



Content



- Introduction and Experiments
 - Accelerators and Cosmic Rays



- Sky Survey for Source Candidates
- The 1st Catalogue of Microquasars
 - MQs as PeV Particle Accelerators
 - Investigation in Depth of a BH+jet system
- Cosmic Ray Source Papulation in "Knee" Energy Range
- Future Experiments

• E_e 2 1 TeV • E_e 2 0.1 TeV

Acceleration



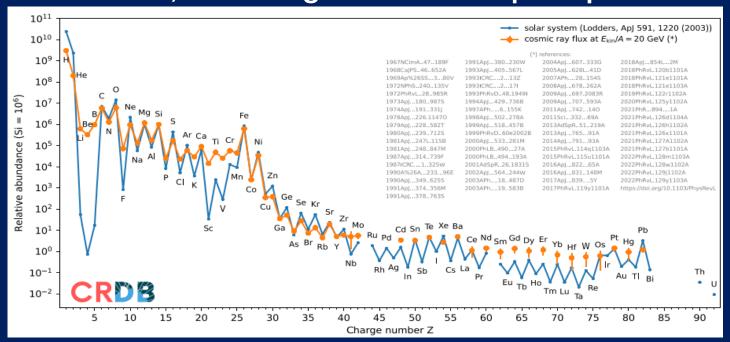
LEP

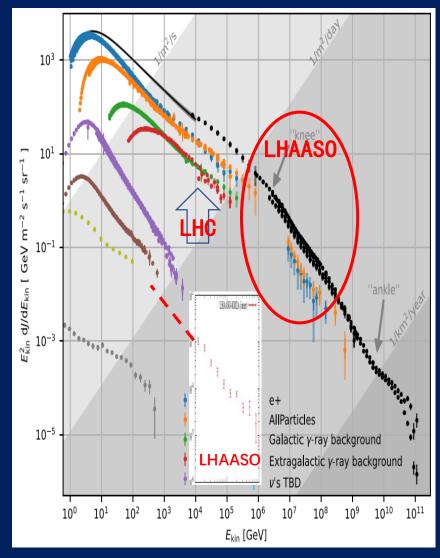
region

- Electron FCC/CEPC accelerator
- LHC
- Proton -accelerator

宇宙加速器: 特性及其存在的证据 Cosmic Accelerators: evidences of existence

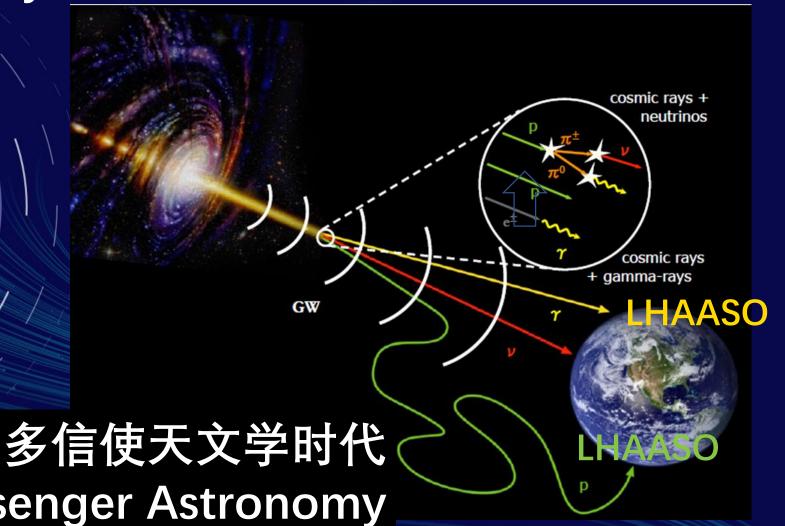
- After 110 years, we have learnt a lot about CRs near the Sun
- Particularly in the era of high precision measurements
- However, their origin is still an open question





宇宙加速器在哪里?怎么才能找到它们? We knew that the accelerators exist. Where are they? How to find them?

LHAASO is
a y-ray Telescope
and
a CR detector

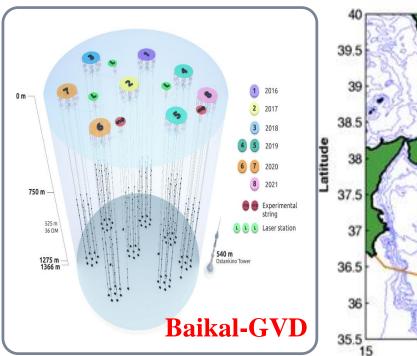


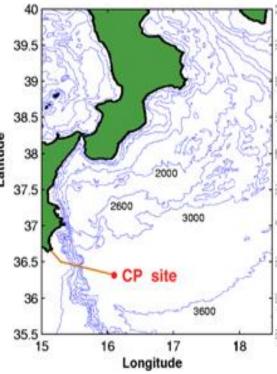
Era of Multi-Messenger Astronomy

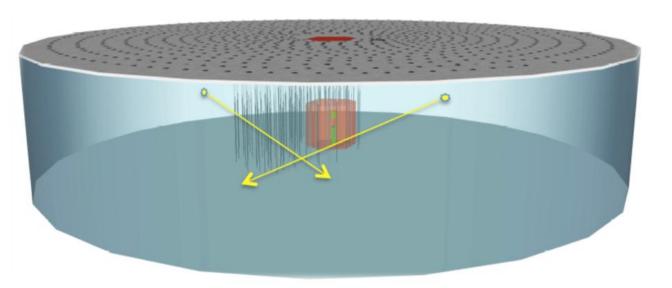
Neutrino 06/11/2014 中微子 Up-going (Decl= 11.5°) v_{μ} $E_{\nu} > E_{\mu} \sim 2.6 \times 10^{15} \text{ eV}$ **Astronomy** "Bert" 1.04 PeV "Big Bird" 2 PeV "Ernie" 1.14 PeV Aug. 2011 Jan. 2012 Dec. 2012 80 Showers ⊢• Earth absorption Tracks ⊢X 60 Declination (degrees) 40 20 DOWNGOING -60 -80 early Time scale 10² 10³ late Deposited EM-Equivalent Energy in Detector (TeV)

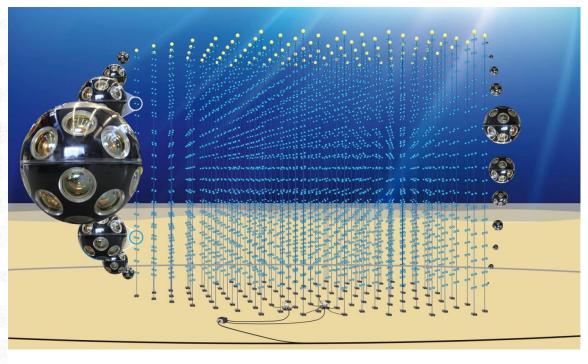
Neutrino Astronomy

- IceCube/Generation-2
- KM3Net
- BaiKal-GVD



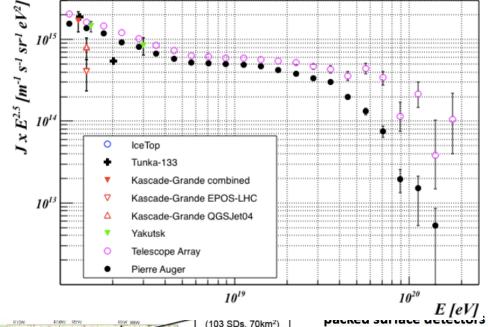


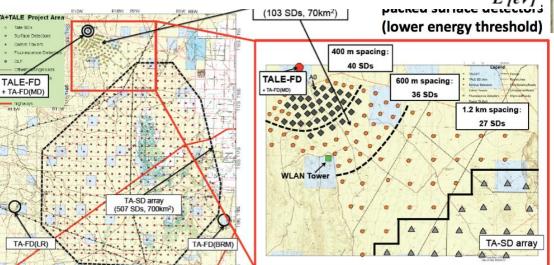


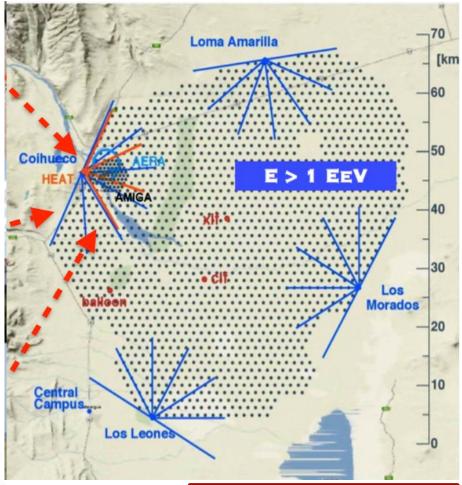


带 电宇宙

UHECR particle astronomy







1400 m a.s.l. [820 g cm⁻²] A ~ 3000 km², 1500 m grid 1660 water Cherenkov SD + 24 FD

1400 m a.s.l. [880 g cm⁻²] 507 SD, 1.2 km grid, 700 km² 16 TALE counters, 400 m grid 3 FD (BR,LR,MD/TALE)

X 2



γ-ray Astronomy

Image Atmospheric Cherenkov Telescopes: pointing observation









终极科学目标: 发现高能宇宙线源。 Scientific Goals

 γ -ray astronomy:

Survey for sources (above 500 GeV)

PeVatrons (above 100 TeV)

All kind of sources: SNR, PWN, MYC, binary,

pulsar, AGN, GRB etc.

Cosmic Ray Physics:

The knees

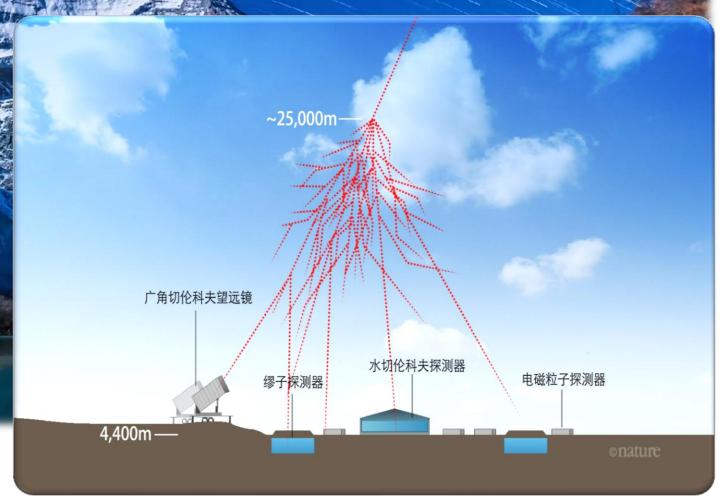
Compositions: individual species H, He and Fe

Anisotropy: (1 TeV to 10 PeV)

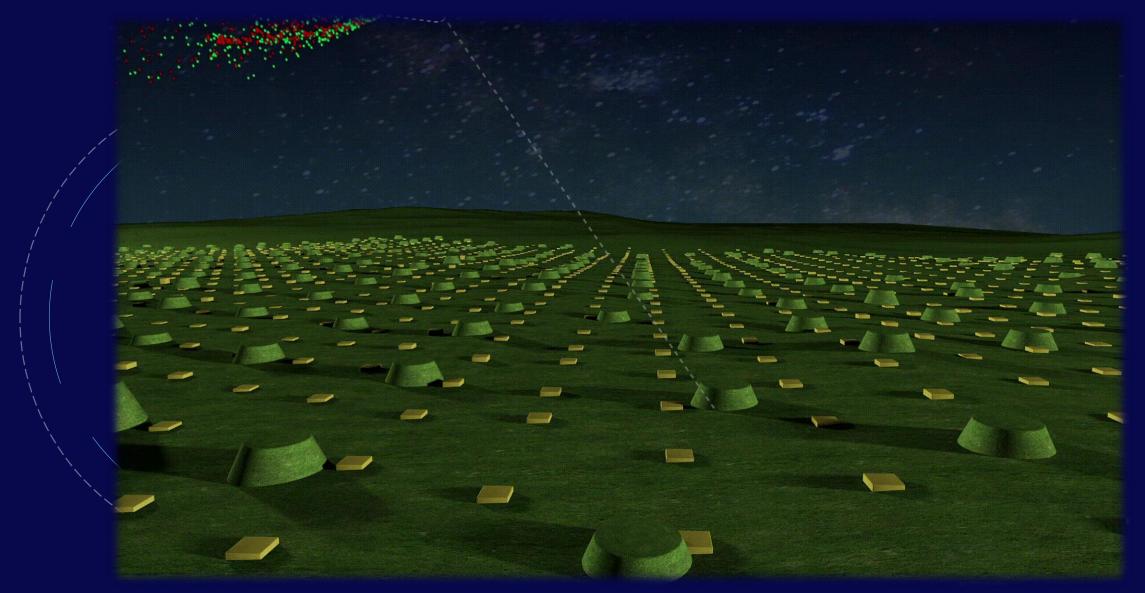
New Physics Front: DM, LIV, etc.



LHAASO



The principle of cosmic ray detection of LHAASO

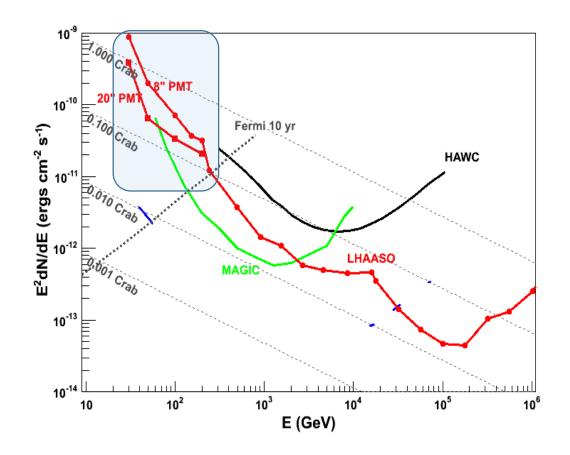


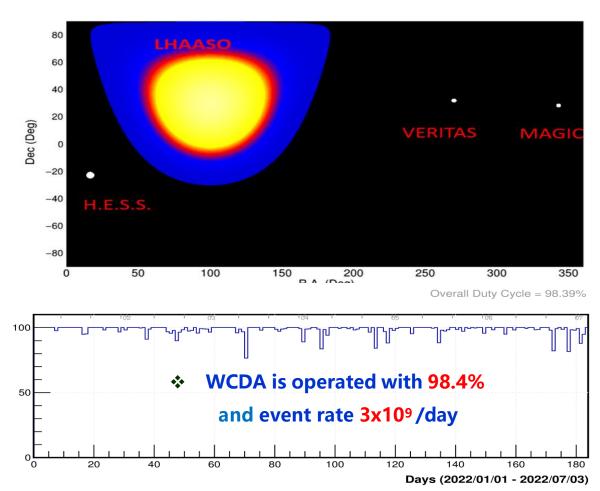
到来的"信使"化作"阵雨"洒落在"拉索"阵列上,一场"雨"持续的时间只有几纳秒

Complementary Observations: survey vs. pointing

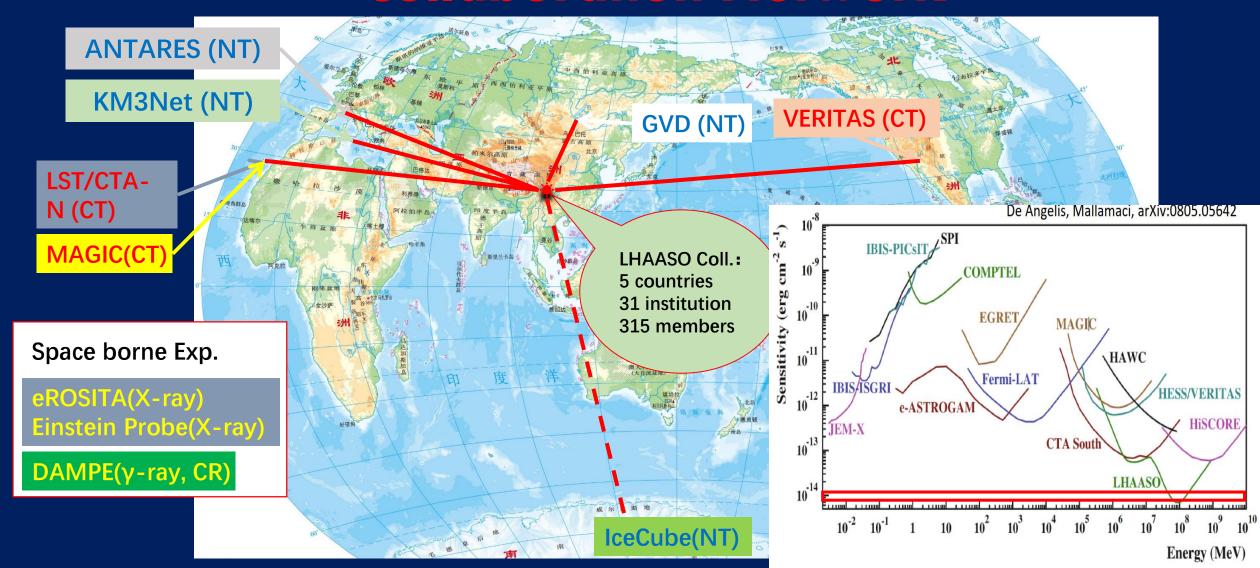
Daily Duty Cycle [%]

- Time-domain astronomy vs. steady radiation
- VHE vs. UHE bands





Multi-Messenger Collaboration Network



Content



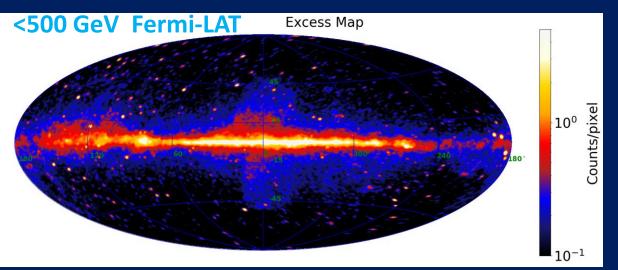
- Introduction and Experiments
- Gamma Ray Source Survey Results

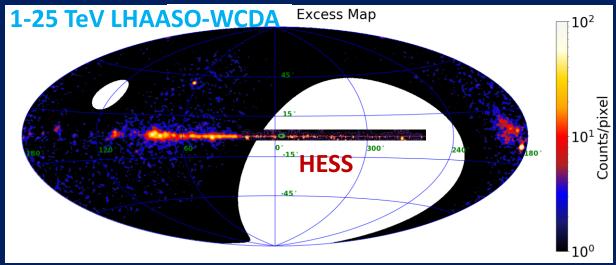


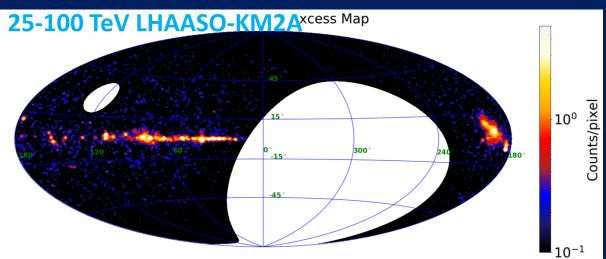
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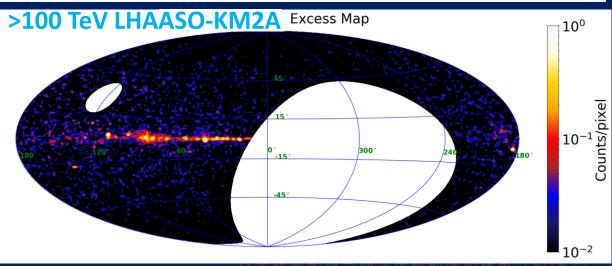
UHE γ-ray Astronomy: survey for sources

> Survey discovered 30+ new sources, 40+ PeVatrons and diffuse γ-ray emission



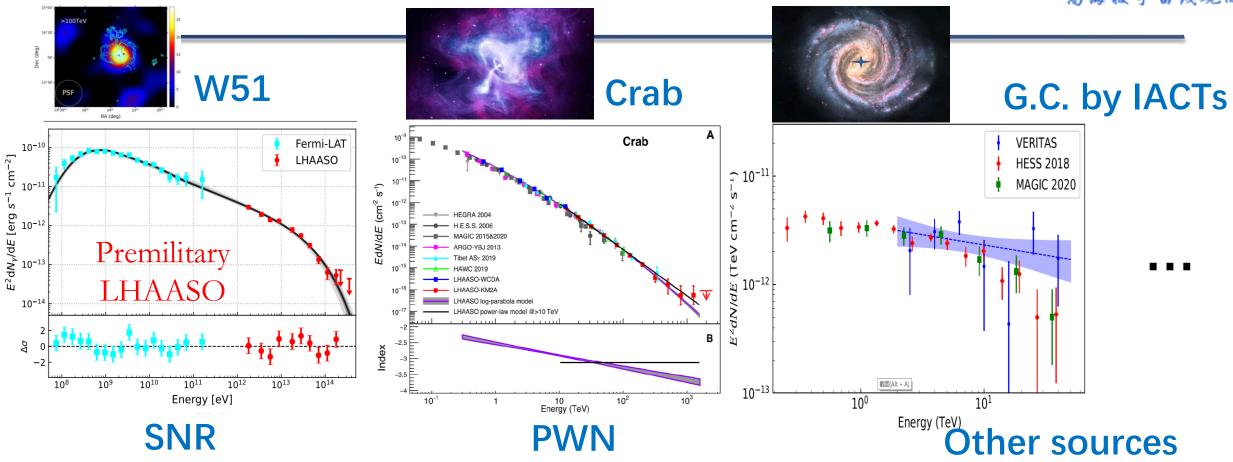






宇宙线源的候选天体Possible Source Candidates





Many types of γ-ray sources have the potential to accelerate particles to 1 PeV and above

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MQs as PeV Particle Accelerators

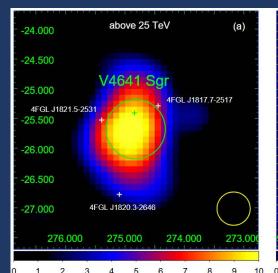


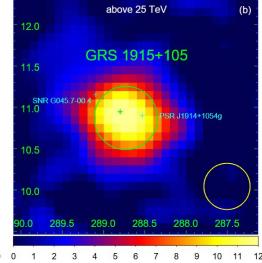
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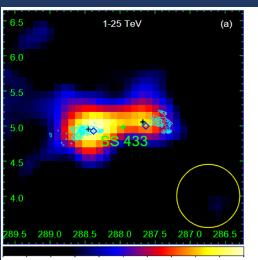
Black Holes and Jets: µQs

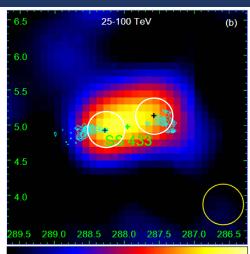


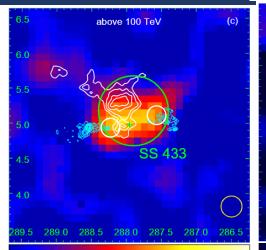
- Very important !!
- New CR source population particularly at energy E >3 PeV

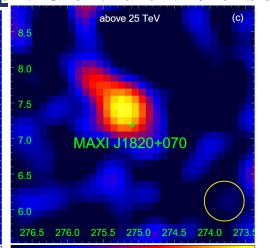


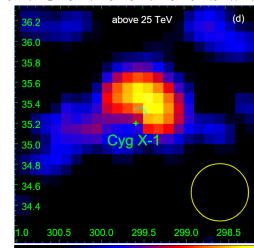






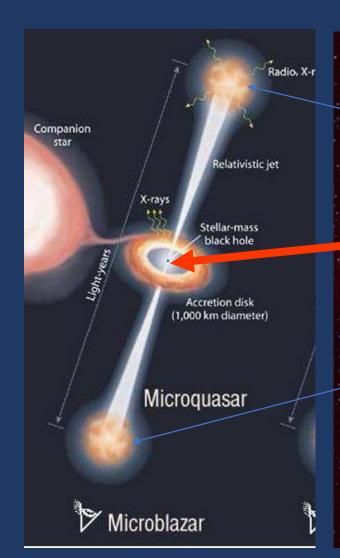


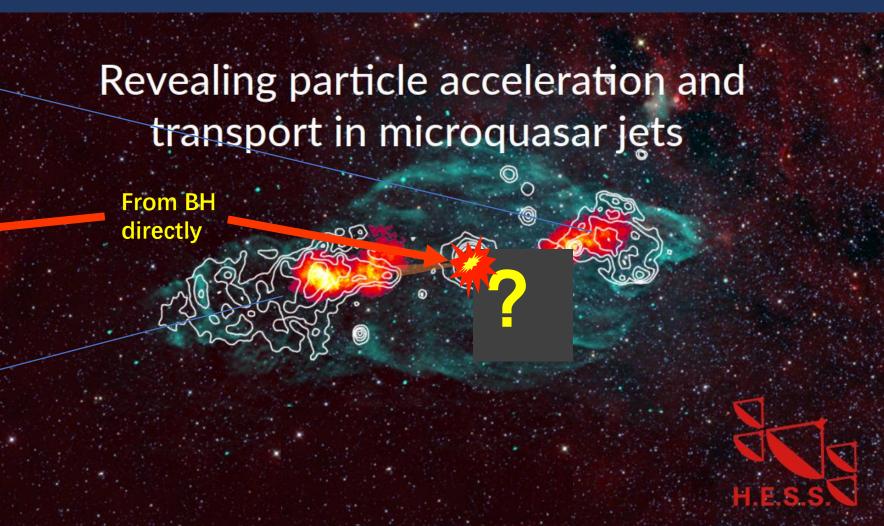




MicroQuasar: SS433 etc

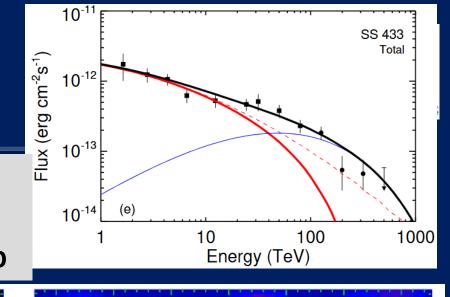


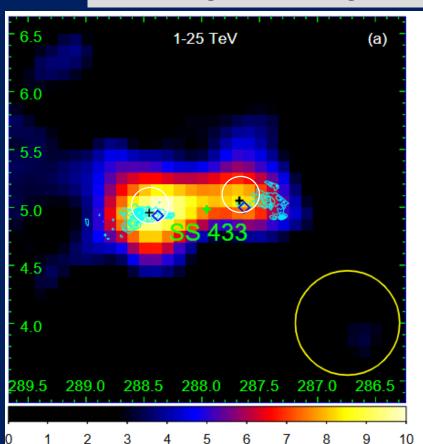


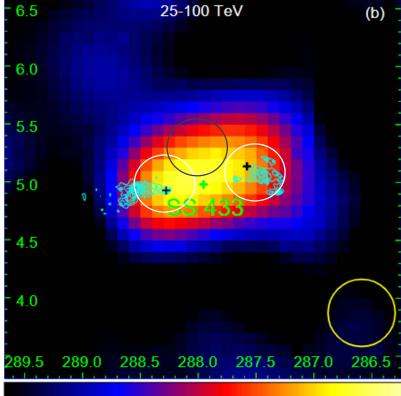


Black Holes and Jets

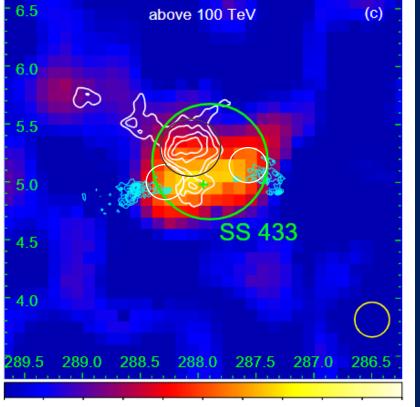
- > LHAASO measured them clearly
- > At low energy, the jets
- > At higher energies, BH itself may have shown up







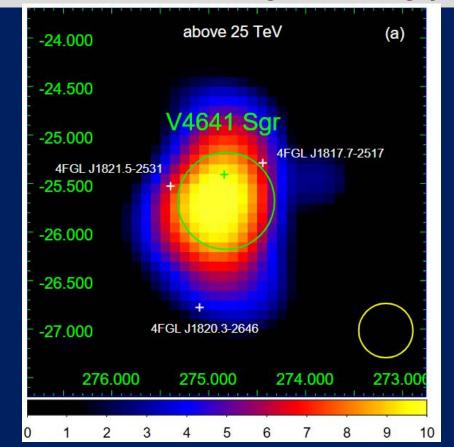
11 12 13 14 15

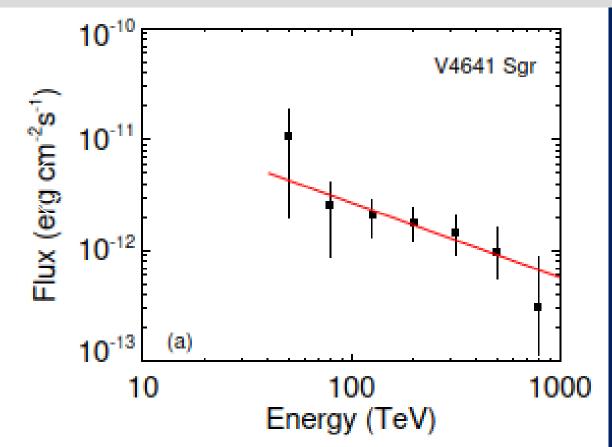


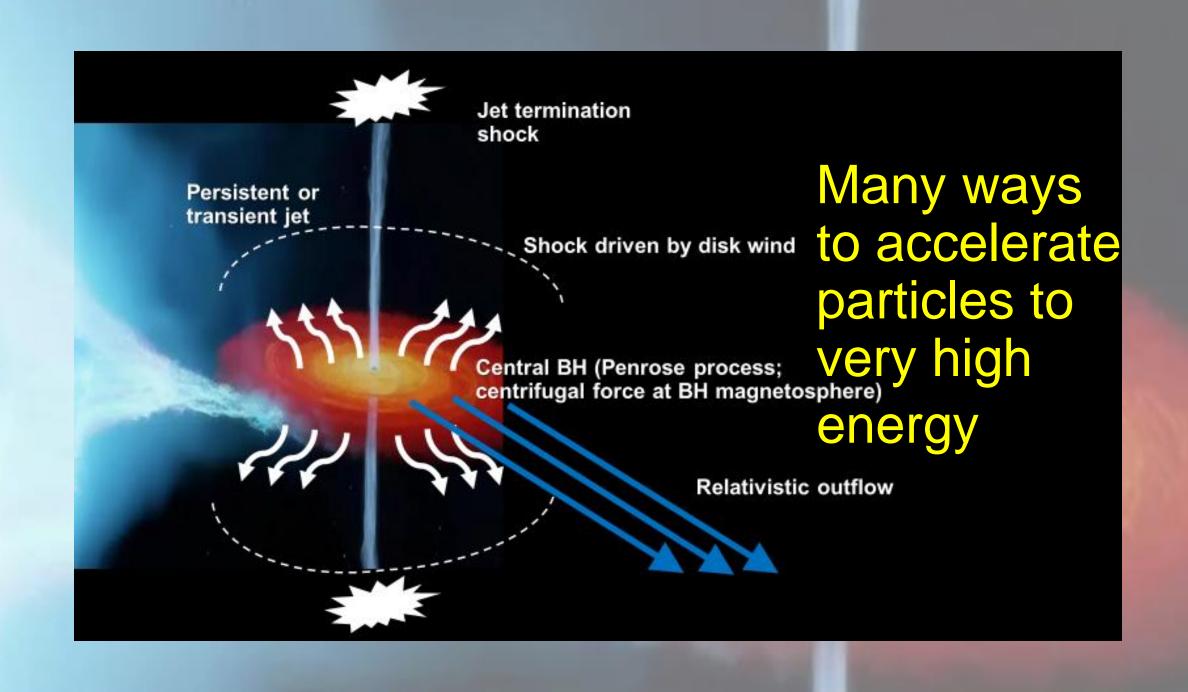
Black Hole as a super-PeVatron?



Very difficult to detect: not only due to the distant: ~20,000 light-year! But also out of main field of view of LHAASO: a source in southern hemisphere Powerful accelerator generating particle at E >10 PeV!!











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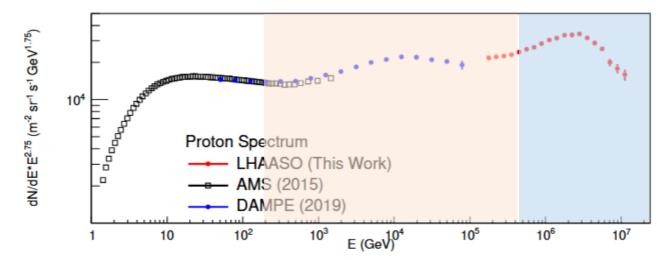


Future Experiments: LACT and HUNT

What LHAASO Observations Tell Us

They are HE Cosmic Ray Sources

- Many evidences collected for the SNRs accelerate CRs to very high energy, W 51, W 44, IC 433, Cas A
- IC 433 amazingly shows the π^0 -decay spectrum extending to tens of TeV without indication of cut-off feature
- Almost all of them cut at energies around 10 TeV or even lower
- Before the SNR contribution completely dying away, the on-set of a new component is observed with the hardening of the spectrum
- Stronger enhancement of the flux than that observed in 10 TeV range



Content



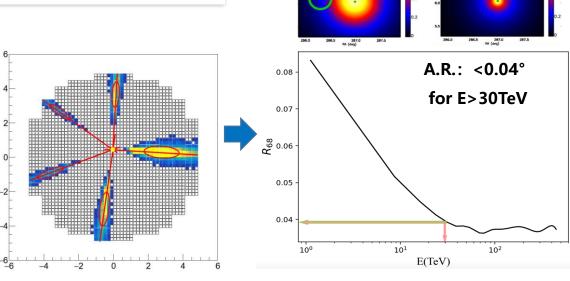
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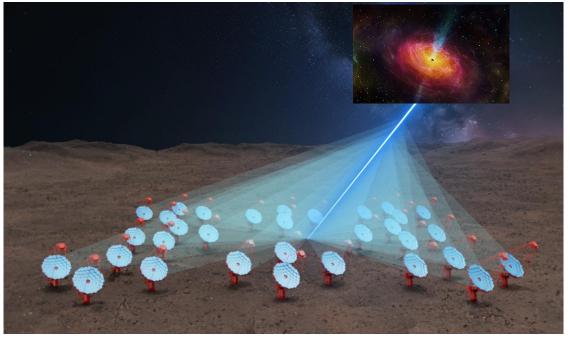


LACT as the upgrading of LHAASO

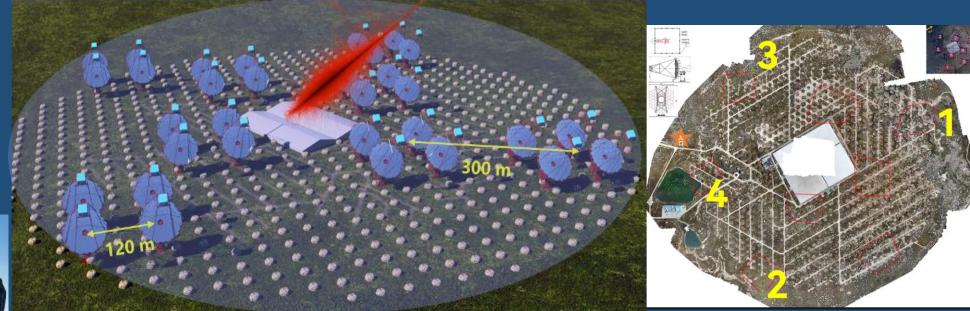
- > Stereo measurement of Cherenkov image
 - **☐** At least 4 telescopes simultaneously
- **Reconstruction**
 - Angular resolution **0.04° for E**>30 TeV







LACT: an IACT array in LHAASO

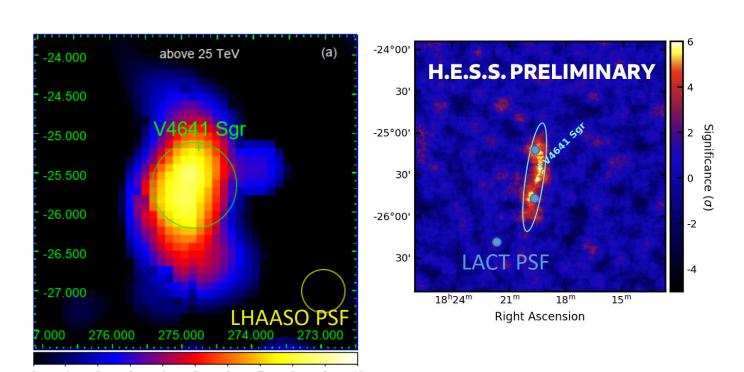


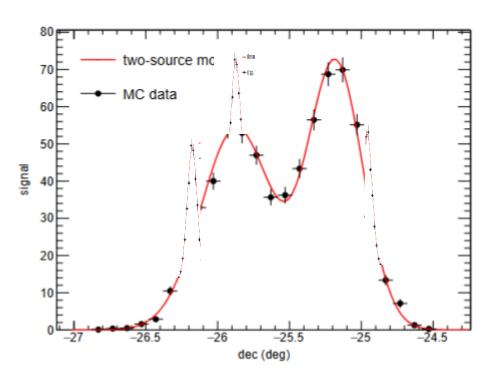


- Funded
- 8X4 array at LHAASO site
- 6-m telescopes
- two proto type telescopes
- First light soon in this year!

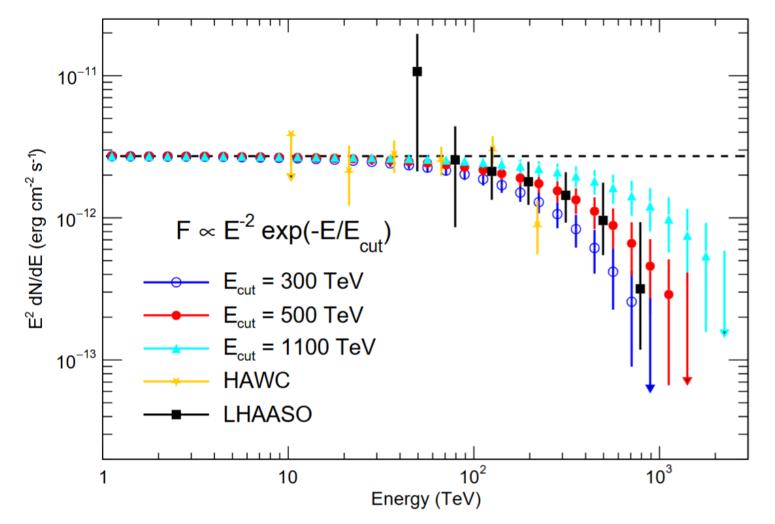
Jet Termination vs central accelerator

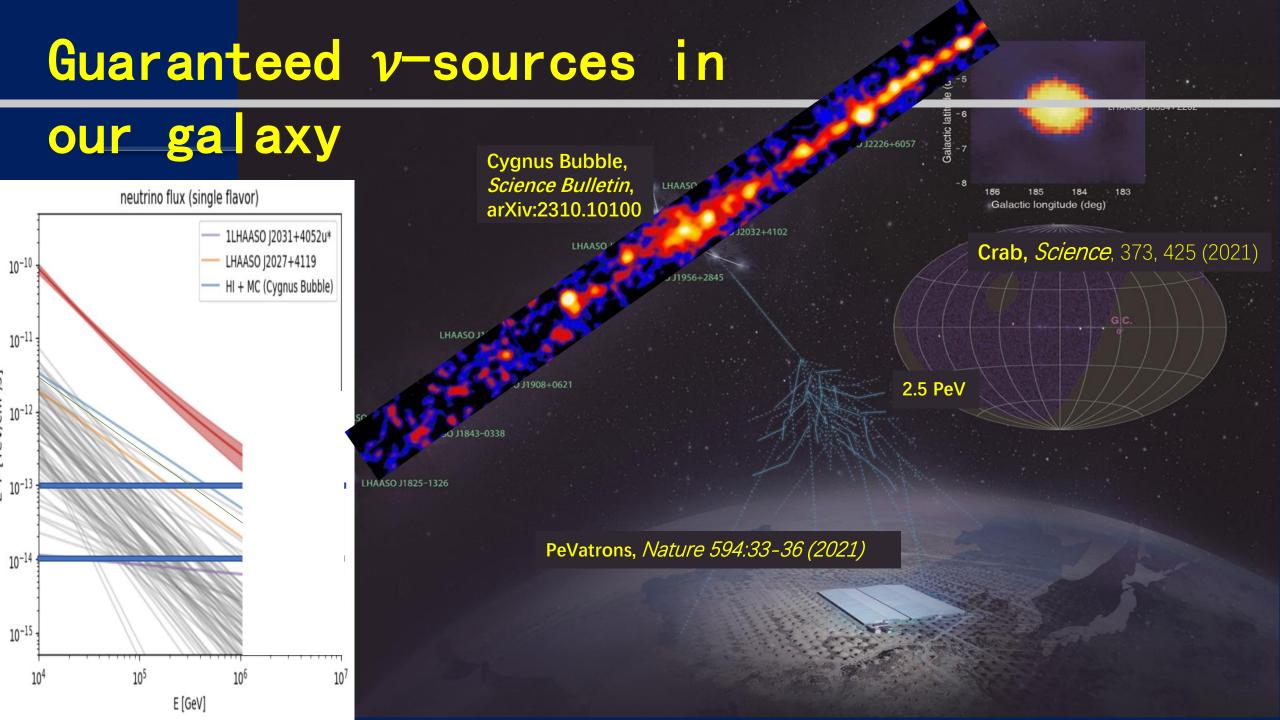
LHAASO vs. LACT (either two point sources or complex morphology)

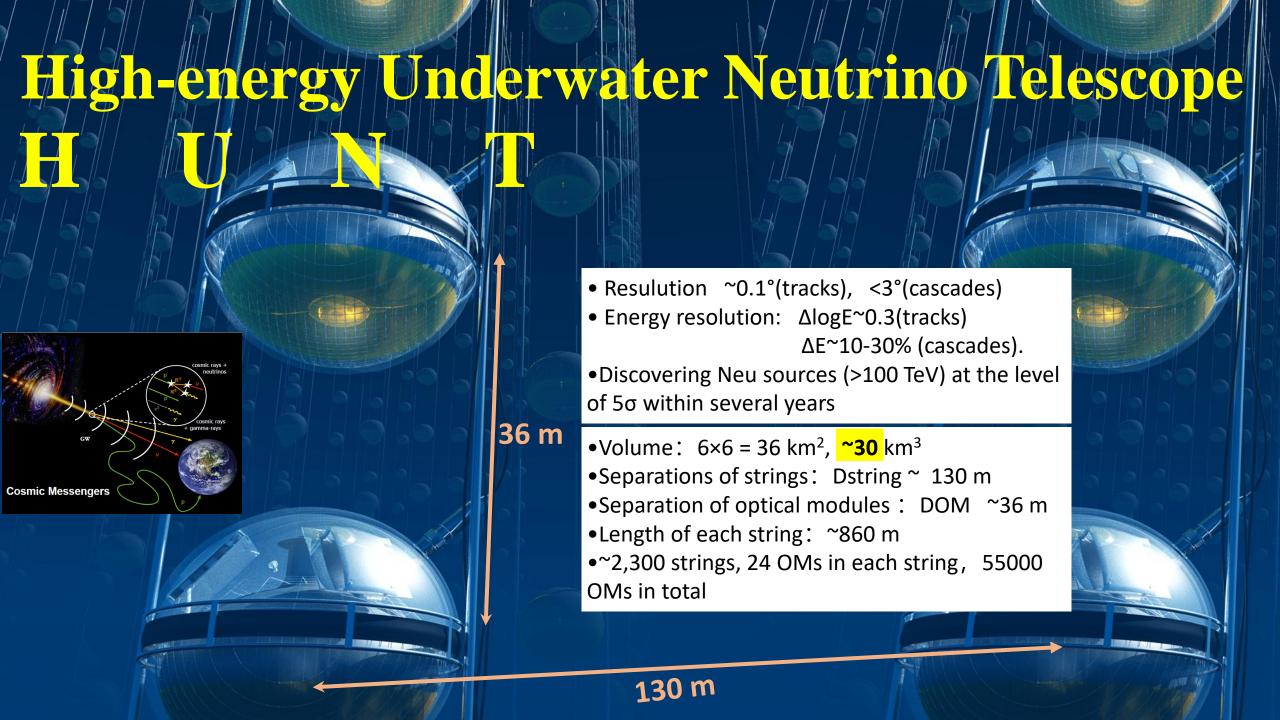




Expected flux measured in one year with the 4 Telescopes of LACT for different assumed cut-off energy $E_{\rm cut}$



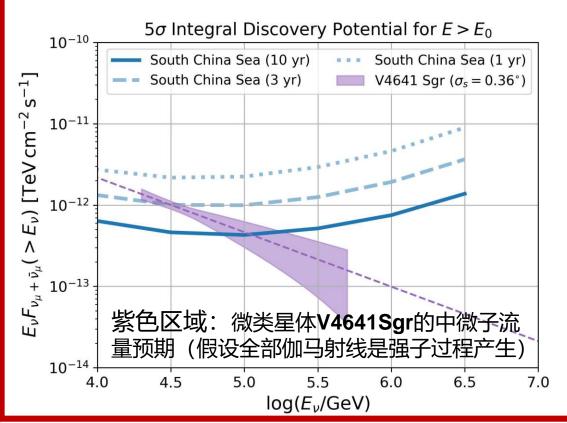


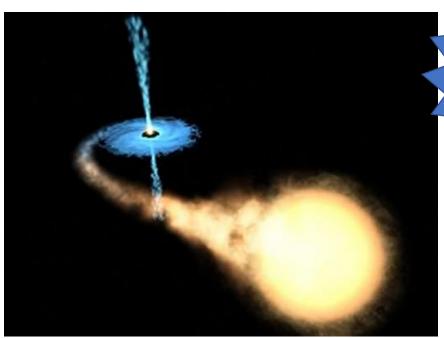


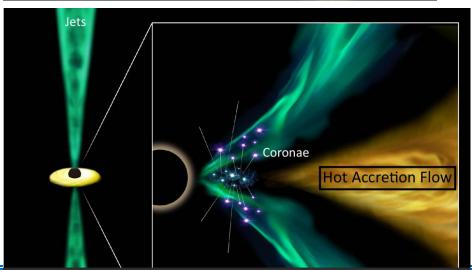


性能预期(确定第一个黑洞宇宙线加速器!最强源)3

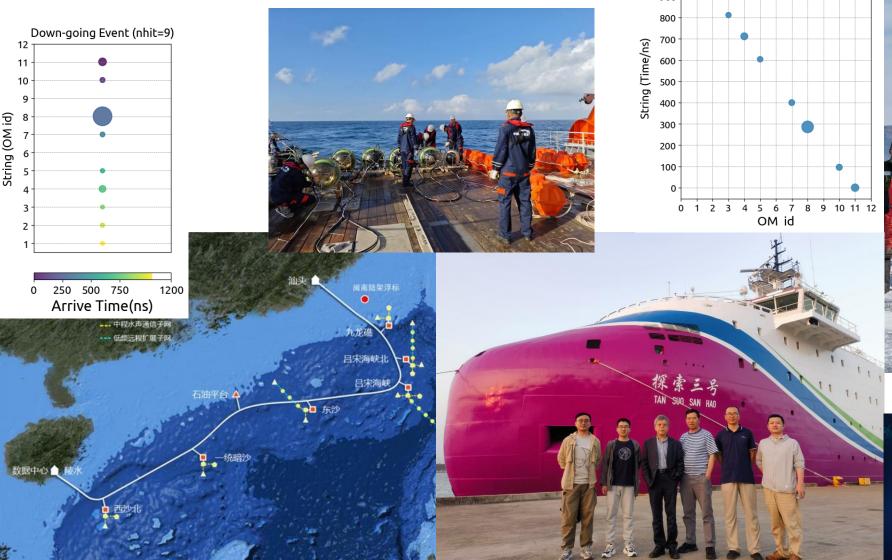
- 2年对10 TeV以上的缪中微子实现5σ观测;
- 10年对100 TeV以上缪中微子实现5σ观测。







Prototype tests at GVD at L. Baikal and in South China sea









- LHAASO has been making discovery of PeVatrons in the Milky Way
- Among the discoveries, the 1st Black-Hole PeVatron catalog unveiled the new potential CR accelerator population
- Cosmic Ray Super-PeVatron is found among the BH+jet systems
- SS433 is found a nearly perfect example of its kind, and they may supply sufficient CRs in the 'knee' region
- Proton/Helium spectra evidence the compact objects may be the major supplier of CR at/above the knee
- In future, the improved spatial resolution will bring the investigations for the mechanism much deep
- Detection of Neutrinos from the PeVatrons will put the last piece of puzzle in