

Status and Prospects of the SNO+ Experiment

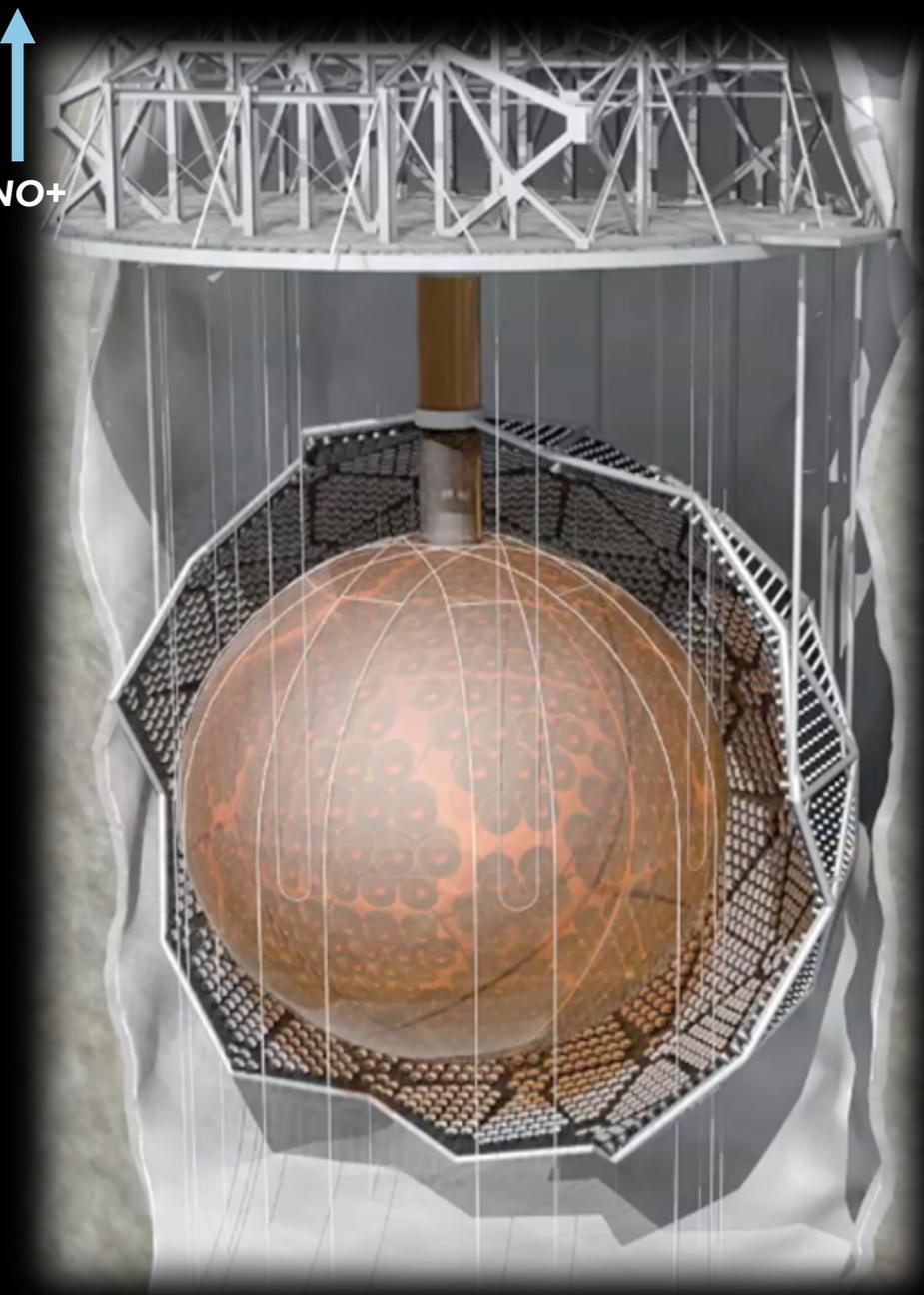
Ana Sofia Inácio, on behalf of the SNO+ Collaboration

University of Lisbon & Laboratório de Instrumentação e Física Experimental de Partículas (LIP), Portugal

SNO+

Multi-purpose neutrino detector located at SNOLAB in Sudbury, Ontario, Canada.

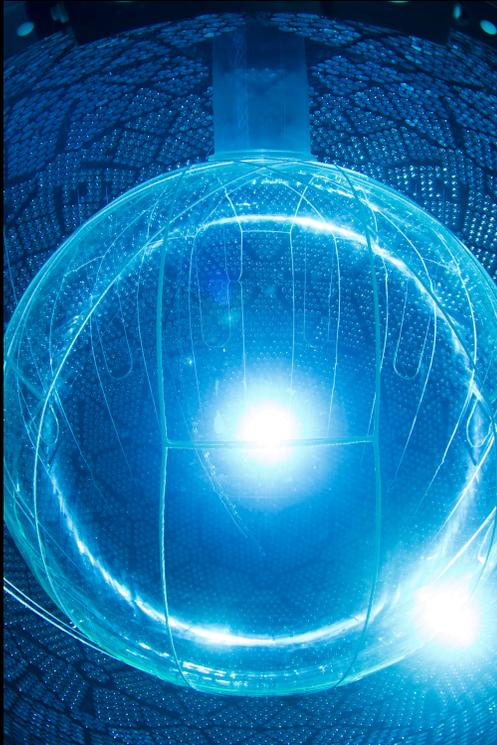
2070 m of rock
overburden,
~70 $\mu\text{l/day}$ in SNO+



The SNO+ Experiment,
[IINST 16 P08059 \(2021\)](#)

SNO+

Multi-purpose neutrino detector located at SNOLAB in Sudbury, Ontario, Canada.

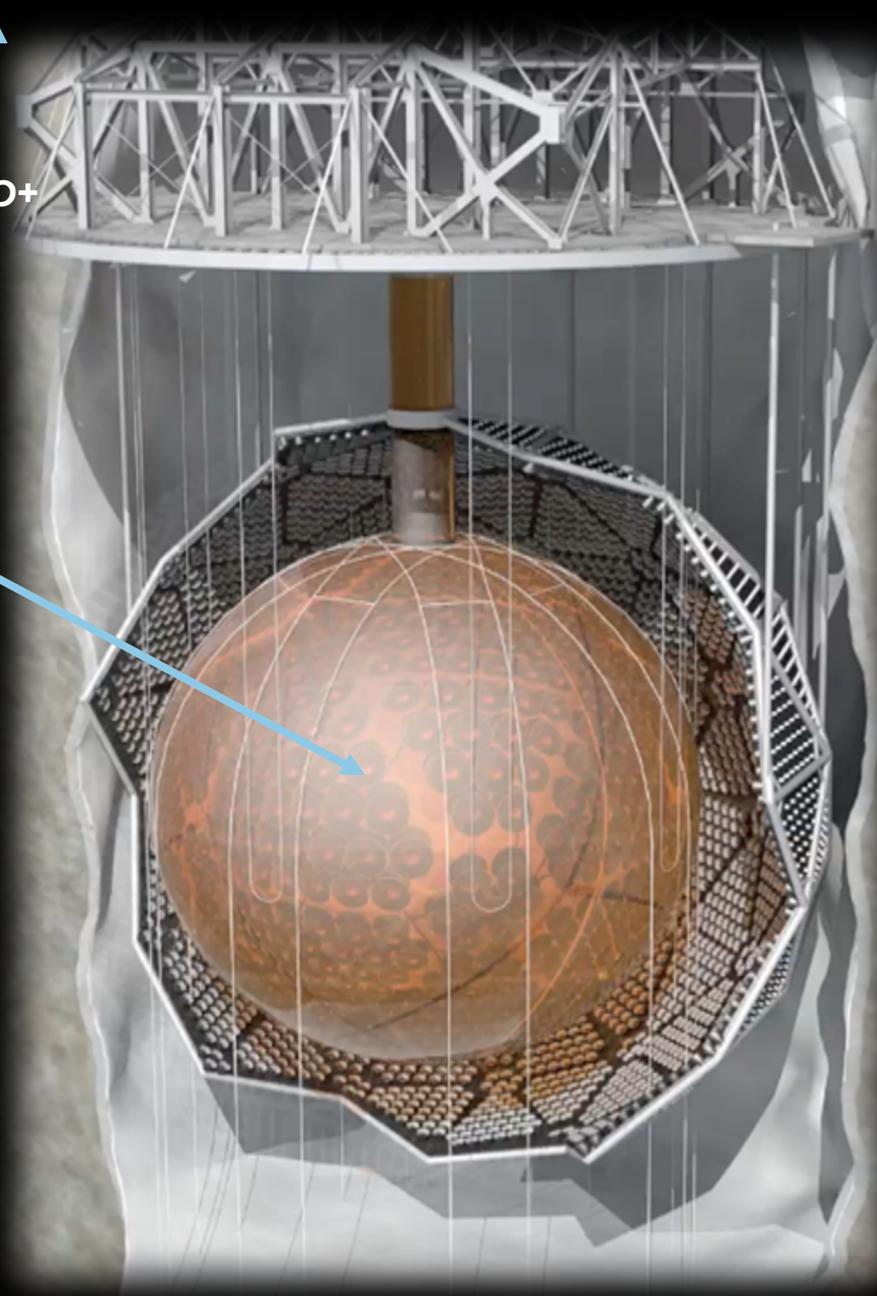


Acrylic vessel (AV)
UV-transparent
6 m radius
5.5 cm thickness

Holds the target medium:

- I. 905 tonnes of UPW
- II. 780 tonnes of LAB+PPO
- III. LAB+PPO + Tellurium cocktail

2070 m of rock
overburden,
~70 μ /day in SNO+



The SNO+ Experiment,
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Multi-purpose neutrino detector located at SNOLAB in Sudbury, Ontario, Canada.

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Acrylic vessel (AV)

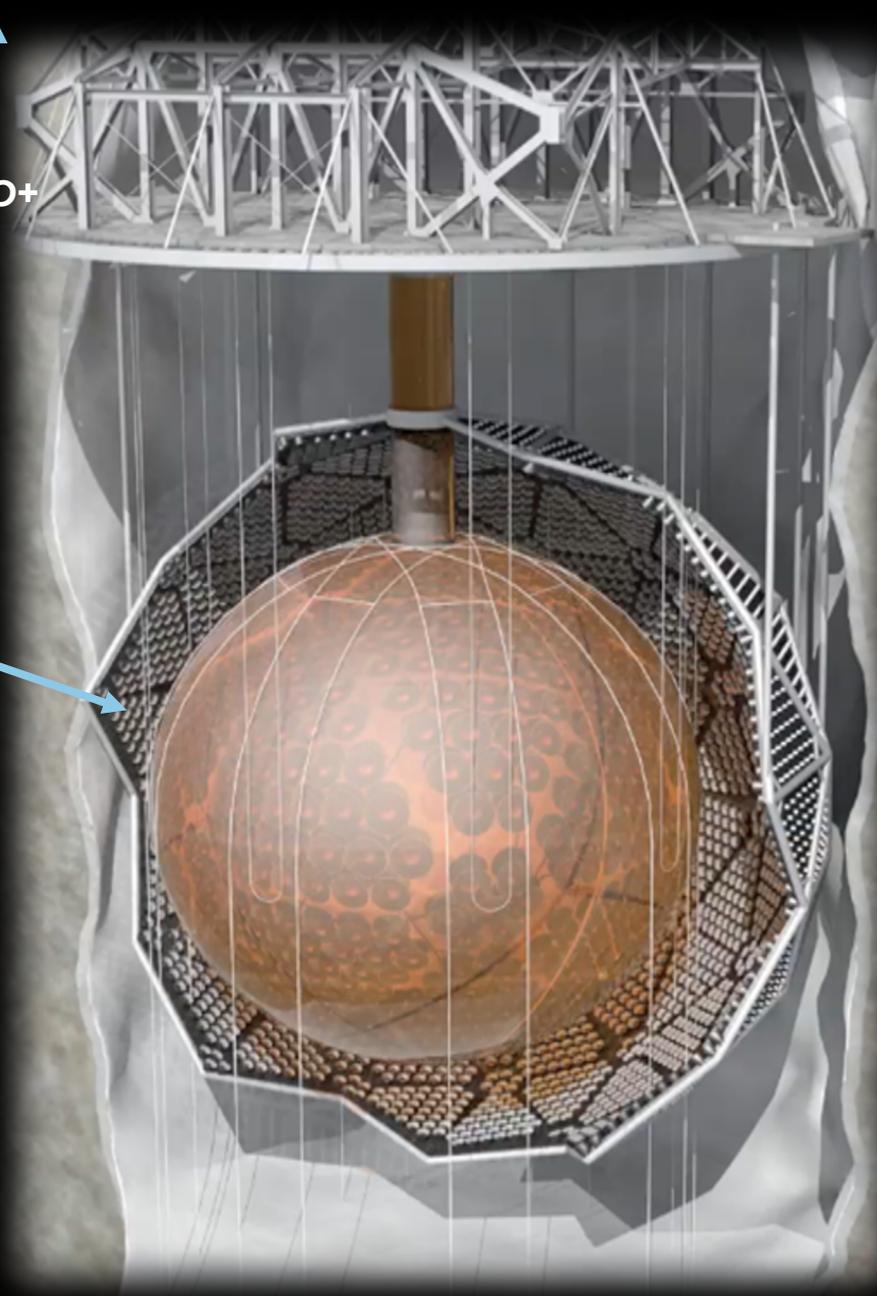
~9400 photomultiplier tubes (PMTs)

54% effective photocoverage

~90 outward looking PMTs for tagging cosmic rays



8" Hamamatsu R1408 PMTs
+
27-cm diameter concentrator



The SNO+ Experiment,
[IINST 16 P08059 \(2021\)](#)

SNO+

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Acrylic vessel (AV)

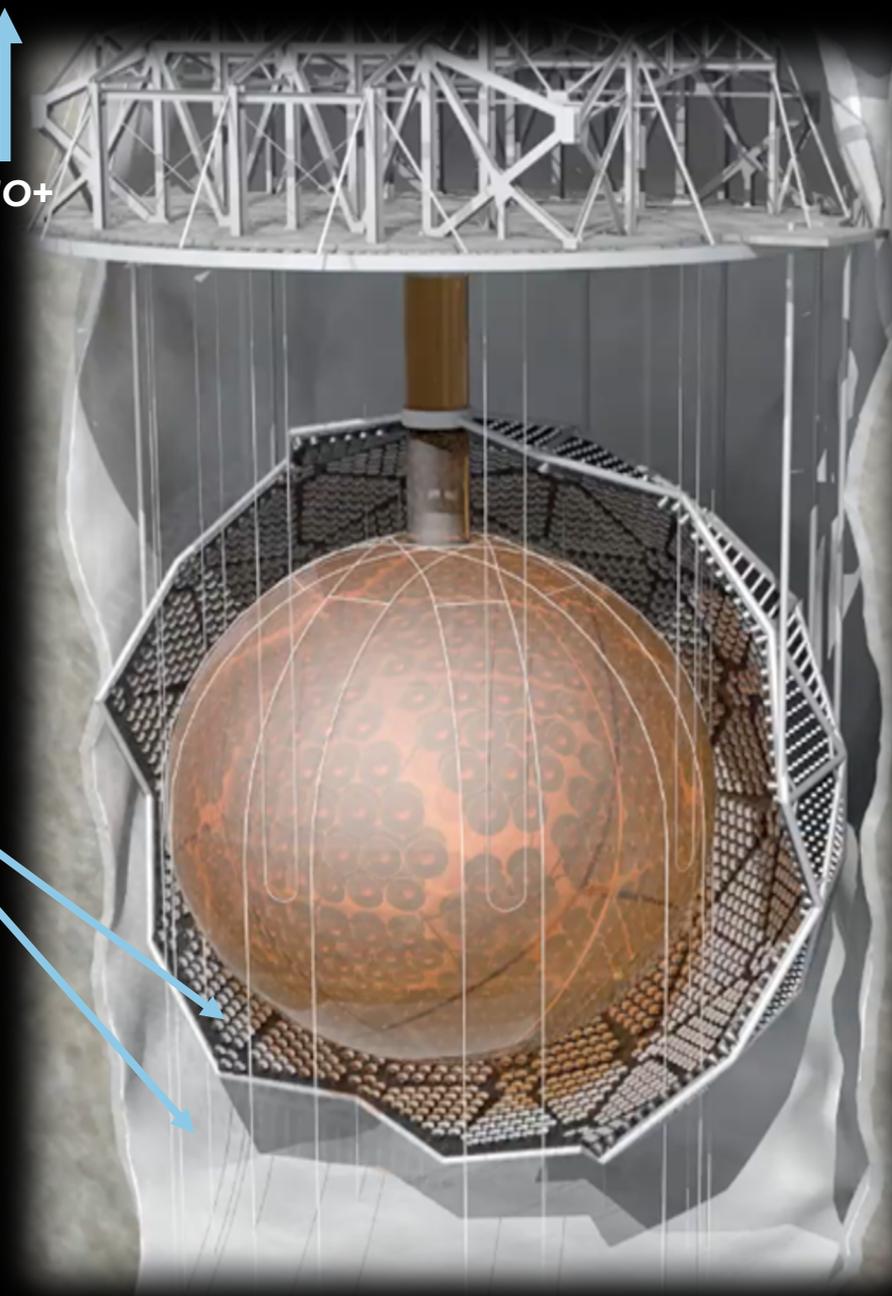
~9400 photomultiplier tubes (PMTs)

Ultra-pure Water shielding

1.7 kt between AV and PMTs,
to reduce backgrounds from PMT materials

5.3 kt between PMTs and cavity,
to reduce background from rock wall

2070 m of rock
overburden,
~70 μ /day in SNO+



SNO+

Multi-purpose neutrino detector located at SNOLAB in Sudbury, Ontario, Canada.

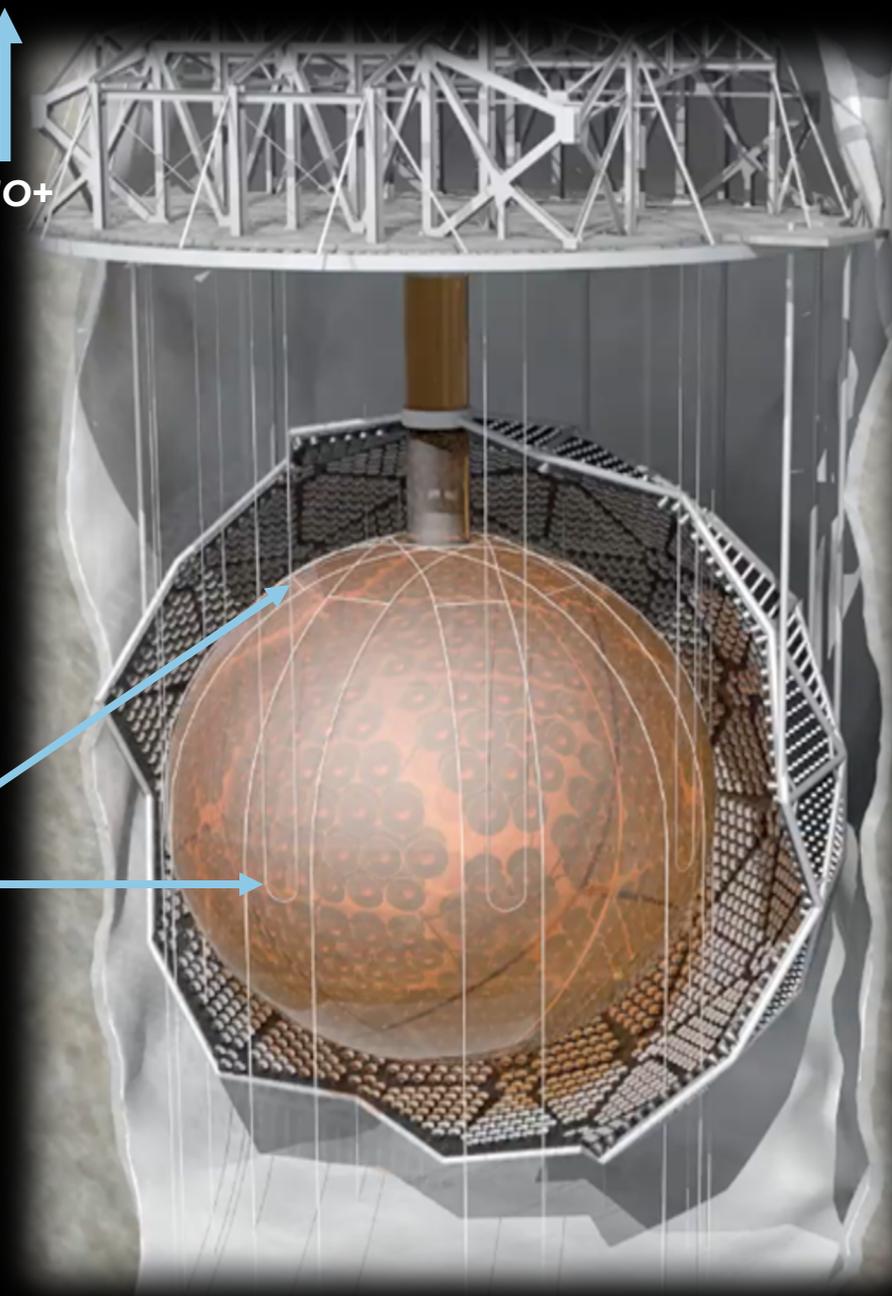
Acrylic vessel (AV)

~9400 photomultiplier tubes (PMTs)

Ultra-pure Water shielding

Low radioactivity hold-up and hold-down rope systems

2070 m of rock
overburden,
~70 μ /day in SNO+



The SNO+ Experiment,
[IINST 16 P08059 \(2021\)](#)

SNO+

Multi-purpose neutrino detector located at SNOLAB in Sudbury, Ontario, Canada.

Acrylic vessel (AV)

~9400 photomultiplier tubes (PMTs)

Ultra-pure Water shielding

Low radioactivity hold-up and hold-down rope systems

**+ Upgraded DAQ/electronics system
for the low threshold and high rate,**

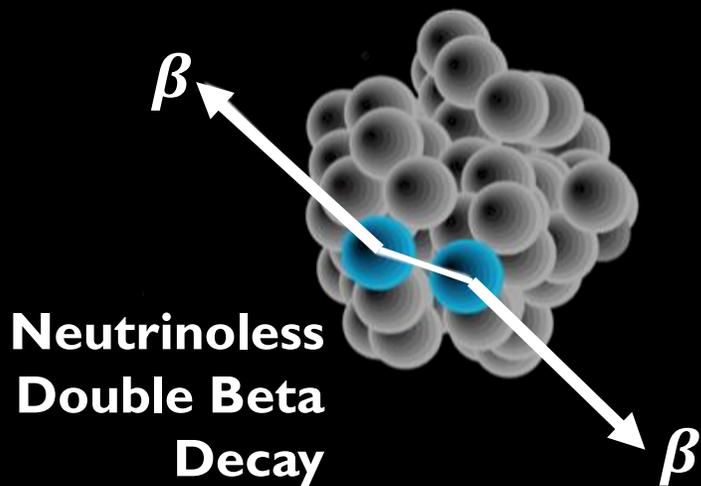
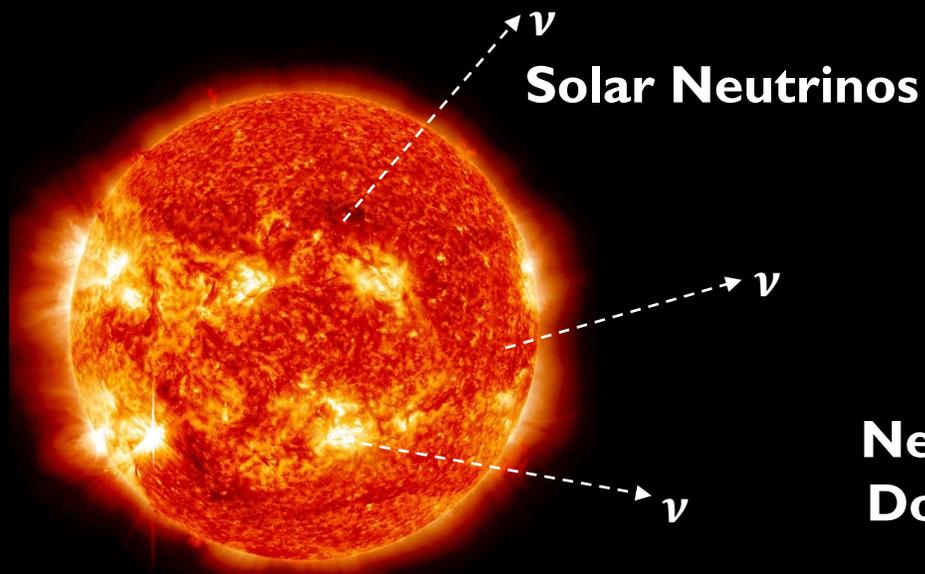
+ New calibration hardware

2070 m of rock
overburden,
~70 μlday in SNO+

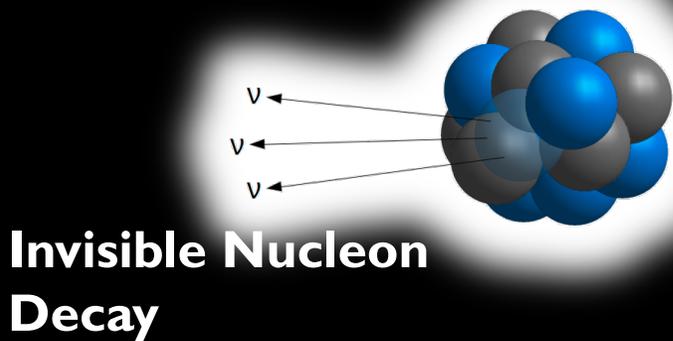


The SNO+ Experiment,
[IINST 16 P08059 \(2021\)](#)

Physics Program



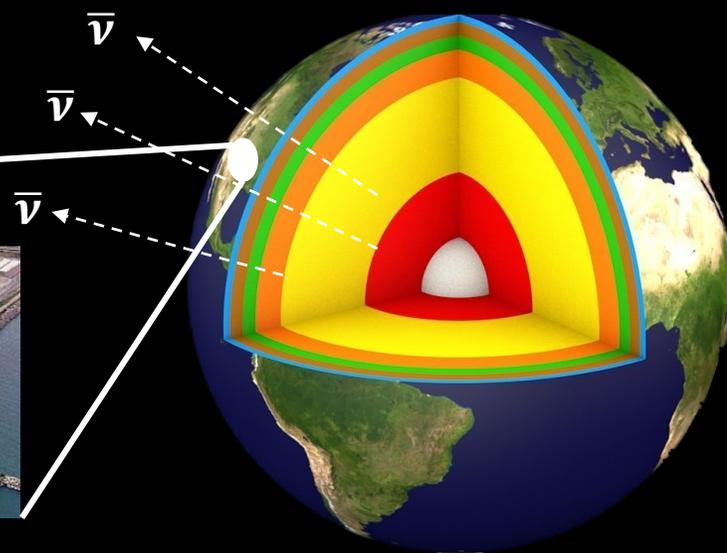
Supernova Neutrinos



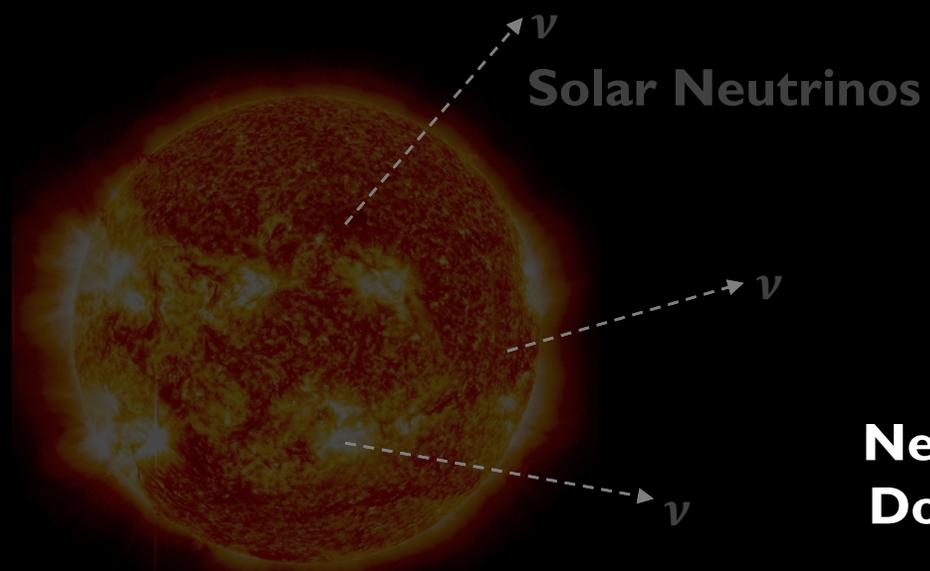
Reactor Anti-Neutrinos



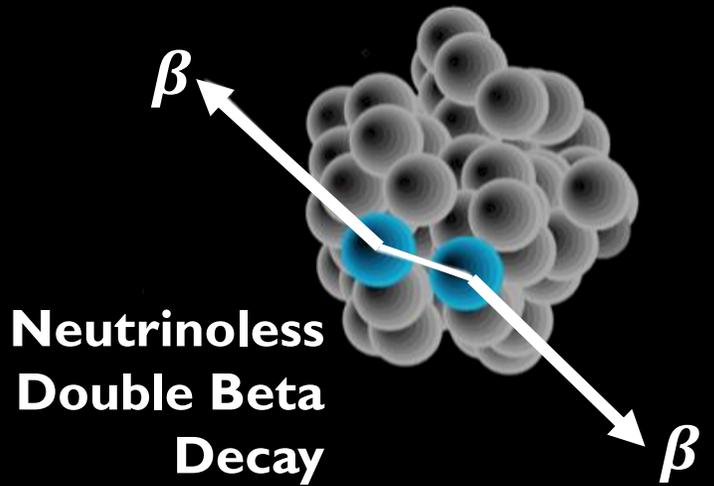
Geo-neutrinos



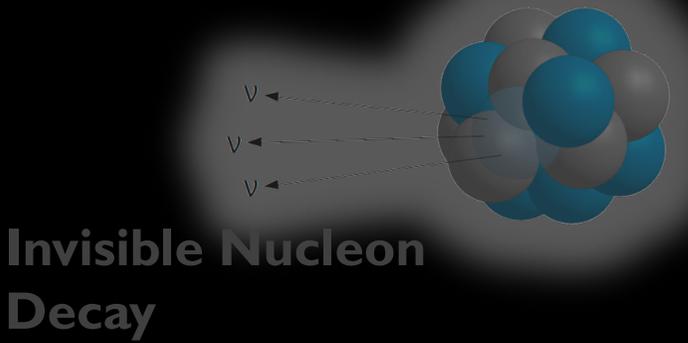
Physics Program



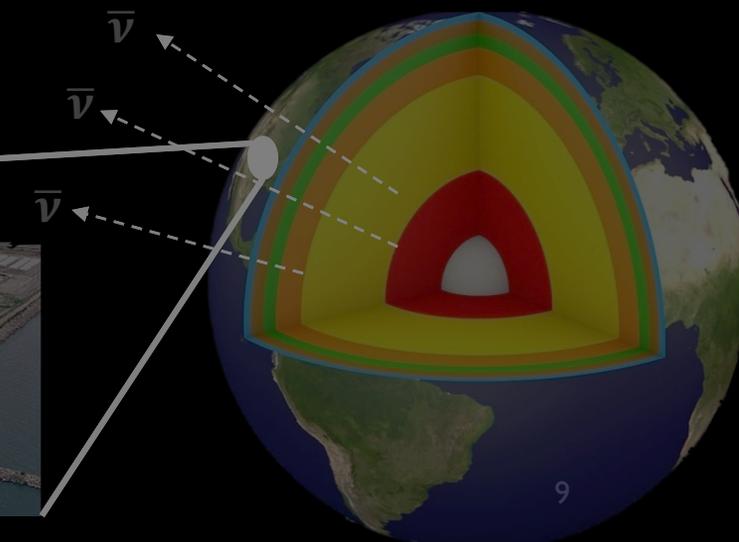
Main Goal



Supernova Neutrinos



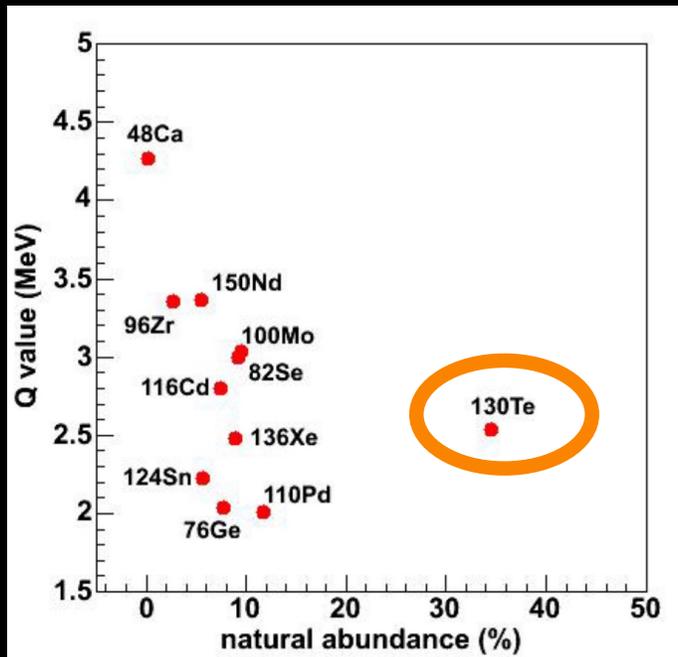
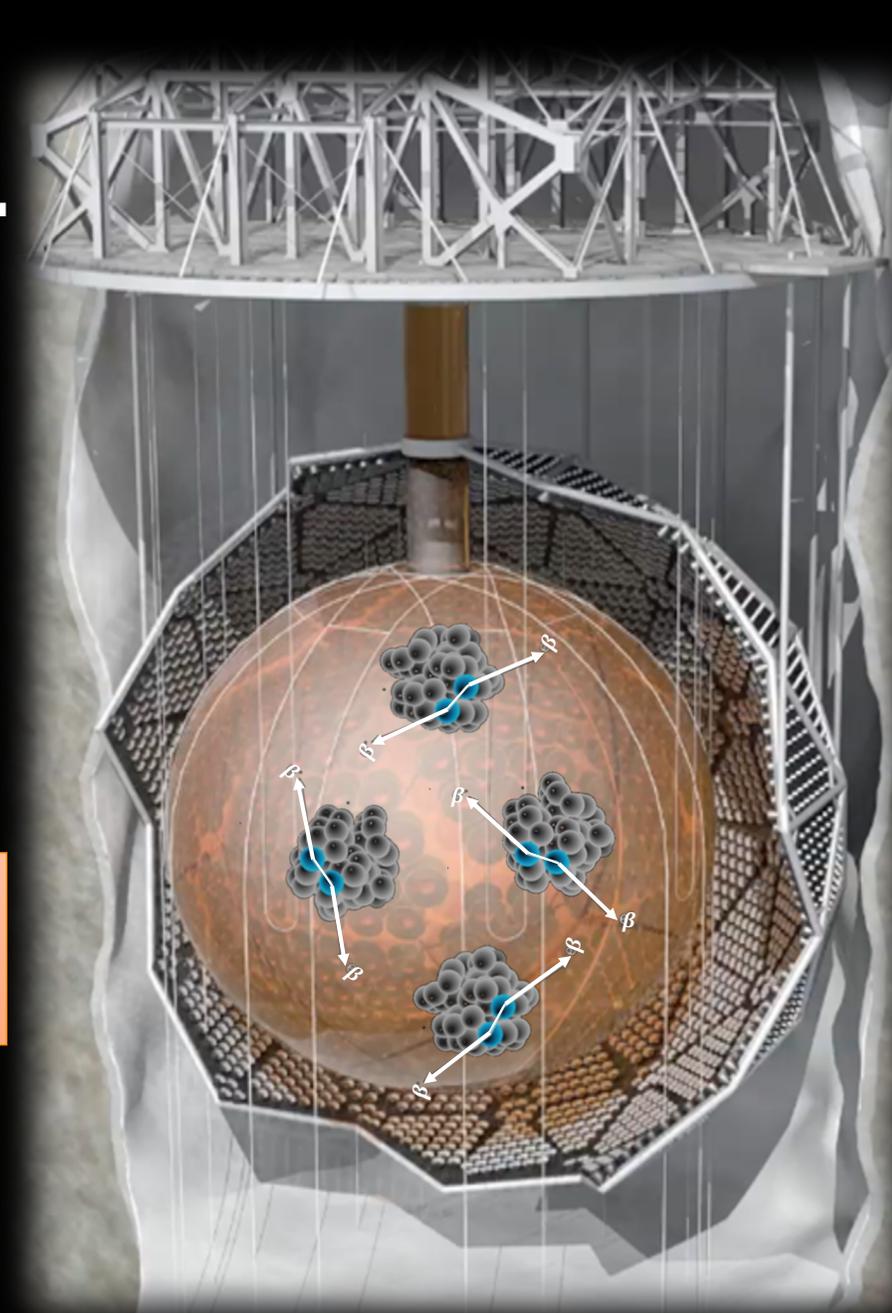
Geo-neutrinos



$0\nu\beta\beta$ Decay Search in SNO+

$0\nu\beta\beta$ decay candidate: ^{130}Te

- Highest natural abundance (34%), no enrichment needed – easily scalable at low cost.
- Q-value at 2.527 MeV – less background from natural radioactivity



**Initial phase loading:
0.5% natural Te by weight
= 1333 kg of ^{130}Te .**

Better (Greater phase space and lower backgrounds)

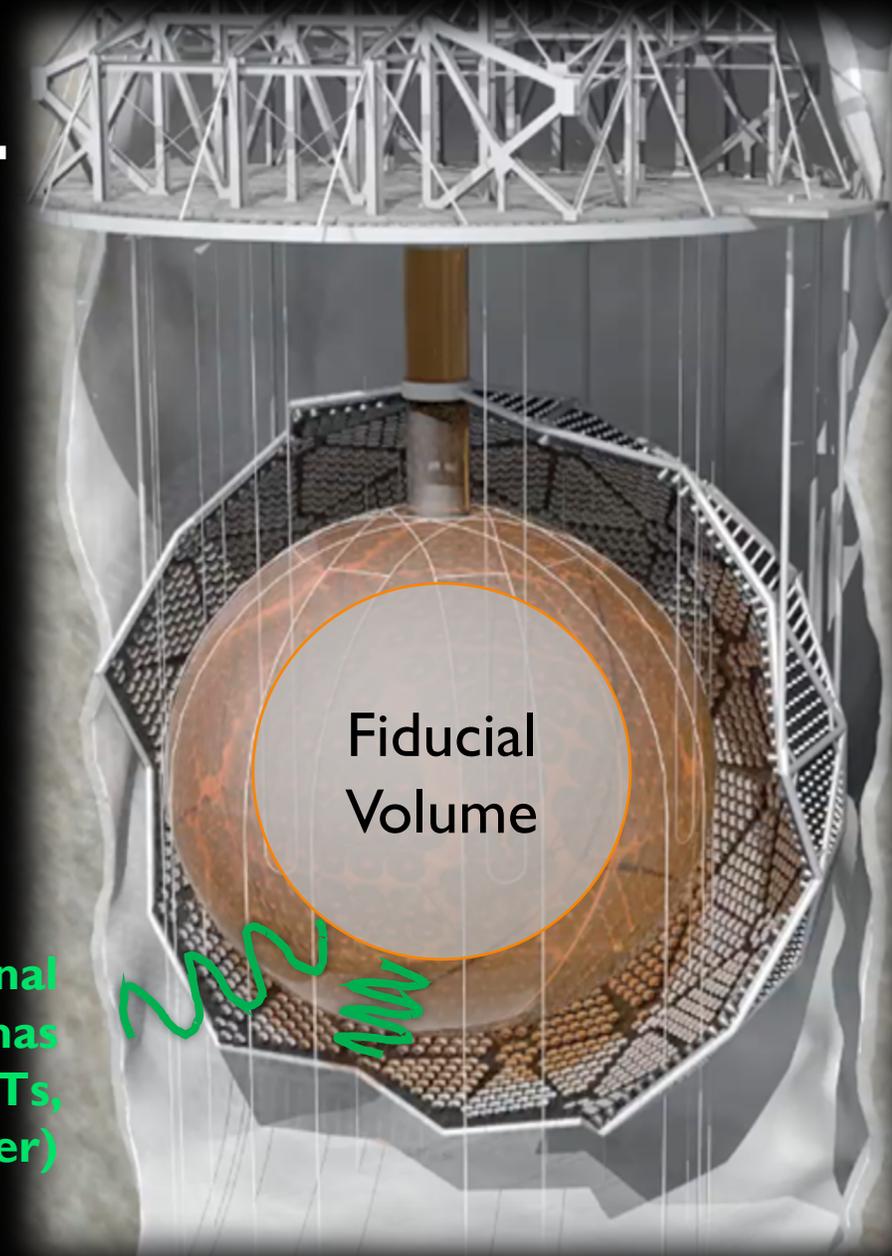
Better (cheaper)

$0\nu\beta\beta$ Decay Search in SNO+

Advantages

- I. Massive detector
 - High statistics
 - Self-shielding from external backgrounds through fiducialization.

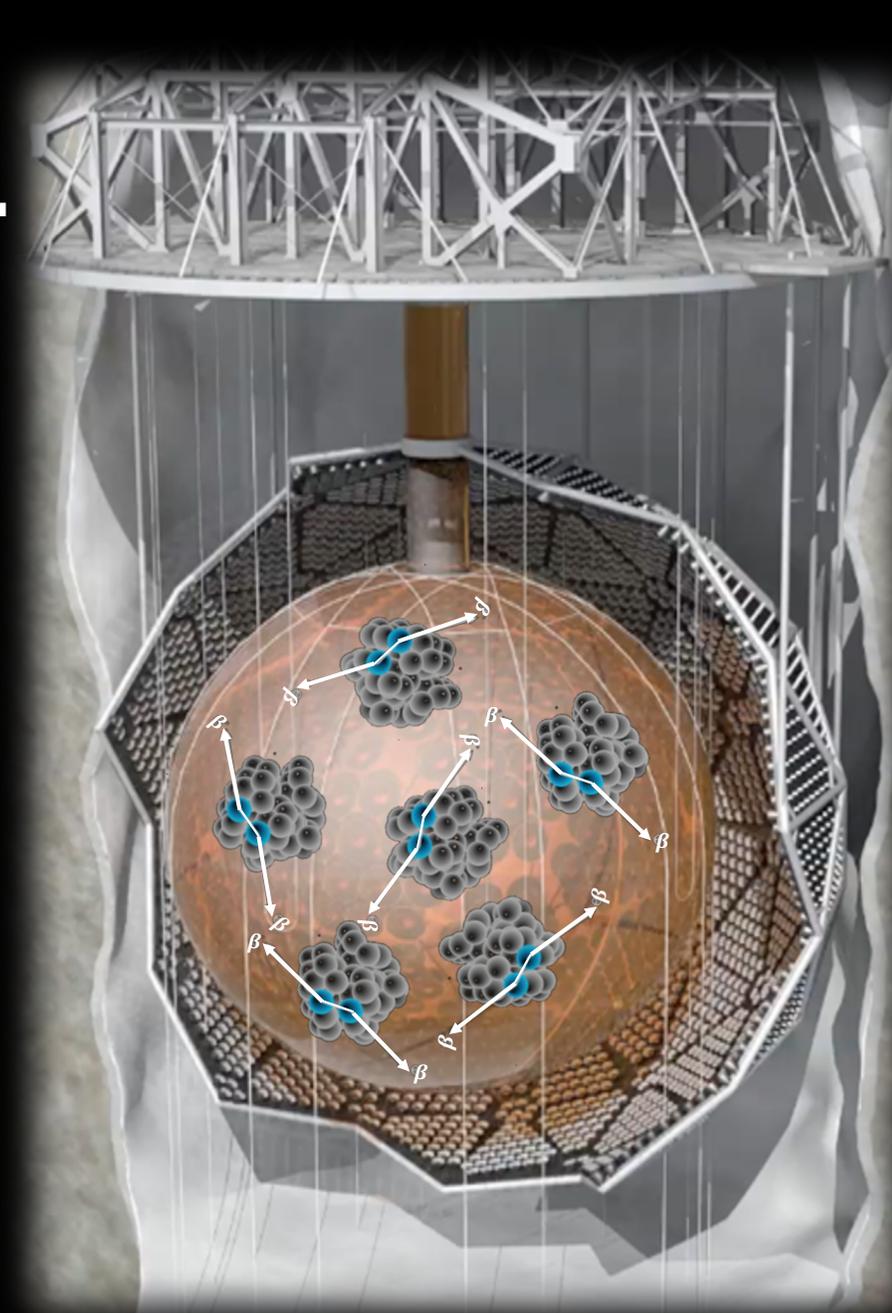
External
gammas
(ropes, PMTs,
rock, water)



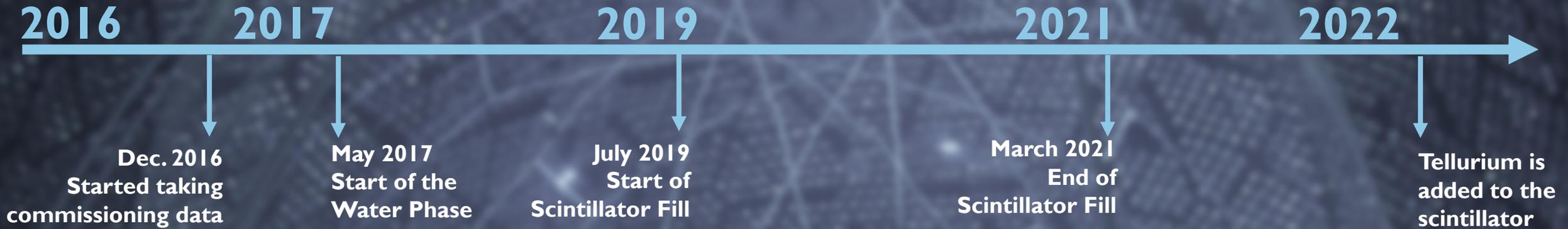
$0\nu\beta\beta$ Decay Search in SNO+

Advantages

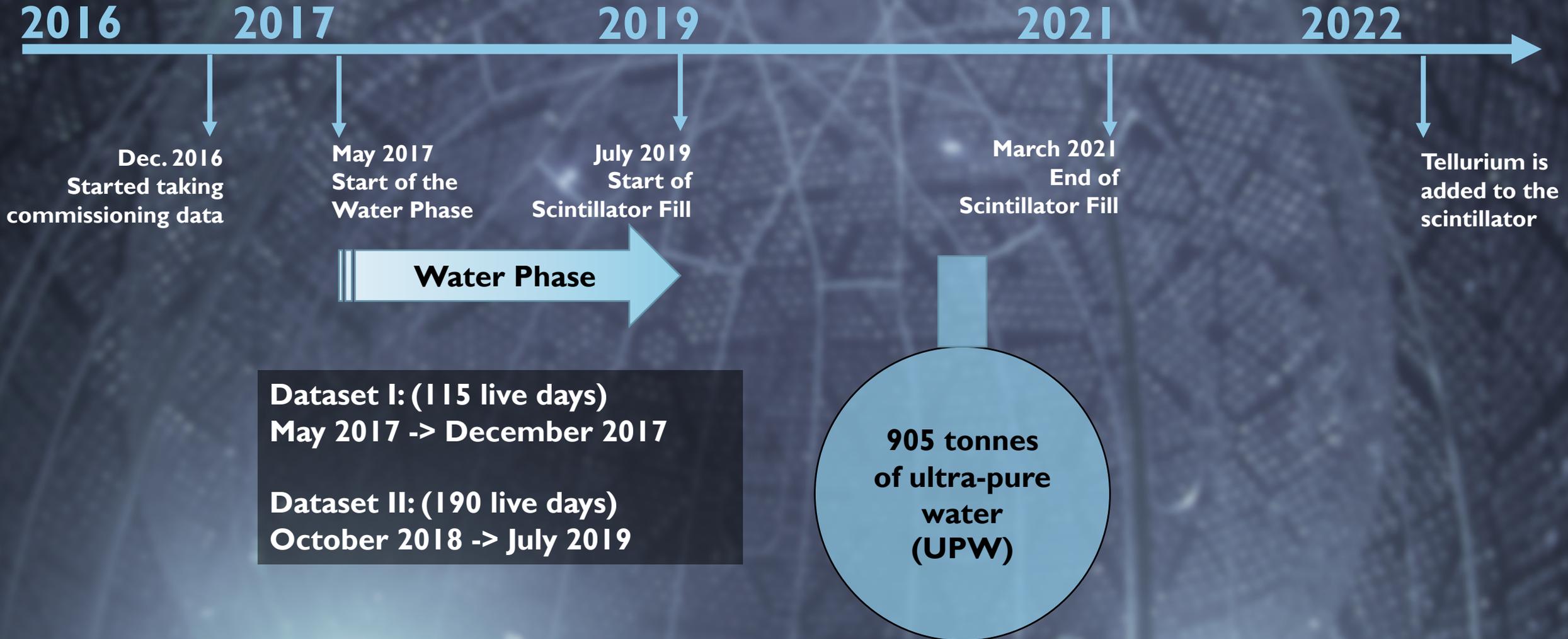
1. Massive detector
 - High statistics
 - Self-shielding from external backgrounds through fiducialization.
2. Liquid scintillator
 - Can be purified
 - Backgrounds can be measured and constrained prior to isotope addition
 - Loading can be scaled



SNO+ Timeline

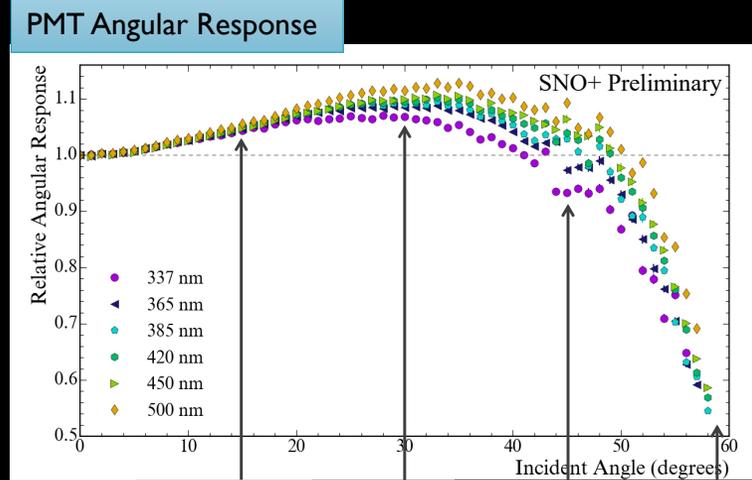
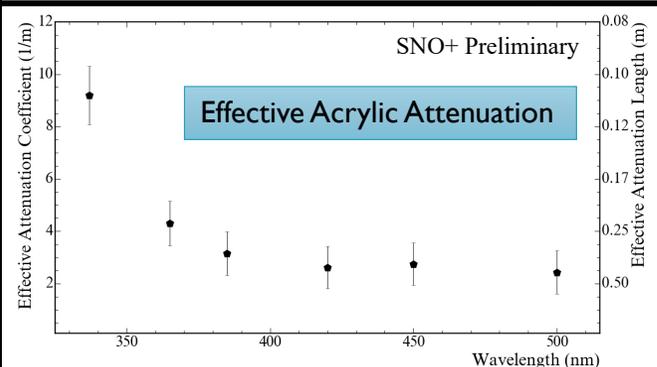
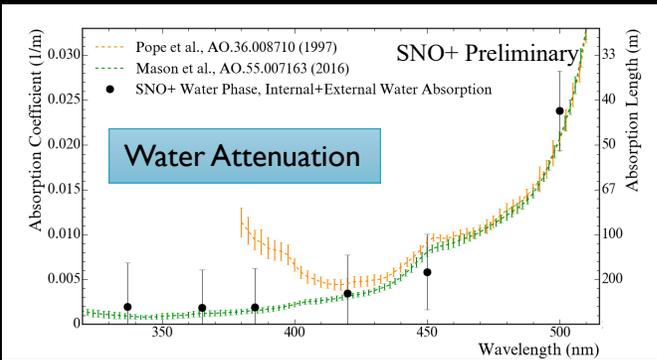
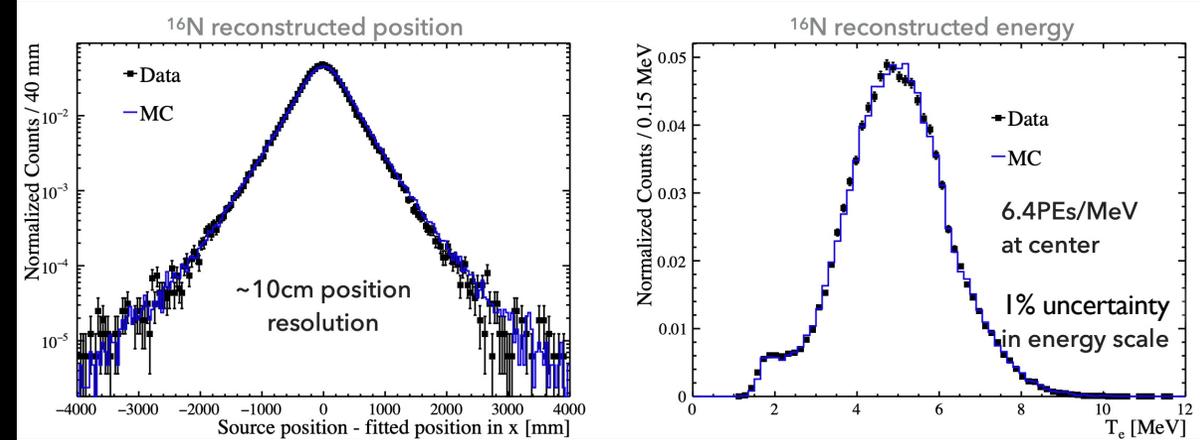


SNO+ Timeline



SNO+ Water Phase

- Detector calibrations:
 - Deployed and in-situ optical and radioactive sources to calibrate PMTs, measure timing and optical properties, calibrate reconstructed energy, position and direction.



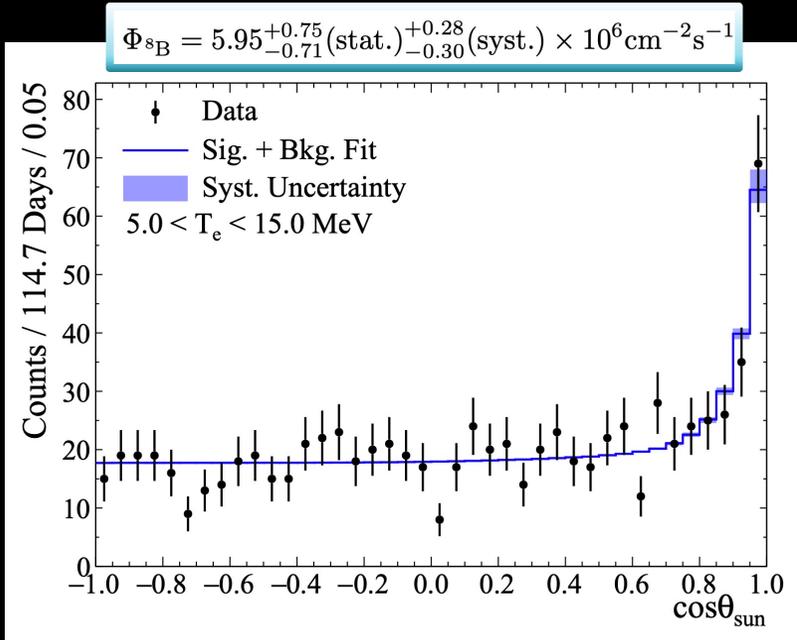
Validation using the ^{16}N source (6.13 MeV gammas), as a function of position

Data – Monte Carlo comparison shows 0.6% agreement, improvement over SNO!

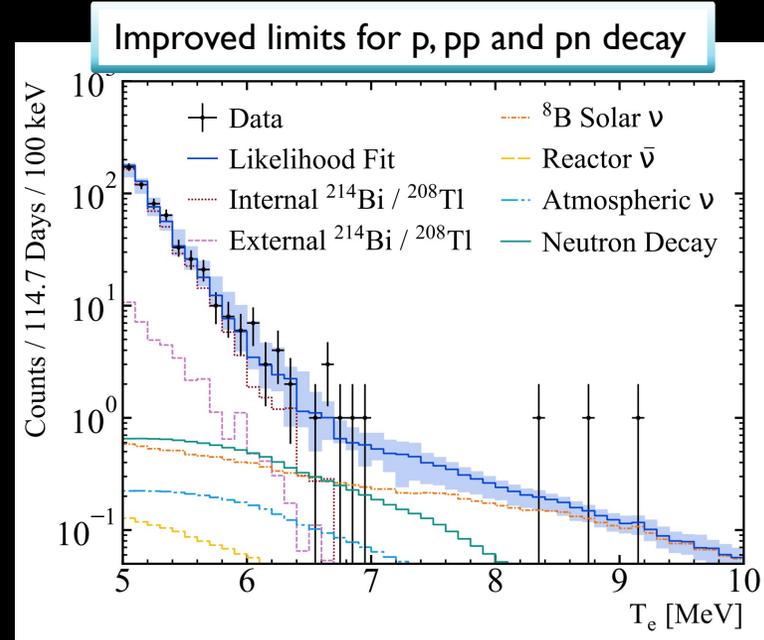


Optical calibration of the SNO+ detector in the water phase with deployed sources, [arXiv:2106.03951](https://arxiv.org/abs/2106.03951)

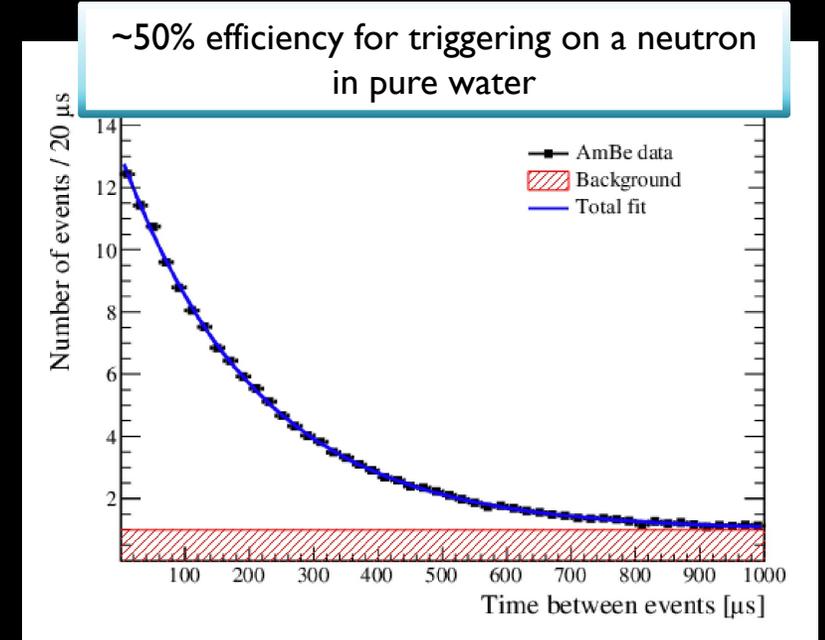
SNO+ Water Phase



^8B Solar Neutrino Flux
[Phys. Rev. D 99, 012012 \(2019\)](#)



Invisible modes of nucleon decay
[Phys. Rev. D 99, 032008 \(2019\)](#)

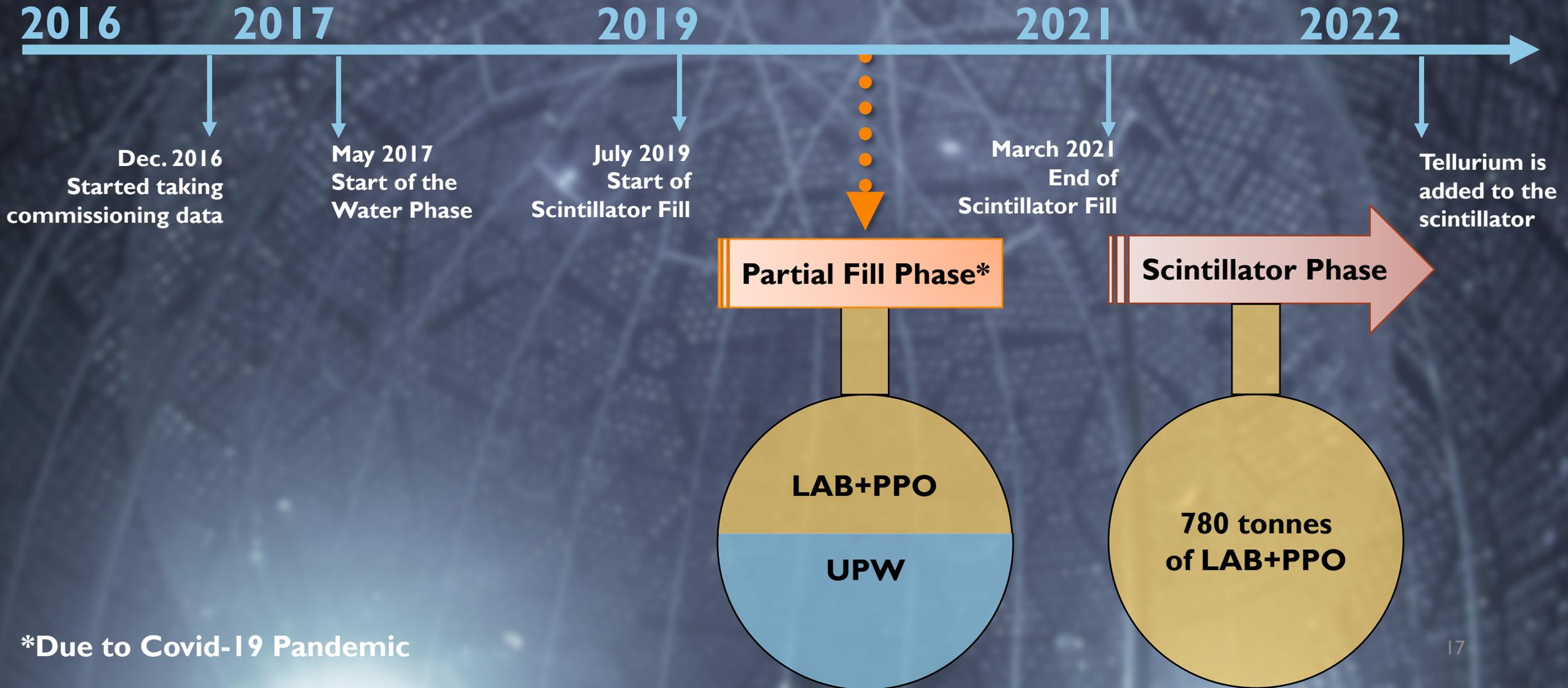


Neutron-proton capture
[Phys. Rev. C 102, 014002 \(2020\)](#)

**Updated results
 coming soon!**

- Higher statistics (additional 190 days of data)
- Including completed optical calibration
- Lower backgrounds due to cover gas installation in October 2018 (Rn levels one order of magnitude lower)

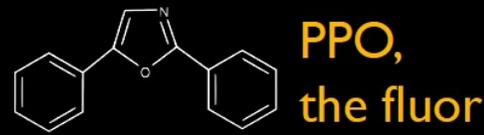
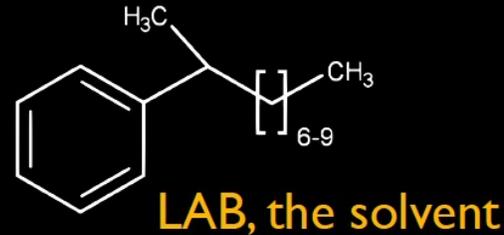
SNO+ Timeline



*Due to Covid-19 Pandemic

Scintillator Filling and Purification

- Linear alkylbenzene (LAB) with 2g/L of PPO.
- Commissioned underground systems and began filling SNO+ in 2019.
- Pause between April and October 2020 due to the pandemic.
- Fill ended in March 2021.



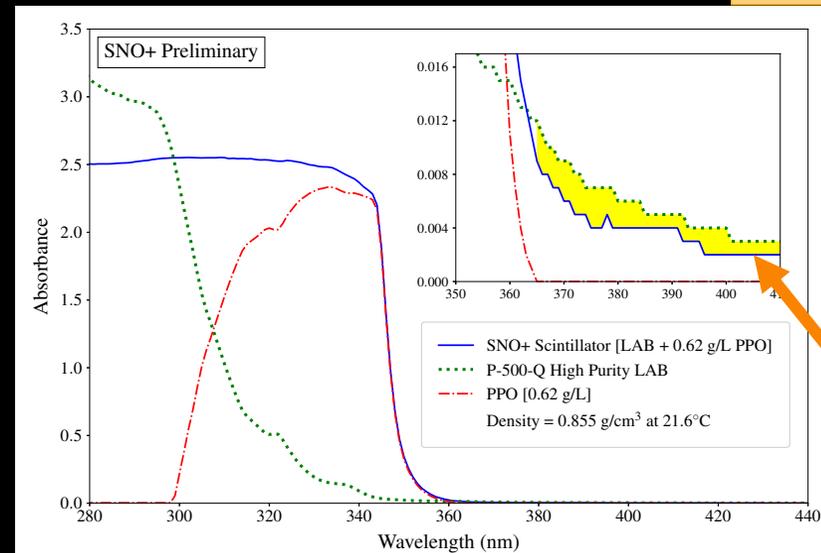
Transfer of LAB from surface to underground in tank railcars at SNOLAB



Purification and Filling Systems

➤ On-going PPO mixing campaign and recirculation.

0.5g/L → 2g/L

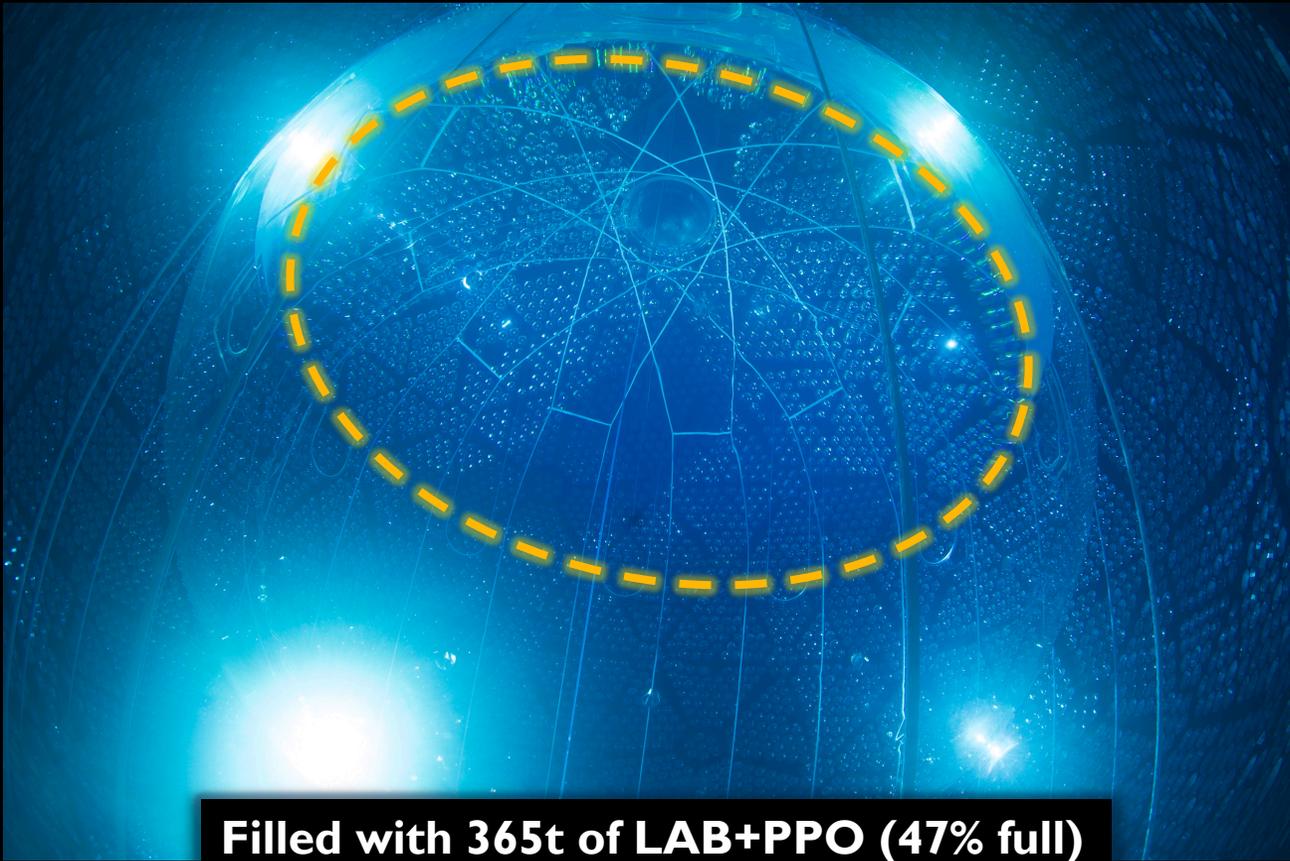


Excellent clarity above PPO absorption region!

Development, characterisation, and deployment of the SNO+ liquid scintillator, [JINST 16 \(2021\) P05009](#)

SNO+ Scintillator Phase

“Partial Fill Phase”



Filled with 365t of LAB+PPO (47% full)

PPO concentration 0.5 g/L
(25% of the nominal value)

- Almost 7 months in a half-filled configuration
- Trigger threshold lowered from 1 MeV (water) to ~40 keV (scint)
- Data during the fill used to **measure and monitor the backgrounds** in the liquid scintillator

Perform a “target out” $\beta\beta$ analysis

→ prepare/test analysis and techniques using real data

→ determine count rate in the ROI in the absence of Te

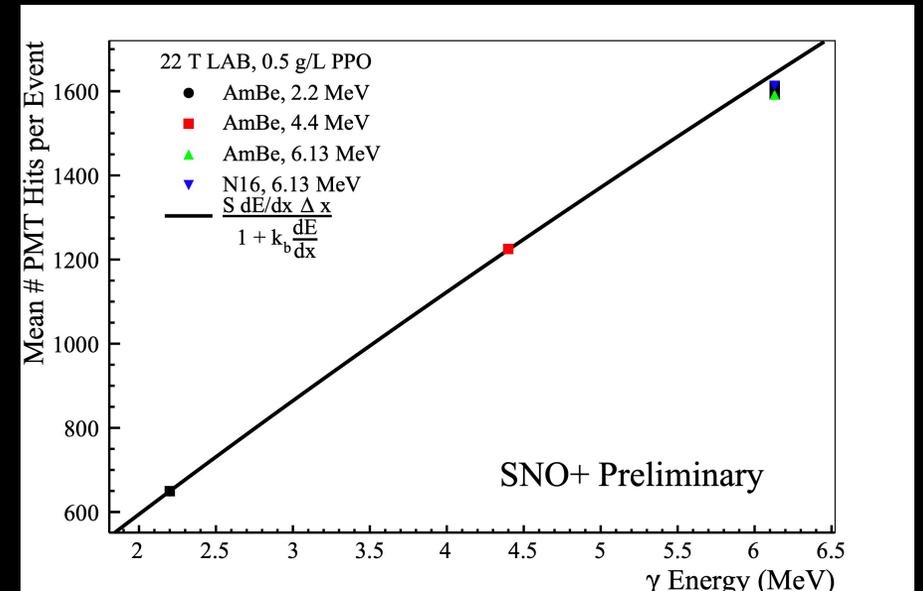
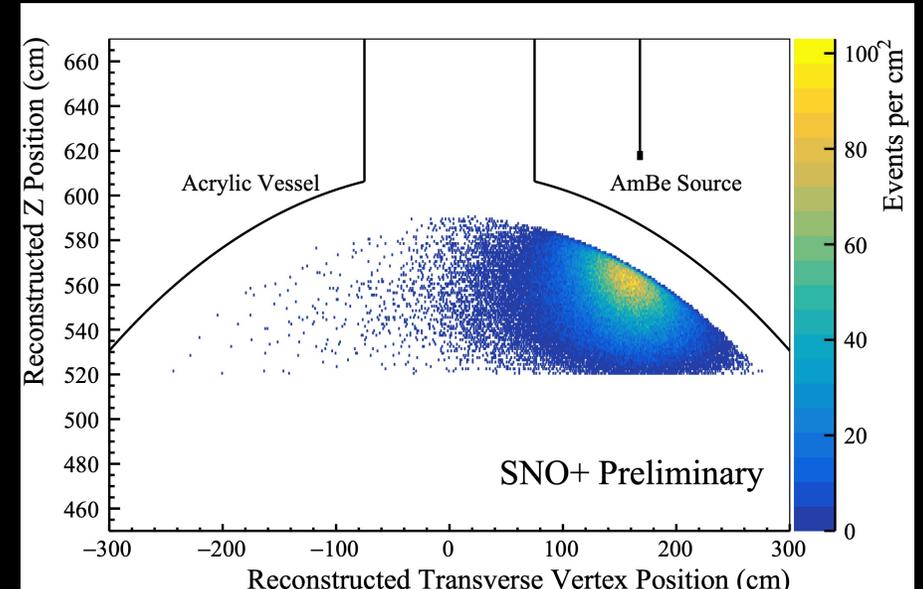
SNO+ Scintillator Phase

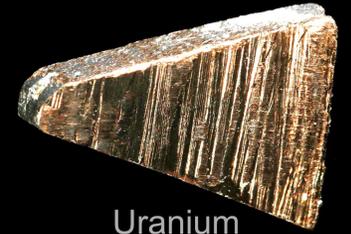
“Partial Fill Phase”

- Detector response during the fill was measured with optical and radioactive source calibrations
- Source deployments performed outside the acrylic vessel
 - Leave the scintillator undisturbed and avoid contamination
 - Demonstrated the capability to reconstruct events in a hybrid LS/water detector

With a PPO concentration of only 0.5 g/L we see a light yield equivalent to ~300 p.e. / MeV
Consistent with expectations!

Extrapolates to ~650 p.e. / MeV at 2.0 g/L PPO



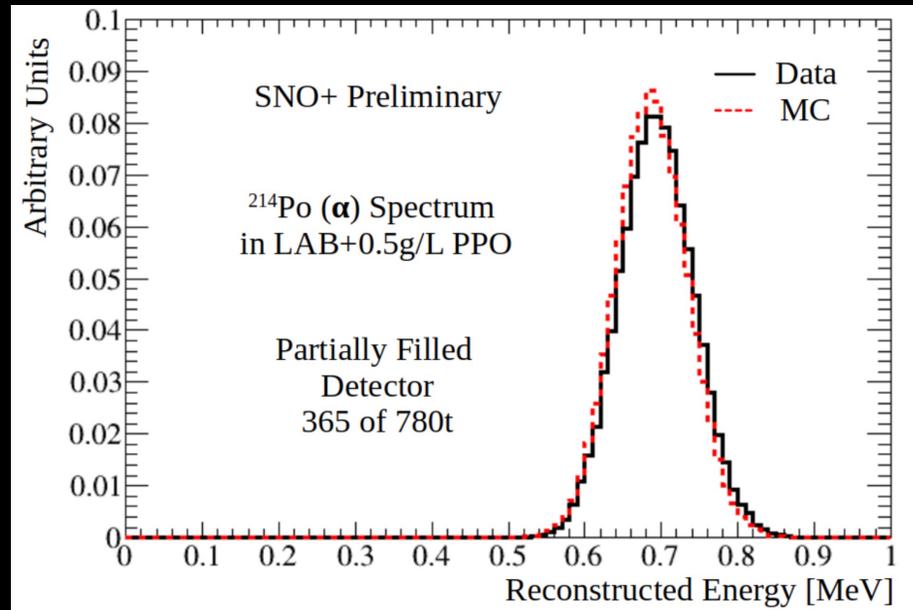
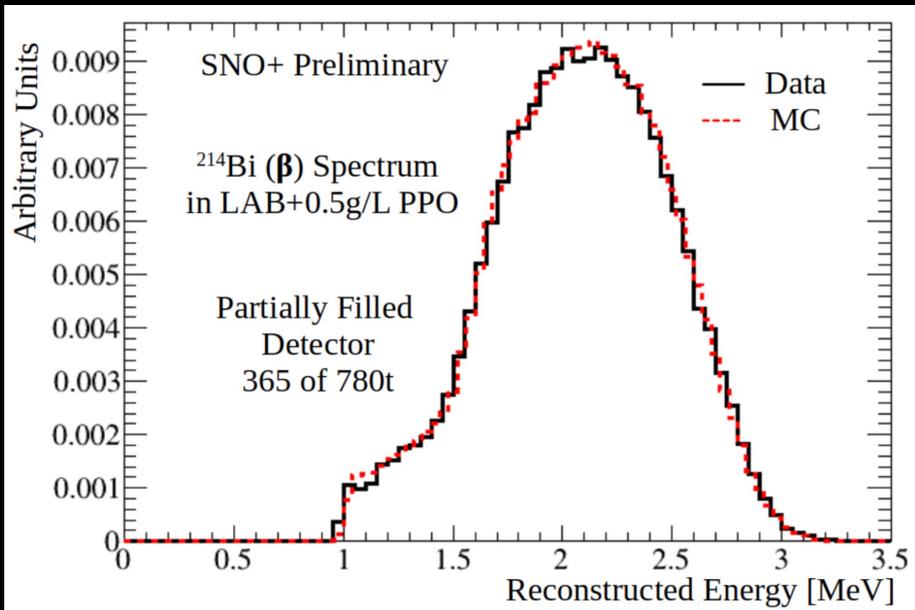
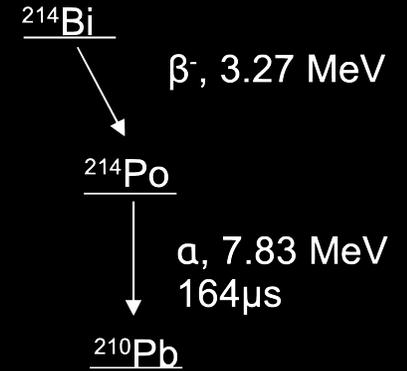


Uranium

SNO+ Scintillator Phase

“Partial Fill Phase”

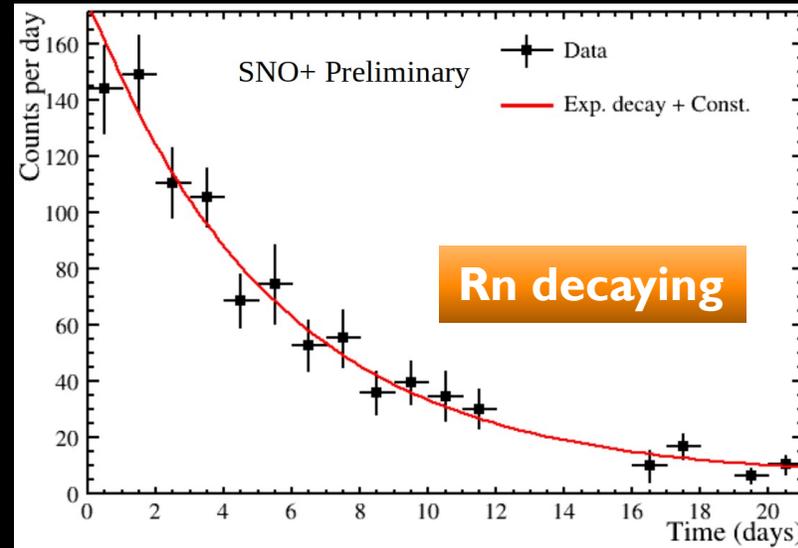
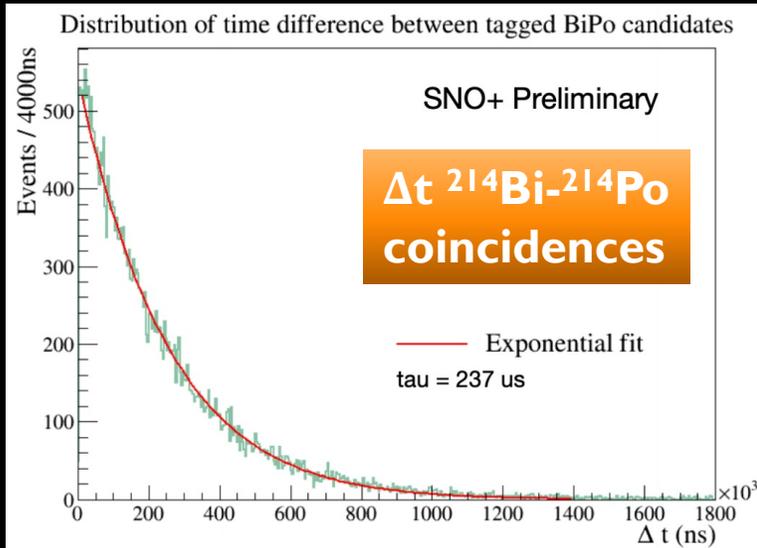
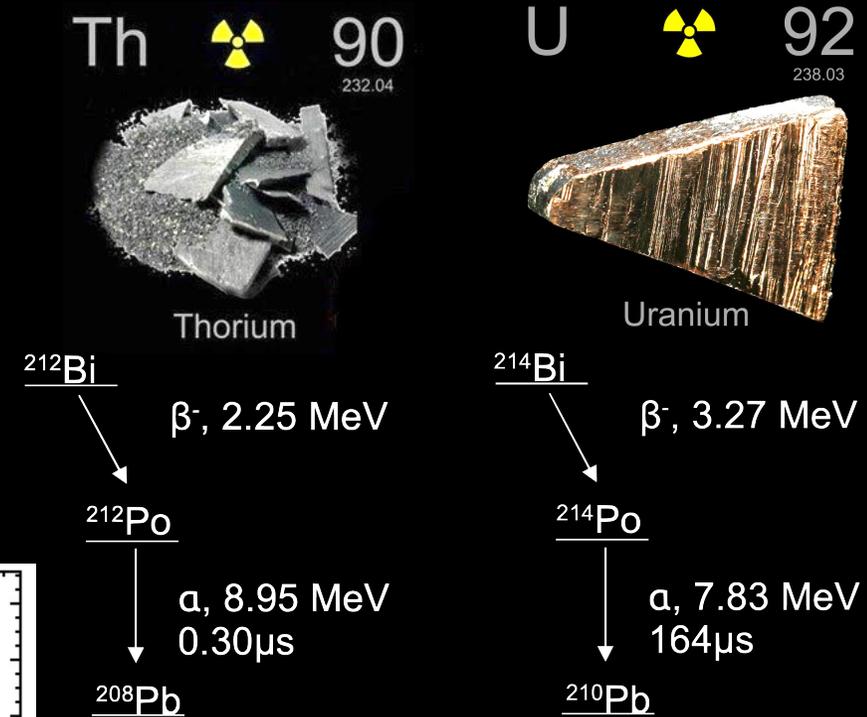
- Used ^{214}Bi and ^{214}Po tagged coincidences to calibrate particle timing profile, light yield and quenching.



SNO+ Scintillator Phase

“Partial Fill Phase”

- Measure intrinsic U and Th levels in the scintillator.
 - From ^{212}Bi and ^{214}Bi , which can be tagged by looking for the Po follower.



U Rate: $4.7 \pm 1.2 \times 10^{-17}$ gU/gScint

Th Rate: $5.3 \pm 1.5 \times 10^{-17}$ gTh/gScint

Measurements using the partial fill data show that our U and Th rates are below the requirements for the $0\nu\beta\beta$ search.

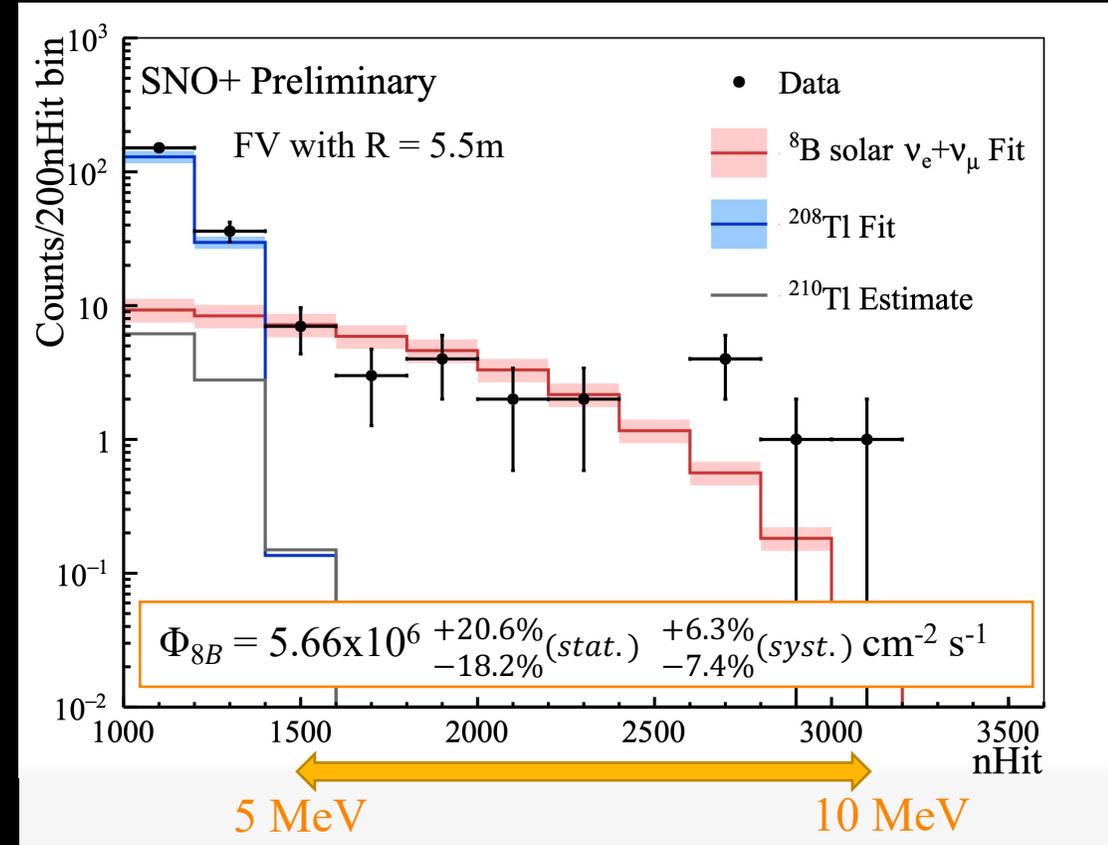
SNO+ Scintillator Phase

“Partial Fill Phase”

- Neutrino Physics program has begun:
 - Live for supernova neutrinos
 - Geo and Reactor anti-neutrinos
 - ^8B solar neutrinos

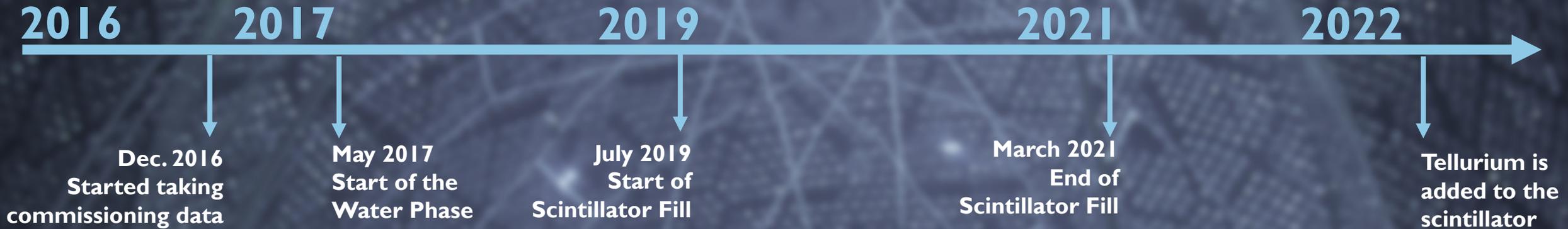
Preliminary results from the Partial Fill phase:

- Proof of concept + preparation of analysis tools for full fill phase.
- Exploring Cherenkov/Scintillation light separation and directionality in the low PPO data.



Preliminary ^8B solar ν + Bkg. Fit to the PF data, including preliminary systematics. Fitted flux compatible with other measurements.

SNO+ Timeline



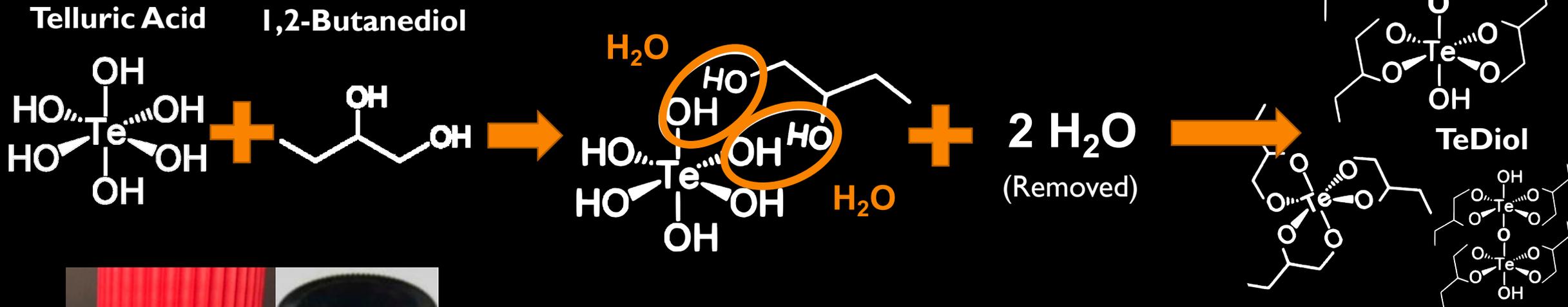
$0\nu\beta\beta$ Search

780 tonnes
of LAB+PPO

+0.5% Te
loading

Tellurium Loading

- Forming an organometallic compound from telluric acid and butanediol:



- TeDiol is mixed directly into the LAB+PPO with 15 mg/L bis-MSB and a stabilizer called Dimethyldodecylamine (DDA).
- Final Te-loaded LS cocktail expected to produce ~ 460 p.e. / MeV in SNO+ for 0.5% Te loading.

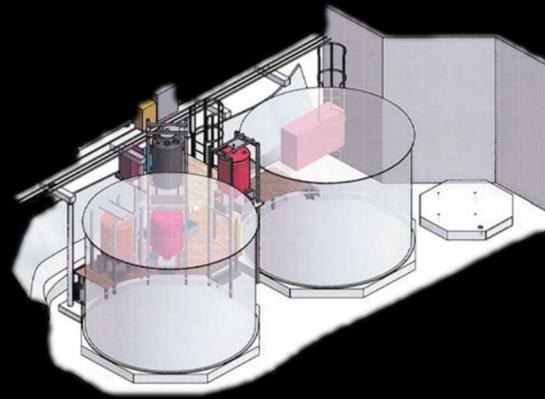
Tellurium Plants

- ~8 tons of telluric acid has been “cooling” underground for several years.
- Ton-scale underground purification of telluric acid for further background reduction.
- Target purification for Te cocktail:
 - ~ 10^{-15} g/g U
 - ~ 10^{-16} g/g Th

Tellurium Purification Plant

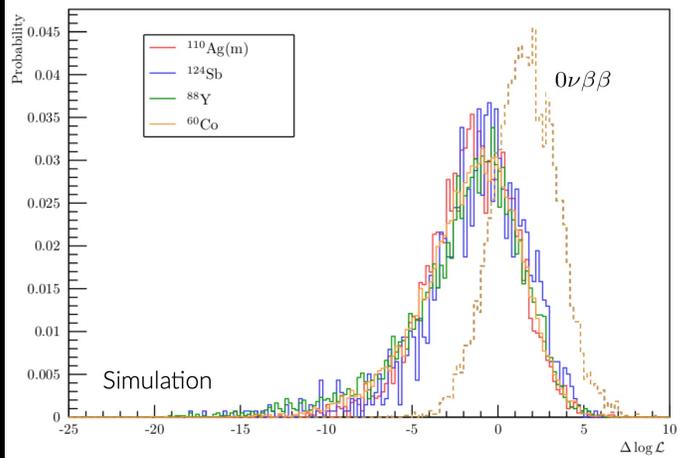


Tellurium Loading Plant



SNO+ Tellurium Phase - Prospects

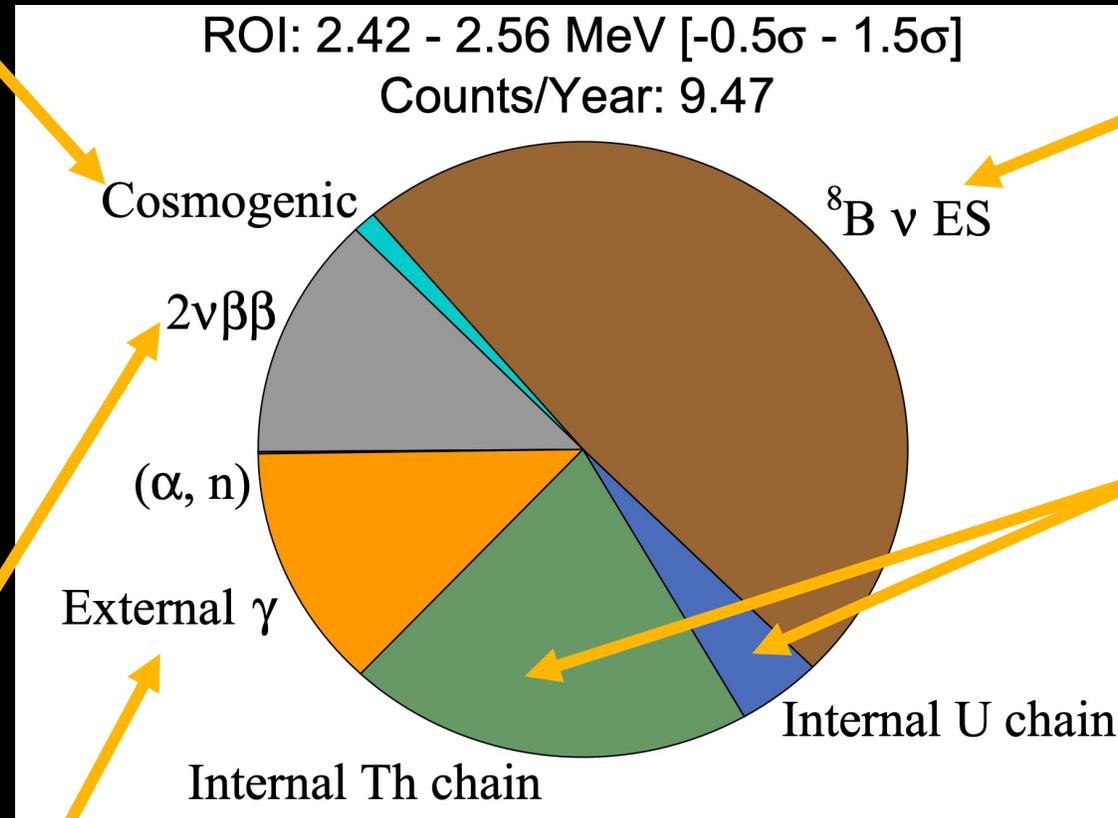
Telluric acid has been “cooling” underground for several years
 + Te purification
 + can be verified with multi-site analysis



Dunger and Biller, NIM 943, 162420, (2019)

Suppressed by asymmetric ROI

Measured in Water Phase, below target



Well known from other measurements

LS contribution measured during the partial fill, below our target for $\beta\beta$
 U and Th from the Te addition to be minimized with the purification systems

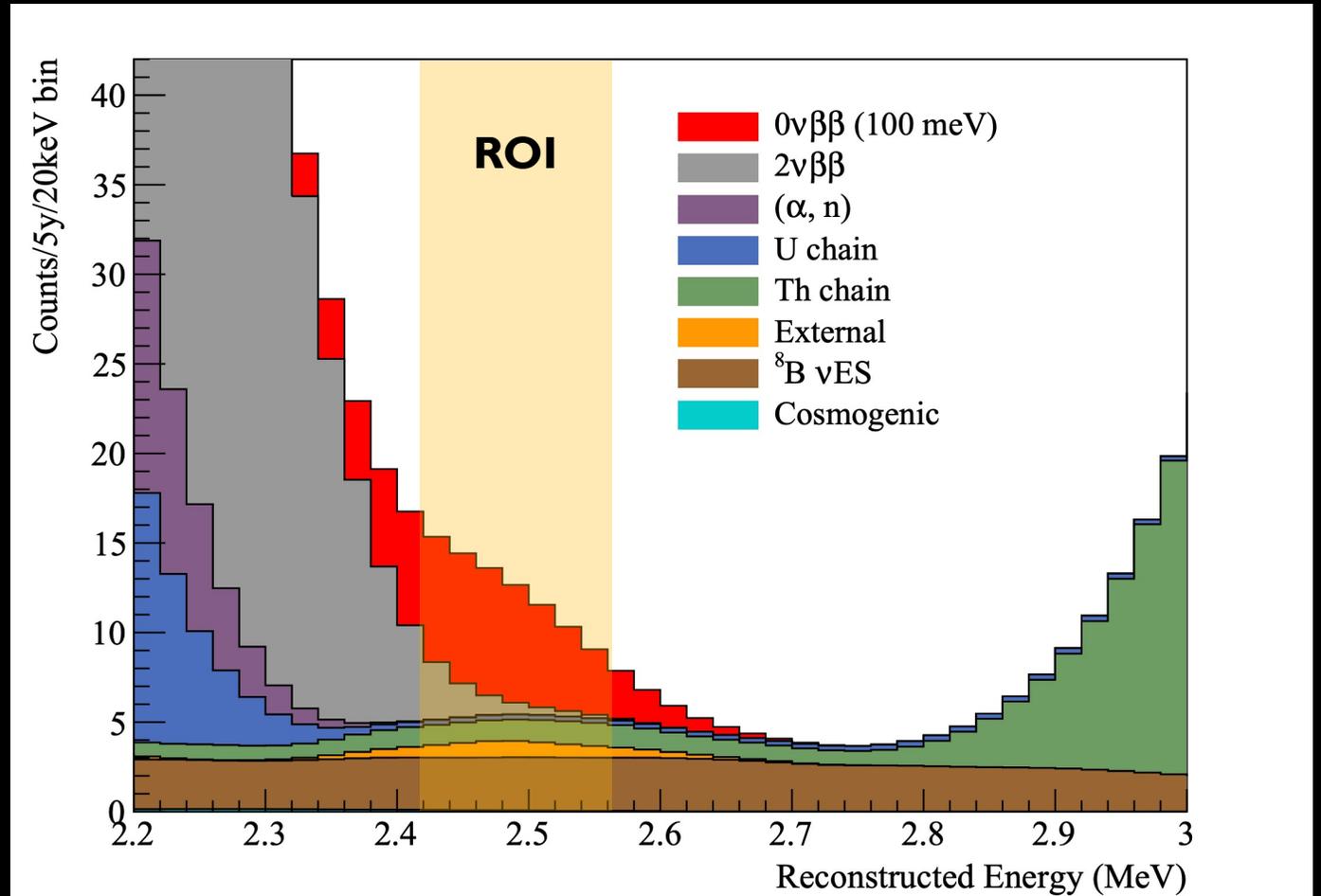
SNO+ Tellurium Phase - Prospects

Expected Energy Spectrum after 5
Years with 0.5% Te loading,
Fiducial Volume of 3.3 m radius

Expect 9.47 events / year in this ROI
(with our target background levels)



Half-Life Sensitivity $> 1.8 \times 10^{26}$ years



SNO+ Tellurium Phase - Prospects

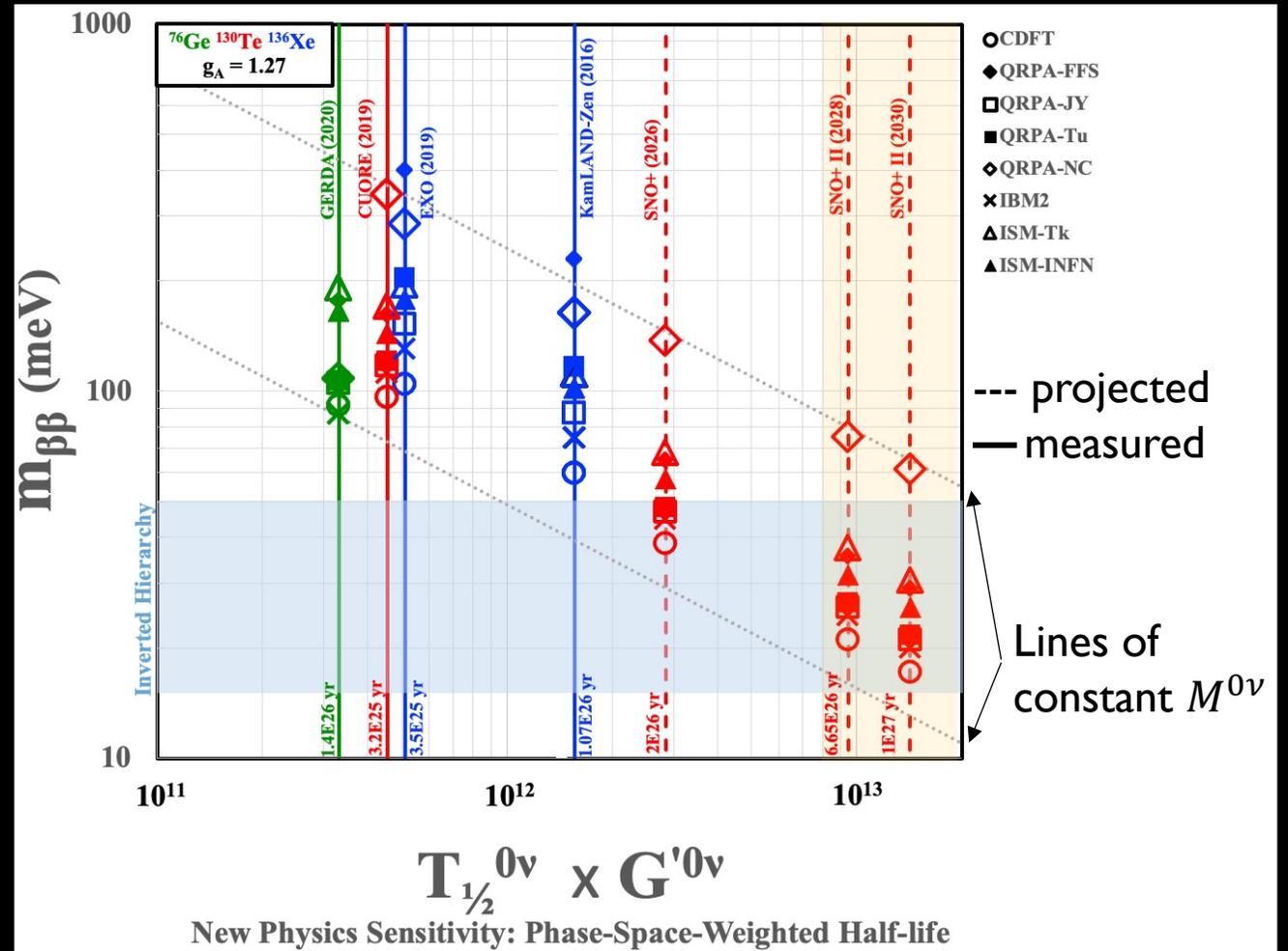
- From a simple counting analysis, for 5 years, in an optimized energy ROI and fiducial volume:

Expected Half-Life Sensitivity $> 2.1 \times 10^{26}$ years

$m_{\beta\beta}$ range 37-89 meV (model dependent)

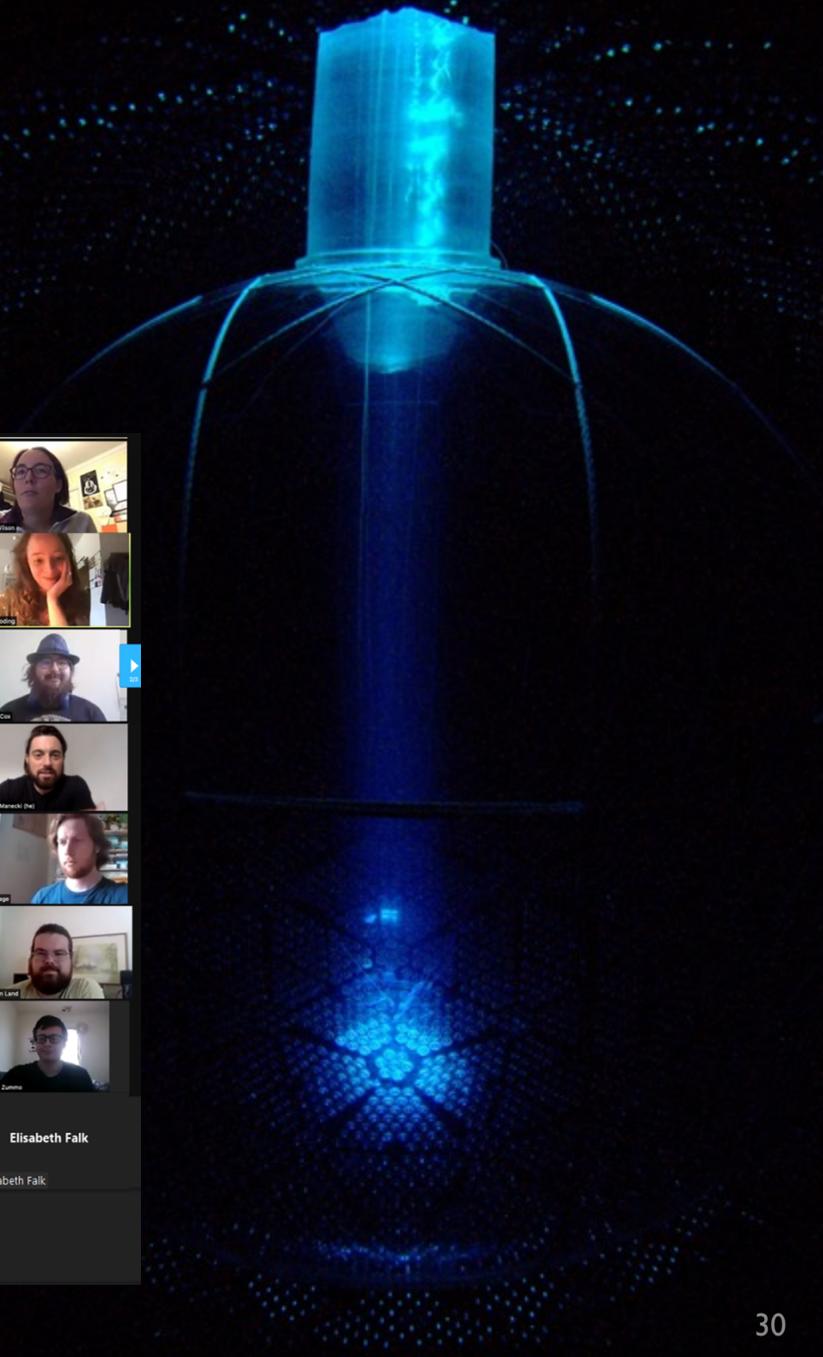
- Will be further optimized with a spectral fit.
- SNO+ approach can be scaled up.
 - Te loading up to 3% with good light yield and stability
 - Cost is relatively very low ($< \$2M$ per ton of decay isotope)

SNO+ Phase II



Thank You!

SNO+ Collaboration 2021

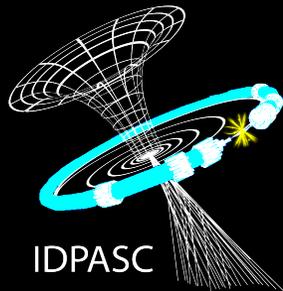


Acknowledgements

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IDPASC