

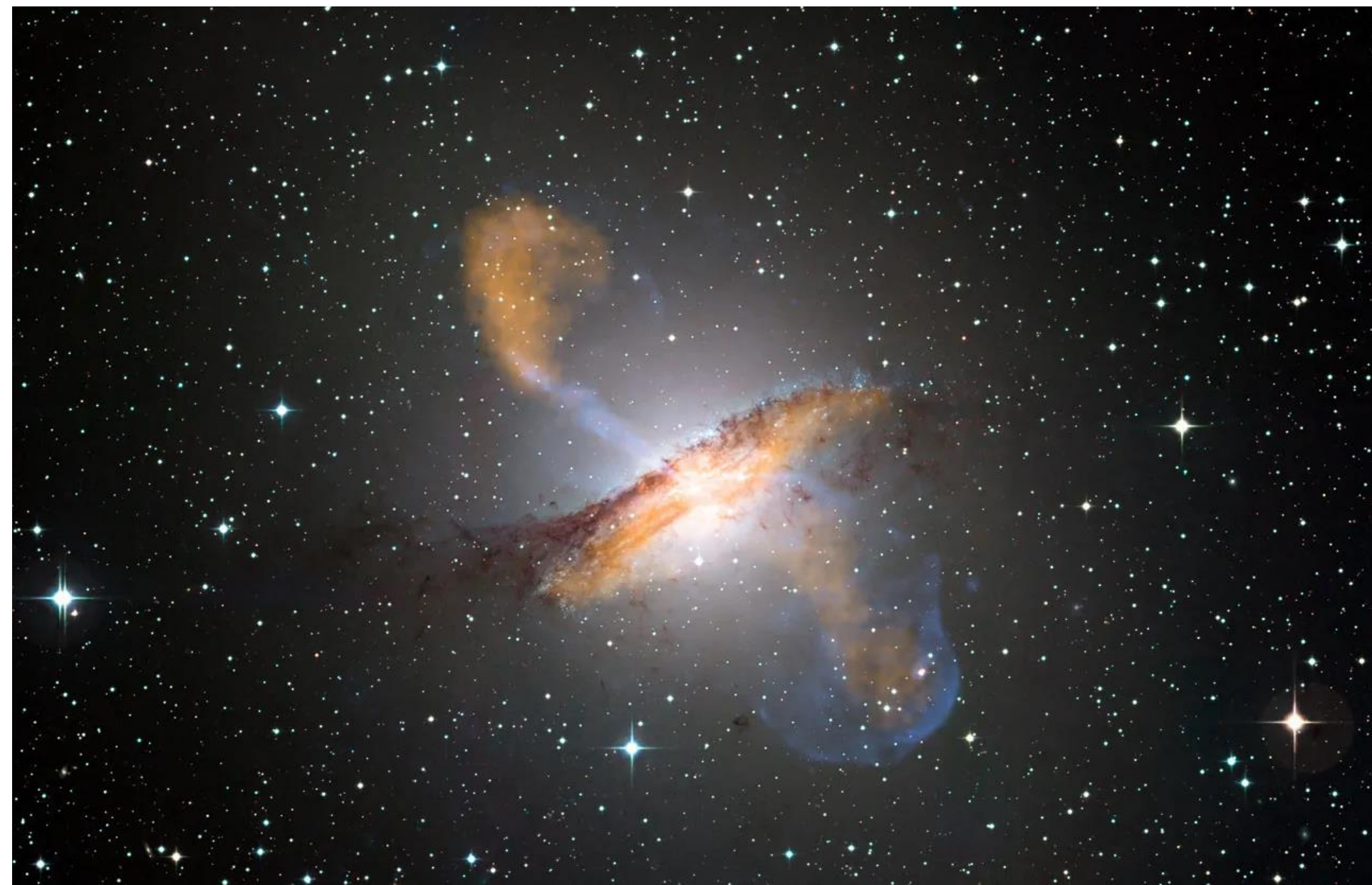
Kinetic Simulations of High-Field Fusion Plasmas

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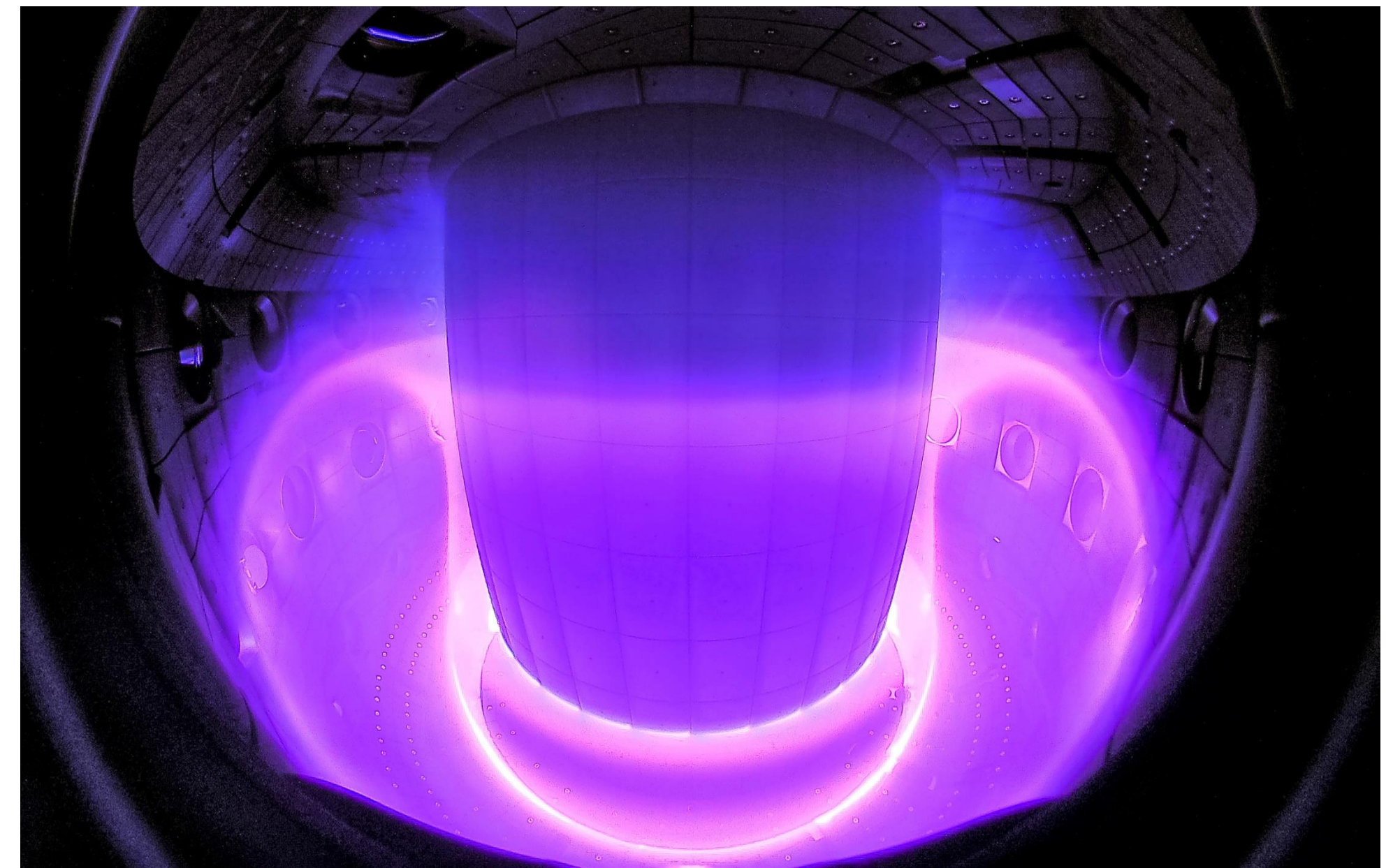
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Why study highly magnetized plasmas?

Astrophysical plasmas

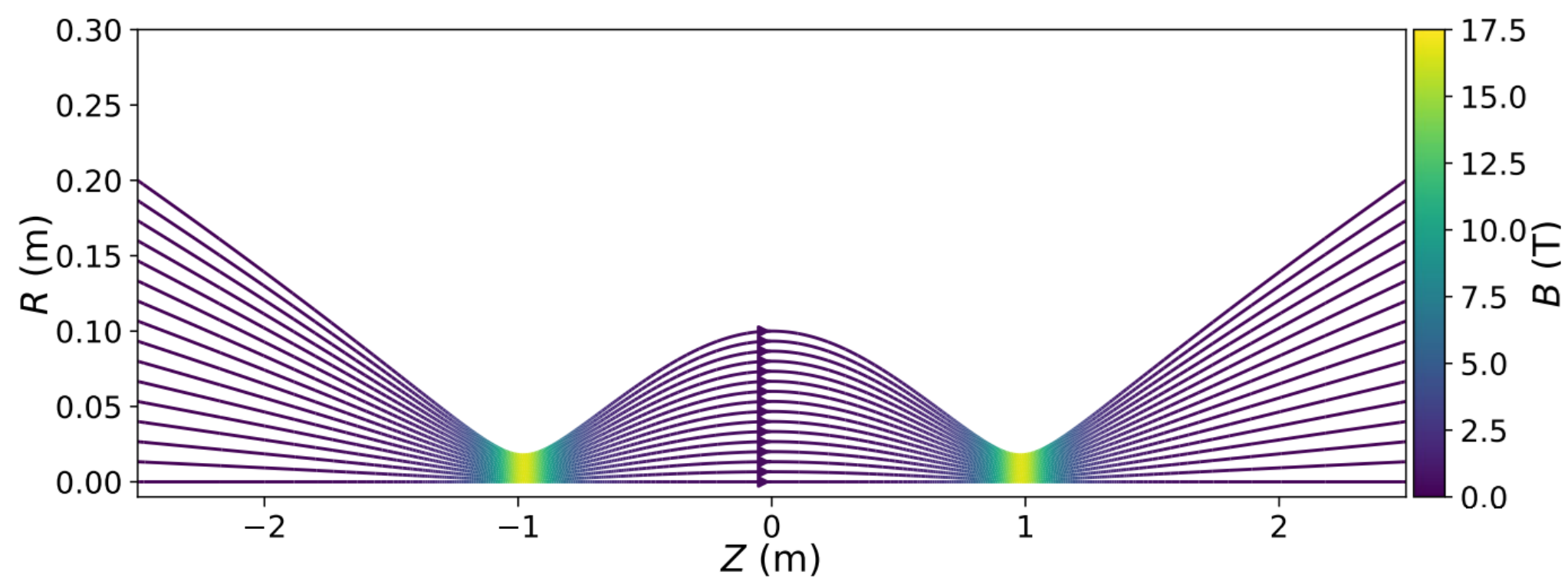
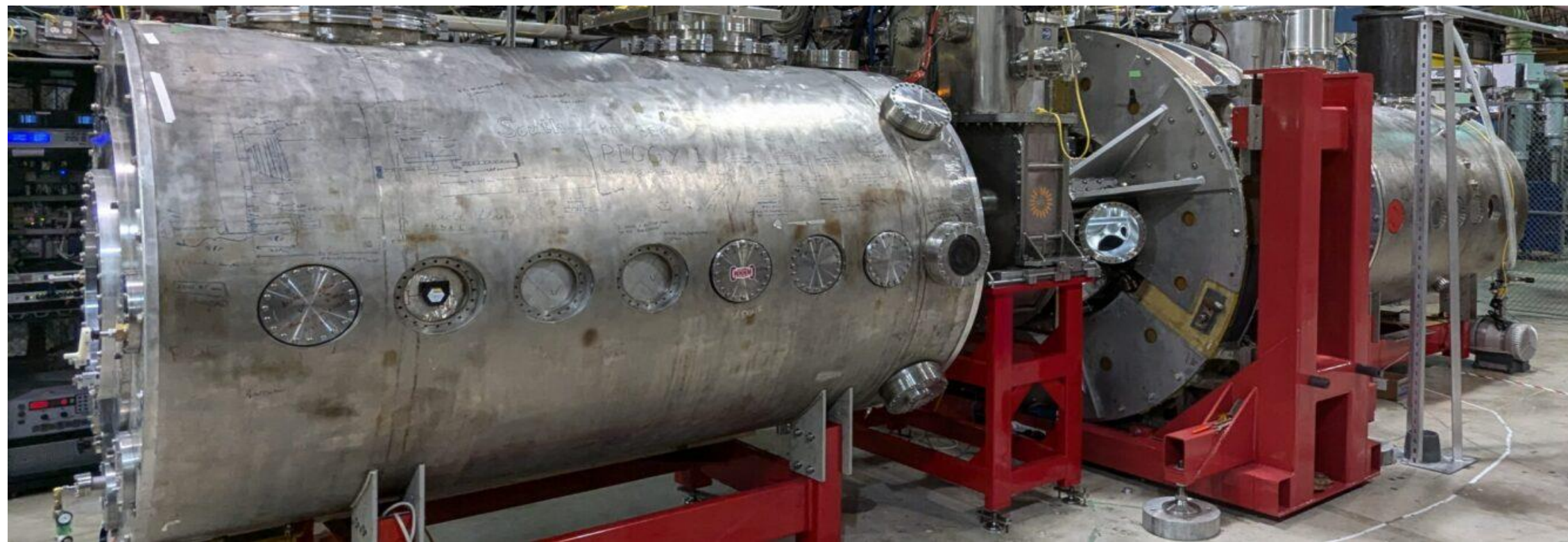


Fusion plasmas



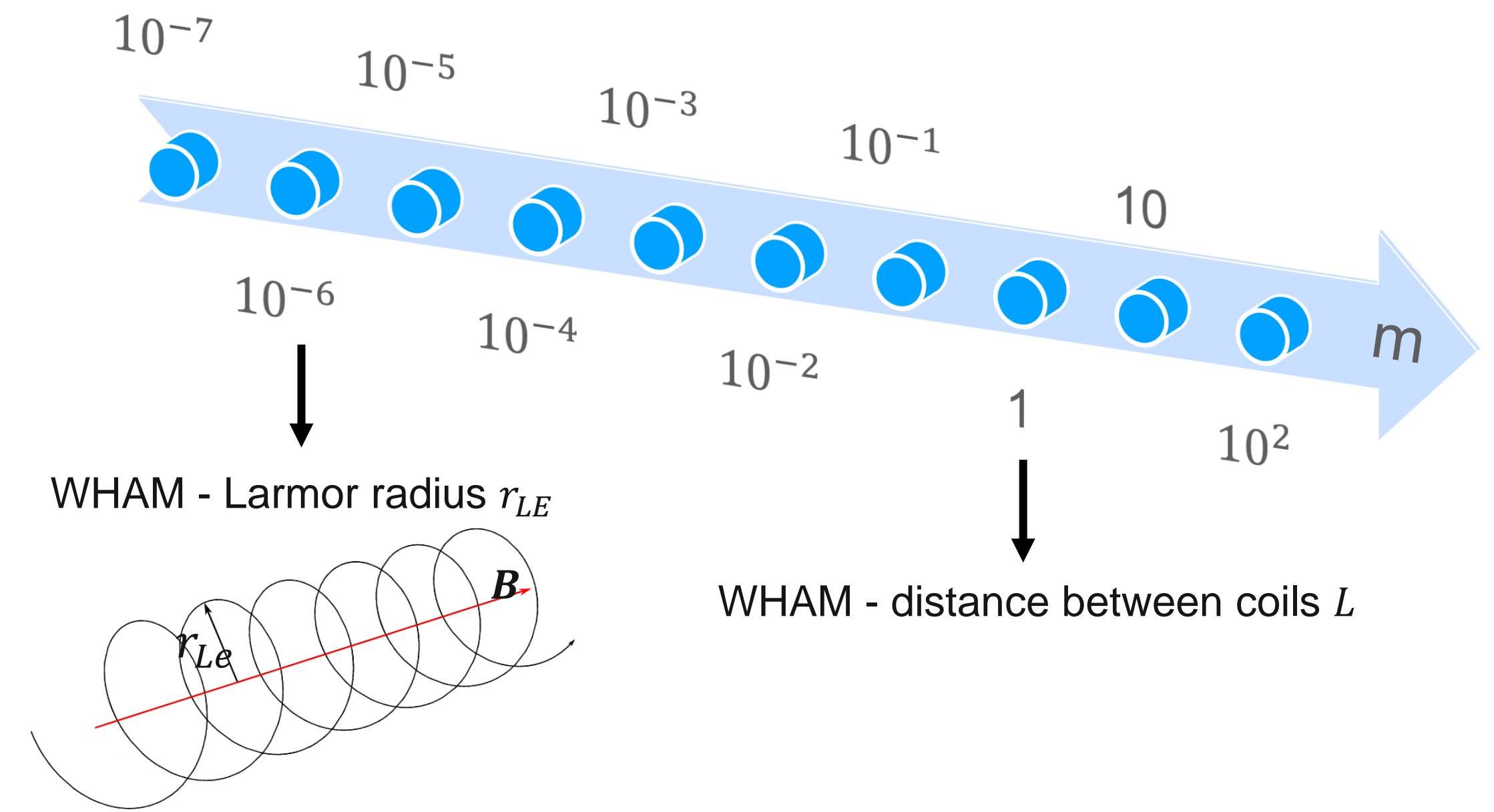
High-field plasmas offer fusion potential, but vast-scale separations pose challenges

WHAM, a high-field magnetic mirror for fusion

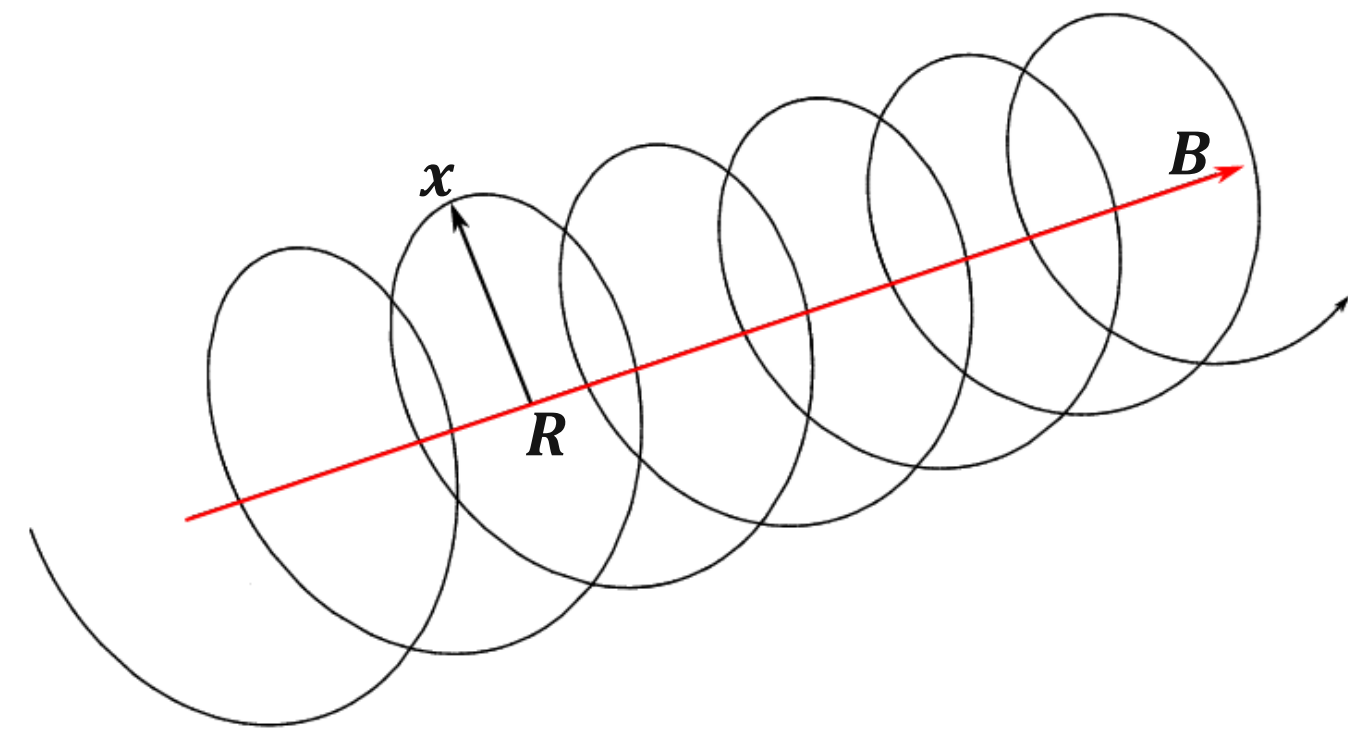


High-field fusion relies on kinetic simulations. However...

Challenges:
Complexity
Non-linearity
Interplay across vast scales



Average out Ω with the Guiding Centre Approximation (GCA)



$$\frac{d\mathbf{R}}{dt} = \frac{\mathbf{u}_{\parallel}}{\Gamma} + \mathbf{v}_E + \mathbf{v}_c,$$

$$\frac{du_{\parallel}}{dt} = \frac{q}{m} E_{\parallel} + u_{\parallel} \mathbf{v}_E \cdot (\mathbf{b} \cdot \nabla) \mathbf{b} + \Gamma \mathbf{v}_E \cdot (\mathbf{v}_E \cdot \nabla) \mathbf{b},$$

$$\frac{d\mu}{dt} = 0,$$

Avoiding resolving Ω , leads to a computational gain

WHAM – important scales

	Magnetic Field [T]	Gyro-frequency [s/rad]	Plasma Frequency [s/rad]
Formula	B	$\frac{1}{\Omega} = \frac{m_e c}{e B}$	$\frac{1}{\omega_{pe}} = \sqrt{\frac{m_e}{4\pi n_e e^2}}$
Centre $z=0\text{m}$	0.86	7e-12	3e-12
Coils $z=\pm 1\text{m}$	17	3e-13	6e-12

20^{D+1} times speed up