#### LATTES: shower reconstruction

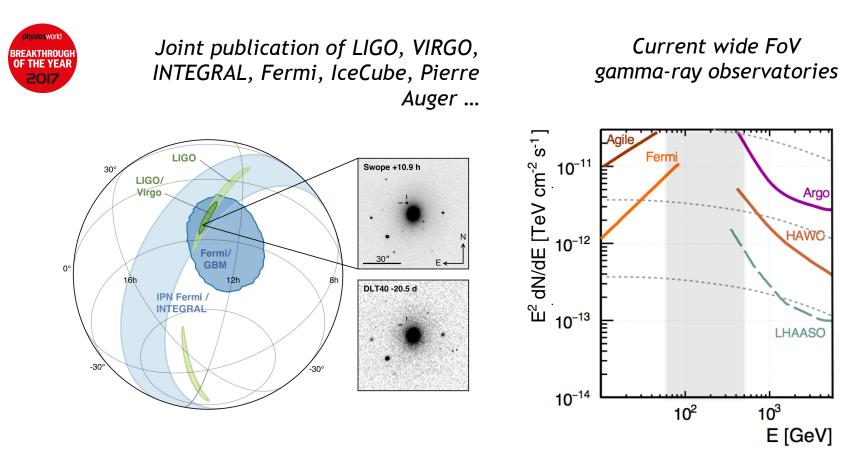
#### Ruben Conceição

#### on behalf of the LATTES team



LATTES: MVA workshop, Granada, September 14<sup>th</sup> 2018

#### The era of multi-messenger observations

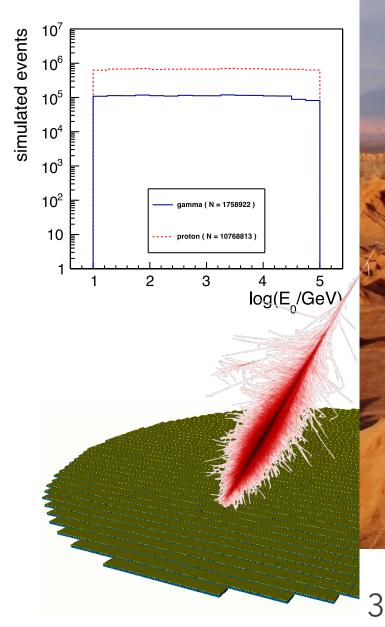


- Simultaneous observation of a Gravitational Wave + electromagnetic counterparts
- Study of transient phenomena in all energy windows is one of the main ingredients

## Simulation Framework

#### End-to-end realistic simulation

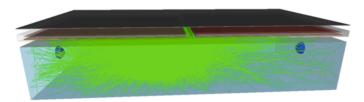
- Extensive Air Showers: CORSIKA
  - v7.6400 with Fluka2011.2c
  - More than 50 000 gamma/proton shower simulated randomly between 10 GeV - 300 TeV
  - Gammas have a fixed zenith angle of 10 degrees
  - Observation level at 5200 m of altitude
- Detector simulation: Geant4
  - ◊ v10.1.3
  - ♦ Core array 20 000 m<sup>2</sup>
  - Each shower is resampled 100 times over a big area containing all the array



#### Simulation Framework

#### Reconstruction

 First order analyses with little optimization only to demonstrate principle



#### Performance and sensitivity

Astroparticle Physics 99 (2018) 34-42



Design and expected performance of a novel hybrid detector for very-high-energy gamma-ray astrophysics

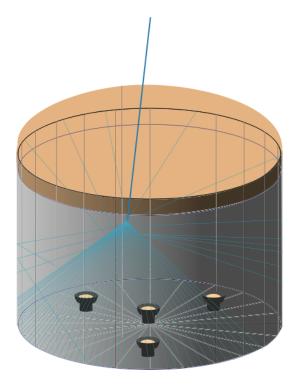


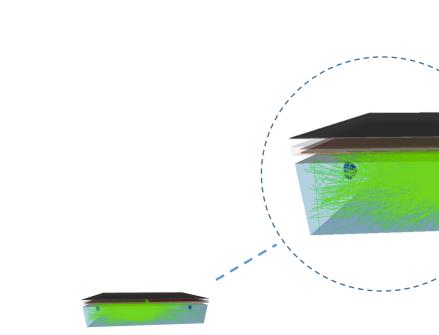
P. Assis<sup>a,b</sup>, U. Barres de Almeida<sup>c</sup>, A. Blanco<sup>d</sup>, R. Conceição<sup>a,b,\*</sup>, B. D'Ettorre Piazzoli<sup>e</sup>, A. De Angelis<sup>f,g,b,a</sup>, M. Doro<sup>h,f</sup>, P. Fonte<sup>d</sup>, L. Lopes<sup>d</sup>, G. Matthiae<sup>i</sup>, M. Pimenta<sup>b,a</sup>, R. Shellard<sup>c</sup>, B. Tomé<sup>a,b</sup>

## LATTES expected performance

- Trigger and effective area
- Energy reconstruction
- Geometry reconstruction
- Gamma/hadron discrimination
- Sensitivity to steady sources

#### Station: HAWC vs LATTES

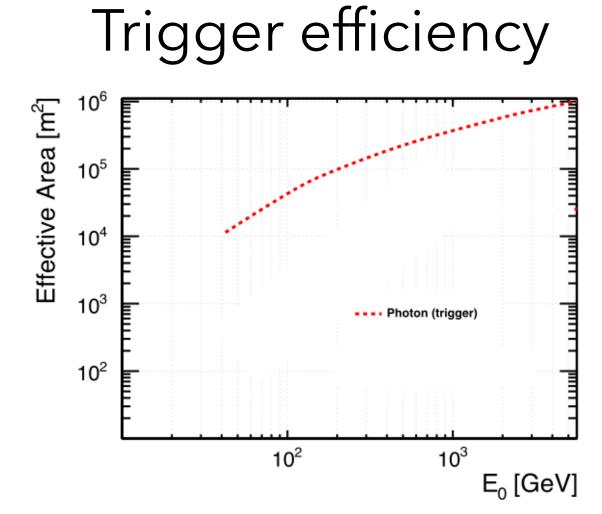




HAWC (present detector) LATTES (next generation)

## LATTES expected performance

- Trigger and effective area
- Energy reconstruction
- Geometry reconstruction
- Gamma/hadron discrimination
- Sensitivity to steady sources



- Use WCD stations to trigger at low energies
  - Trigger condition
    - ♦ Station: require more than 5 p.e. in each PMT
    - Event: require 3 triggered stations

## LATTES expected performance

Trigger and effective area

Core reconstruction

Energy reconstruction

Geometry reconstruction

Gamma/hadron discrimination

Sensitivity to steady sources

### Shower core reconstruction

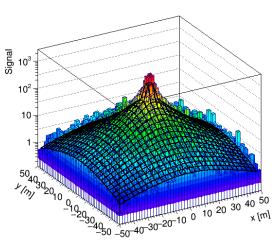
Average LDF

- ♦ Use the WCD signal
- ♦ Barycenter
  - Initial guess
  - Works but the core is always reconstructed inside the array

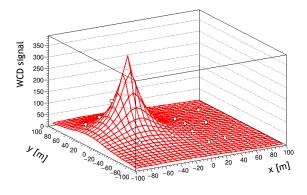
#### Fit the WCD LDF

 Fit photon average LDF to fix the shape

- ♦ Function inspired in HAWC
- Nearly no evolution with energy
- ♦ Use this form to find the maximum, i.e. the shower core



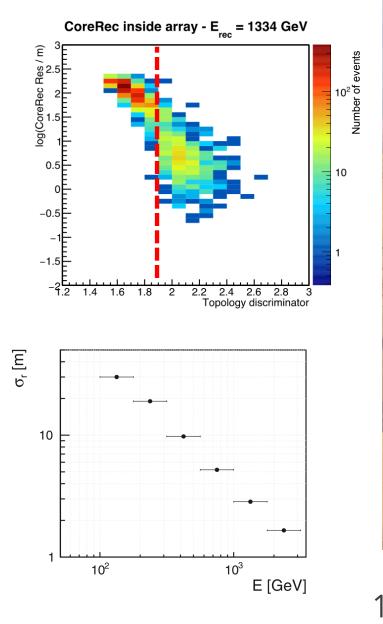
Single event



$$S_i = S(A, \vec{x}, \vec{x}_i) = A \Big( \frac{1}{2\pi\sigma^2} e^{-|\vec{x}_i - \vec{x}|^2/2\sigma^2} + \frac{N}{(0.5 + |\vec{x}_i - \vec{x}|/R_m)^3} \Big)$$
R. Conceição

### Shower core reconstruction

- Test whether the shower is inside/outside the array
  - Explore LDF topology
  - Is maximum observed inside of array?
  - Currently exploring the quality of the fit
  - Fixed cut for all energies
- Resolution better than 10 meters for showers above 300 GeV



## LATTES expected performance

- Trigger and effective area

#### Energy reconstruction

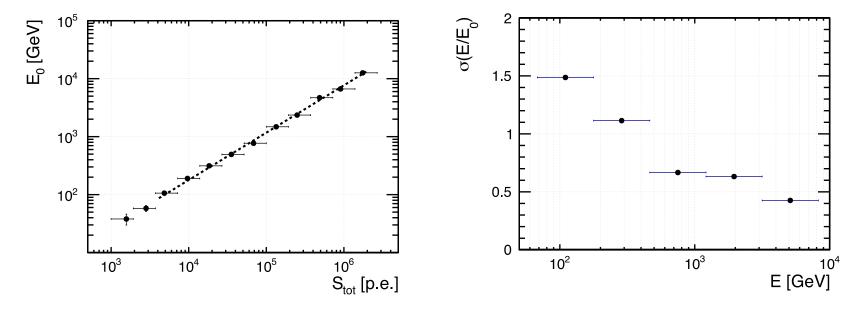
- Geometry reconstruction
- Gamma/hadron discrimination
- Sensitivity to steady sources

### Energy reconstruction

- $E_0 \rightarrow$  Simulated energy
  - $E \rightarrow$  Reconstructed energy







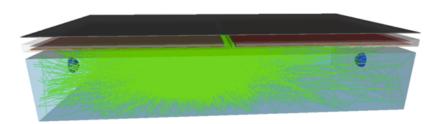
- Use as energy estimator the total signal recorded by WCDs
   Use only shower cores reconstructed inside array
- Energy resolution at low energy dominated by shower fluctuations

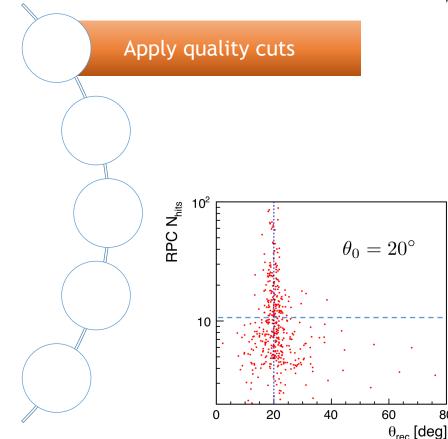
## LATTES expected performance

- Trigger and effective area
- Energy reconstruction
- Geometry reconstruction
- Gamma/hadron discrimination
- Sensitivity to steady sources

♦ Use RPC hit time information

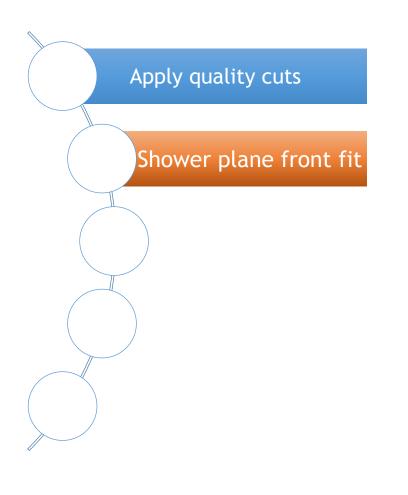
- Take advantage of high spatial and time resolution
- Used time resolution of 1 ns





- ♦ Use RPC hit time information
  - Apply previous shower rec quality cuts
  - Apply cuts on the number of registered hits on the RPCs
  - ♦ Consider only RPCs in triggered WCD stations

80

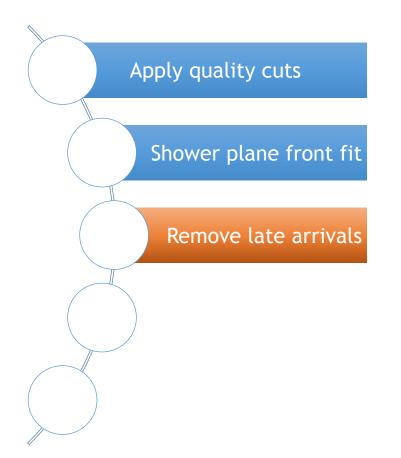


- ♦ Use RPC hit time information
- Perform shower
   reconstruction
- Use shower front plane approximation

stations

Analytical procedure

barycenter b

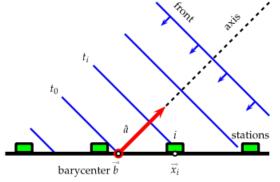


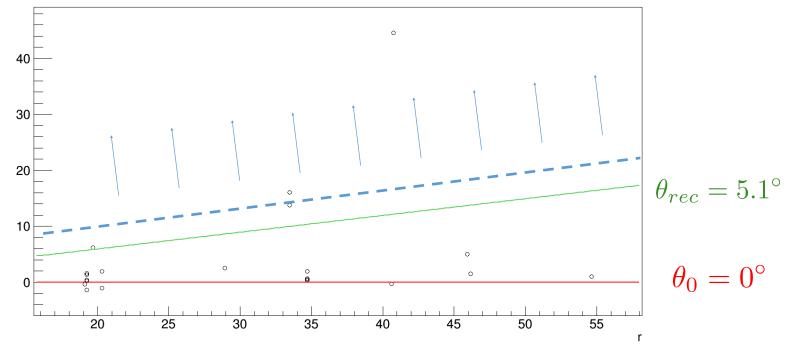
♦ Use RPC hit time information

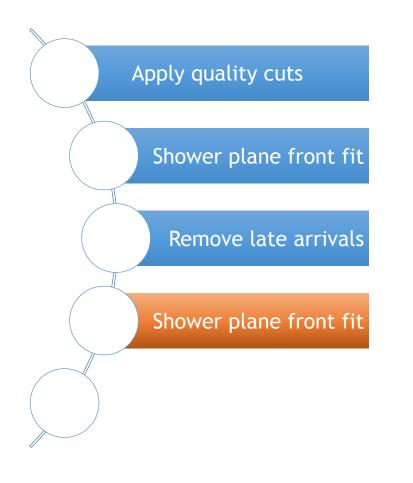
- Identify late arrivals
   with respect to Rec
   Shower Front
- Mainly low energy electrons that lost correlation with shower front

### Removal of late arrivals

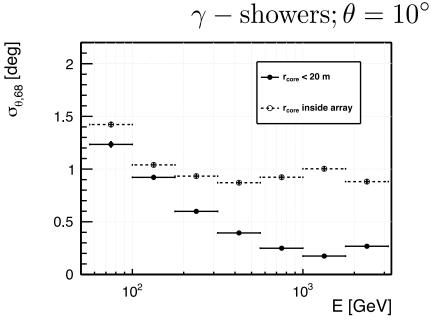
- Example of a vertical gamma shower
- Plot depicts arrival time (ns) distance to simulated shower core (m)

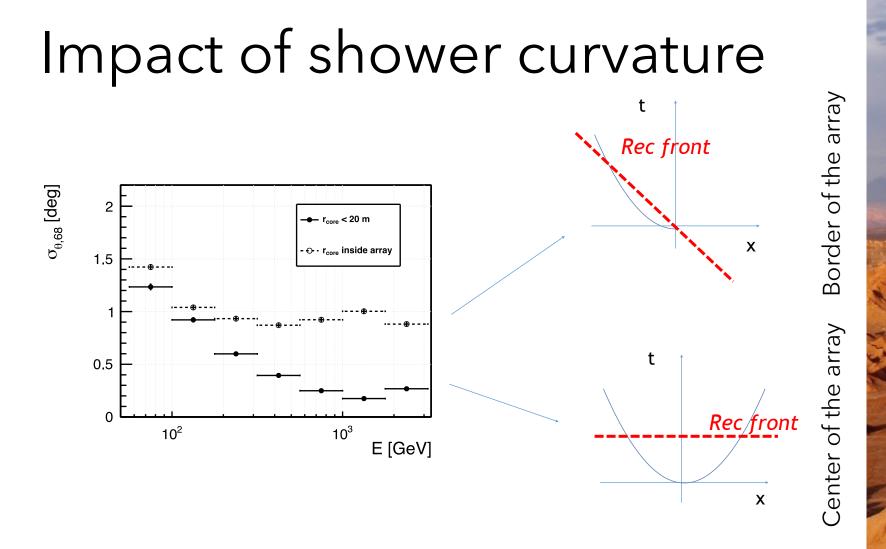






- ♦ Use RPC hit time information
  - Repeat fit without arrivals
  - ♦ Initial guess for next step

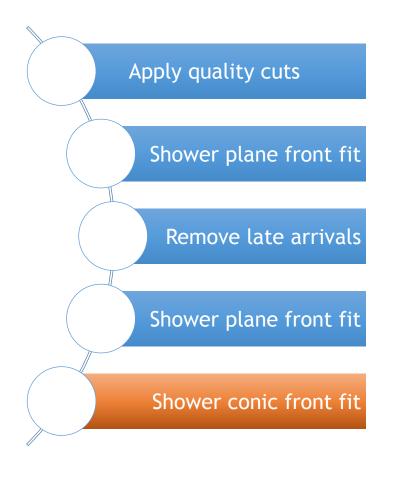




Solution: implement a conic fit instead of fitting a plane

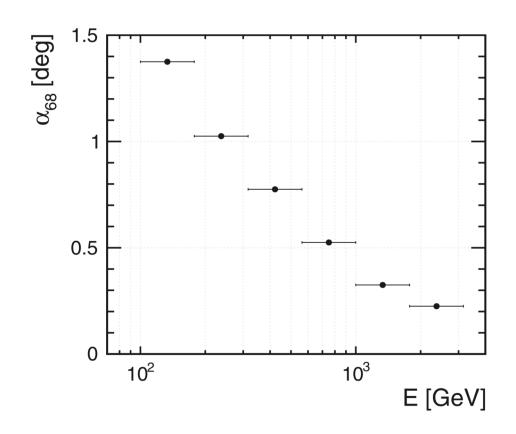
$$\chi^2 = \sum \left( c \cdot (T_n - T_0) - X_n \cdot - Y_n \cdot m - (R_n \cdot \alpha) \right)^2$$

R. Conceição



- ♦ Use RPC hit time information
  - ♦ Fit the shower
     geometry using a
     shower conic front
     model

## Shower geometry reconstruction



A good angular resolution can be achieved for all events reconstructed inside the array

## LATTES expected performance

Trigger and effective area

Energy reconstruction

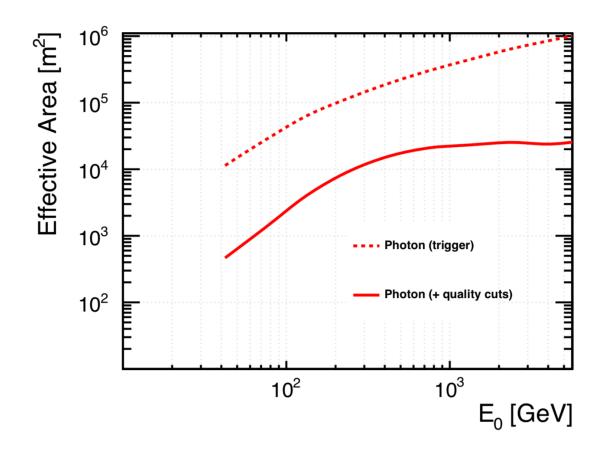
Geometry reconstruction

Gamma/hadron discrimination

Sensitivity to steady sources

Shower rec quality cuts

#### Effective Area

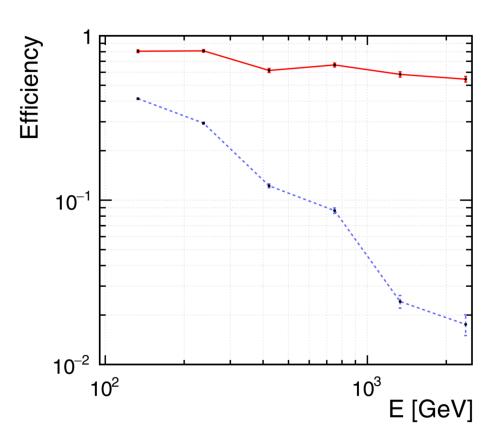


Even applying all quality cuts LATTES gets an effective area of ~1000 m<sup>2</sup> for E = 100 GeV

## LATTES expected performance

- Trigger and effective area
- Energy reconstruction
- Geometry reconstruction
- Gamma/hadron discrimination
- Sensitivity to steady sources

## LATTES g/h discrimination



(see next talk)

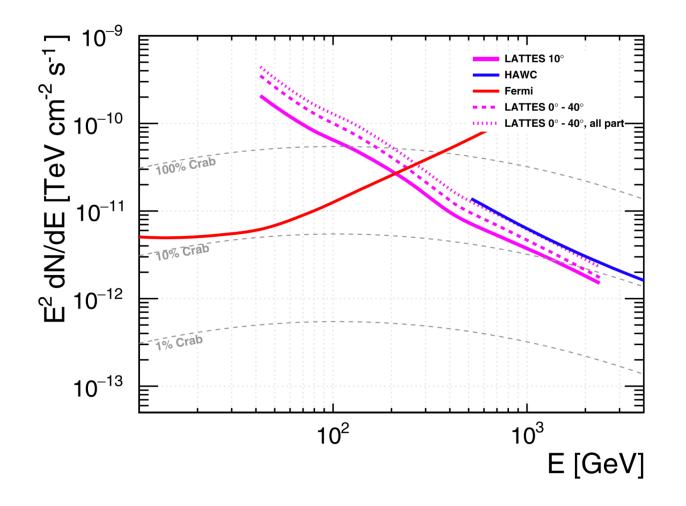
Although not optimized the gamma/hadron discrimination results are already very encouraging

## LATTES expected performance

- Trigger and effective area
- Energy reconstruction
- Geometry reconstruction
- Gamma/hadron discrimination

#### Sensitivity to steady sources

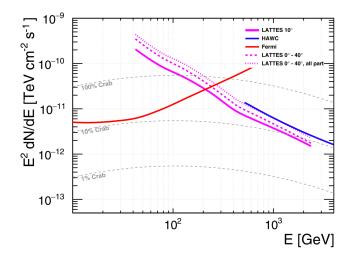
#### Sensitivity to steady sources

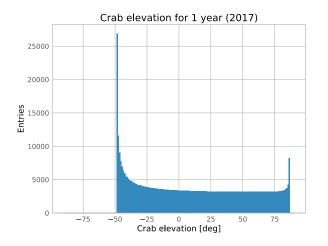


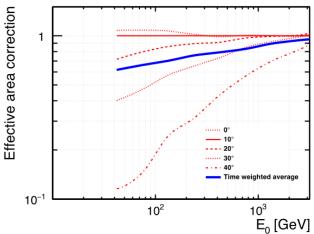
♦ Full line: full MC calculation for a source at 10 degrees in zenith

## Sensitivity to steady sources

- ♦Dashed line: Crab transit as seen by HAWC
  - Degradation of effective area with zenith angle estimated from electromagnetic energy at ground



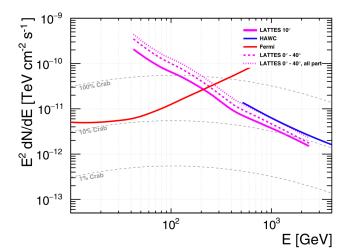


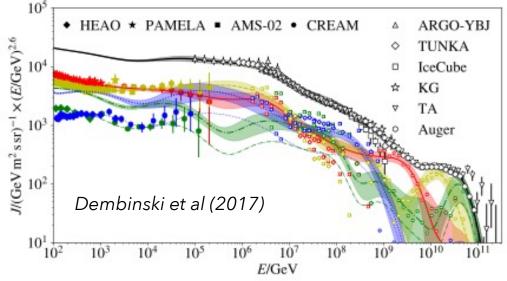


## Sensitivity to steady sources

♦ Dotted line: CR all-spectrum

- ♦ Additional elements (He, N, Fe...)
- Assume that LATTES cannot distinguish gammas from irons





## Summary

- LATTES shower reconstruction performance has been evaluated yielding very good results
  - Shower trigger (effective area)
  - Shower core reconstruction
  - Shower energy reconstruction
  - Shower geometry reconstruction
  - Gamma/hadron discrimination
- ATTES capabilities are far from being fully explored
  - Possible improvements already identified
    - Sparse array to veto far away high-energy showers (main background source)
    - ♦ Use RPC patterns to discriminate g/h
    - ♦ Better assess LATTES ability to reconstruct
      - Inclined showers
      - ♦ Heavier primaries induced showers
    - ÷ ...

### Acknowledgements



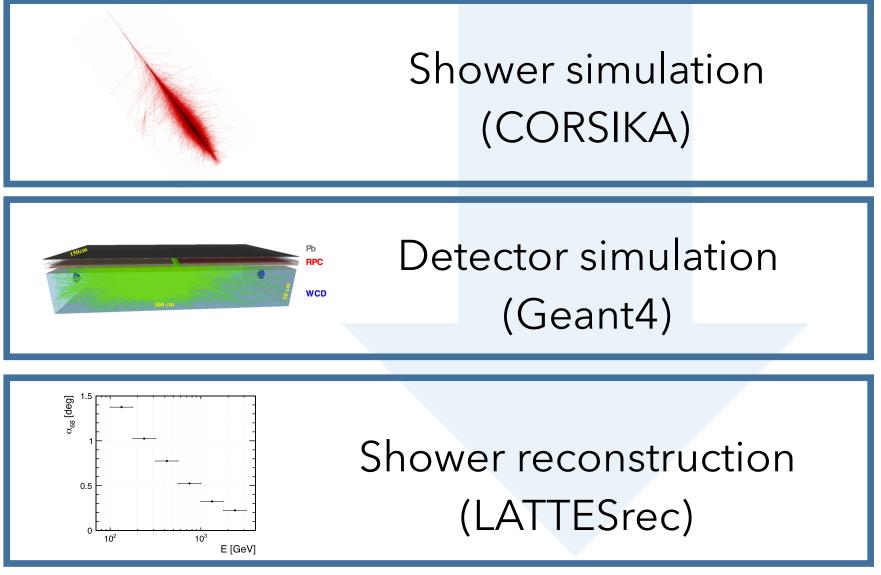






# Backup slides

#### Towards LATTES sensitivity...



## LATTES: a hybrid detector

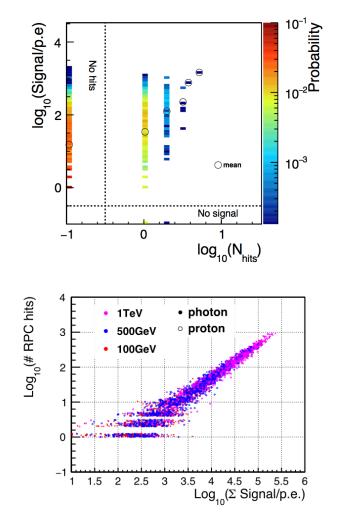
#### ♦ Thin lead plate

- To convert the secondary photons
- Improve geometric reconstruction
- Resistive Plates Chamber
  - Sensitive to charged particles
  - Good time and spatial resolution
  - Improve geometric reconstruction
  - ♦ Explore shower particle patterns at ground

#### Water Cherenkov Detector

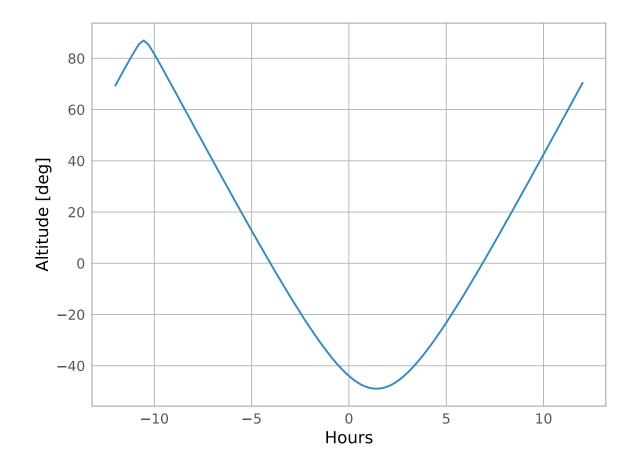
- Sensitive to secondary photons and charged particles
- Measure energy flow at ground
- Improve trigger capability
- Improve gamma/hadron discrimination

#### WCD vs RPC (station level)



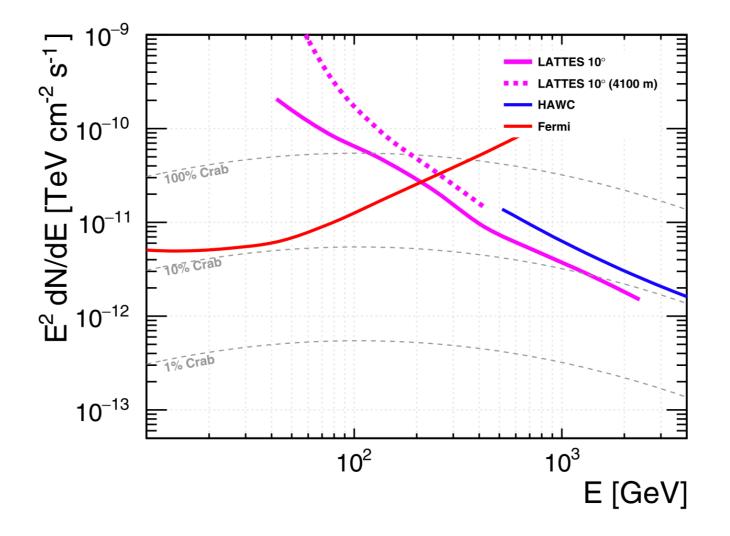
Inter-calibration

#### Crab



57

#### Impact of altitude



#### Accidentals contamination

Considering a time window D, the mean number of stations that randomly trigger within D is :

$$n_s = N_s \times R \times D$$

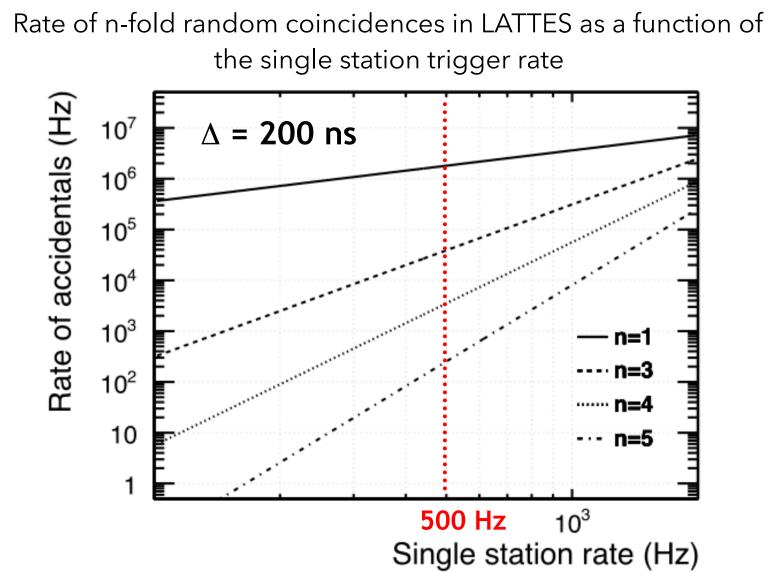
with  $N_s$  the # of stations in the array and R the single station trigger rate.

For LATTES  $N_s = 3600$  and R was estimated from MC simulations to be of the order of 500 Hz; taking D ~ 200 ns yields :

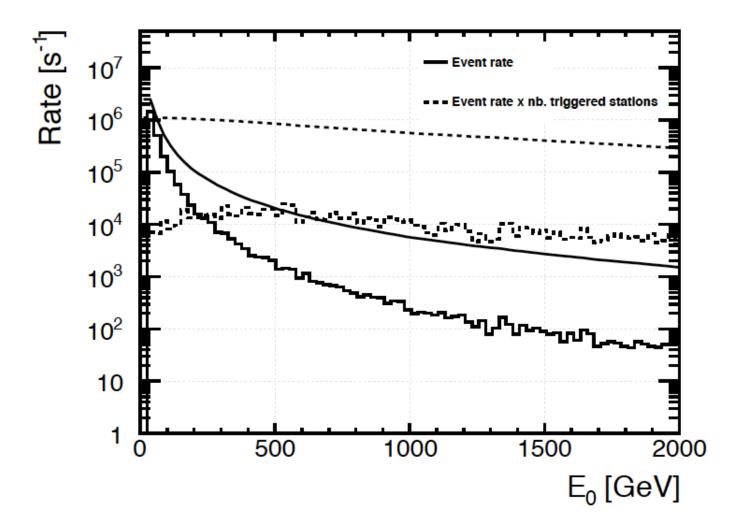
to be compared with the minimum of stations required in a shower trigger,  $n_s=3$ .

In any case a detailed MC simulation of the impact of the accidentals should be performed !

## Random triggers



#### Cosmic rays and station trigger rate



#### Reconstruction efficiency

