CEvNS nuclear physics aspects

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Outline



• Experimental status and motivations of $CE\nu NS$

CEvNS in the Standard Model

- Reactor, accelerator and astrophysical-induced event rates
- Impact of nuclear form factors
- Beyond CEvNS: Incoherent calculations

Electromagnetic neutrino properties

- Neutrino magnetic moments & charge radii
- Active-sterile transitions

Summary

Physics Motivations of $CE\nu NS$



talks by O. Miranda, N. Cargioli

$CE\nu NS$ experiments worldwide



slide from Carla Bonifazi: TAUP 2021, Valencia

talks by Scholberg (COHERENT), Hakenmüller (CONUS)

Standard Model physics & nuclear structure

Standard Model CE ν NS cross section

 $CE\nu NS$ cross section expressed through the nuclear recoil energy T_A

$$\left(\frac{d\sigma}{dT_A}\right)_{\rm SM} = \frac{G_F^2 m_A}{\pi} \left[\mathcal{Q}_V^2 \left(1 - \frac{m_A T_A}{2E_\nu^2} \right) + \mathcal{Q}_A^2 \left(1 + \frac{m_A T_A}{2E_\nu^2} \right) \right] F^2(Q^2)$$

[DKP, Kosmas: PRD 97 (2018)]

- E_{ν} : is the incident neutrino energy
- *m_A*: the nuclear mass of the detector material
- Z protons and N = A Z neutrons
- vector Q_V and axial vector Q_A contributions
- $F(Q^2)$: is the nuclear form factor

$$\begin{aligned} \mathcal{Q}_{V} &= \left[2(g_{u}^{L} + g_{u}^{R}) + (g_{d}^{L} + g_{d}^{R}) \right] Z + \left[(g_{u}^{L} + g_{u}^{R}) + 2(g_{d}^{L} + g_{d}^{R}) \right] N, \\ \mathcal{Q}_{A} &= \left[2(g_{u}^{L} - g_{u}^{R}) + (g_{d}^{L} - g_{d}^{R}) \right] (\delta Z) + \left[(g_{u}^{L} - g_{u}^{R}) + 2(g_{d}^{L} - g_{d}^{R}) \right] (\delta N), \end{aligned}$$

• $(\delta Z) = Z_+ - Z_-$ and $(\delta N) = N_+ - N_-$, where $Z_+ (N_+)$ and $Z_- (N_-)$ refers to total number of protons (neutrons) with spin up or down [Barranco et al.: JHEP 0512 (2005)]

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$$\begin{aligned} \mathcal{Q}_V &= \left[\frac{1}{2} - 2\sin^2\theta_W\right] Z - \frac{1}{2}N, \\ \mathcal{Q}_A &= \frac{1}{2}(\delta Z) + \frac{1}{2}(\delta N), \end{aligned}$$

• weak mixing angle: $\sin^2 \theta_W$ not well measured at low energies



Nuclear rms radius

Phenomenological studies: form factor written in terms of the rms radius

The CEnNS process as unique probe of the neutron density distribution of nuclei

ron

The CEnNS process itself can be used to provide the first model independent measurement of the neutron distribution radius, which is basically unknown for most of the nuclei.

Even if it sounds strange, spatial distribution of neutrons inside nuclei is basically unknown!

The neutron distribution radius Rn and the difference between Rn and the rms radius Rp of the proton distribution (the socalled "neutron skin")

Scattered neutrino The Z boson couples preferentially with 2 Boson

neutrons

Muleorrecoil

slide from: M. Cadeddu @ NuFact 2018

Phenomenological form factors (Klein-Nystrand)

Follows from the convolution of a Yukawa potential with range $a_k = 0.7$ fm over a Woods-Saxon distribution, approximated as a hard sphere with radius R_A .

$$F_{\rm KN} = 3 \frac{j_1(QR_A)}{qR_A} \left[1 + (Qa_k)^2\right]^{-1}$$

The rms radius is: $\langle R^2 \rangle_{\rm KN} = 3/5R_A^2 + 6a_k^2$ [Klein, Nystrand, PRC 60 (1999) 014903]

- CEvNS data provides: a data driven determination of the neutron rms radius
- COHERENT (Csl) + APV (Cs): can disentangle the Cs and I contributions



talk by N. Cargioli

Nuclear structure models: BCS calculations

Within the context of the quasi-particle random phase approximation (QRPA) method, the form factors for protons (neutrons) are obtained as

$$F_{N_n} = \frac{1}{N_n} \sum_{j} \hat{j} \langle g.s. || j_0(|\mathbf{q}|r) || g.s. \rangle \left(v_j^{p(n)} \right)^2$$

where $\hat{j} = \sqrt{2j+1}$, $N_n = Z$ (or N), $v_j^{p(n)}$ are the BCS probability amplitudes, determined by solving iteratively the BCS equations.

[Kosmas, Vergados, Civitarese, Faessler, NPA 570 (1994) 637]

After choosing the active model space the following important parameters must be properly adjusted

- the harmonic oscillator (h.o.) size parameter b
- the two pairing parameters g^{p(n)}_{pair} for proton (neutron) pairs that renormalise the monopole (pairing) residual interaction of the Bonn C-D two-body potential (describing the strong two-nucleon forces)



- Realistic proton and neutron form factors
- The Bonn C-D residual interaction is mediated via one-meson exchange [Machleidt, PRC 63 (2001) 024001]

Nuclear structure models: deformed shell model

Kota & Sahu: Structure of Medium Mass Nuclei: Deformed Shell Model and Spin-Isospin Interacting Boson Model, CRC Press

- Assume axial symmetry
- Model space: a set single-particle (sp) orbitals + an effective two-body Hamiltonian
- Lowest-energy intrinsic states: by solving the HF single-particle equation self-consistently.
- Excited intrinsic configurations: via particle-hole excitations over the lowest intrinsic state.
- **Problem!** Intrinsic states $\chi_{\mathcal{K}}(\eta)$: do not have definite angular momenta

States of good angular momentum, projected from an intrinsic state $\chi_{K}(\eta)$

$$|\psi_{MK}^{J}(\eta)\rangle = rac{2J+1}{8\pi^{2}\sqrt{N_{JK}}}\int d\Omega D_{MK}^{J^{*}}(\Omega)R(\Omega)|\chi_{K}(\eta)\rangle,$$

where N_{JK} is the normalization constant given by

$$N_{JK} = \frac{2J+1}{2} \int_0^\pi d\beta \sin\beta d^J_{KK}(\beta) \langle \chi_K(\eta) | e^{-i\beta J_y} | \chi_K(\eta) \rangle .$$

- $R(\Omega) = \exp(-i\alpha J_z) \exp(-i\beta J_y) \exp(-i\gamma J_z)$: general rotation operator
- Ω : the Euler angles (α , β , γ)
- $|\psi_{MK}^{J}(\alpha)\rangle$ projected from different intrinsic states are not in general orthogonal to each other

Band mixing calculations are performed after appropriate orthonormalization. The resulting eigenfunctions are of the form

$$|\Phi^{J}_{M}(\eta)\rangle = \sum_{K,\alpha} S^{J}_{K\eta}(\alpha) |\psi^{J}_{MK}(\alpha)\rangle, \qquad S^{J}_{K\eta}(\alpha): \text{ expansion coefficients}$$

The nuclear matrix elements occurring in the calculation of magnetic moments, elastic and inelastic spin structure functions etc. are evaluated using the wave functions $|\Phi_M^J(\eta)\rangle$. $_{11/25}$

Comparison of the nuclear methods



Impact of form factor on CE ν NS: COHERENT exp.







Impact of form factor on CE ν NS: COHERENT exp.



[DKP, Kosmas, Sahu, Kota, Hota PLB 800 (2020) 135133]



Up to 8% difference in the theoretical event rates

Naumov Bednyakov formalism

- 0

$$\frac{d\sigma_{\text{inc}}}{dT_A} = \frac{4G_F^2 m_A}{\pi} \sum_{f=n,p} g_{\text{inc}}^f \left(1 - |F_f|^2\right) \\
\times \left[A_+^f \left(\left(g_{L,f} - g_{R,f} a b^2\right)^2 + g_{R,f}^2 a b^2 (1-a) \right) + A_-^f g_{R,f}^2 (1-a) \left(1 - a + a b^2\right) \right].$$
(1)

$$a = \frac{q^2}{q_{\min}^2} \simeq \frac{T_A}{T_A^{\max}}, \quad b^2 = \frac{m_f^2}{s}.$$
 (2)

Here, $A_{\pm}^{p} \equiv Z_{\pm}$ $(A_{\pm}^{n} \equiv N_{\pm})$ represents the number of protons (neutrons) with spin $\pm 1/2$ and $s = (p + k)^{2}$ is the total energy squared in the center-of-mass frame (p denotes an effective 4-momentum of the nucleon).

[Bednyakov, Naumov, PRD 98 (2018) 053004]

For a more detailed study see: [Hoferichter, Menéndez, Schwenk PRD 102, 074018]

Incoherent vs. Coherent rates: πDAR and reactors

reactor neutrinos

 πDAR neutrinos



Incoherent vs. Coherent rates: solar neutrinos



Solar neutrinos

Incoherent calculations: multipole expansion

Interaction Hamiltonian for neutral-current (NC) neutrino-nucleus scattering

$$\langle f | \hat{H}_{eff} | i \rangle = \frac{G_F}{\sqrt{2}} \int d^3 \mathbf{x} \, \langle \ell_f | \hat{j}_{\mu}^{lept}(\mathbf{x}) | \ell_i \rangle \langle J_f | \hat{\mathcal{J}}^{\mu}(\mathbf{x}) | J_i \rangle$$

with $\langle \ell_f | \hat{j}^{jept}_{\mu} | \ell_i \rangle = \bar{\nu}_{\alpha} \gamma_{\mu} (1 - \gamma_5) \nu_{\alpha} e^{-i\mathbf{q}\cdot\mathbf{x}}, \quad \mathbf{q}: 3 - \text{momentum transfer}$

 In the Donnelly-Walecka multipole decomposition method, the NC, double diff. SM cross section from an initial |*J_i* to a final |*J_f* nuclear state (constructed explicitly through QRPA realistic nuclear structure calculations), reads

$$\frac{d^2\sigma_{i\to f}}{d\Omega\,d\omega} = \frac{G_F^2}{\pi} \frac{\varepsilon_i \varepsilon_f}{(2J_i+1)} \left(\sum_{J=0}^{\infty} \sigma_{\rm CL}^J + \sum_{J=1}^{\infty} \sigma_{\rm T}^J \right) \,,$$

ε_i (ε_f) is the initial (final) neutrino energy and ω is the nucleus excitation energy.
 Contributions to σ^J_{CL} (Coulomb-longitudinal) and σ^J_T (transverse electric-magnetic) components [Donnelly and Peccei, Phys. Rept. 50 (1979) 1]

$$\begin{aligned} \sigma_{\mathrm{CL}}^{J} = & (1 + a\cos\theta) |\langle J_{f} || \hat{\mathcal{M}}_{J}(\kappa) || J_{i} \rangle|^{2} + (1 + a\cos\theta - 2b\sin^{2}\theta) |\langle J_{f} || \hat{\mathcal{L}}_{J}(\kappa) || J_{i} \rangle|^{2} \\ &+ \left[\frac{\omega}{\kappa} (1 + a\cos\theta) + d \right] 2\Re |\langle J_{f} || \hat{\mathcal{L}}_{J}(\kappa) || J_{i} \rangle || \langle J_{f} || \hat{\mathcal{M}}_{J}(\kappa) || J_{i} \rangle|^{*} , \\ \sigma_{\mathrm{T}}^{J} = & (1 - a\cos\theta + b\sin^{2}\theta) \left[|\langle J_{f} || \hat{\mathcal{T}}_{J}^{mag}(\kappa) || J_{i} \rangle|^{2} + |\langle J_{f} || \hat{\mathcal{T}}_{J}^{el}(\kappa) || J_{i} \rangle|^{2} \right] \\ &+ \left[\frac{(\varepsilon_{i} + \varepsilon_{f})}{\kappa} (1 - a\cos\theta) - d \right] 2\Re e |\langle J_{f} || \hat{\mathcal{T}}_{J}^{mag}(\kappa) || J_{i} \rangle || \langle J_{f} || \hat{\mathcal{T}}_{J}^{el}(\kappa) || J_{i} \rangle|^{*} \end{aligned}$$

from kinematics: $a=1,~b=arepsilon_iarepsilon_f/\kappa^2,~d=0$ and $\kappa=|\mathbf{q}|$

Evaluation of the Nuclear Matrix Elements

Seven new operators are defined (proton-neutron representation) as

$$\begin{split} T_1^{JM} &\equiv & M_M^J(\kappa\mathbf{r}) = \delta_{LJ} j_L(\kappa\mathbf{r}) Y_M^L(\hat{\mathbf{r}}), \\ T_2^{JM} &\equiv & \Sigma_M^J(\kappa\mathbf{r}) = \mathbf{M}_M^{JJ} \cdot \boldsymbol{\sigma}, \\ T_3^{JM} &\equiv & \Sigma_M^{'J}(\kappa\mathbf{r}) = -i \left[\frac{1}{\kappa} \nabla \times \mathbf{M}_M^{JJ}(\kappa\mathbf{r}) \right] \cdot \boldsymbol{\sigma}, \\ T_4^{JM} &\equiv & \Sigma_M^{'J}(\kappa\mathbf{r}) = \left[\frac{1}{\kappa} \nabla M_M^J(\kappa\mathbf{r}) \right] \cdot \boldsymbol{\sigma}, \\ T_5^{JM} &\equiv & \Delta_M^{J}(\kappa\mathbf{r}) = \mathbf{M}_M^{JJ}(\kappa\mathbf{r}) \cdot \frac{1}{\kappa} \nabla, \\ T_6^{JM} &\equiv & \Delta_M^{'J}(\kappa\mathbf{r}) = -i \left[\frac{1}{\kappa} \nabla \times \mathbf{M}_M^{JJ}(\kappa\mathbf{r}) \right] \cdot \frac{1}{\kappa} \nabla \\ T_7^{JM} &\equiv & \Omega_M^{J}(\kappa\mathbf{r}) = M_M^{J}(\kappa\mathbf{r}) \boldsymbol{\sigma} \cdot \frac{1}{\kappa} \nabla. \end{split}$$

Closed compact analytic formulae for the single-particle reduced ME (upper) and many-body reduced ME (lower) for QRPA calculations, are deduced.

$$\langle (n_1\ell_1)j_1 || T_i^J || (n_2\ell_2)j_2 \rangle = e^{-y} y^{\beta/2} \sum_{\mu=0}^{n_{max}} \mathcal{P}_{\mu}^{i, J} y^{\mu}, \quad y = (\kappa b/2)^2, \quad n_{max} = (N_1 + N_2 - \beta)/2, \quad N_i = 2n_i + \ell_i$$

$$\langle f || \hat{T}^J || 0_{gs}^+ \rangle = \sum_{j_2 \ge j_1} \frac{\langle j_2 || \hat{T}^J || j_1 \rangle}{\hat{j}} \left[X_{j_2 j_1} u_{j_2}^{p(n)} v_{j_1}^{p(n)} + Y_{j_2 j_1} v_{j_2}^{p(n)} u_{j_1}^{p(n)} \right]$$

[Chasioti, Kosmas NPA 829 (2009) 234] P.G. Giannaka, D.K. Papoulias, T.S. Kosmas, unpublished (for any configuration $(j_1, j_2)J$)

19/25

Electromagnetic neutrino properties

Electromagnetic contribution to $CE\nu NS$ cross section

The Electromagnetic CEVNS cross section reads [Vogel, Engel.: PRD 39 [1989] 3378]

$$\left(\frac{d\sigma}{dT_A}\right)_{\rm EM} = \frac{\pi a_{\rm EM}^2 \mu_\nu^2 Z^2}{m_e^2} \left(\frac{1 - T_A/E_\nu}{T_A}\right) F^2(Q^2) \,.$$

• can be dominant for sub-keV threshold experiments

• may lead to detectable distortions of the recoil spectrum

The helicity preserving SM cross section adds incoherently with the helicity-violating EM cross section

$$\left(\frac{d\sigma}{dT_A}\right)_{\rm tot} = \left(\frac{d\sigma}{dT_A}\right)_{\rm SM} + \left(\frac{d\sigma}{dT_A}\right)_{\rm EM}$$

 μ_{ν}^2 is the effective neutrino magnetic moment in the mass basis relevant to a given neutrino beam (reactor, SNS, etc.)

- Experimental measurements usually constrain some process-dependent effective parameter combination
- needs to be expressed in terms of fundamental parameters (TMMs + CP phases + mixing-angles)
- Even in the case of laboratory neutrino experiments, where the initial neutrino flux is fixed to have a well determined given flavor, there is no sensitivity to the final neutrino state, 1/25

Analysis of the COHERENT data: EM properties

Neutrino magnetic moment contrbution

$$\left(\frac{d\sigma}{dT_N}\right)_{\rm SM+EM} = \mathcal{G}_{\rm EM}(E_\nu\,,\,T_N) \frac{d\sigma_{\rm SM}}{dT_N}\,,$$

$$\mathcal{G}_{\rm EM} = 1 + \frac{1}{G_F^2 M} \left(\frac{\mathcal{Q}_{\rm EM}}{\mathcal{Q}_W^V}\right)^2 \frac{\frac{1 - T_N / E_\nu}{T_N}}{1 - \frac{M T_N}{2E_\nu^2}}$$

• EM charge: $Q_{EM} = \frac{\pi^{a} E_{M} \mu_{\nu_{\alpha}}}{m_{e}} Z$ [Vogel, Engel: PRD 39 [1989] 3378]

Neutrino charge radius

$$\sin^2 \theta_W \to \sin^2 \overline{\theta_W} + \frac{\sqrt{2\pi a_{\rm EM}}}{3G_F} \langle r_{\nu_{\alpha}}^2 \rangle \,.$$



Miranda et al. JHEP 05 (2020) 130



Active-sterile transitions via magnetic moment





stay tuned! [Miranda, DKP, Sanders, Tórtola, arXiv:2109.XXXX]

poster ID#58 (room E) by Oscar Sanders

CE ν **NS**

$\nu_{\alpha} + (A, Z) \rightarrow \nu_{\alpha} + (A, Z), \quad \alpha = (e, \mu, \tau)$

- Finally observed on Csl (2017) and LAr (2020) (other: MINER, TEXONO, CONNIE, Ricochet, νGEN, ν-cleus etc.)
- CONUS, CONNIE (hints)
- Very high experimental sensitivity (low detector threshold) is required

- largest theoretical uncertainty coming nuclear physics
- sensitive to nuclear charge radii
- can probe electromagnetic neutrino properties





Thank you for your attention !