a peak of excess corresponding to beam



Excess plateaus ~right width correspond to neutrinos interacting in detector from the beams.

# ICARUS at FNAL: Activation & Commissioning

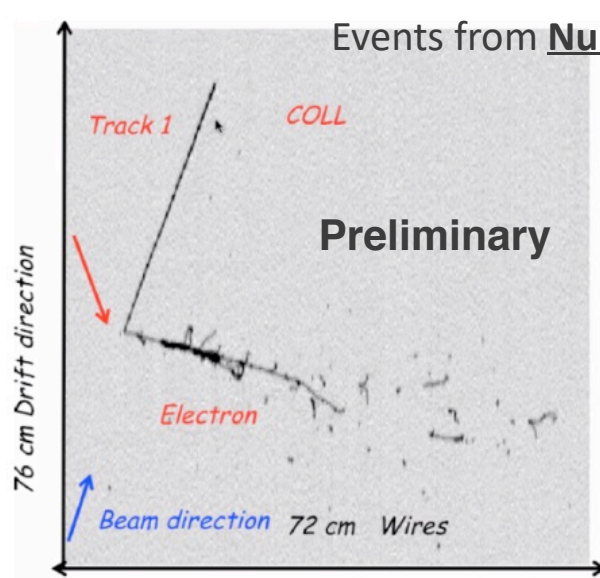
- Trigger commissioning included confirmation of a peak of excess corresponding to beam
- Visual scanning effort to find neutrino candidates in early collected data



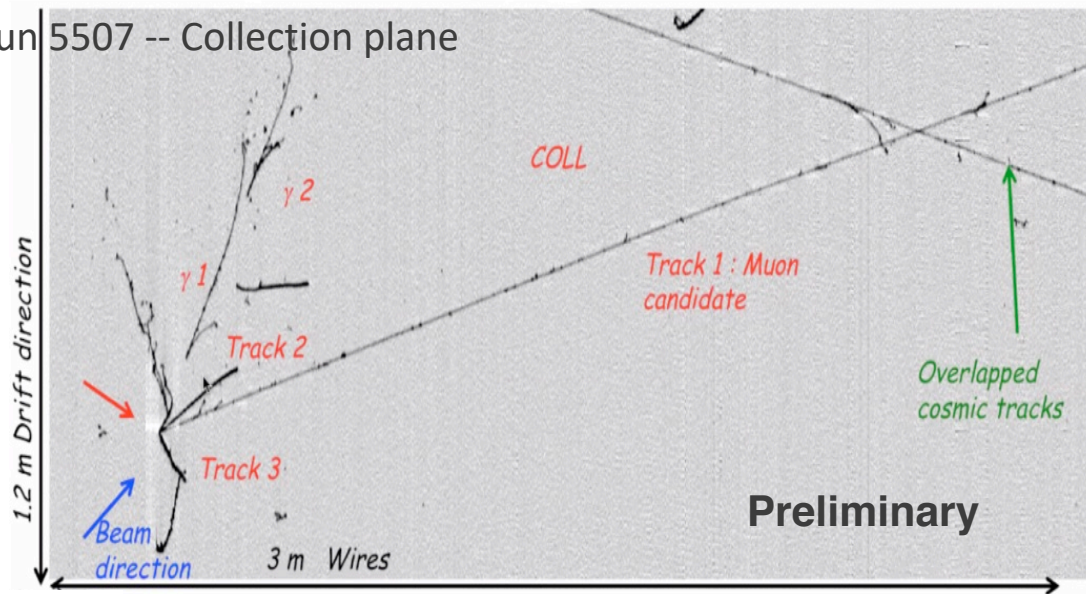
One of the first candidate  $\nu_\mu$  interactions found in BNB. Electron lifetime at this point was  $\sim 1.2$  ms

# ICARUS at FNAL: Activation & Commissioning

- Trigger commissioning included confirmation of a peak of excess corresponding to beam
- Visual scanning effort to find neutrino candidates in early collected data



Candidate  $\nu_e$  interaction from NuMI. Electron lifetime  $\sim 3.2$  ms

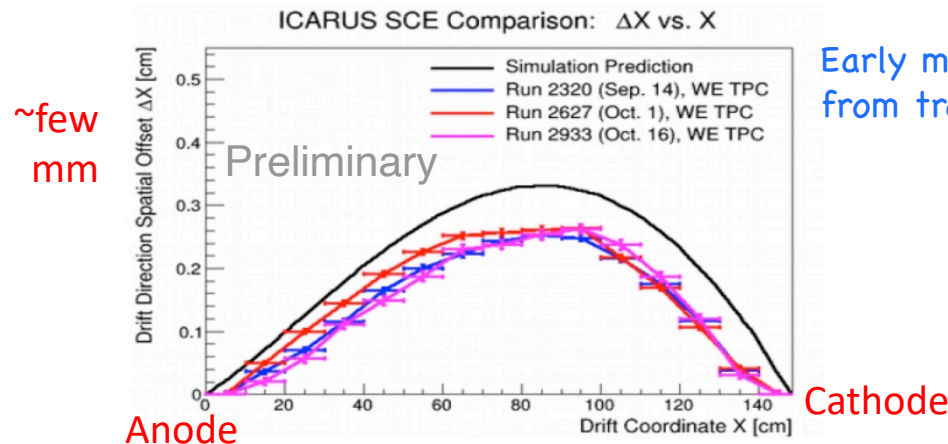


Candidate  $\nu_\mu$  interaction from NuMI.



# ICARUS at FNAL: Activation & Commissioning

- Trigger commissioning included confirmation of a peak of excess corresponding to beam
- Visual scanning effort to find neutrino candidates in early collected data
  - Input to status of trigger and readout. e.g. fix any TPC channel mis-mapping
  - Provides set of commissioning events that can be studied for their properties
  - Currently using these events to better understand data reconstruction
- Calibration campaign with early data to understand detector effects, calorimetry, etc.



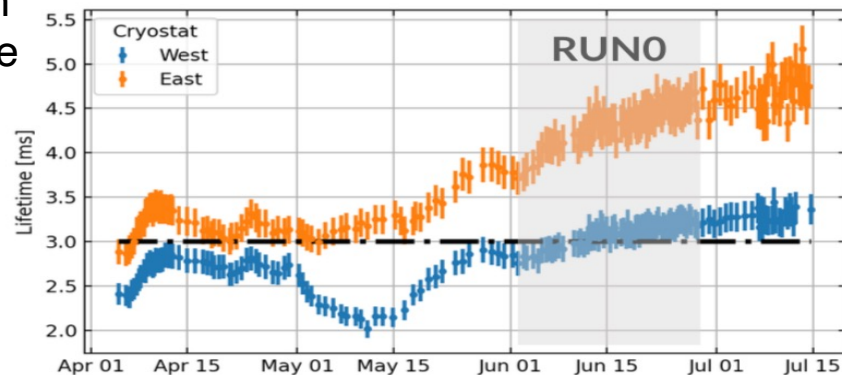
Early measurement of space charge effect  
from tracks extending from anode to cathode

Also drift velocity, calorimetry  
measurements (i.e.  $e/\text{ADC}$ ), look  
at recombination, etc.

# ICARUS at FNAL: Activation & Commissioning

- Trigger commissioning included confirmation of a peak of excess corresponding to beam
- Visual scanning effort to find neutrino candidates in early collected data
  - Input to status of trigger and readout. e.g. fix any TPC channel mis-mapping
  - Provides set of commissioning events that can be studied for their properties
  - Currently using these events to better understand data reconstruction
- Calibration campaign with early data to understand detector effects, calorimetry, etc.
- Just before summer, a dry run (Run0) was taken as an emulation of physics data mode to prepare for when the beam begins running in fall again

Electron lifetime showed improvement over the course of Run0, above the desired minimum of 3ms. Further improvements to the cryogenic system are planned.



# ICARUS at FNAL: Collaborative Effort

P. Abratenko<sup>19</sup>, A. Aduszkiewicz<sup>20</sup>, J. Asaadi<sup>23</sup>, M. Babicz<sup>2</sup>, W. F. Badgett<sup>5</sup>, L.F. Bagby<sup>5</sup>, B. Behera<sup>4</sup>, V. Bellini<sup>7</sup>, O. Beltramello<sup>2</sup>, R. Benocci<sup>13</sup>, S. Berkmann<sup>52</sup>, S. Bertolucci<sup>6</sup>, M. Betancourt<sup>5</sup>, S. Biagi<sup>11</sup>, K. Biery<sup>5</sup>, M. Bonesini<sup>13</sup>, T. Boone<sup>4</sup>, B. Bottino<sup>8</sup>, A. Braggiotti<sup>15</sup>, J. Bremer<sup>2</sup>, S. Brice<sup>5</sup>, J. Brown<sup>5</sup>, H. Budd<sup>22</sup>, H. Carranza<sup>23</sup>, S. Centro<sup>15</sup>, G. Cerati<sup>5</sup>, M. Chalifour<sup>2</sup>, A. Chatterjee<sup>21</sup>, D. Cherdack<sup>21</sup>, S. Cherubini<sup>11</sup>, T. Coan<sup>18</sup>, A. Cocco<sup>14</sup>, M. Convery<sup>17</sup>, S. Copello<sup>8</sup>, A. De Roeck<sup>2</sup>, L. Di Noto<sup>8</sup>, C. Distefano<sup>11</sup>, M. Diwan<sup>1</sup>, S. Dytman<sup>22</sup>, S. Dolan<sup>2</sup>, L. Domine<sup>17</sup>, R. Doubnik<sup>5</sup>, F. Drielsma<sup>17</sup>, V. Fabbri<sup>6</sup>, C. Fabre<sup>2</sup>, A. Falcone<sup>13</sup>, C. Farnese<sup>15</sup>, A. Fava<sup>5</sup>, A. Ferrari<sup>12</sup>, F. Ferraro<sup>8</sup>, N. Gallice<sup>12</sup>, M. Geynisman<sup>5</sup>, D. Gibin<sup>15</sup>, W. Gu<sup>1</sup>, M. Guerzoni<sup>6</sup>, A. Guglielmi<sup>15</sup>, S. Hahn<sup>5</sup>, C. Hilgenberg<sup>4</sup>, B. Howard<sup>5</sup>, R. Howell<sup>23</sup>, C. James<sup>5</sup>, W. Jang<sup>23</sup>, D. H. Kao<sup>17</sup>, W. Ketchum<sup>5</sup>, U. Kose<sup>2</sup>, J. Larkin<sup>1</sup>, G. Laurenti<sup>6</sup>, Q. Lin<sup>17</sup>, G. Lukhanin<sup>5</sup>, N. Mauri<sup>6</sup>, A. Mazzacane<sup>5</sup>, K.S. McFarland<sup>23</sup>, D. P. Mendez<sup>1</sup>, G. Meng<sup>15</sup>, A. Menegolli<sup>16</sup>, O. G. Miranda<sup>3</sup>, D. Mladenov<sup>2</sup>, N. Moggi<sup>6</sup>, A. Montanari<sup>6</sup>, C. Montanari<sup>5,b</sup>, M. Mooney<sup>4</sup>, G. Moreno Granados<sup>3</sup>, J. Mueller<sup>4</sup>, M. Nessi<sup>2</sup>, T. Nichols<sup>5</sup>, S. Palestini<sup>2</sup>, M. Pallavicini<sup>8</sup>, V. Paolone<sup>21</sup>, R. Pappaleo<sup>11</sup>, L. Pasqualini<sup>6</sup>, L. Patrizii<sup>6</sup>, G. Petrillo<sup>17</sup>, C. Petta<sup>7</sup>, V. Pia<sup>6</sup>, F. Pietropaolo<sup>2,a</sup>, F. Poppi<sup>6</sup>, M. Pozzato<sup>6</sup>, A. Prosser<sup>5</sup>, G. Putnam<sup>20</sup>, X. Qian<sup>1</sup>, A. Rappoldi<sup>16</sup>, R. Rechenmacher<sup>5</sup>, G. Riccobene<sup>11</sup>, F. Resnati<sup>2</sup>, A. Rigamonti<sup>2</sup>, G.L. Raselli<sup>16</sup>, M. Rossella<sup>16</sup>, C. Rubbia<sup>9</sup>, P. Sala<sup>12</sup>, P. Sapienza<sup>11</sup>, G. Savage<sup>5</sup>, A. Scaramelli<sup>16</sup>, A. Scarpelli<sup>1</sup>, D. Schmitz<sup>20</sup>, A. Schukraft<sup>5</sup>, F. Sergiampietri<sup>2</sup>, G. Sirri<sup>6</sup>, A. Soha<sup>5</sup>, L. Stanco<sup>15</sup>, J. Stewart<sup>1</sup>, C. Suter<sup>7</sup>, H. Tanaka<sup>17</sup>, M. Tenti<sup>6</sup>, K. Terao<sup>17</sup>, F. Terranova<sup>13</sup>, D. Torretta<sup>5</sup>, M. Torti<sup>13</sup>, F. Tortorici<sup>7</sup>, Y.T. Tsai<sup>17</sup>, S. Tufnali<sup>2</sup>, T. Usher<sup>17</sup>, M. Vicenzi<sup>8</sup>, B. Viren<sup>1</sup>, D. Warner<sup>4</sup>, Z. Williams<sup>24</sup>, P. Wilson<sup>5</sup>, R.J. Wilson<sup>4</sup>, T. Wongjirad<sup>19</sup>, A. Wood<sup>21</sup>, E. Worcester<sup>1</sup>, M. Worcester<sup>1</sup>, M. Wospakrik<sup>5</sup>, H. Yu<sup>1</sup>, J. Yu<sup>24</sup>, F. Varanini<sup>15</sup>, S. Ventura<sup>15</sup>, C. Vignoli<sup>10</sup>, A. Zani<sup>12</sup>, C. Zhang<sup>1</sup>, J. Zennaro<sup>5</sup>, J. Zettlemoyer<sup>5</sup>, S. Zucchelli<sup>6</sup>, M. Zuckerbrot<sup>5</sup>

a On Leave of Absence from INFN Padova

b On Leave of Absence from INFN Pavia

1. Brookhaven National Lab., USA
2. CERN, Switzerland
3. CINVESTAV, Mexico
4. Colorado State University, USA
5. Fermi National Accelerator Lab., USA
6. INFN Bologna and University, Italy
7. INFN Catania and University, Italy
8. INFN Genova and University, Italy
9. INFN GSSI, L'Aquila, Italy
10. INFN LNGS, Assergi (AQ), Italy
11. INFN LNS, Catania, Italy
12. INFN Milano, Milano, Italy
13. INFN Milano Bic. and University, Italy
14. INFN Napoli, Napoli, Italy
15. INFN Padova and University, Italy
16. INFN Pavia and University, Italy
17. SLAC National Accelerator Lab., USA
18. Southern Methodist University, USA
19. Tufts University, USA
20. University of Chicago
21. University of Houston, USA
22. University of Pittsburgh, USA
23. University of Rochester, USA
24. University of Texas (Arlington), USA

11 INFN groups, 11 USA institutions,  
1 Mexican institution, CERN

Spokesperson: C. Rubbia, GSSI

2

# Summary

- ICARUS is a liquid argon TPC originally operated at LNGS in Italy
- After a successful refurbishment and upgrade, it was moved to Fermilab in the USA where it will see neutrinos from both the BNB and NuMI neutrino beams at Fermilab
- Question of sterile neutrinos remains open, both related to the previous measurements of LSND and MiniBooNE and the more recent claim of oscillation signature from Neutrino-4
  - Short Baseline Neutrino Program at FNAL will use the BNB beam to probe the LSND and MiniBooNE excess. ICARUS serves as the Far Detector of this study.
  - ICARUS is well situated to probe the Neutrino-4 parameters using NuMI and BNB
- ICARUS also able to measure cross-sections, search for light dark matter, etc.
- ICARUS was filled and the TPC activated in 2020 and has been continuing toward full installation/commissioning of all systems since
- A set of commissioning neutrino candidates have been visually identified and work on the various pieces leading towards analysis are ongoing
- No beam at moment (summer shutdown) but excited to collect neutrinos again ~October