TOPOLOGICAL MODES IN 2D MAGNETIZED DIRAC MATERIALS

2nd Cycle Integrated Project in Engineering Physics Vasco Santos - 100372

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Vasco, what's really your project?

How am I approaching the problem?

- The problem is approached both theoretically, and numerically;
- The basis for this is an hydrodynamical model for the electron liquid;

What have I done already?

- Utilized a simplified, linearized´model;
- Obtained good analytical results for interface modes;
- Numerical simulations show good agreement with thẹ̃ theoretical predictions;

What do I plan to do in the future?

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- Improve the already existing framework for simulations;
- Attempt to arrive at new analytical solutions, using more comprehensive models, reaching new physics;

How am I doing it?

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How am i doing it?

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How am I doing it?

 $\sim V(y)$

 $\sim \Omega_B(y)$

 $\Omega_B(y) = \Omega_B \,\text{sgn}(y)$

- Only allows one bound state;
- Discontinuous function;
- May lead to nonphysical results;

- Allows a finite number of bound states;
- Continuous function, assures continuity of solutions;
- Other two functions can be seen as limits of this one;

 $\overline{V(y)}$ $\sim\!\!\Omega_B(y)$ $\Omega_B(y) = \frac{\Omega_B}{\epsilon} y$

- Allows an infinite number of bound states;
- Does not contemplate the stabilization of the magnetic field;
- Could only be useful for a very localized approximation;

What have I done?

Linearizing the equations yields:

$$
\frac{\partial n}{\partial t} + n_0 \frac{\partial u}{\partial x} + n_0 \frac{\partial v}{\partial y} = 0
$$
\n
$$
\frac{\partial u}{\partial t} + \frac{S^2}{n_0} \frac{\partial n}{\partial x} - \Omega_B v - \eta_o \nabla^2 v = 0
$$
\ncyclotron frequency: Ω_B
\n
$$
\frac{\partial v}{\partial t} + \frac{S^2}{n_0} \frac{\partial n}{\partial y} + \Omega_B u + \eta_o \nabla^2 u = 0
$$
\nEquilibrium density: n_0
\nPlasmon sound velocity: S

 $= e \sqrt{\frac{d_0 n_0}{\varepsilon m_0}}$

The only term that is retained from the stress tensor is the Hall viscosity term

What have I got?

Kelvin Mode

 η_o = 0

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What have I got?

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What have I got?

If you were not convinced, yes, there is a very good agreement between the theory, and the numerical results.

What will I do?

- Where are the Rossby modes, when considering the Hall viscosity?
	- Ans: They could exist as shown by numerical results, but more analytical work needs to be performed.
- Isn't the linearized model a very simplified regime?
	- Ans: Yes, and in the future, nonlinearities must be considered to have a more complete understanding of the system.
- What can be improved in the numerical simulations?
	- Ans: So far, some extra considerations regarding time scaling need to be reviewed, which that can be crucial to understand how the modes evolve.

Questions?

References

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Adapted from: Gurzhi, R. N., 1968 Sov. Phys. Usp. 11 255 Ho, D. Y. H. et al., Phys. Rev. B 97, 121404(R)

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Appendix - Interface Modes

The system changes, and now we can write it as:

$$
(\omega - Sk)Q + iSL_A = 0
$$

$$
(\omega + Sk)R + iSL_A = 0
$$

$$
\omega A + i\frac{1}{2}SL_AQ + i\frac{1}{2}SL_A = 0
$$

Appendix - What have I done?

We employ the changes of variables: (essential for chirality)

$$
q = \frac{S}{n_0}n + u \quad r = \frac{S}{n_0}n - u
$$

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We assume transversal modulation, and longitudinal propagation

$$
\begin{pmatrix} q(x, y, t) \\ r(x, y, t) \\ v(x, y, t) \end{pmatrix} = \text{Re}\left[\begin{pmatrix} Q(y) \\ R(y) \\ A(y) \end{pmatrix} e^{i(kx - \omega t)}\right]
$$

Majda, A. , Courant Institute of Mathematical Sciences, 2003

Appendix - Inviscid Regime General Modes

Solving the top two equations for A, and plugging in the third one, we obtain a Schödinger equation for the amplitude A:

$$
-\frac{\partial^2 A}{\partial y^2} + \left[\frac{\Omega_B^2(y)}{S^2} + \frac{k}{\omega} \frac{\partial \Omega_B(y)}{\partial y} \right] A = \frac{(\omega^2 - S^2 k^2)}{S^2} A
$$

Depending on the modulation assumed for the transition, the system will present different sets of solutions!

Appendix - Inviscid Regime Chiral Modes

We force R=0, to ensure a unidirectional propagation, making the modes chiral

$$
(\omega - Sk)Q + iSL_A = 0
$$

\n
$$
\begin{array}{ccc}\n& & \text{Trivial solution (A=0):} \\
& & \text{Kelvin mode} \\
& & \omega A + iS\frac{1}{2}L_{+}Q = 0\n\end{array}
$$
\n
\n
$$
\begin{array}{ccc}\n& & \text{Trivial solution (A=0):} \\
& & \text{Kelvin mode} \\
& & \text{Non-trivial solution:} \\
& & \text{Yanai mode}\n\end{array}
$$

Appendix - Hall-Viscous Regime Chiral Modes

In the Hall-viscous regime, we will only search for chiral modes

$$
(\omega - Sk)Q + iS\mathcal{L}_{-}A = 0
$$

\n
$$
iS\mathcal{L}_{+}A = 0
$$

\n
$$
\omega A + i\frac{1}{2}S\mathcal{L}_{+}Q = 0
$$

\n
$$
\mathcal{L}_{\pm} = S\frac{\partial}{\partial y} \pm \Omega_{B}(y) \mp \eta_{o}(y)k^{2} \pm \eta_{o}(y)\frac{\partial^{2}}{\partial y^{2}}
$$

\nNon-trivial solution:
\n
$$
i\frac{\partial^{2}}{\partial y^{2}}
$$

Appendix - Magnetic Inversions

The electrons in the 2D hydrodynamic regime are only sensitive to out-of-plane variations of the magnetic field. Therefore, we only consider the z-component of the field.

Appendix - Magnetic Inversions

$$
\Omega_B(y) = \Omega_B \operatorname{sgn}(y)
$$

$$
\frac{\partial^2 A}{\partial y^2} + \frac{k \Omega_B}{\omega} \delta(y) A = \frac{(\omega^2 - S^2 k^2 - \Omega_B^2)}{S^2} A
$$

$$
\frac{\partial}{\partial B(y)} = \Omega_B \tanh\left(\frac{y}{\epsilon}\right)
$$
\n
$$
\frac{\partial^2 A}{\partial y^2} - \left[\frac{k\Omega_B}{\epsilon\omega} + \frac{\Omega_B^2}{S^2}\right] \operatorname{sech}^2\left(\frac{y}{\epsilon}\right) A = \frac{(\omega^2 - S^2 k^2 - \Omega_B^2)}{S^2} A
$$
\n
$$
\frac{\left(\frac{k\Omega_B}{\epsilon\omega} + \frac{\Omega_B^2}{S^2}\right) \operatorname{sech}^2\left(\frac{y}{\epsilon}\right) A}{\frac{\Omega_B(y)}{\omega^2} - \frac{\partial^2 A}{\partial y^2} - \left(\frac{\Omega_B y}{S\epsilon}\right)^2 A} = \left[\frac{\omega^2}{S^2} - k^2 - \frac{\Omega_B k}{\epsilon\omega}\right] A
$$

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We map the problem as an eigenvalue problem

$$
\Psi = \begin{pmatrix} n(x, y, t) \\ u(x, y, t) \\ v(x, y, t) \end{pmatrix} = \begin{pmatrix} N \\ U \\ V \end{pmatrix} e^{i(k_x x + k_y y - \omega t)}
$$

$$
\omega \Psi = \hat{\mathcal{H}} \Psi \qquad \hat{\mathcal{H}} = \begin{bmatrix} 0 & i k_x & i k_y \\ i k_x & 0 & -(\Omega_B - \eta_o k^2) \\ i k_y & (\Omega_B - \eta_o k^2) & 0 \end{bmatrix}
$$

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Appendix - Eigenstates

 $\sigma = \eta_o(y)/\Omega_B(y)$

General
\n**Eigenstates:**
\n
$$
\Psi_0 = \begin{bmatrix} \Omega_B(1-\sigma k) \\ -ik_y \\ ik_x \end{bmatrix}, \quad \Psi_{\pm} = \begin{bmatrix} \pm \frac{k_x}{k} \pm \frac{i\Omega_B k_y}{k|\omega_{\pm}|} (1-\sigma k^2) \\ \pm \frac{k_y}{k} \mp \frac{i\Omega_B k_x}{k|\omega_{\pm}|} (1-\sigma k^2) \end{bmatrix}
$$
\n**Poles in the**
\n**Q**
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Appendix - Berry Curvature

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Appendix - Berry Curvatures

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Appendix - Inviscid Regime

Trivial Modes - Rossby and Poincaré Modes

 $C_{+} = +2$

Profile:

$$
A(y) = P_l^{\mu} \left(\tanh\left(\frac{y}{\epsilon}\right) \right)
$$

\n
$$
\mu = \frac{\epsilon \sqrt{k^2 S^2 + \Omega^2 - \omega^2}}{S}
$$

\n
$$
= \frac{\epsilon \sqrt{k^2 S^2 + \Omega^2 - \omega^2}}{S}
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= \frac{\epsilon \sqrt{k^2 S^2 + \Omega^2 - \omega^2}}{S}
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= \frac{\epsilon \sqrt{k^2 S^2 + \Omega^2 - \omega^2}}{S}
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= -\frac{\epsilon \sqrt{k^2 S^2 + \Omega^2 - \omega^2}}{S}
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= -\frac{\epsilon \sqrt{k^2 S^2 - \omega^2 + \Omega^2}}{S}
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= -\frac{\epsilon \sqrt{k^2 S^2 - \omega^2 + \Omega^2}}{S}
$$

\n
$$
= -\frac{\epsilon \sqrt{k^2 S^2 - \omega^2
$$

Polyanin, A. D., Zaitsev, V. F., (2004) Handbook of Nonlinear Partial Differential Equations,Chapman and Hall/CRC

Appendix - Rossby Modes

Cosme, P., Terças, H., & Santos, V. (2023), in proceedings of 2023 IEEE NMCD Conference

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Appendix - Inviscid Regime Chiral Modes - Kelvin Mode

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Appendix - Kelvin Mode

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Appendix - Inviscid Regime

Chiral Modes - Yanai Modes

Profile:

$$
A(y) = A_0 \operatorname{sech}^{\alpha}\left(\frac{y}{\epsilon}\right) \text{ with } \alpha = \frac{\Omega_B \epsilon}{S}
$$

$$
Q(y) = i \frac{2A_0 \Omega_B}{\omega - Sk} \operatorname{sech}^{\alpha}\left(\frac{y}{\epsilon}\right) \tanh\left(\frac{y}{\epsilon}\right)
$$

Dispersion relation:

$$
\omega_{\pm}=\frac{Sk}{2}\pm\frac{S}{2}\sqrt{k^2+4\frac{\Omega_B}{S\epsilon}}
$$

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Appendix - Yanai Mode

Cosme, P., Terças, H., & Santos, V. (2023), in proceedings of 2023 IEEE NMCD Conference

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Appendix - Hall-Viscous Regime

Chiral Modes - Kelvin Mode

Polyanin, A. D., Zaitsev, V. F., (2004) Handbook of Nonlinear Partial Differential Equations,Chapman and Hall/CRC

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Appendix - Viscid Kelvin Modes

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Appendix - Hall-Viscous Regime

Chiral Modes - Yanai Mode

Polyanin, A. D., Zaitsev, V. F., (2004) Handbook of Nonlinear Partial Differential Equations,Chapman and Hall/CRC

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We consider the change of variables to z, and the ansatz w(z)

$$
z = \operatorname{sech}^2(y/\epsilon) \qquad \qquad Q(z) = (-z)^{\lambda \epsilon/2} w(z)
$$

$$
\eta_o \lambda^2 - \lambda S - \eta_o k^2 + \Omega_B = 0
$$

$$
z(1-z)w''(z) - \frac{\lambda\epsilon}{2} \cdot \frac{\lambda\epsilon + 1}{2}w(z) + \left[1 + \lambda\epsilon - \frac{S\epsilon}{2\eta_o} - z\left(\frac{\lambda\epsilon}{2} + \frac{\lambda\epsilon + 1}{2} + 1\right)\right]w'(z) = 0
$$

$$
Q(z) = (-z)^{\lambda\epsilon/2} {}_2F_1\left(\frac{\frac{\epsilon\lambda}{2}}{\frac{S\epsilon}{2\nu_o} + \frac{1}{2}}; 1 - z\right)
$$

Polyanin, A. D., Zaitsev, V. F., (2004) Handbook of Nonlinear Partial Differential Equations,Chapman and Hall/CRC T. Morita, (1996) Interdisciplinary Information Sciences

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Appendix - Ordinary Hyperbolic Function

 B lue: Q(y); Green: $\text{Red: } {}_2F_1(1-z);$

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Appendix - Viscid Yanai Modes

Cosme, P., Terças, H., & Santos, V. (2023), in proceedings of 2023 IEEE NMCD Conference

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DEDALUS is a library that uses spectral methods:

$$
a(x) = \sum_{k=0}^{N} a_k \phi_k(x)
$$

Burns, K. J. , Vasil, G. M., Oishi, J. S., Lecoanet, D., Brown, B. P., Phys. Rev. Res. (2020)

Appendix - Numerical Dispersion Relation

The problem is projected as an eigenvalue problem for ω .

 $\omega M \cdot \Psi + L \cdot \Psi = 0$ $-i\omega n + iku + J\frac{dv}{dx} + \text{lift}(\tau_1) = 0$ $\begin{pmatrix} n(x, z, t) \\ u(x, z, t) \\ v(x, z, t) \end{pmatrix} = \begin{pmatrix} N(z) \\ U(z) \\ V(z) \end{pmatrix} e^{i(kx - \omega t)}$ $-i\omega u + ikn - \alpha zv - \beta \left(-k^2v + J\frac{d^2v}{dz^2}\right) = 0$ $-i\omega v + J\frac{dn}{dz} + \alpha z u + \beta \left(-k^2 u + J\frac{d^2 u}{dz^2} \right) + \text{lift}(\tau_2) = 0$ $\frac{d}{dt} = -i\omega$, $\frac{d}{dx} = ik$ $D_u - J\frac{du}{dz} + \text{lift}(\tau_3) = 0$ $D_v - J\frac{dv}{dz} + \text{lift}(\tau_4) = 0$ $J = dz/dy$ $\alpha = \frac{\Omega_B \epsilon}{S}$ $\beta = \frac{\eta_o \Omega_B}{S^2}$ $z \in [-1, 1]$ $n(z = \pm 1) = 0$

Burns, K. J. , Vasil, G. M., Oishi, J. S., Lecoanet, D., Brown, B. P., Phys. Rev. Res. (2020)

 $\frac{dn}{dz}(z=\pm 1)=0$

The residual of the calculations is fitted by the program as **a** polynomial function with a coefficient τ .

$$
\mathcal{L}(x_i) + \tau P(x_i) = f(x_i)
$$

The lift operator functions as a n-order "anti-derivative".

Tau variables are employed to force boundary conditions, if necessary.

Burns, K. J. , Vasil, G. M., Oishi, J. S., Lecoanet, D., Brown, B. P., Phys. Rev. Res. (2020)

Appendix - Mode Simulation

This is set as an initial value problem, where we provide the initial configuration:

$$
M \cdot \partial_t \Psi + L \cdot \Psi = F(\Psi, t)
$$

$$
\Psi(x, y, 0) = \begin{pmatrix} n(x, y, 0) \\ u(x, y, 0) \\ v(x, y, 0) \end{pmatrix} = \text{Re}\left[\begin{pmatrix} N(y) \\ U(y) \\ V(y) \end{pmatrix}e^{ikx}\right]
$$

$$
\frac{\partial n}{\partial t} + n_0 \frac{\partial u}{\partial x} + n_0 \frac{\partial v}{\partial y} = 0
$$

$$
\frac{\partial u}{\partial t} + \frac{S^2}{n_0} \frac{\partial n}{\partial x} - \Omega_B v - \eta_o \nabla^2 v = 0
$$

$$
k = n\pi/L
$$

$$
\frac{\partial v}{\partial t} + \frac{S^2}{n_0} \frac{\partial n}{\partial y} + \Omega_B u + \eta_o \nabla^2 u = 0
$$

Burns, K. J. , Vasil, G. M., Oishi, J. S., Lecoanet, D., Brown, B. P., Phys. Rev. Res. (2020)