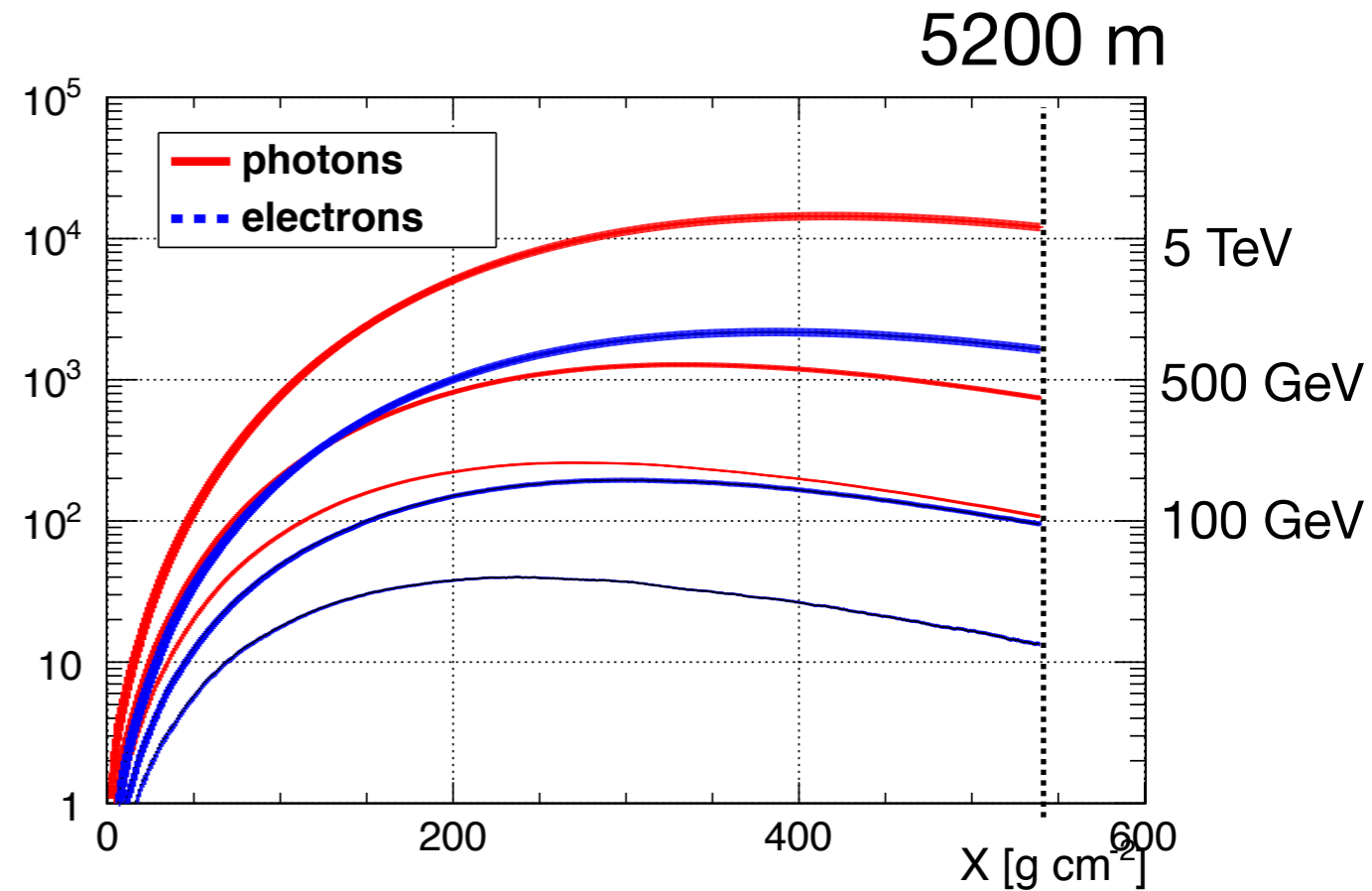
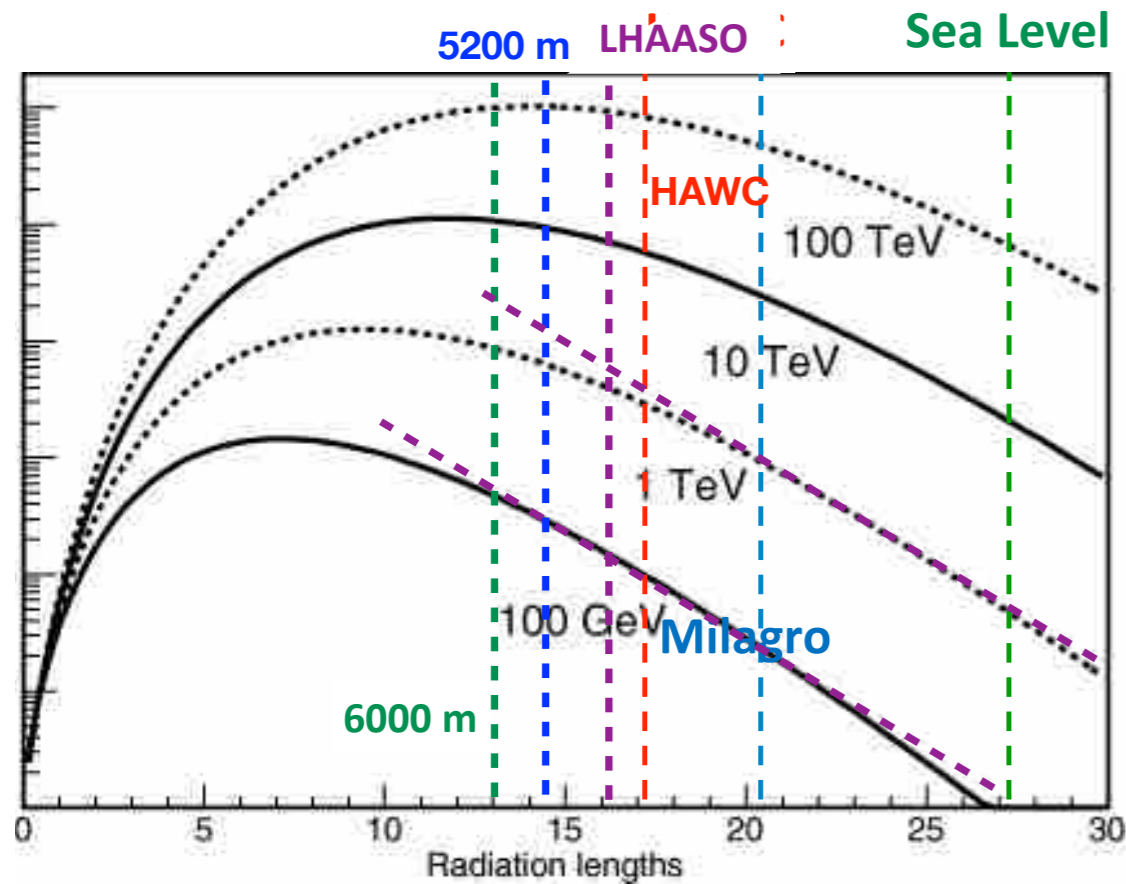


# *Compact WCD with a matrix of SiPM*

Ruben Conceição, Giovanni La Mura, Cédric Perennes,  
Mário Pimenta, Bernardo Tomé

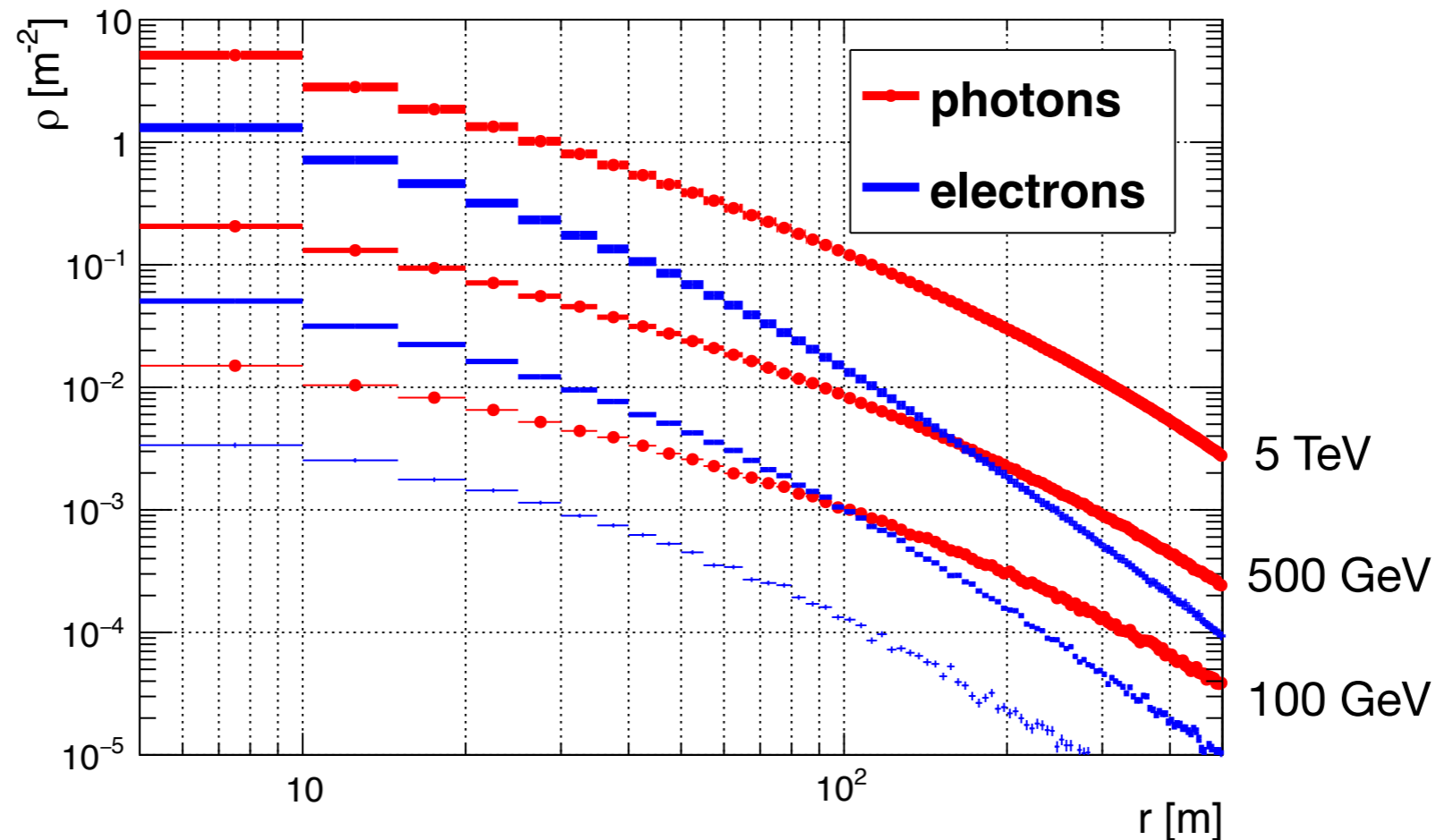


# The low energy challenge



- Even at high altitudes (reasonable!) low energy photon showers are still detected at  $> 5 X_0$  below  $X_{\text{max}}$
- For  $E = 100$  GeV vertical showers,  $\sim 100$  e.m. particles reach the ground at  $\sim 5000$  m a.s.l
- $\sim 90\%$  of the shower secondaries are photons

# Mean LDF at 5200 m altitude

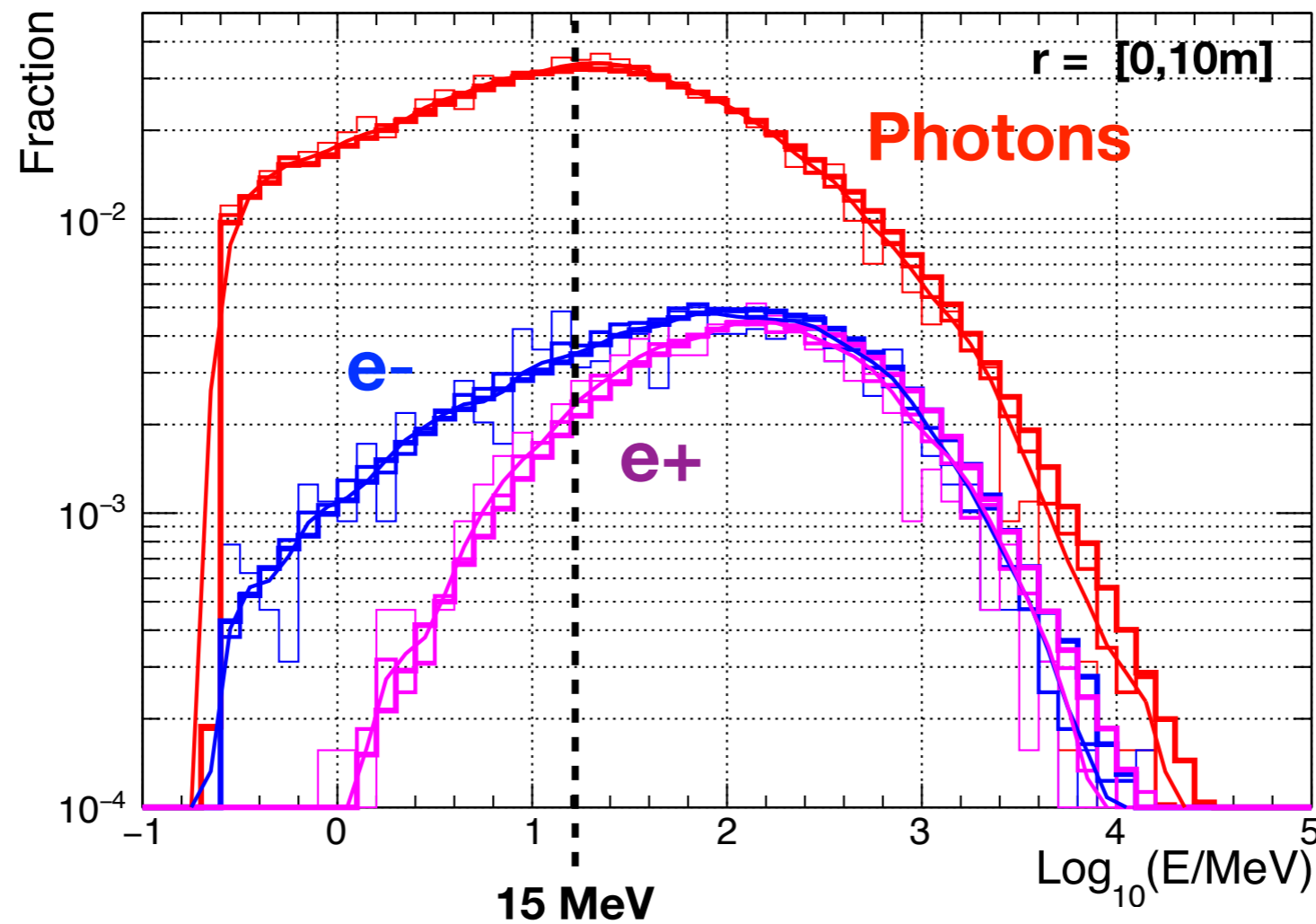


@ 100 GeV :

- $r = 100 \text{ m} \rightarrow \langle \rho \rangle \sim 0.001 \text{ photons/m}^2$
- $r = 10 \text{ m} \rightarrow \langle \rho \rangle \sim 0.015 \text{ photons/m}^2$
- Increasing the station area does not help to increase the individual station signal...

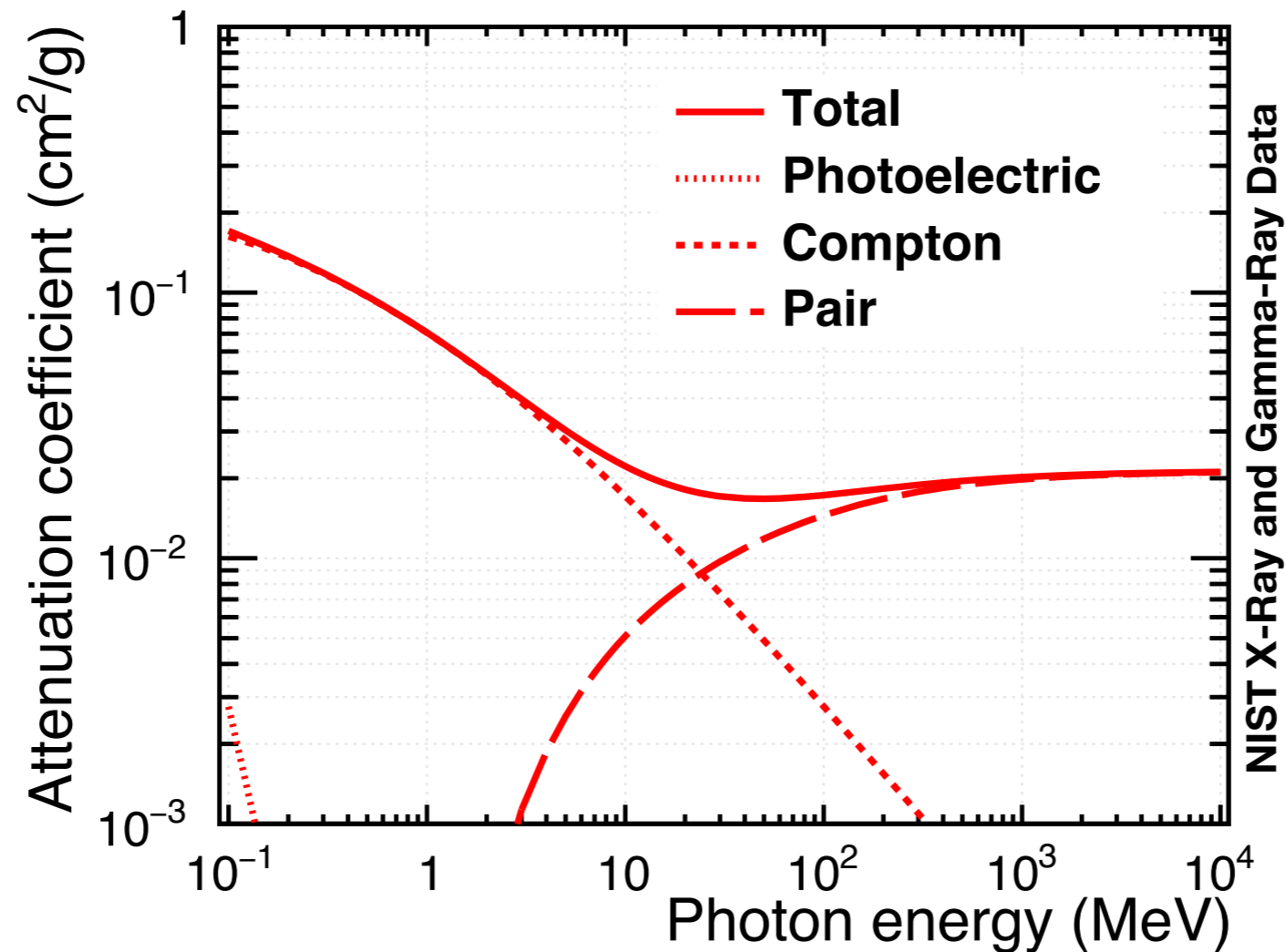
Need single-particle station sensitivity !

# Secondary particle spectra @ 5200 m altitude



- WCD sensitivity to single  $\sim 10$  MeV - 15 MeV photons is a challenge !

# Photon interactions in H<sub>2</sub>O



- At  $> 10$  MeV :  $L_{\text{Compton}} + L_{\text{Pair}} \sim 50 \text{ g/cm}^2$  ( $\sim 9/7 X_0$ )
- 1 m H<sub>2</sub>O  $\rightarrow P_{\text{int}} = 0.86$ ; 2 m H<sub>2</sub>O  $\rightarrow P_{\text{int}} = 0.98$
- 2 x water depth increases the interaction probability by 1.14 ...

# A compact WCD concept

Addressing new (and old...) challenges

- **Station signal**

- WCD should be white to lower energy threshold!

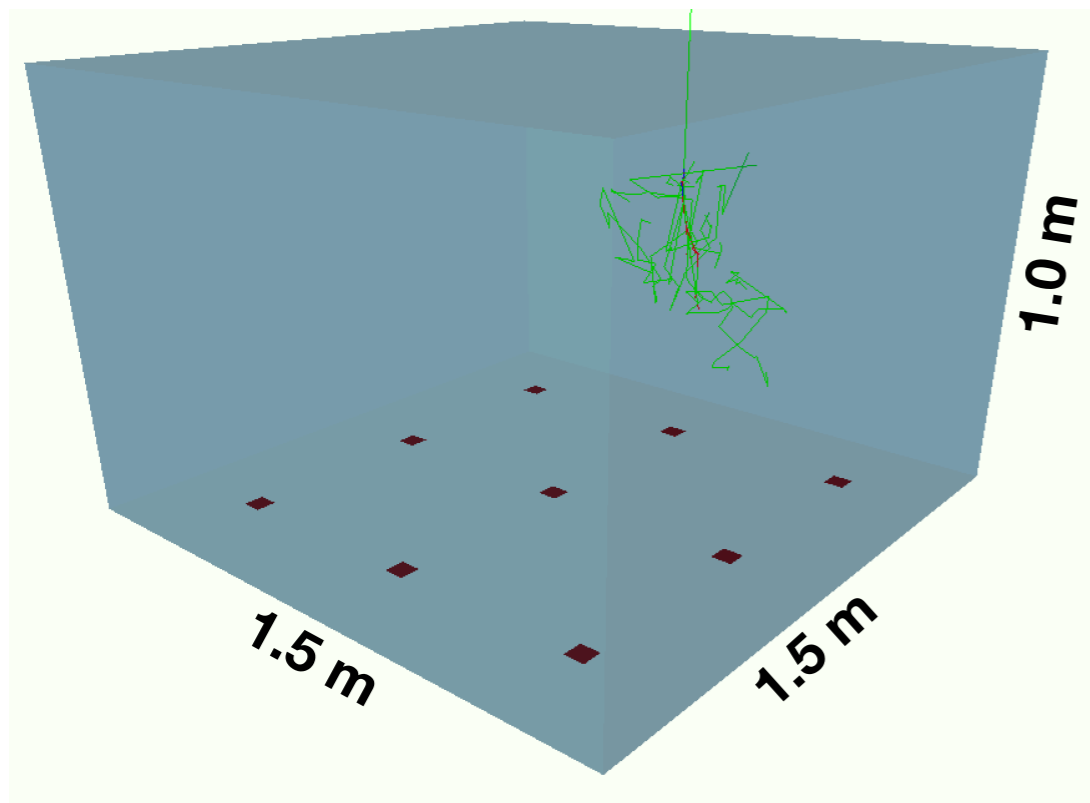
- **Timing**

- Access direct Cherenkov light pulse

- **Gamma/hadron discrimination**

- WCD signal patterns at ground as potential discriminator
- At higher energies muon identification is a powerful discriminant variable
  - Explore timing characteristics to enhance muon id

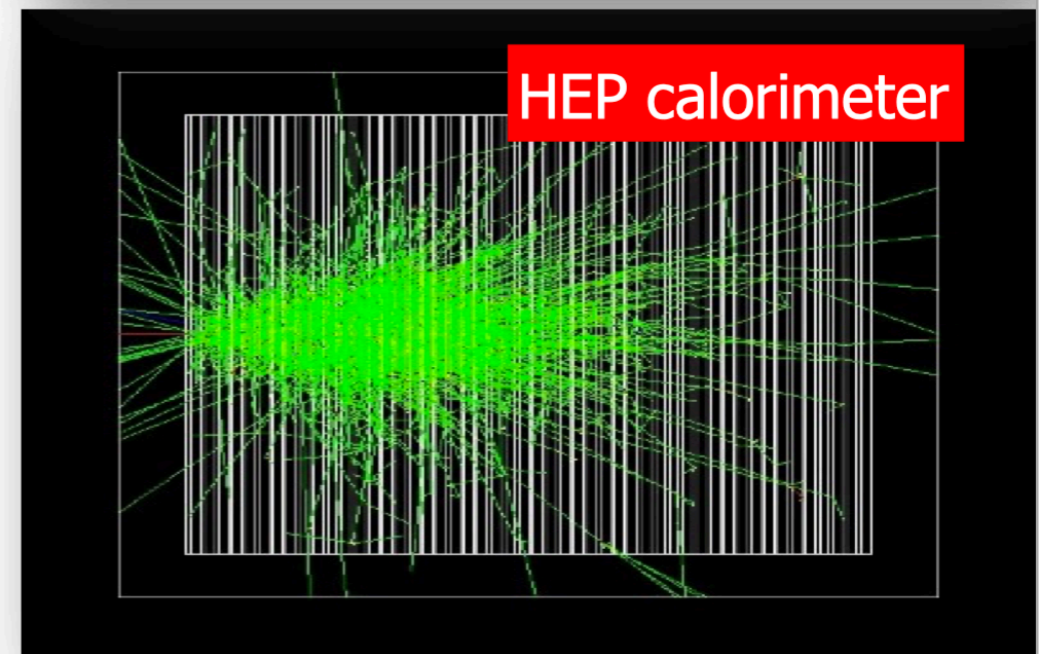
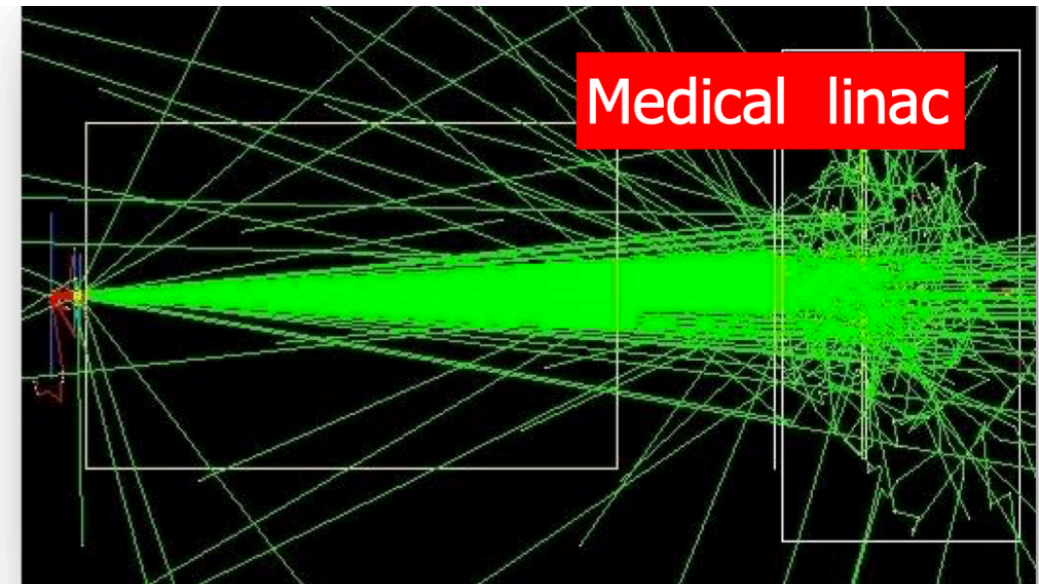
# Station layout



- Water Cherenkov Detector
- 1.5 x 1.5 x 1 m<sup>3</sup>
- 3 x 3 matrix of SiPMs @ WCD bottom
- 3x3cm<sup>2</sup> SiPMs
- Inner walls covered with white diffusing Tyvek

# Geant4 simulation toolkit

- Geant4 provides a wide variety of physics components for use in simulation
  - **Photon processes**
    - ▣  $\gamma$  conversion into  $e^+e^-$  pair
    - ▣ Compton scattering
    - ▣ Photoelectric effect
    - ▣ Rayleigh scattering
    - ▣ *Gamma-nuclear interaction in **hadronic sub-package***
  - **Electron and positron processes**
    - ▣ Ionisation
    - ▣ Coulomb scattering
    - ▣ Bremsstrahlung
    - ▣ Positron annihilation
    - ▣ Production of  $e^+e^-$  pairs
    - ▣ *Nuclear interaction in **hadronic sub-package***
  - **Suitable for HEP & many other Geant4 applications with electron and gamma beams**

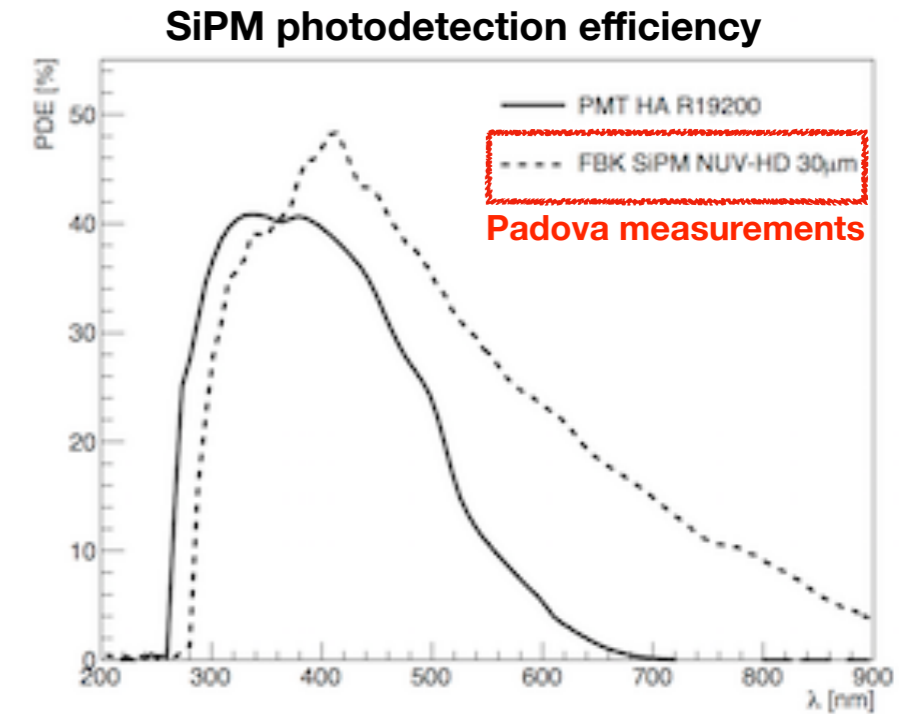


+ equivalent processes for  $\mu$ ; + hadronic processes, ...



# WCD simulation using Geant4

- Explore Geant4 functionalities to simulate **optical photon production and propagation** :
  - Cherenkov emission, refraction and reflection at boundaries, bulk absorption, Rayleigh scattering, etc.
- **$\lambda$  dependence** of all the relevant optical processes and material properties also implemented :
  - transmission efficiencies, dielectric constants, surface properties, photodetection efficiency,...



# WCD simulation using Geant4

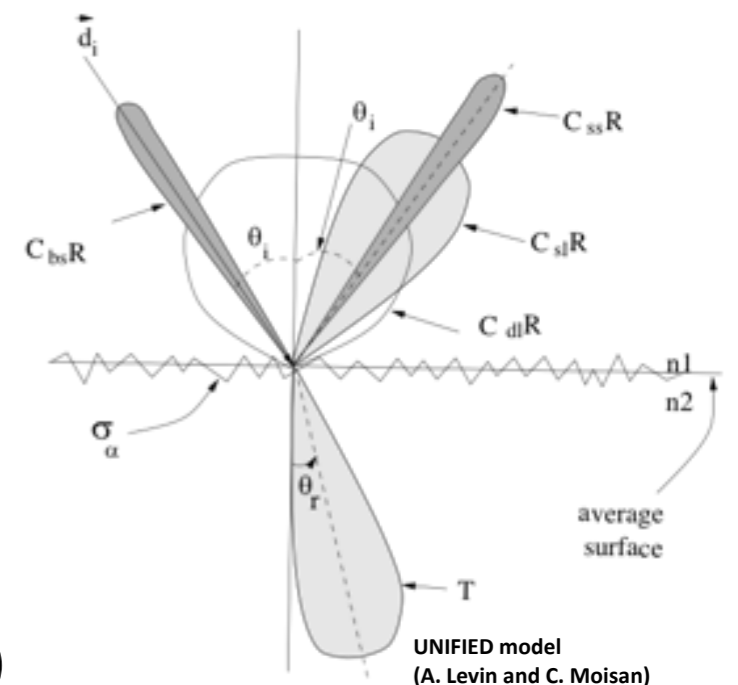
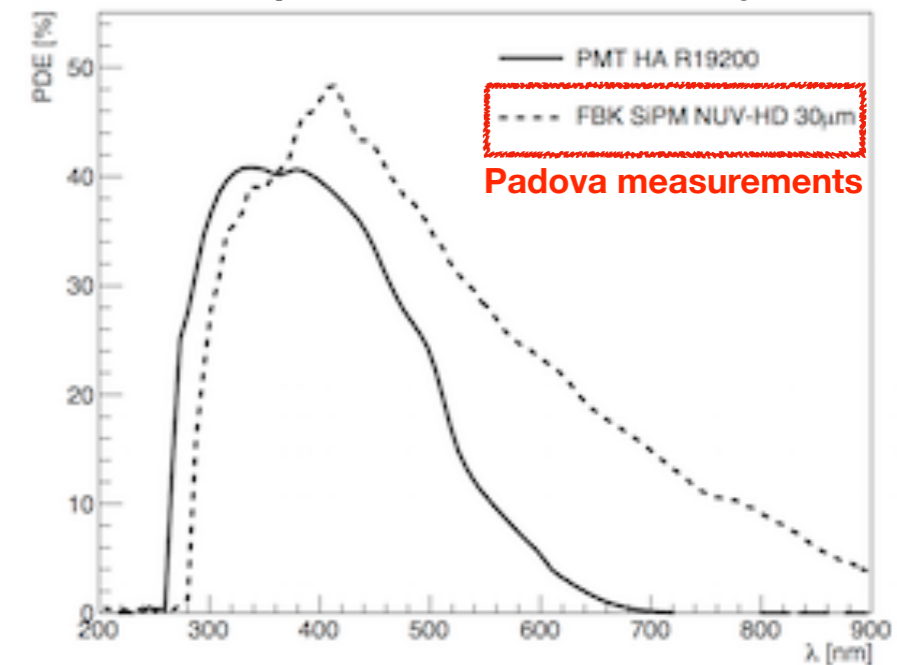
- Explore Geant4 functionalities to simulate **optical photon production and propagation** :
  - Cherenkov emission, refraction and reflection at boundaries, bulk absorption, Rayleigh scattering, etc.
- **$\lambda$  dependence** of all the relevant optical processes and material properties implemented :
  - transmission efficiencies, dielectric constants, surface properties, photodetection efficiency,...

- **Tyvek**

- Described using the **G4 UNIFIED optical model**;
- Specular and diffusive properties;
- $R \sim 95\%$ , for  $\lambda > 450$  nm
  - 80% of which is diffusively reflected;
  - 20% is reflected around the specular reflection direction, with  $\sigma_\alpha \sim 0.2^\circ$

- Tyvek properties “inspired” in Auger simulation parameters :-)

SiPM photodetection efficiency

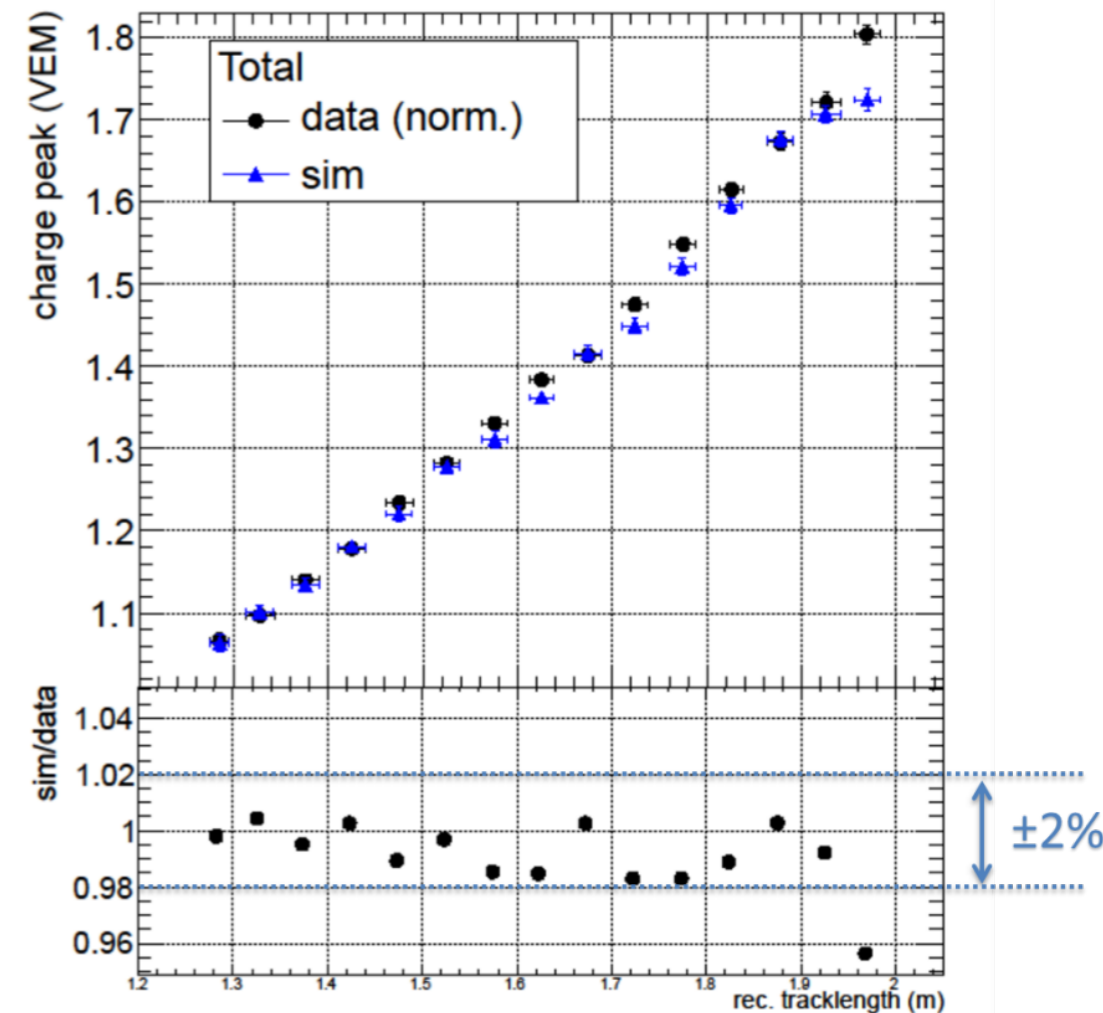
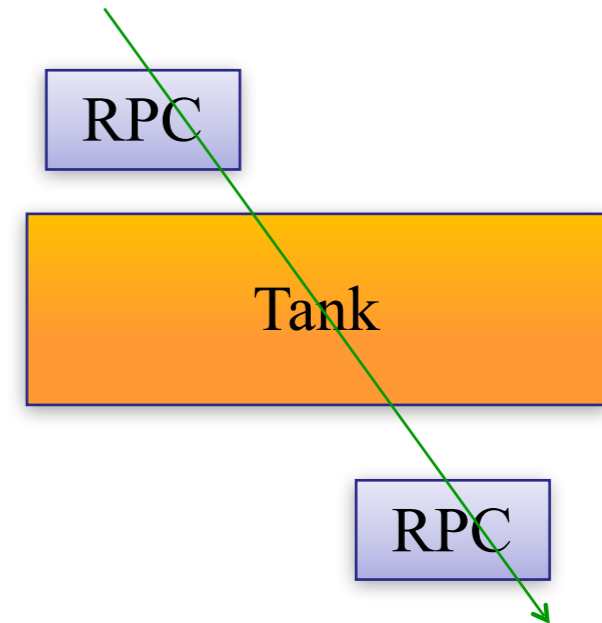


# RPC based $\mu$ -hodoscope in a test WCD at Auger Observatory



$\mu$

Data/Simulation agreement at % level !

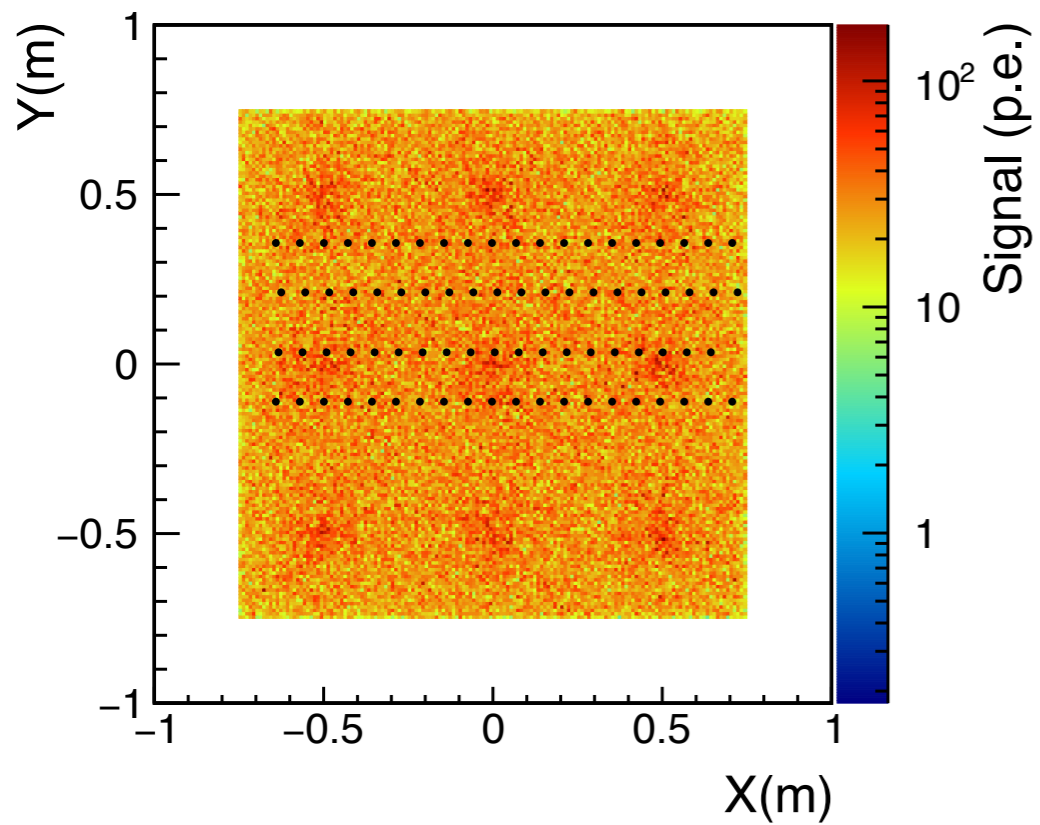


# First simulation studies

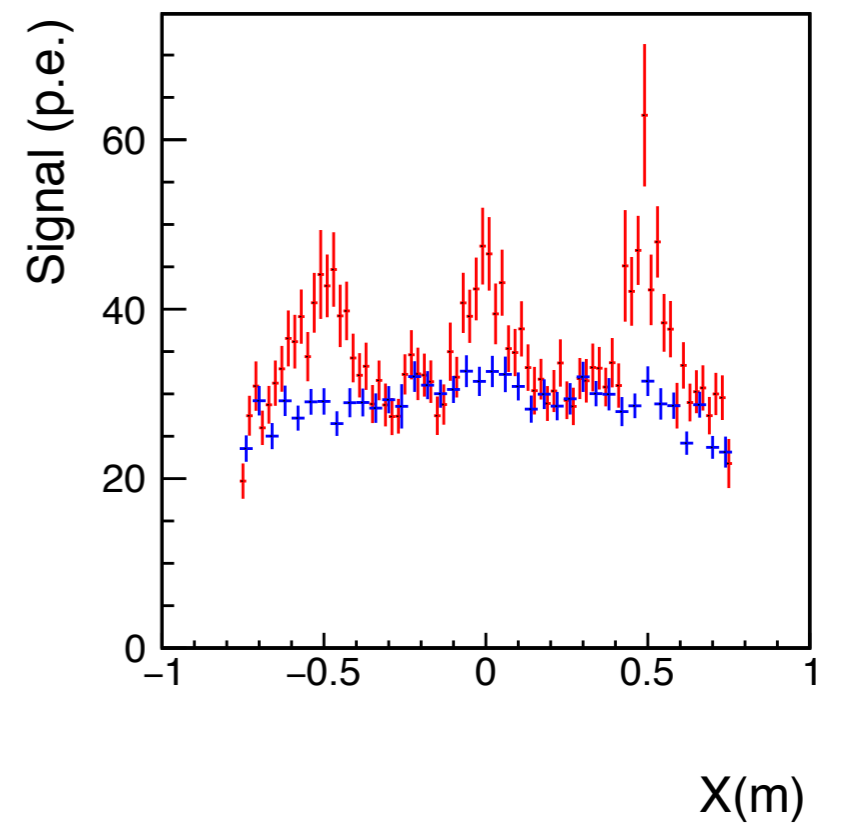
- **Standalone Geant4**
  - Single particles uniformly injected at top surface
  - Photons with fixed energy and with shower energy spectra
- Shower simulations

# Signal size and uniformity

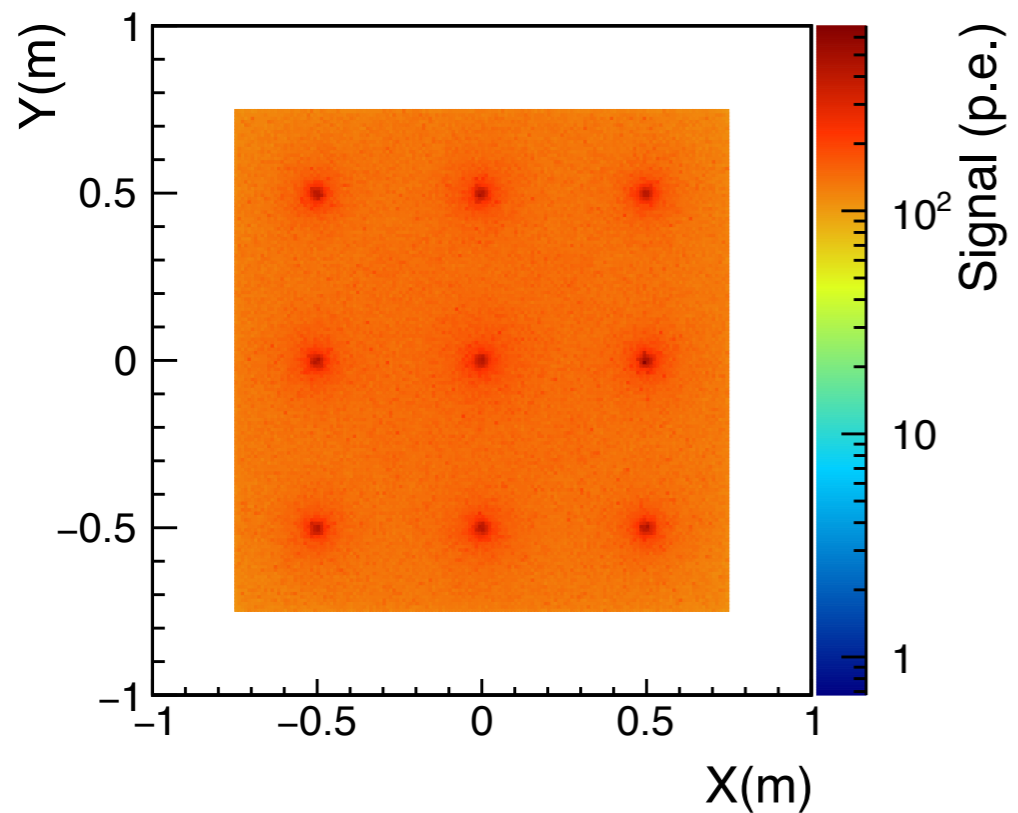
Vertical incidence



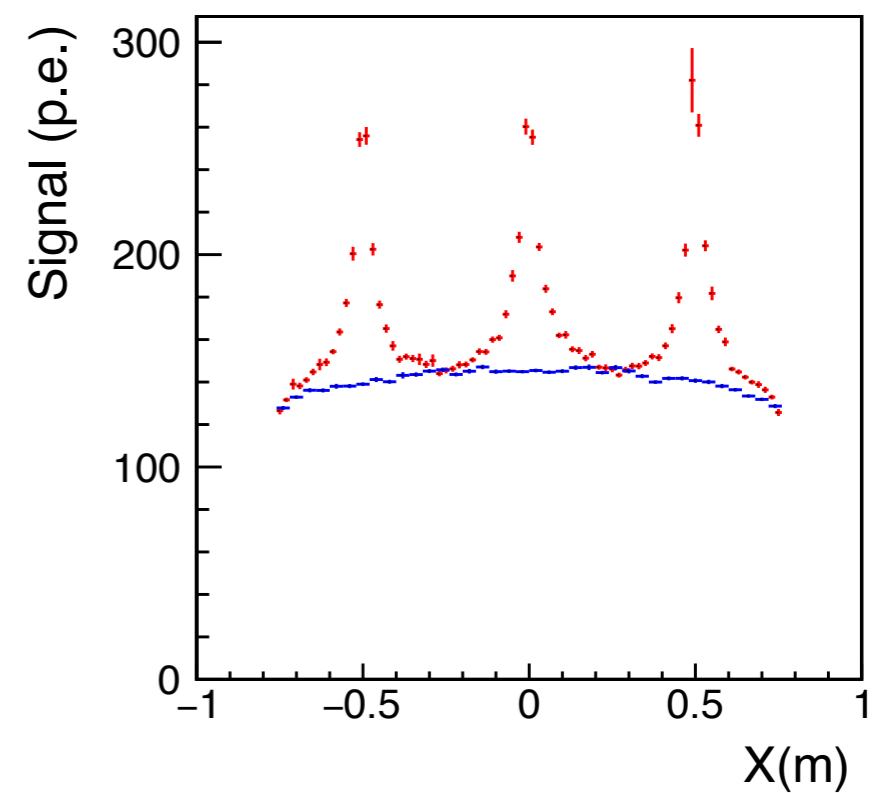
Photons



Using gamma spectrum in  $r = [0m, 10m]$

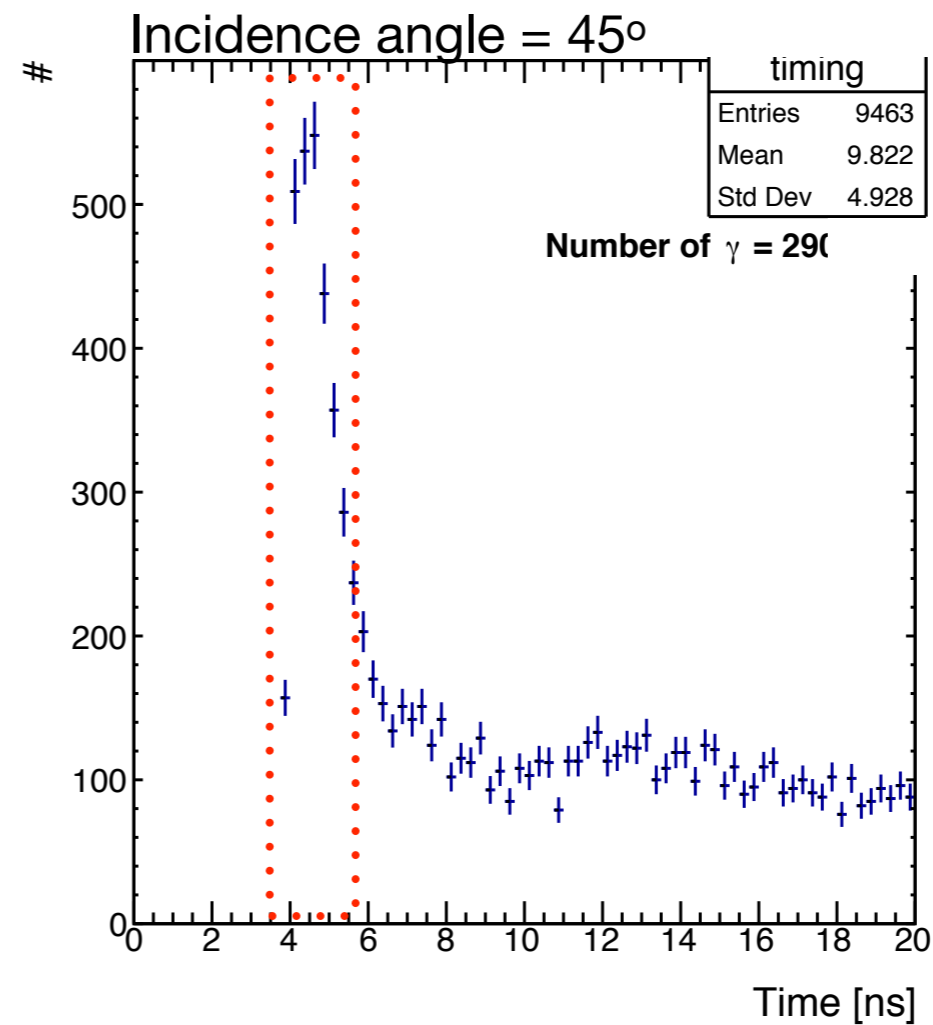
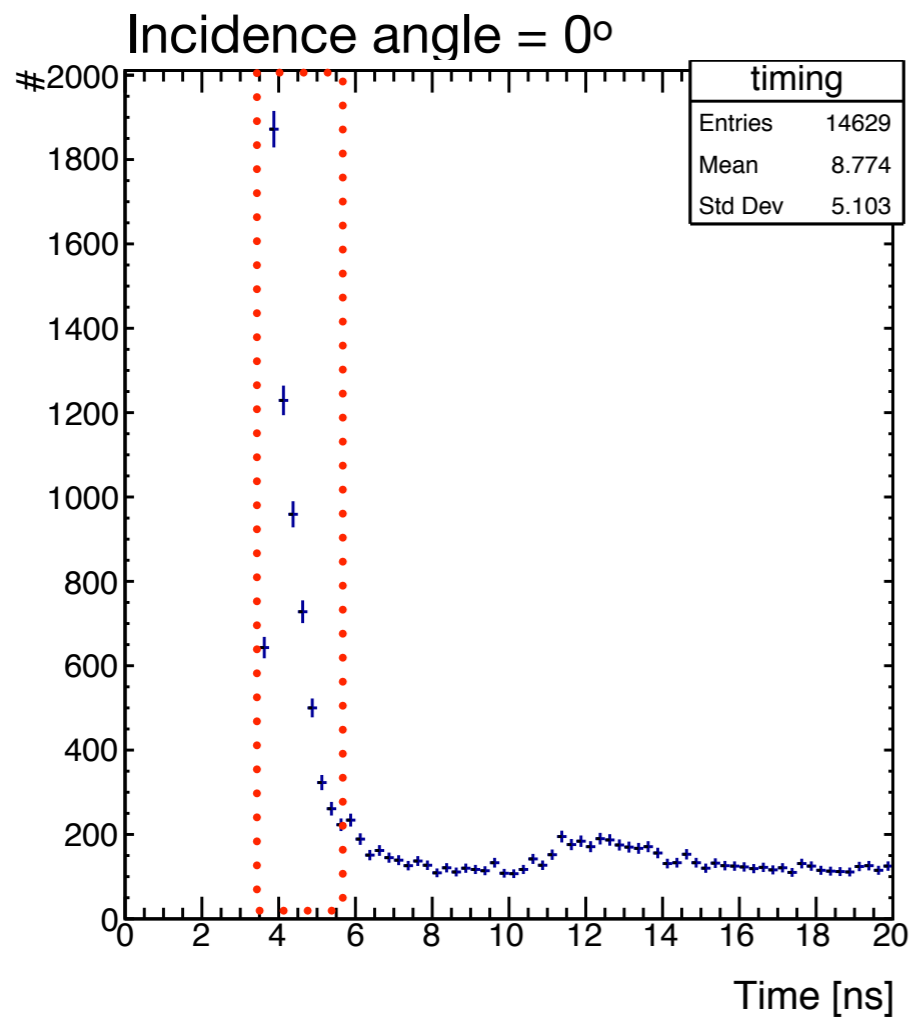
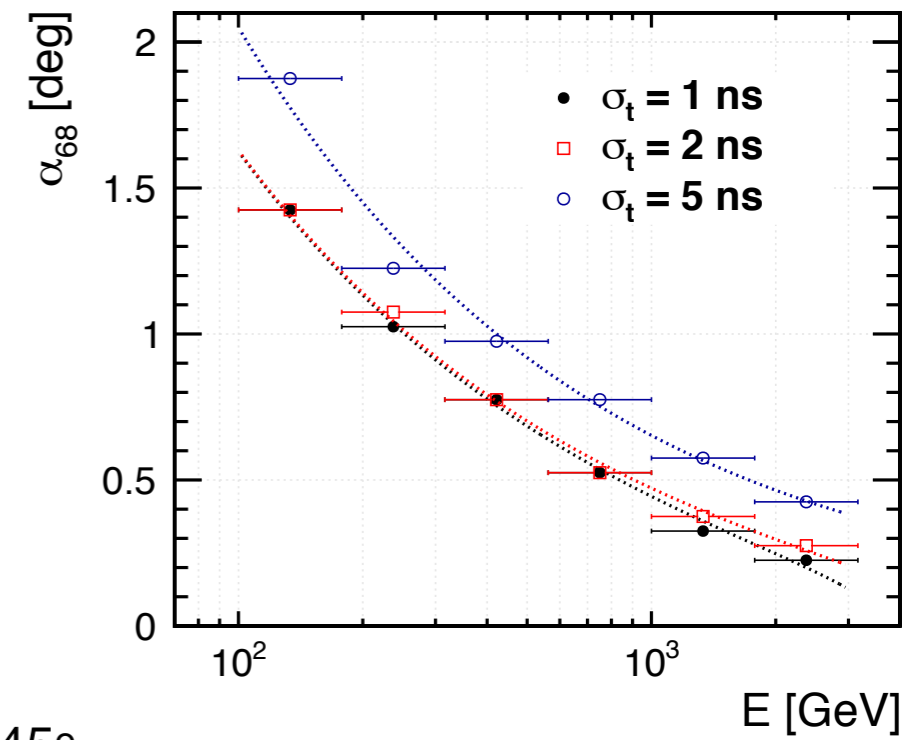


Muons



# Station time trace

Photons with  $E = 15 \text{ MeV}$



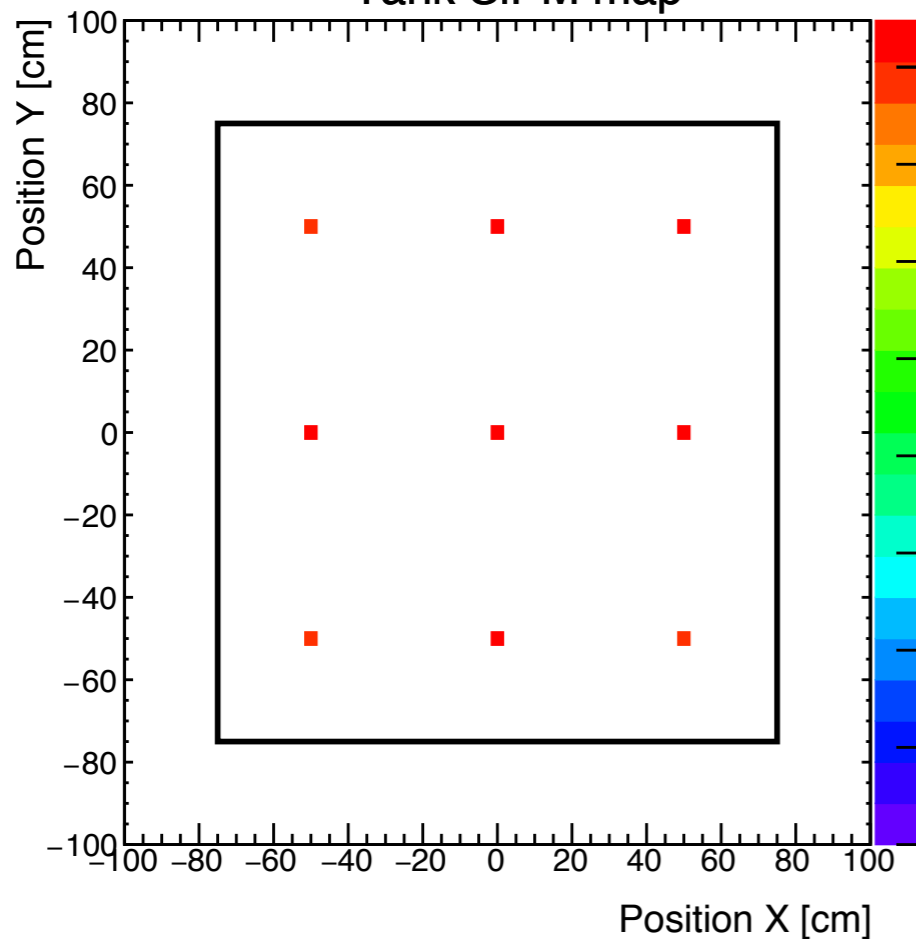
The fast, direct light pulse, arrives within 2 ns !

# SiPM signal vs angle

Photons with  $E = 15$  MeV

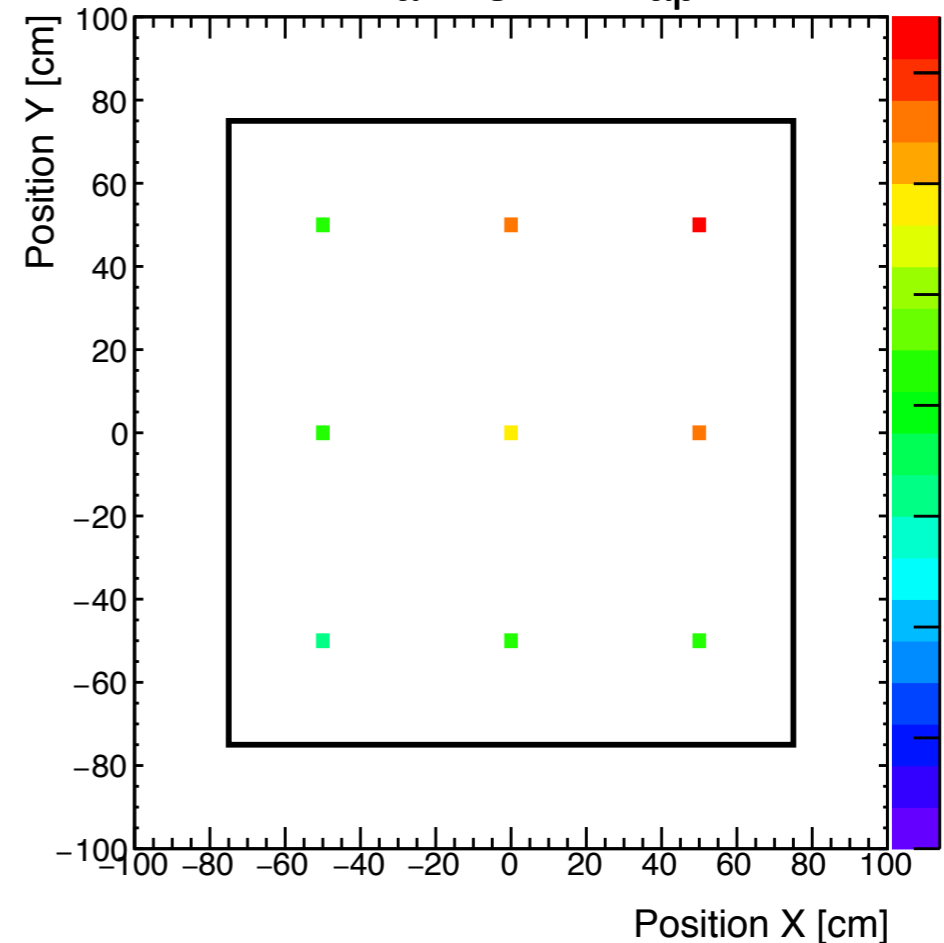
Incidence angle =  $0^\circ$

Tank SiPM map



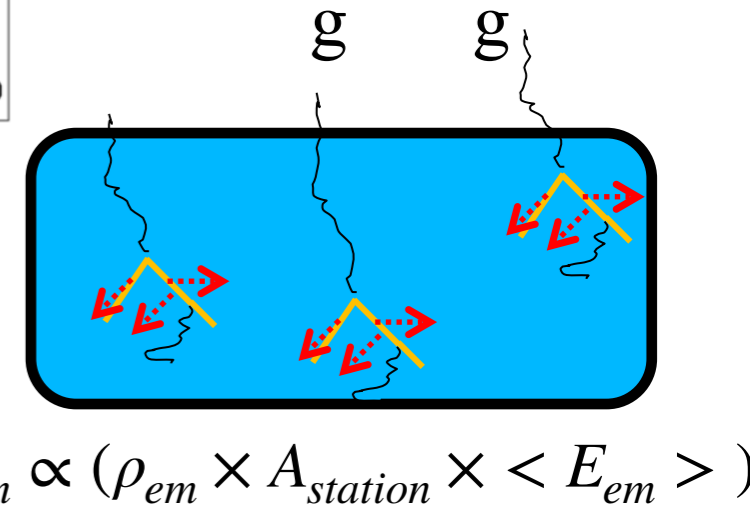
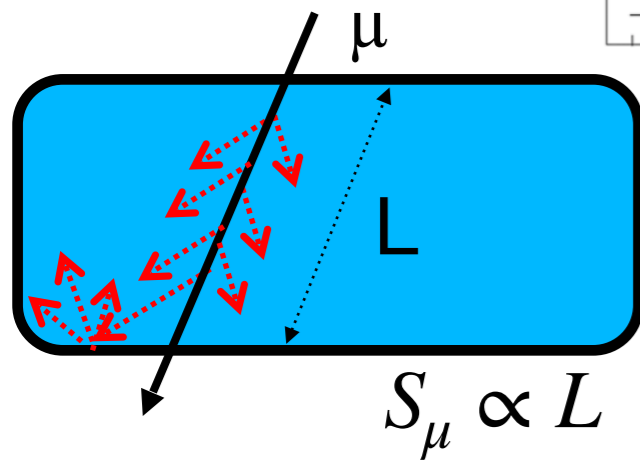
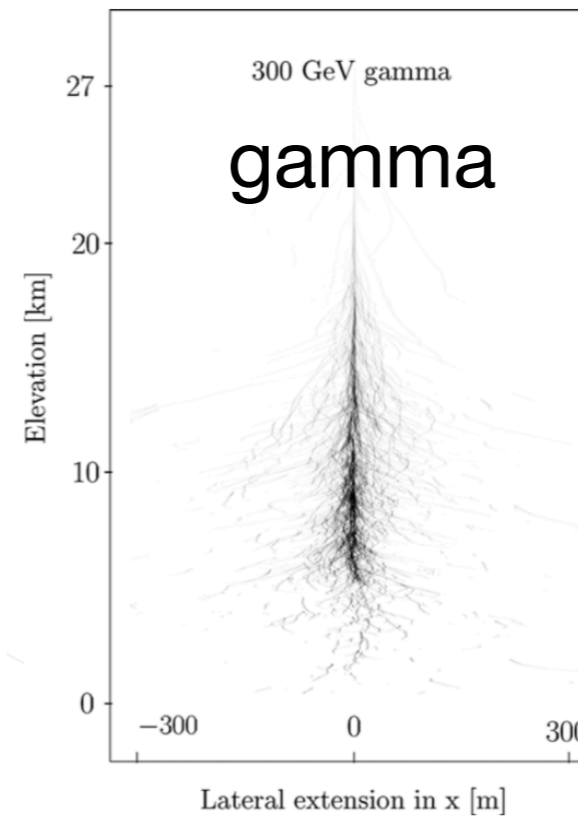
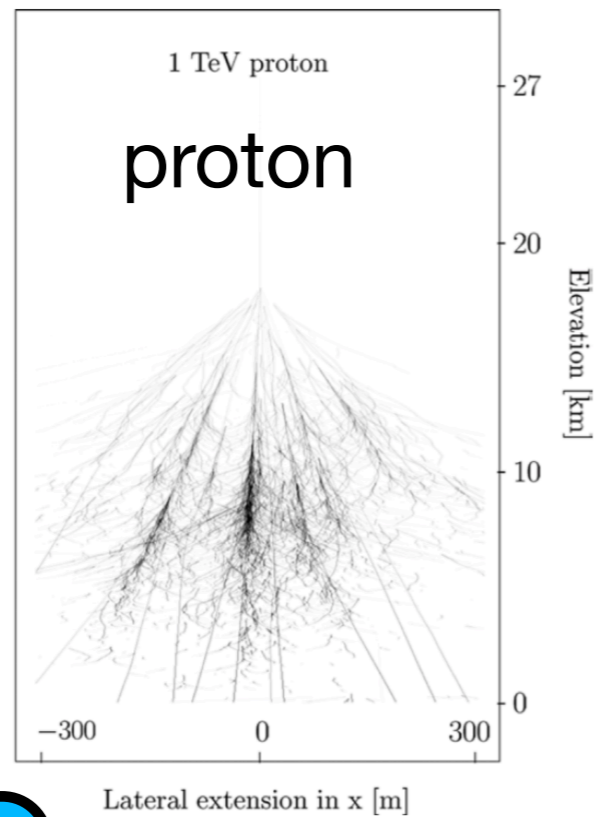
Incidence angle =  $45^\circ$

Tank SiPM map



SiPM signal asymmetry gives additional information; to be explored ?

# $\gamma/h$ separation through muon id



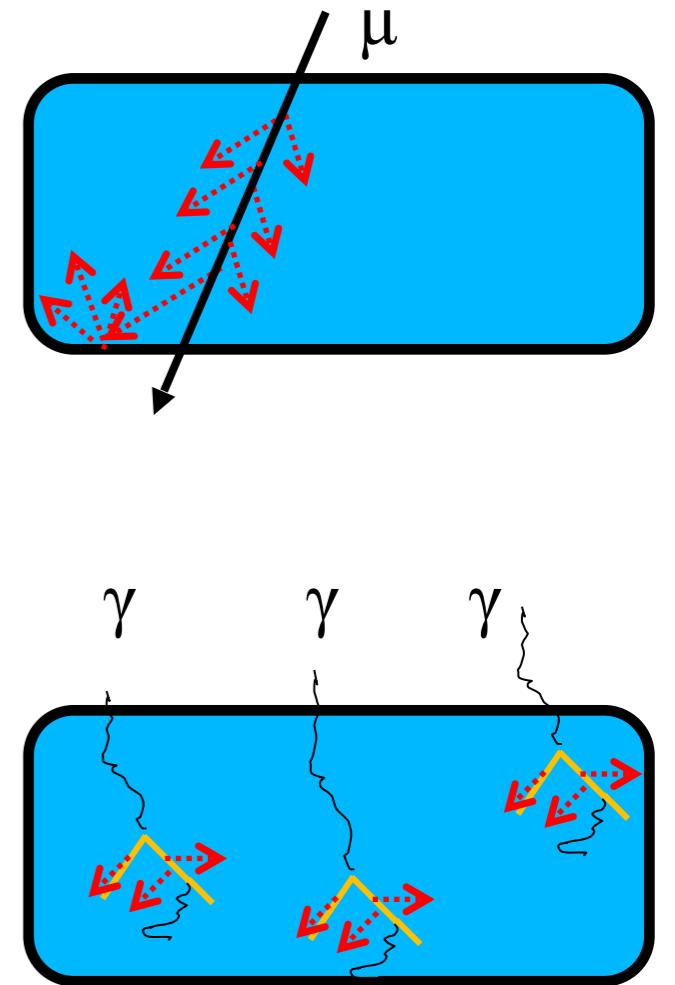
- At high energies muon counting is a powerful gamma/hadron discriminator;
- Muon id is usually based on its large Cherenkov signal in the WCD;
- Increasing  $S_\mu/S_{em}$  :
  - Increase muon track length  $L$ ; reduce station area  $A_{station}$
  - Consider only stations at large distances to shower core :  $\rho_{em}$  and  $\langle E_{em} \rangle$  decrease with  $r_{Core}$



# Muon id in a compact WCD

Exploring the timing and topological characteristics :

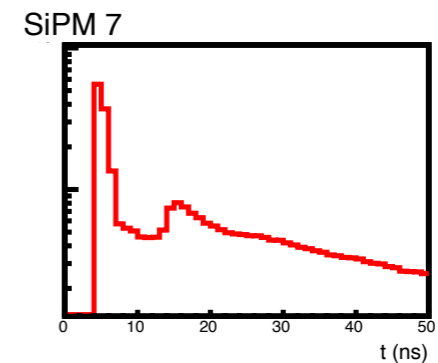
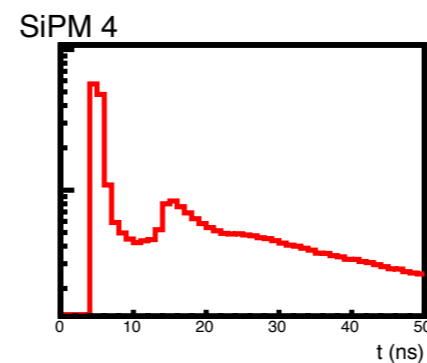
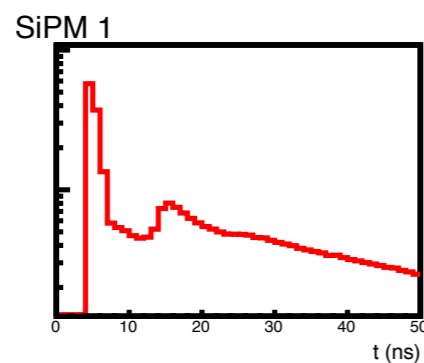
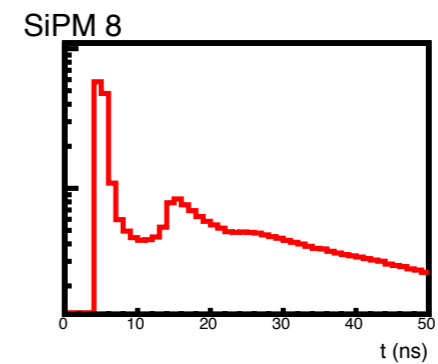
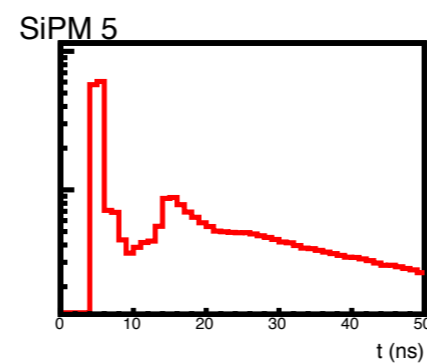
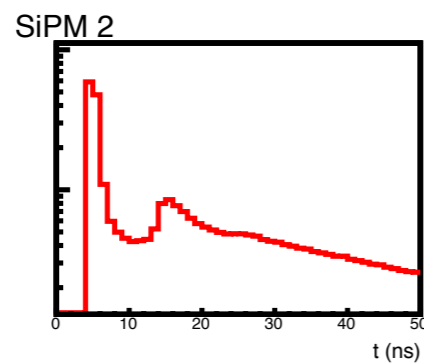
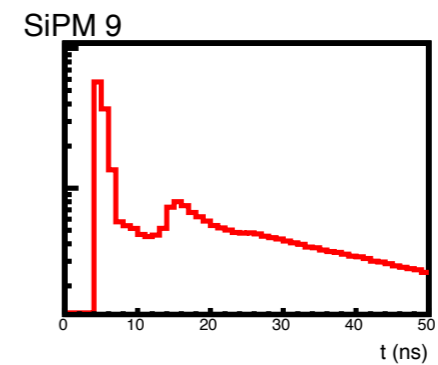
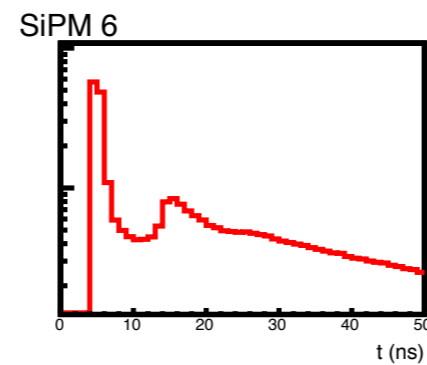
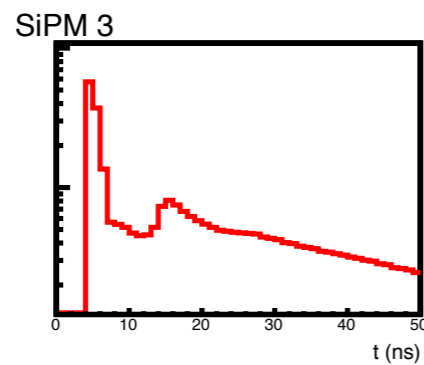
- Fast direct Cherenkov light pulse of a **single muon** is seen mainly in only a part of the readout matrix
- The spreaded signal of **several photons/electrons** entering at random positions in the WCD is seen across the whole readout matrix



# First tests ...

- For  $r > 40$  m :  $S_{\mu} \sim 25 \times S_{\gamma}$
- Each “photon event” consists of 25 photons injected uniformly in the WCD;
- Total signal  $\sim$  signal from a single muon

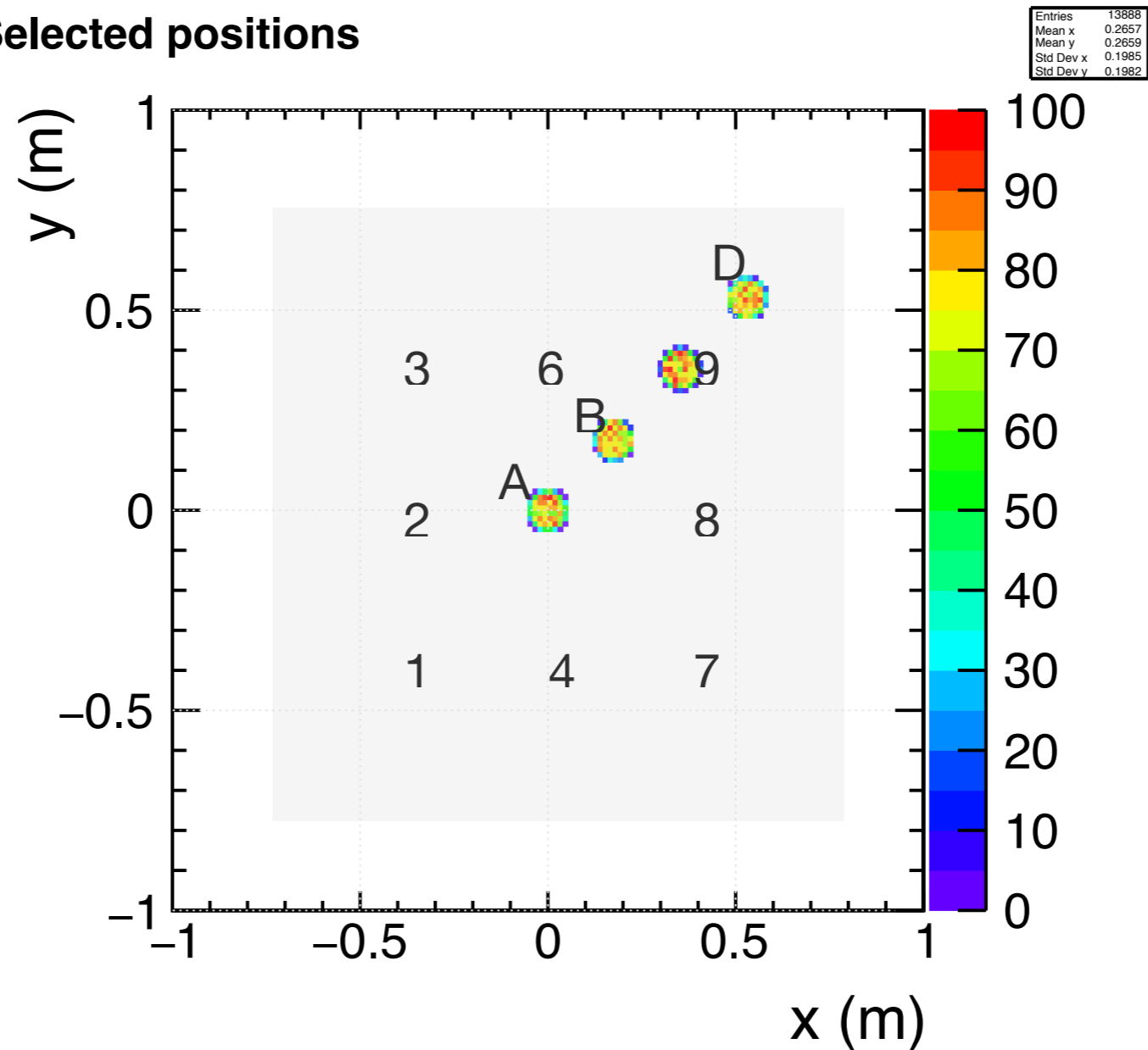
Average traces for photons



# First tests ...

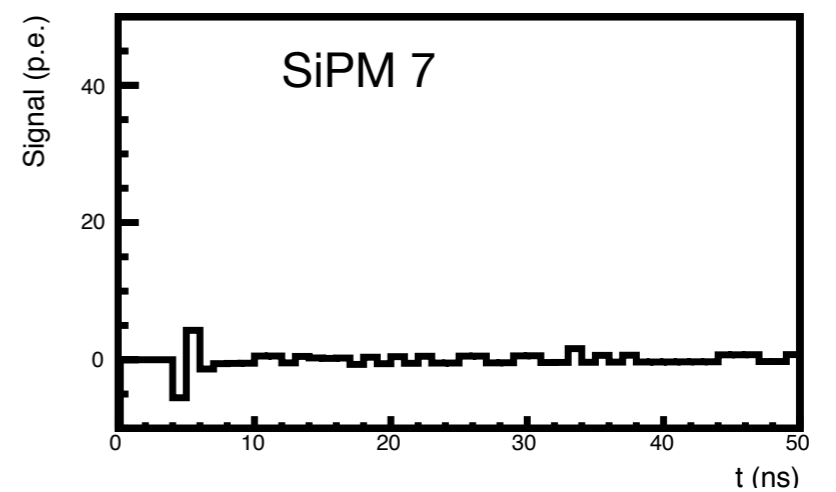
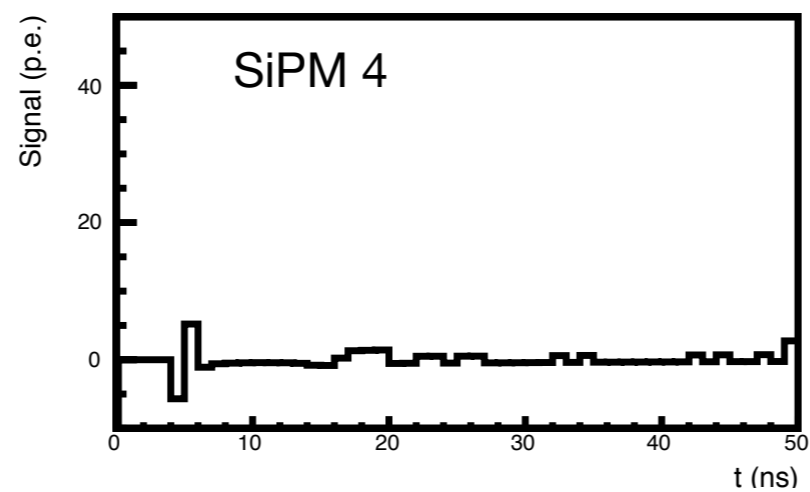
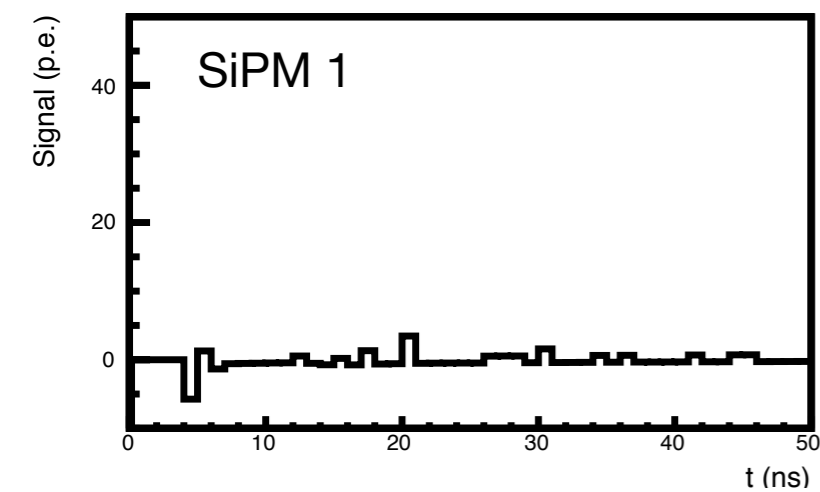
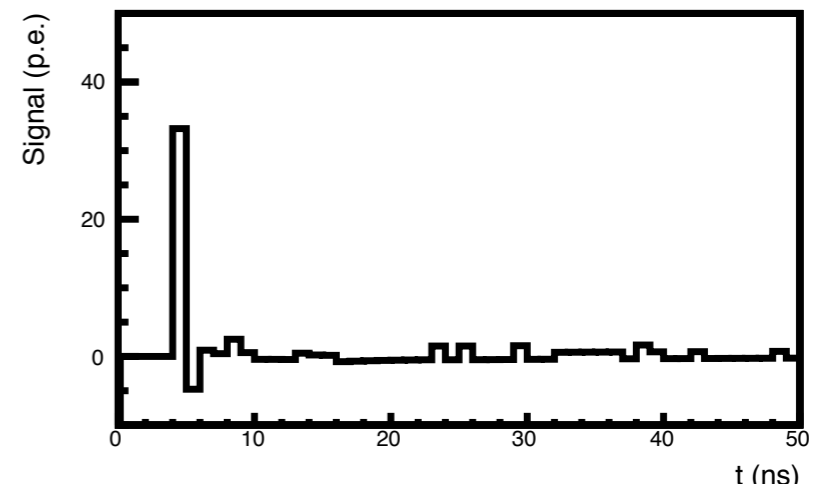
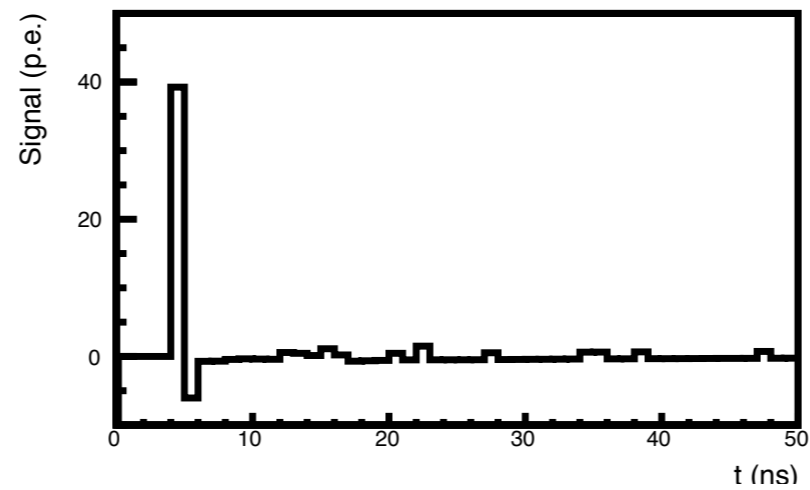
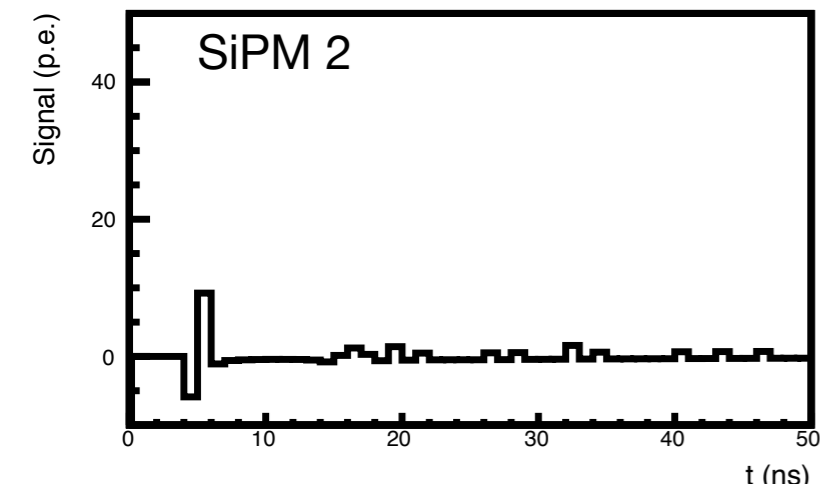
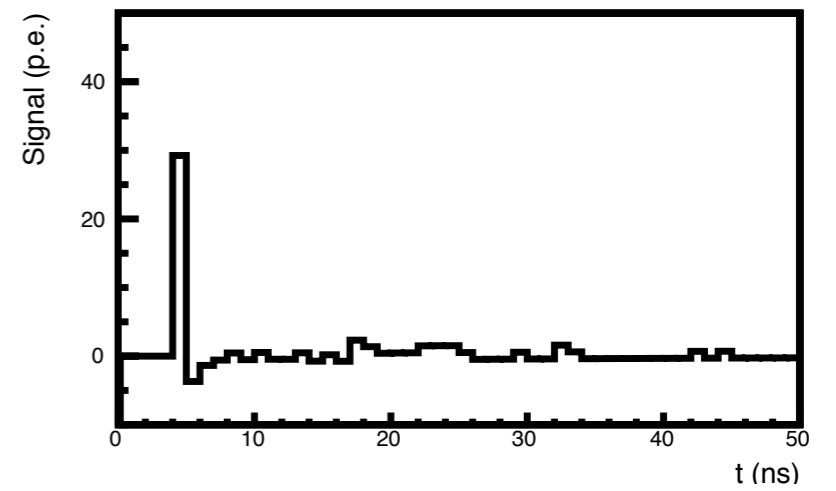
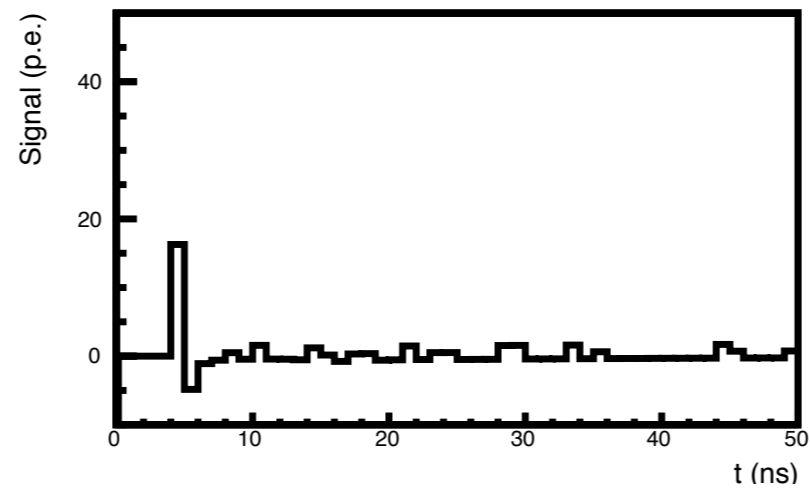
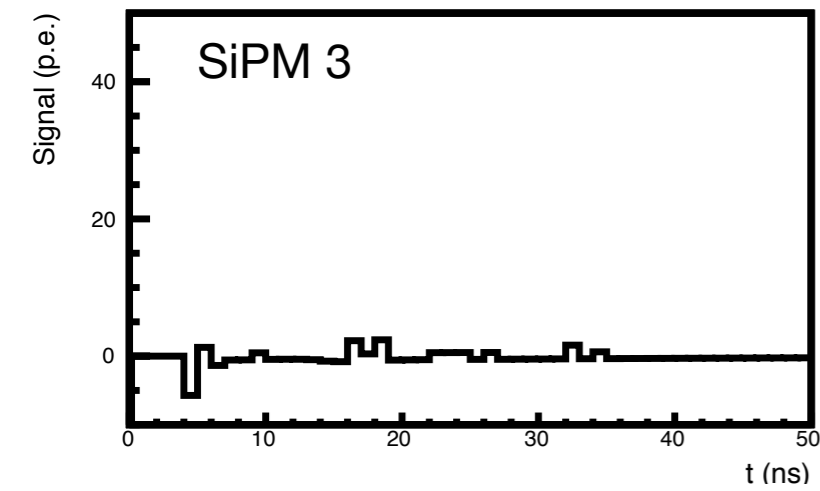
Compute *event by event differences of traces*: **muon - <gamma25>**

Selected positions

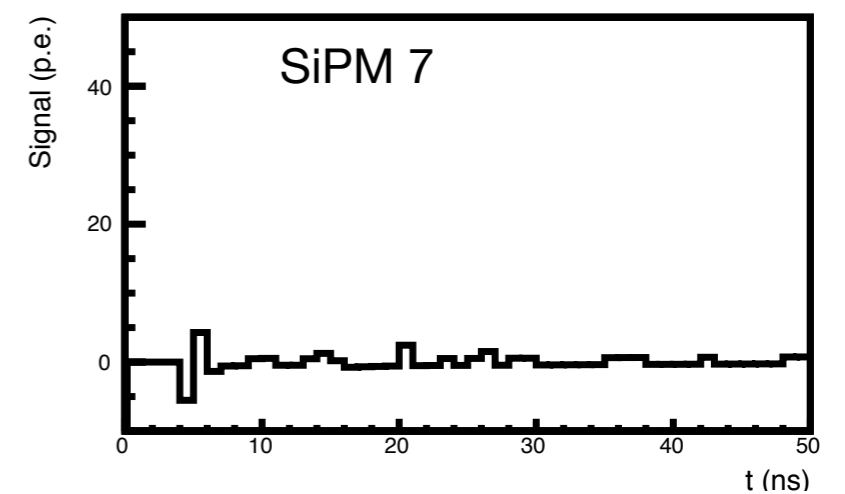
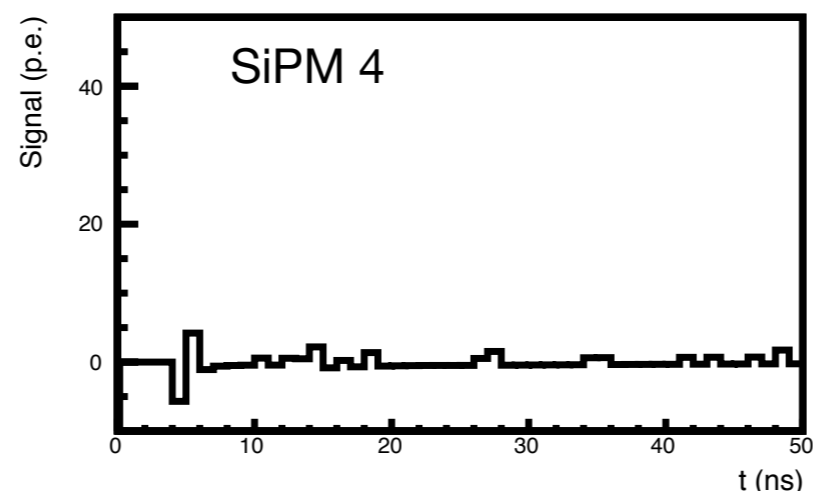
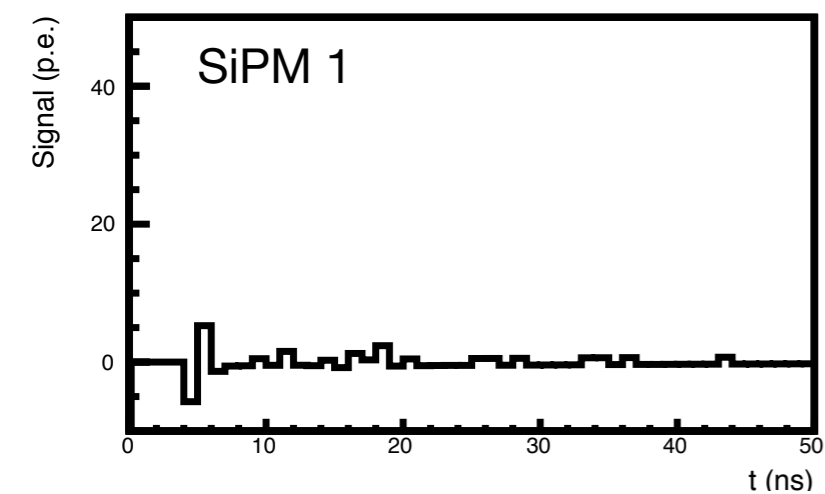
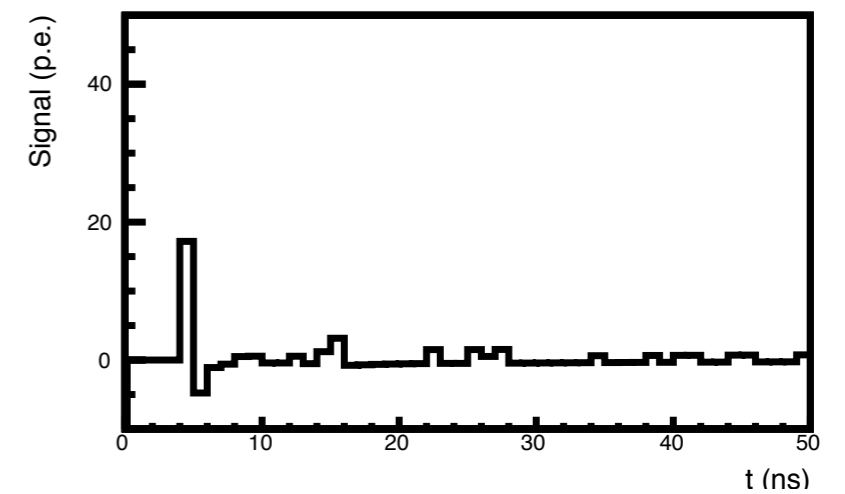
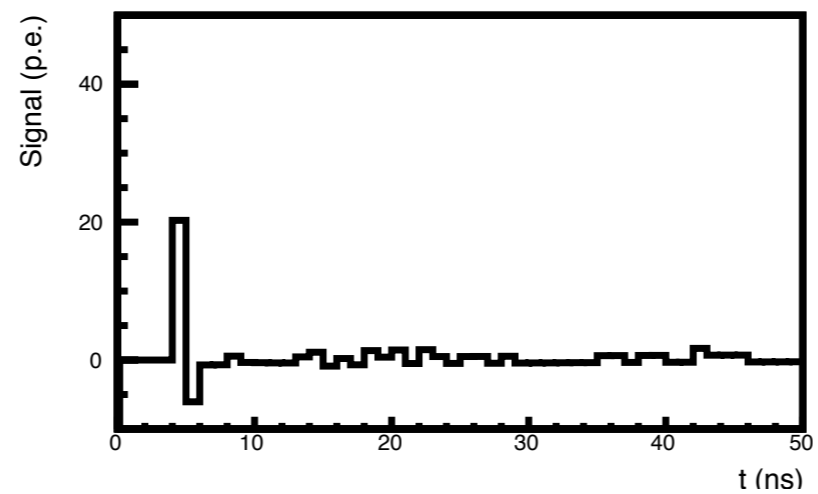
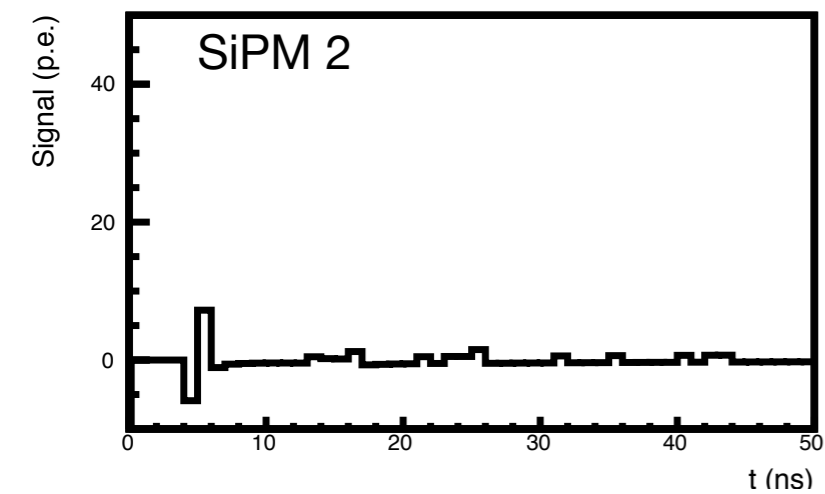
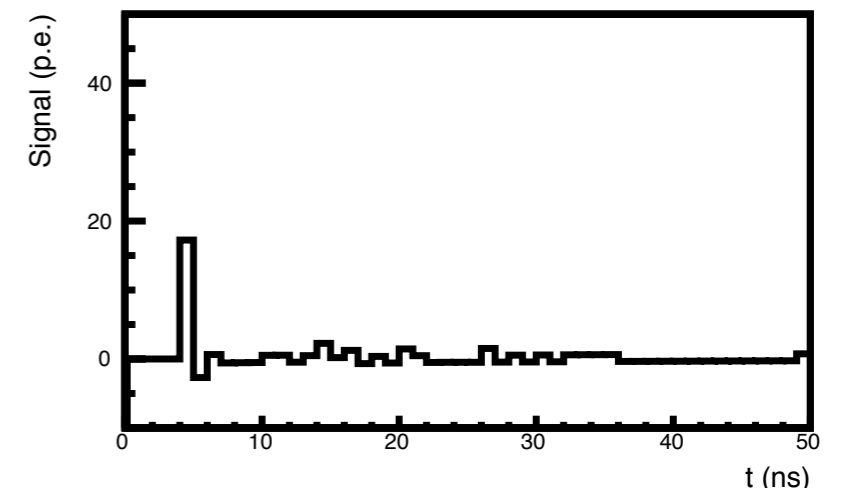
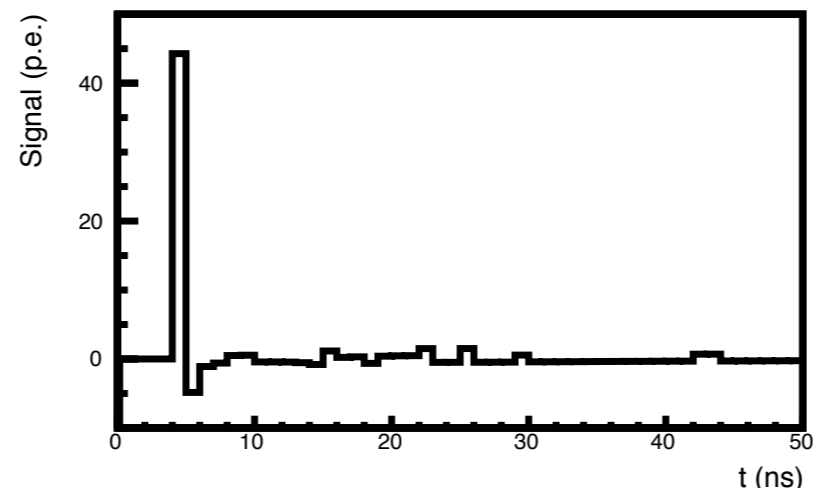
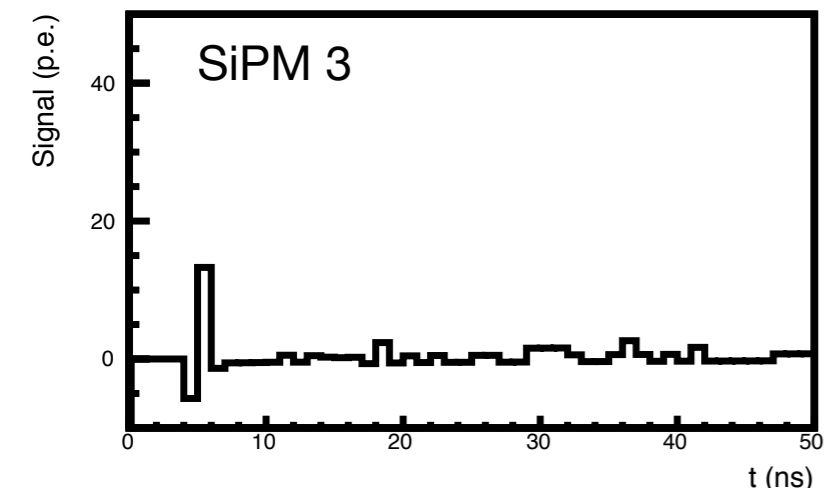


Single muons @ position B →

# Differences of traces : muon - $\langle \text{gamma25} \rangle$

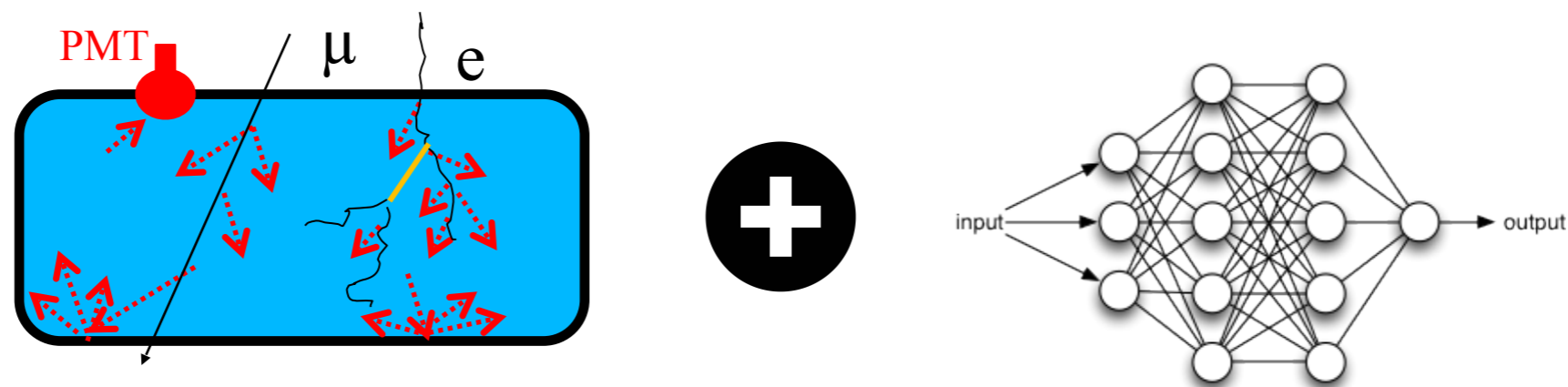


# Differences of traces : muon - $\langle \text{gamma25} \rangle$



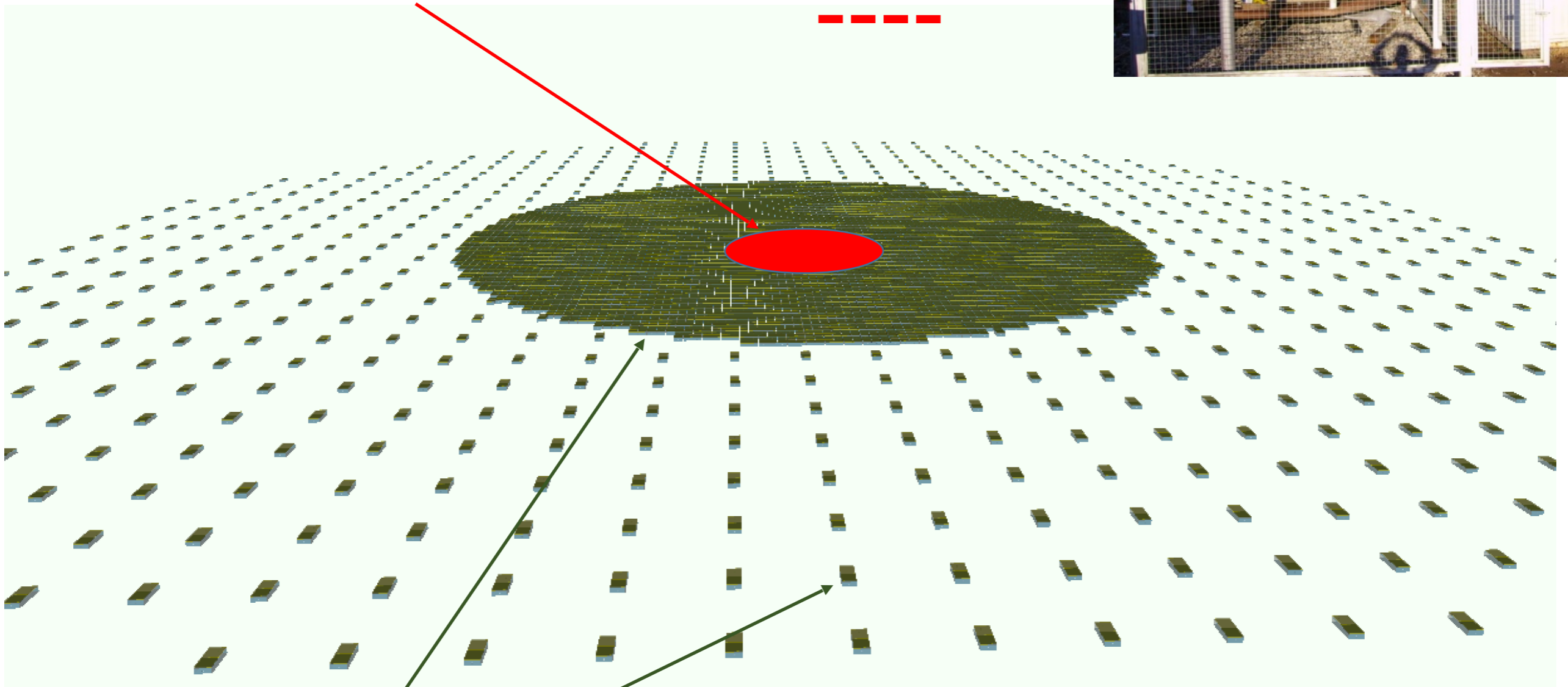
# Towards Muon id with a compact WCD

- Preliminary, single-station, assessment quite promising
- Starting studies with shower events
  - Explore ANN techniques
  - Granada group starting to look into simulations
- Stations close to the shower core could also be used ? ...



# New layout

“ $\mu$  Telescopes” – WCDs+ RPCs



WCDs



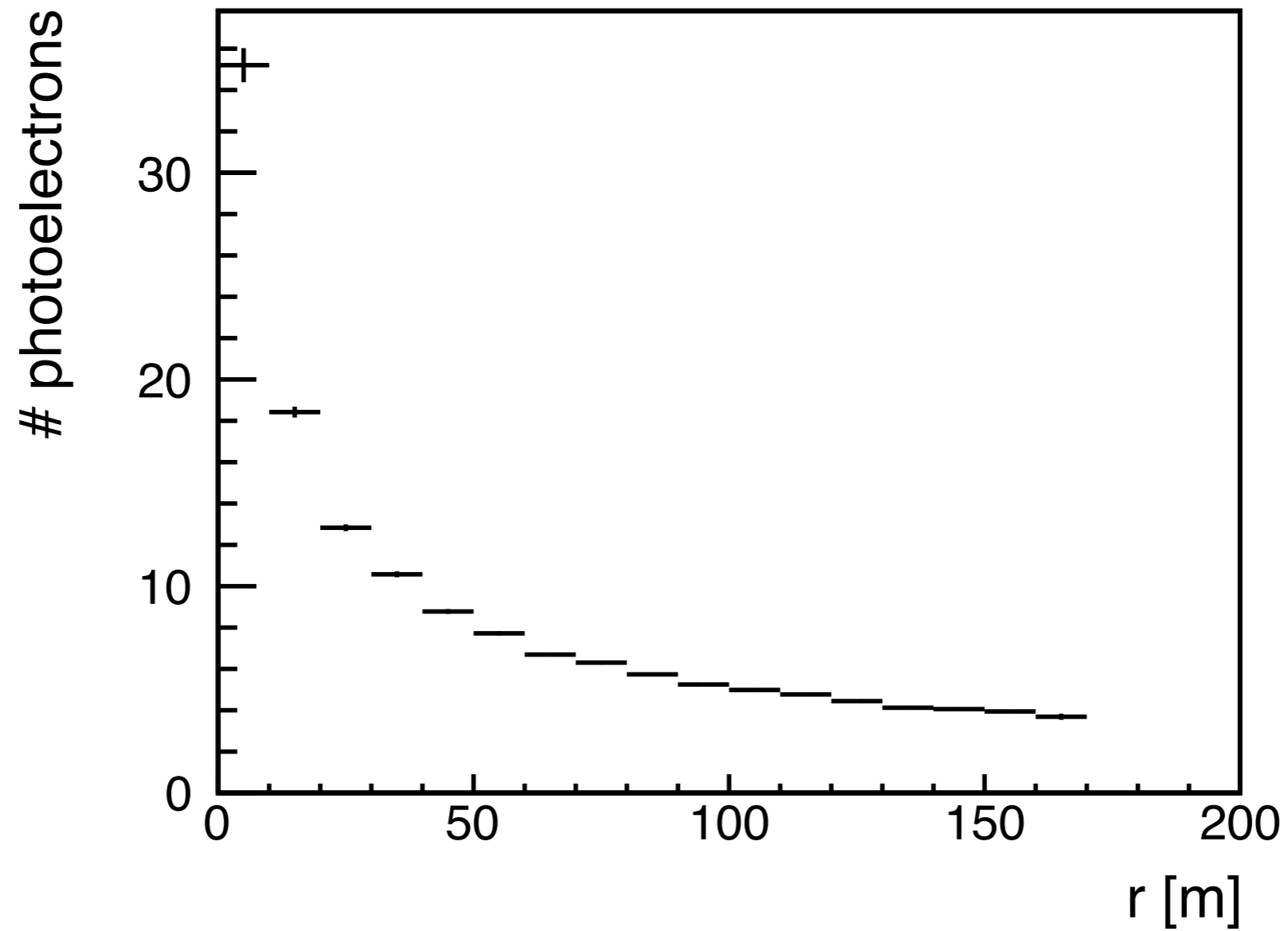
- Train and validate ANN techniques with data
- Study + monitor + calibrate WCD response

# First simulation studies

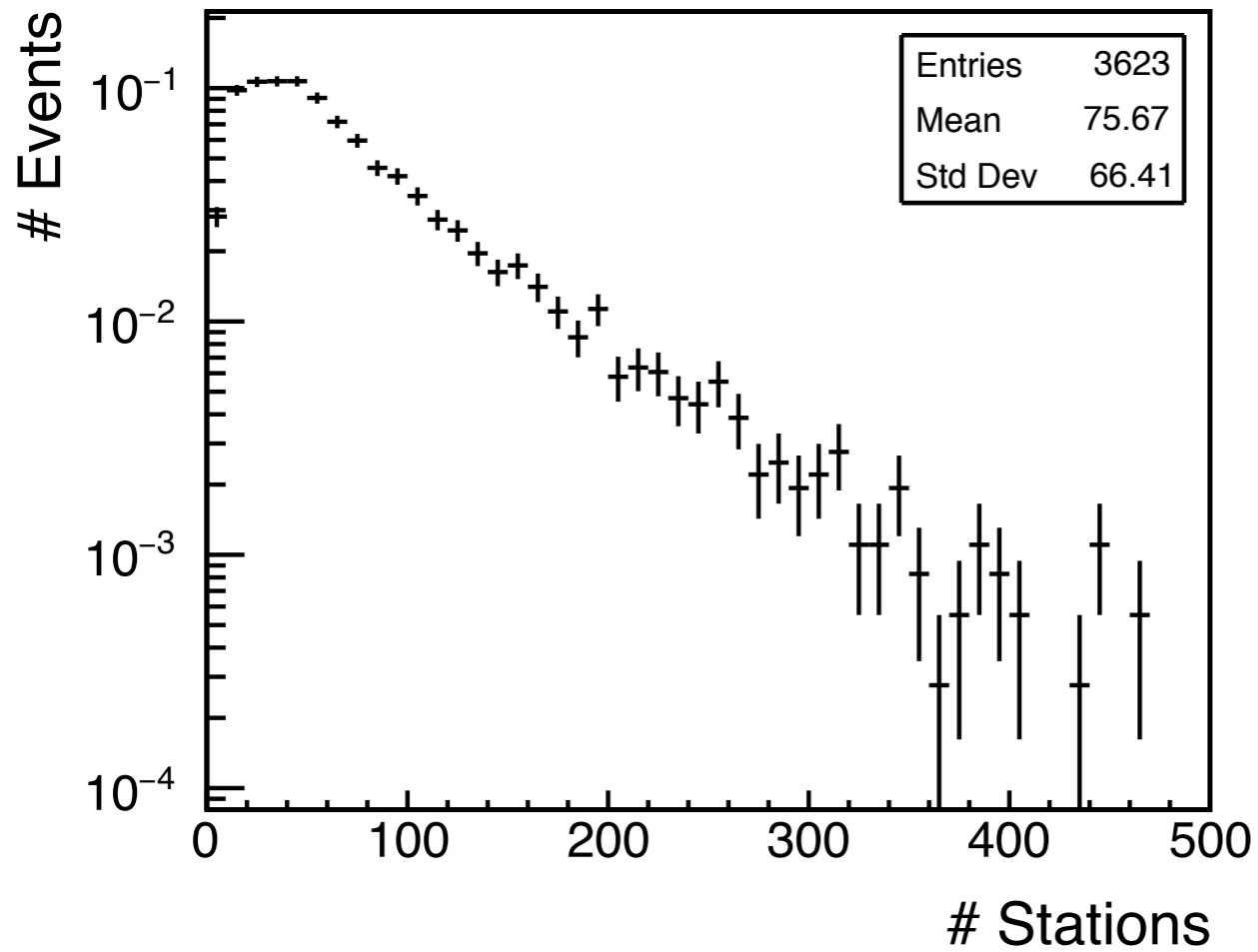
- Standalone Geant4
  - Single particles uniformly injected at top surface
  - Photons @ fixed E and with shower energy spectra
- **Shower simulations**
  - Circular array with  $R = 160$  m ( $80000$  m<sup>2</sup>) @ 5200 m altitude
  - Corsika showers @  $\text{Log}E$  in  $[2.0, 2.2]$  ;  $\theta = 10^\circ$



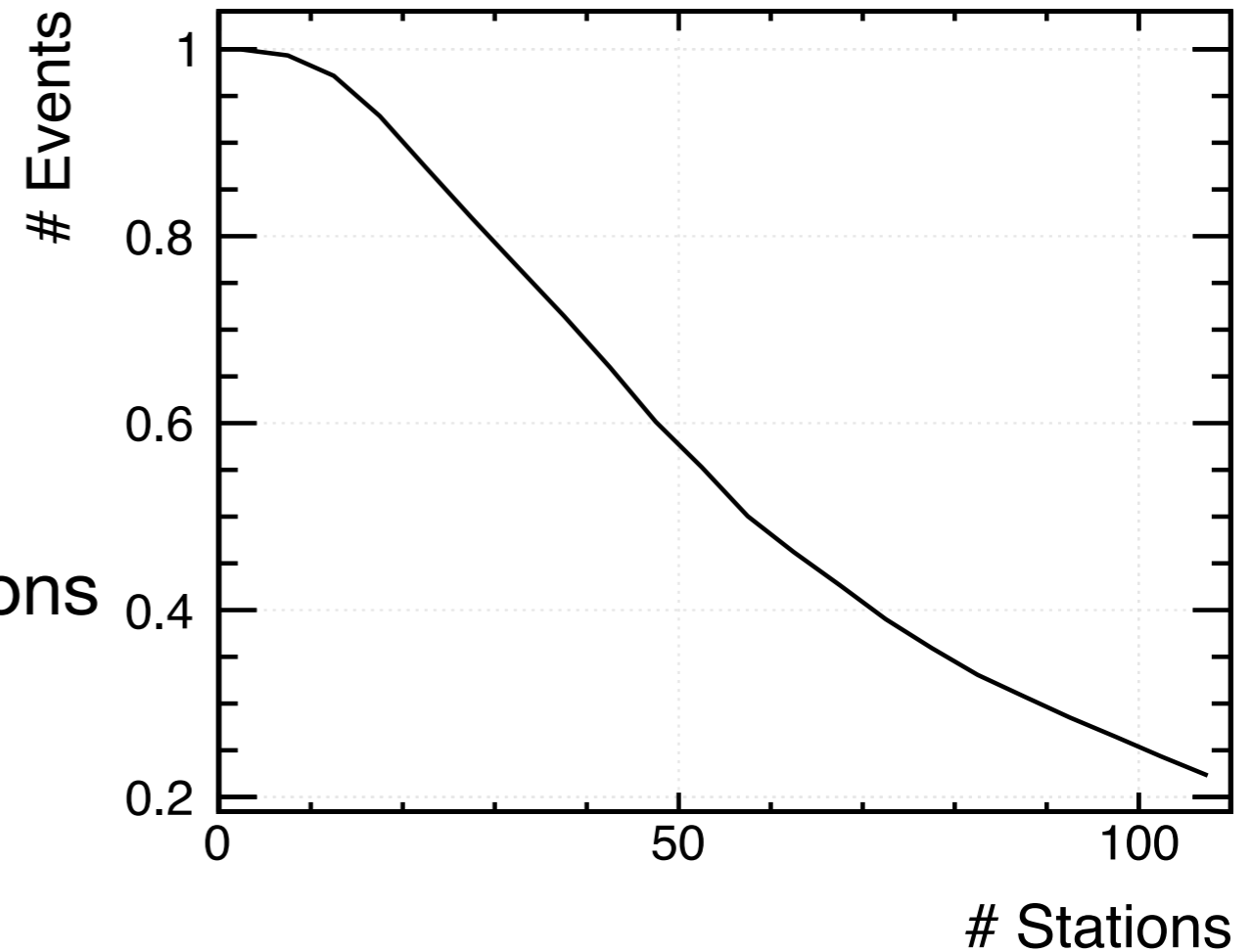
# Station signal mean LDF @ 100 GeV



# Stations with detected signal

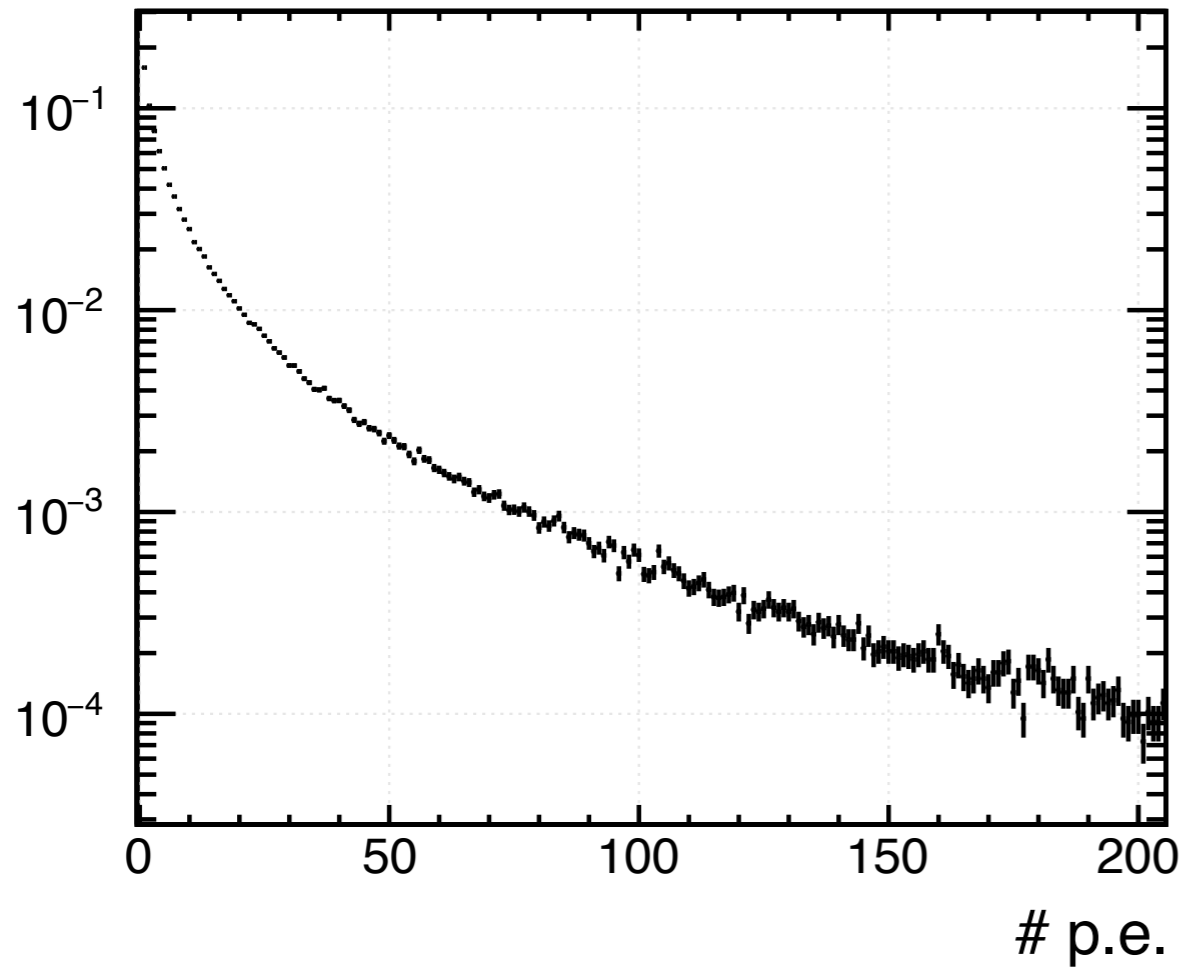


@ 100 GeV

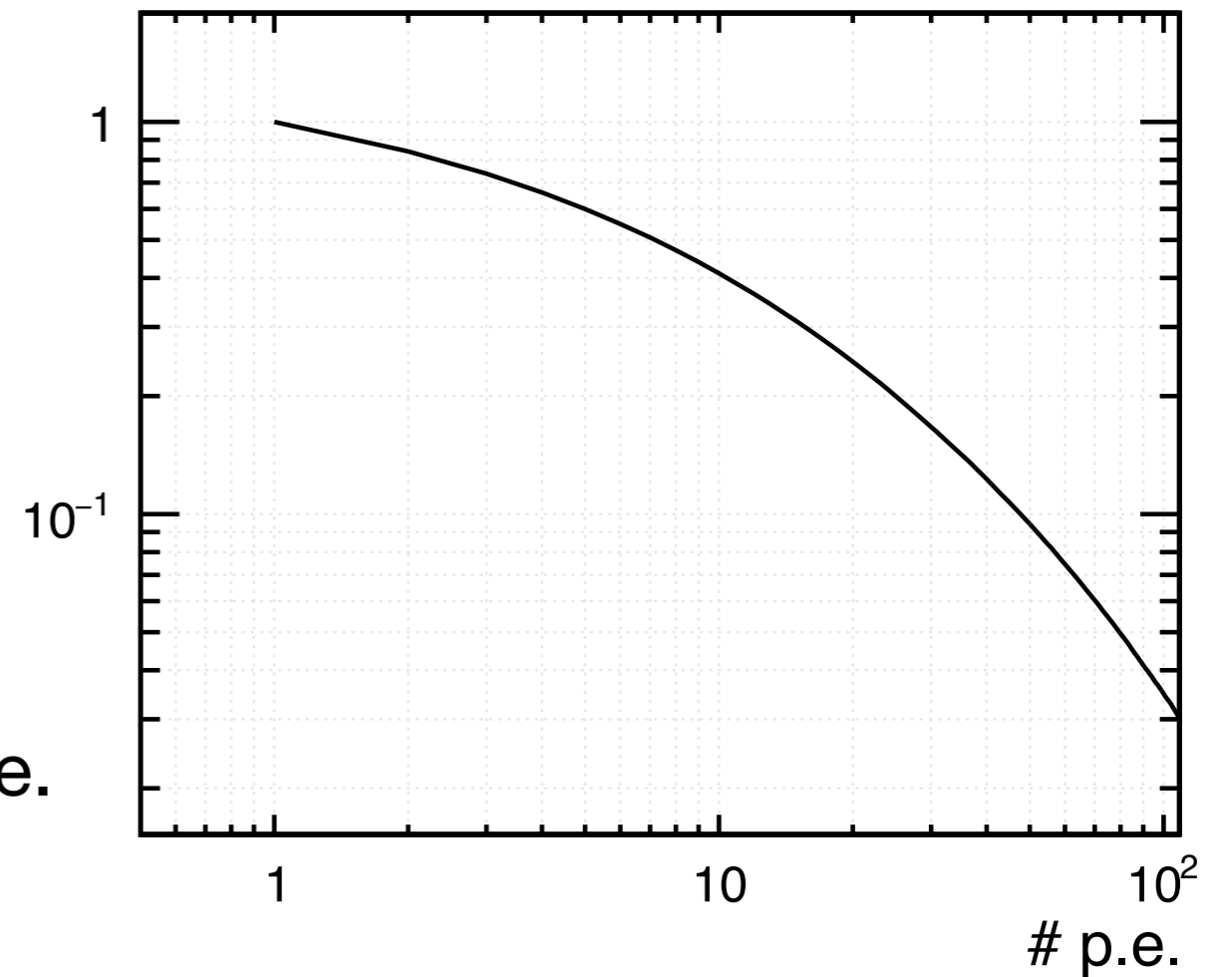


90% of the events have  $> 20$  active stations

# Station signal

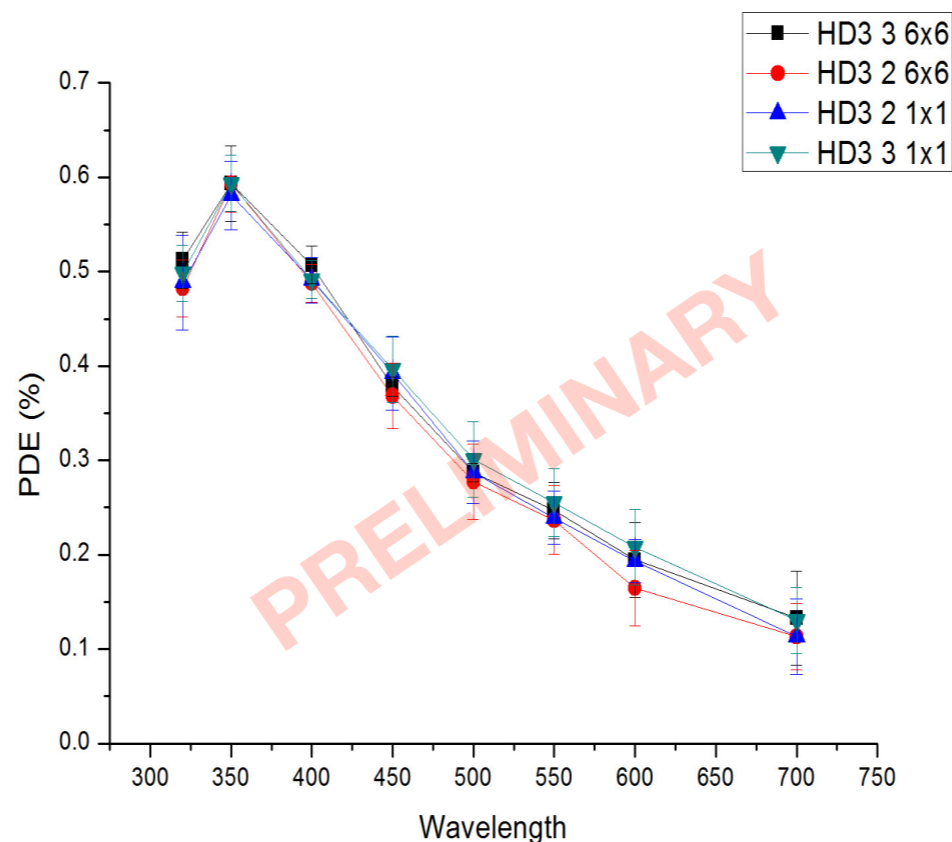


@ 100 GeV



~ 40% of the active stations have > 10 p.e.

- All simulations based on an old SiPM PDE curve;
- Recent measurements by the Padova group on new SiPMs show better performance, namely towards small wavelenghts;
- Simulations to be updated with new PDE curve...

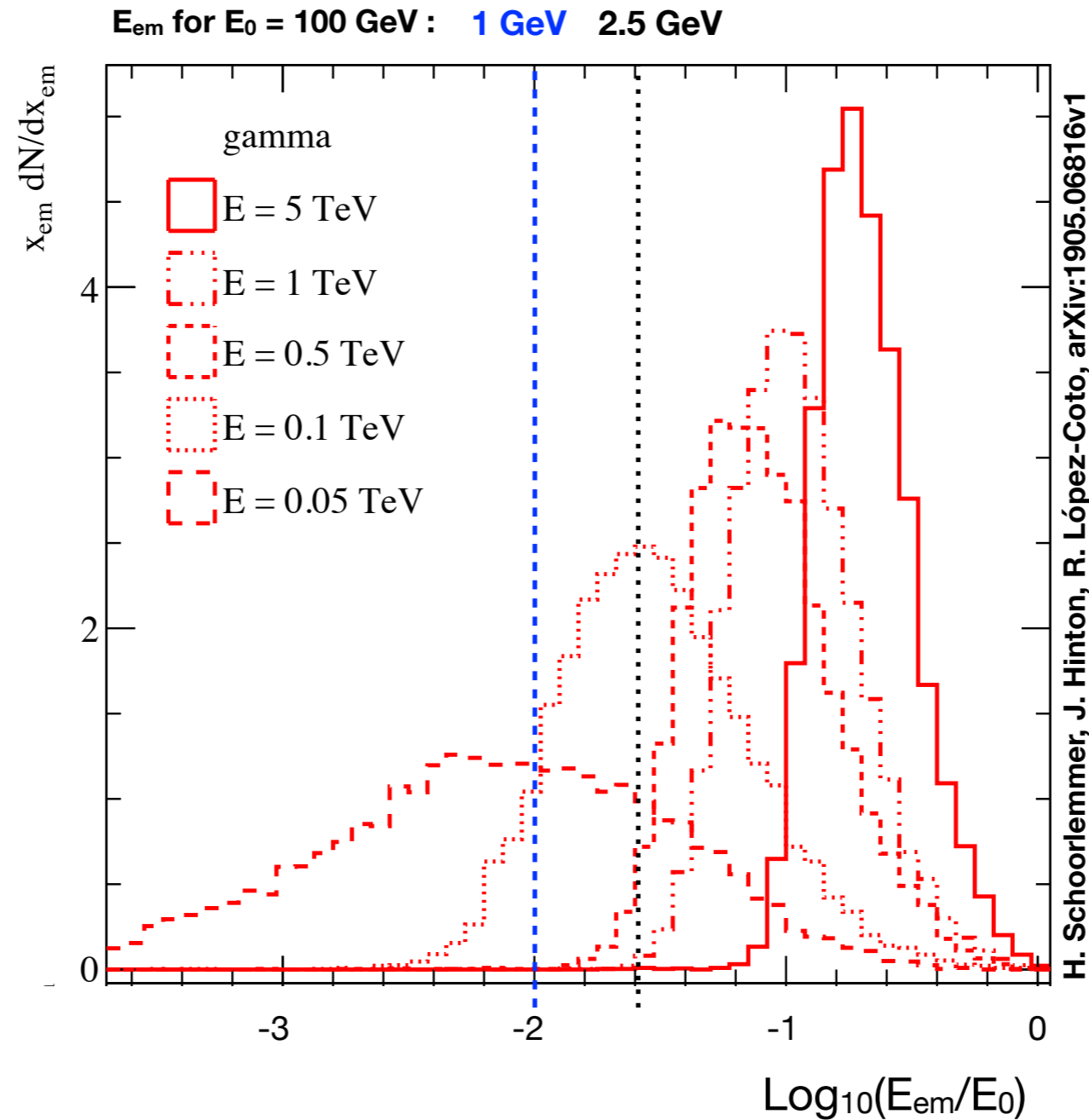


# Summary

- First considerations on a **compact WCD with SiPM readout**
- Triggering at **100 GeV** difficult, but may be achievable...
- Timing, using fast direct light pulse, at the level of 2 ns
- Muon id exploring time trace and using ANN is promising  
→ Small core (about 100 stations) complemented with RPC hodoscopes
- Nothing yet optimised... WCD dimensions, light readout, ...

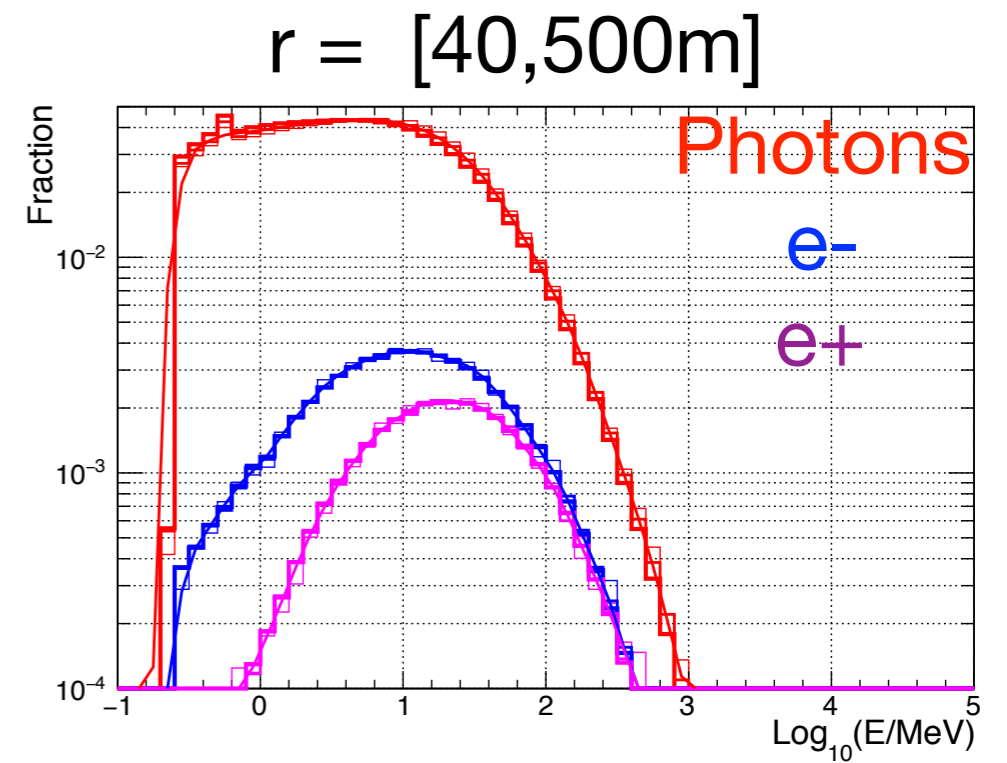
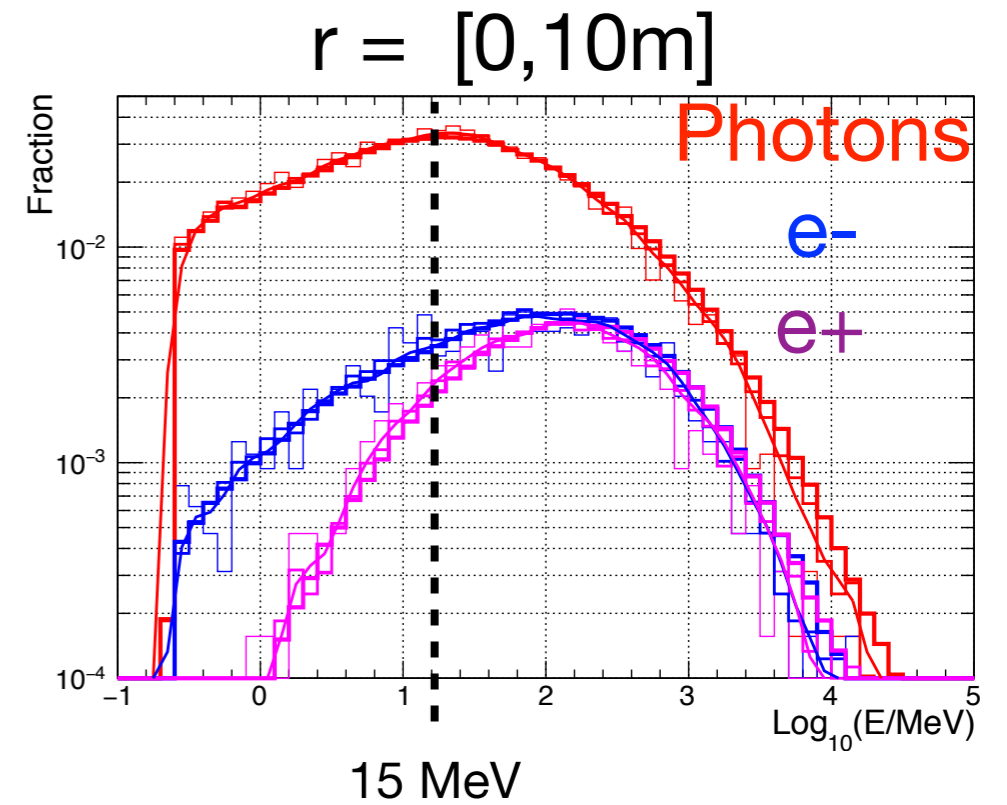
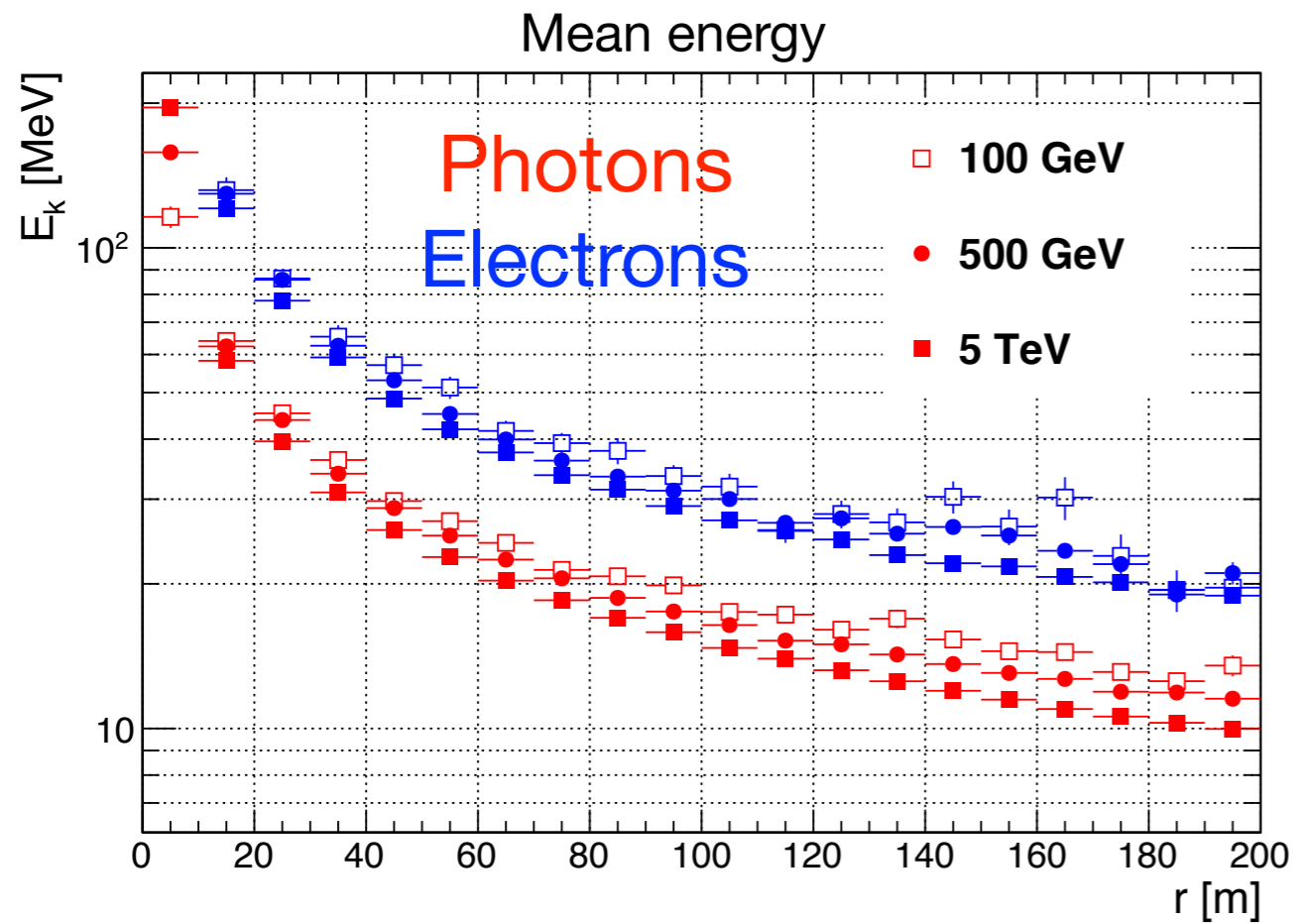
END

# EM energy @ 5000 m



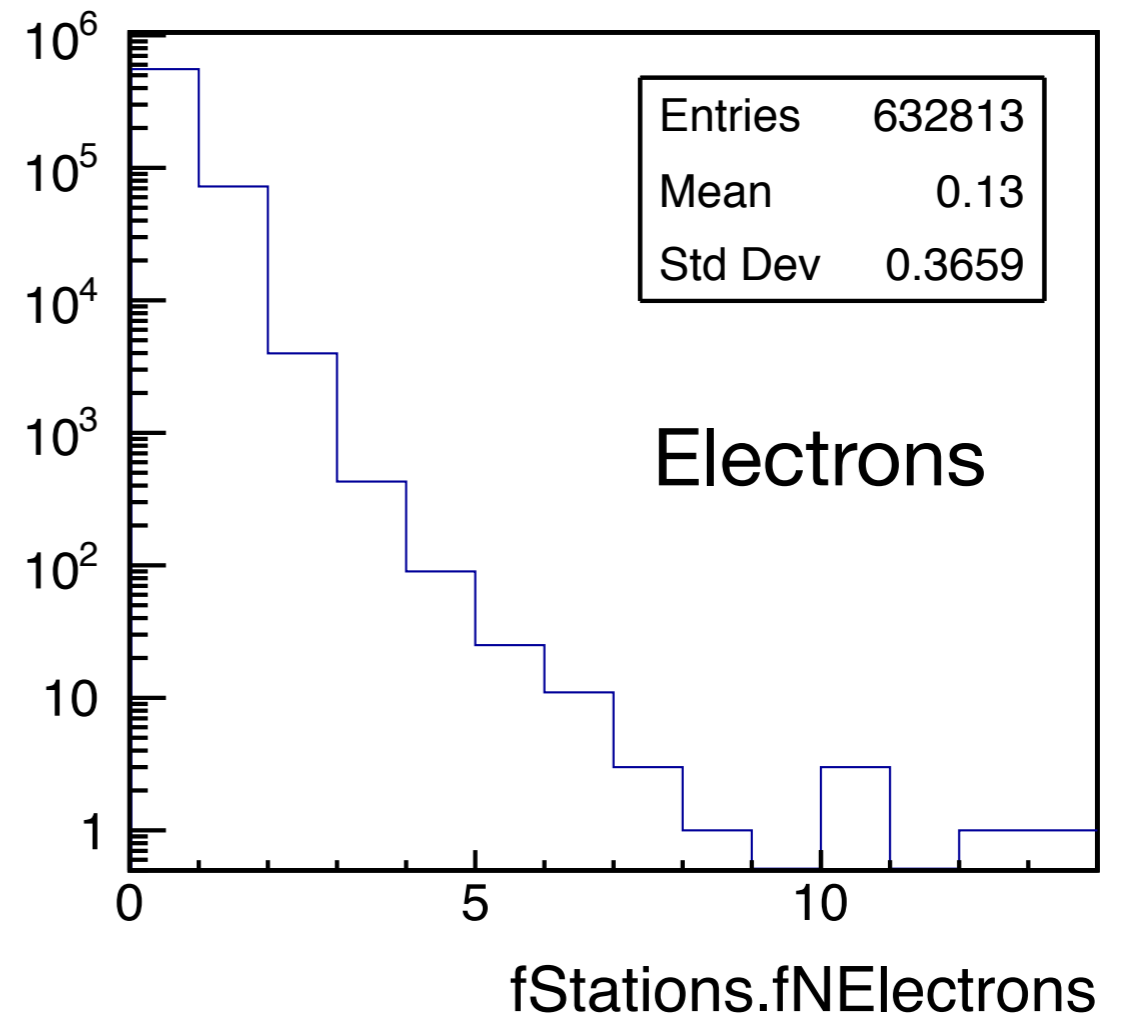
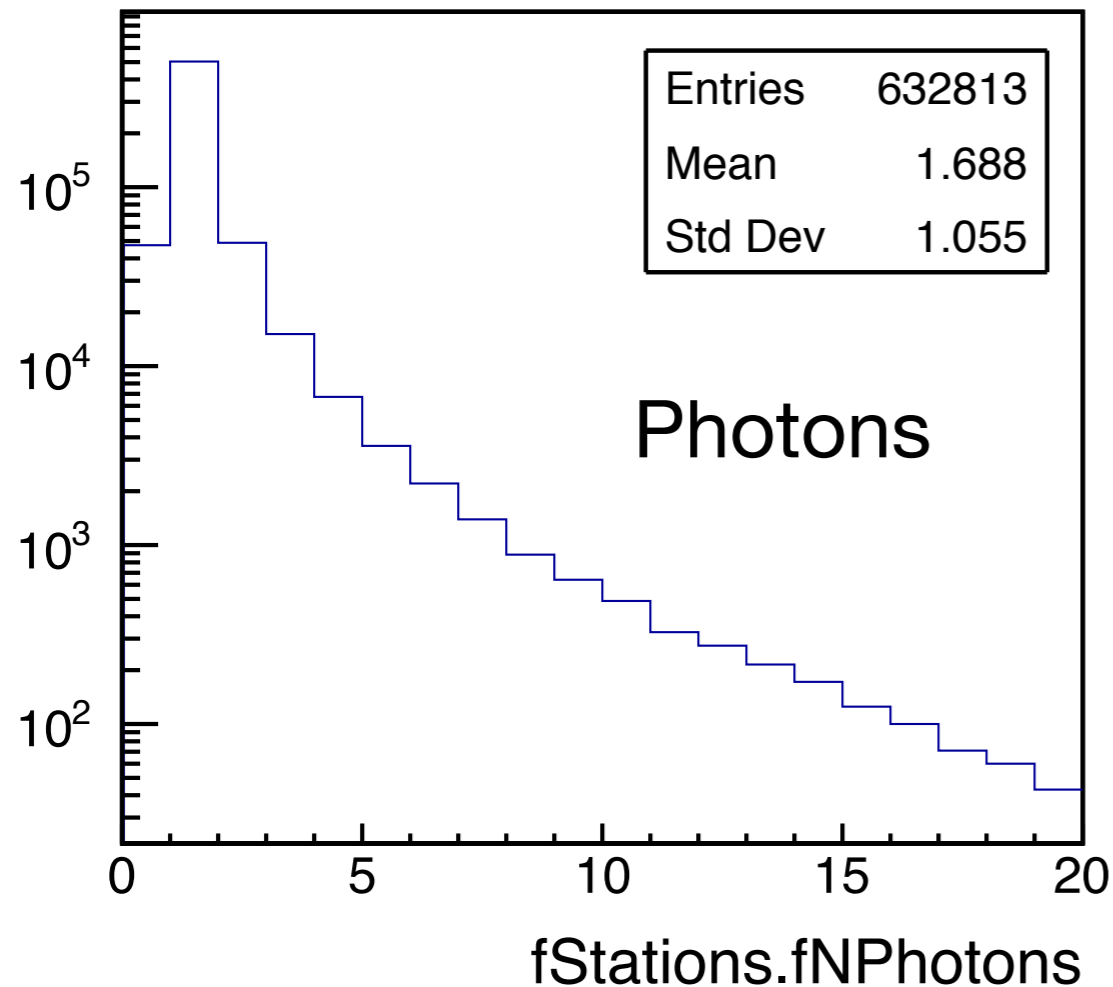
- For  $E_{em} \sim 1$  GeV  $\rightarrow \sim 100$  photons @ ground, each with a typical energy of 10 MeV - 15 MeV
- WCD sensitivity to single  $\sim 10$  MeV photons is a challenge !

# Secondary particle energy at ground (5200m) for gamma showers





# # particles per station @ 100 GeV



~ 80% of the stations have only 1 photon