

The Lunar Ionising Radiation Environment A Benchmark Model

January 28, 2025

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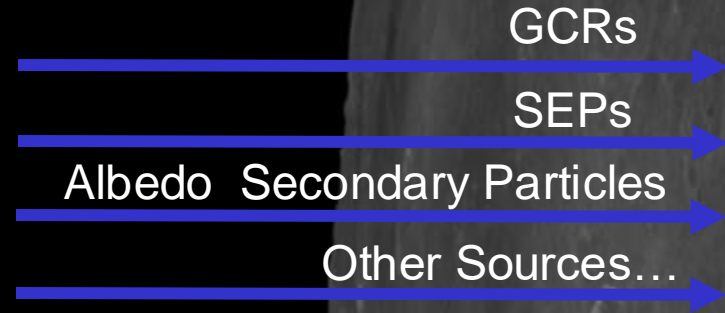
LABORATÓRIO DE INSTRUMENTAÇÃO
E FÍSICA EXPERIMENTAL DE PARTÍCULAS
partículas e tecnologia



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Lunar Radiation Environment

Particle	Energy (eV)
GCRs	10^8 to 10^{20}
SEPs	10^8 to 10^9
Albedo Particles	Up to 10^8
Other sources	



SEPs fluxes can exceed background GCRs fluxes by factors of 10^3 or more!

Figure 2: Juice NavCam view of the Moon
Credits: ESA/Juice/NavCam
Acknowledgements: Airbus

From Mars to the Moon

dMEREM

dMEREM
(Mars Model)

Adapting to the Moon:

Developed by LIP's **SpaceRad** group
using **Geant4**.

Simulates radiation environment at different
Mars locations.

Figure 3: "Mars true-color generated image using OSIRIS"

CREDIT: ESA & MPS for OSIRIS Team
MPS/UPD/LAM/IAA/RSSD/INTA/UPM/DASP/IDA

- Replace Mars' atmospheric/soil models with lunar-specific data.
- Include lunar-specific radiation sources.
- Validate with existing mission data.



From Mars to the Moon

dLEREM

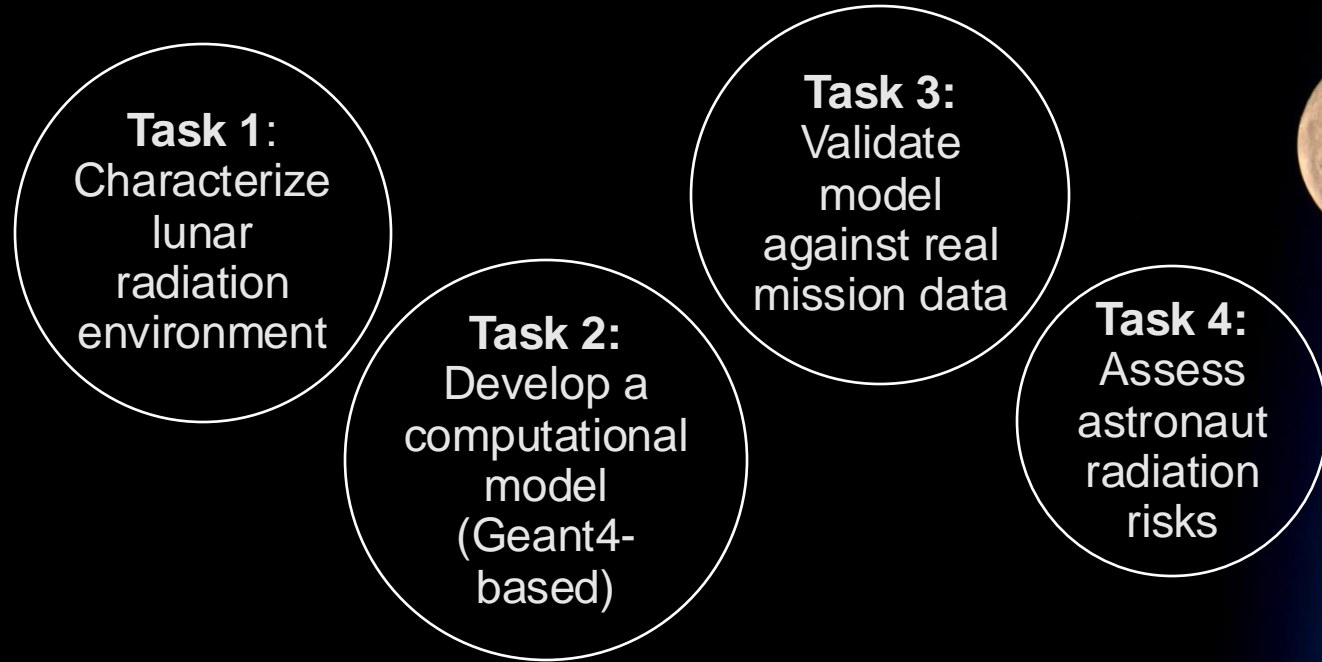


Figure 4: “Full Moon as photographed from on board the International Space Station
CREDIT: NASA/Astronaut Jeff Williams

From Mars to the Moon

dLEREM

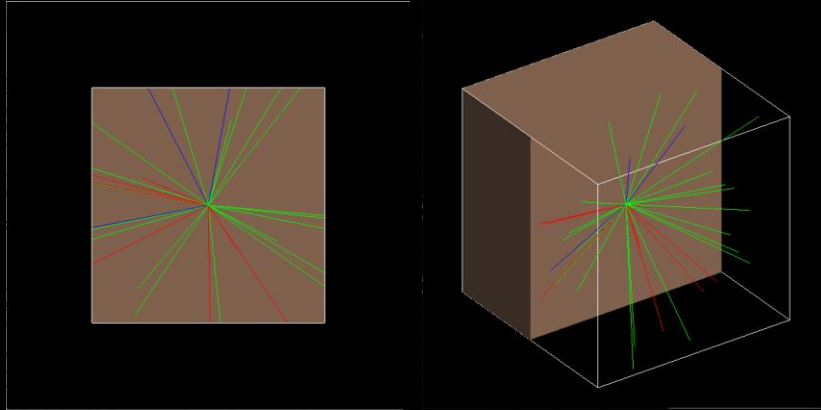


Figure 5: Geant4 Simulation Example with Moon Regolith and Proton Flux of 100 GeV

```

Position - x (mm) : -9.120109676571092e-12-9.120132388436999e-12
Position - y (mm) : 1.083960349821405e-111.083963049215006e-11
Position - z (mm) : 1.479007345873014e-111.479011029867858e-11
Global Time (ns) : 0.01143614791999963 0.01143614791999963
Local Time (ns) : -7.00145453324725e-14-7.001471969092333e-14
Proper Time (ns) : -1.534098534712316e-14-1.534102355087841e-14
Momentum Direct - x : 0.4453226691461333 0.4453226691461333
Momentum Direct - y : -0.5292833130974934 -0.5292833130974934
Momentum Direct - z : -0.7221820371769878 -0.7221820371769878
Momentum - x (MeV/c) : 1.013318641254679 1.013318641254679
Momentum - y (MeV/c) : -1.204368618141755 -1.204368618141755
Momentum - z (MeV/c) : -1.643303993605095 -1.643303993605095
Total Energy (MeV) : 2.332142006503581 2.332142006503581
Kinetic Energy (MeV) : 1.821143096503581 1.821143096503581
Velocity (mm/ns) : 292.5074555687947 292.5074555687947
Volume Name : Regolith Regolith
Safety (mm) : 5e-10 5e-10
Polarization - x : 0 0
Polarization - y : 0 0
Polarization - Z : 0 0
Weight : 1 1
Step Status : AlongStep Proc. AlongStep Proc.
Process defined Step: eIoni eIoni
-----
++List of secondaries generated (x,y,z,kE,t,PID): No. of secondaries = 0
**PostStepDoIt (after all invocations):
++List of invoked processes
1) Transportation

**G4Step Information
Address of G4Track : 0x139f1d3f0
Step Length (mm) : -5.10009201886219e-17
Energy Deposit (MeV) : 0
-----
StepPoint Information PreStep PostStep
-----
Position - x (mm) : -9.120109676571092e-12-9.120132388436999e-12
Position - y (mm) : 1.083960349821405e-111.083963049215006e-11
Position - z (mm) : 1.479007345873014e-111.479011029867858e-11
Global Time (ns) : 0.01143614791999963 0.01143614791999963
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Momentum - z (MeV/c) : -1.643303993605095 -1.643303993605095

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Objectives of dLEREM

Preliminary Validation & Expected Results

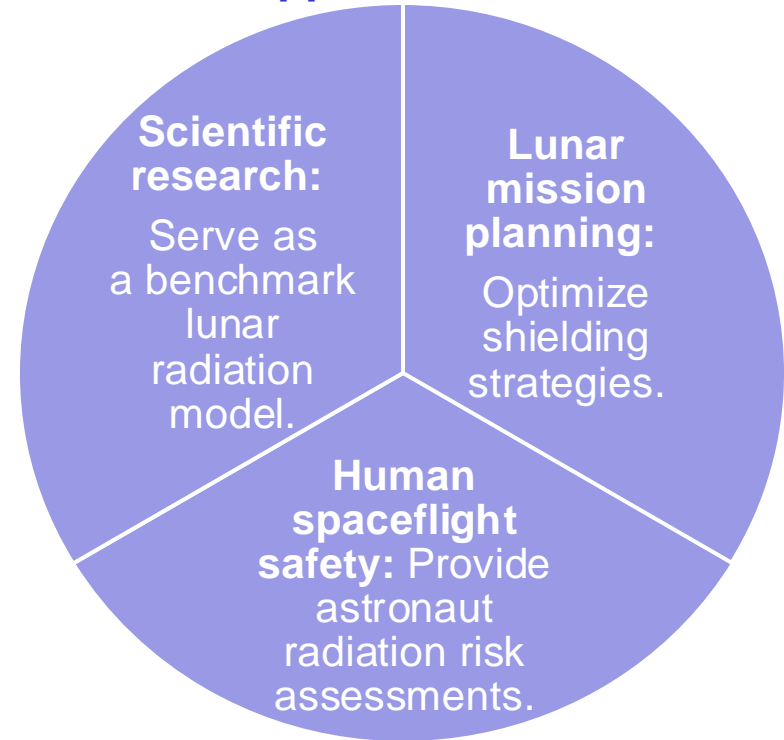
Validation Approach: Compare simulated dose rates and spectra to:

- CRaTER (LRO)
- Apollo radiation measurements
- Chandrayaan RADOM instrument
- JUICE RADEM August Earth-Moon Fly by

Expected Findings: Benchmark dLEREM against real lunar measurements.

Assess Accuracy: in predicting realistic lunar exposure conditions.

Applications:



Conclusion

Building a Safer Future for Space Exploration

Unify fragmented radiation data for better **understanding** and **accessibility**.

Figure 6: LUNA recreates the Moon's surface on Earth, located next to ESA's Astronaut Centre (EAC) in Cologne, Germany.
CREDITS: ESA-L. Breggion



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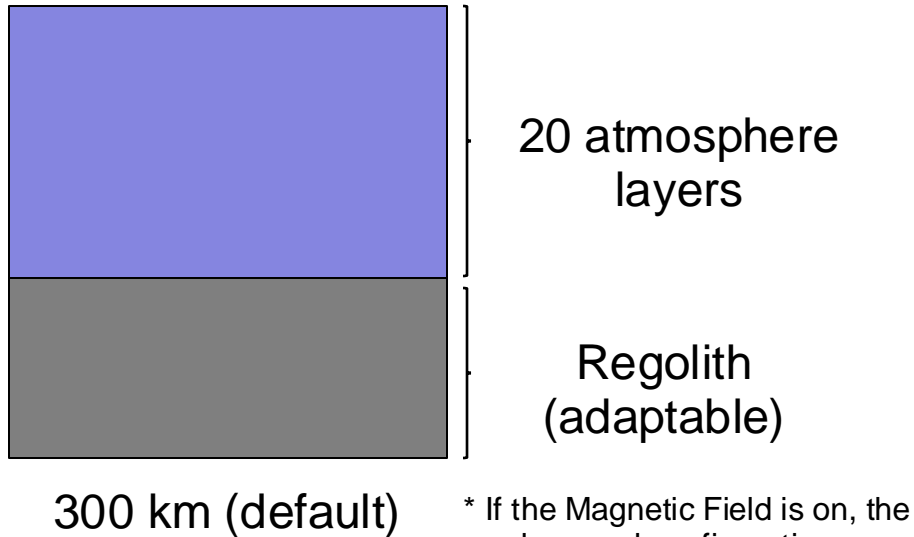
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Extra Slides

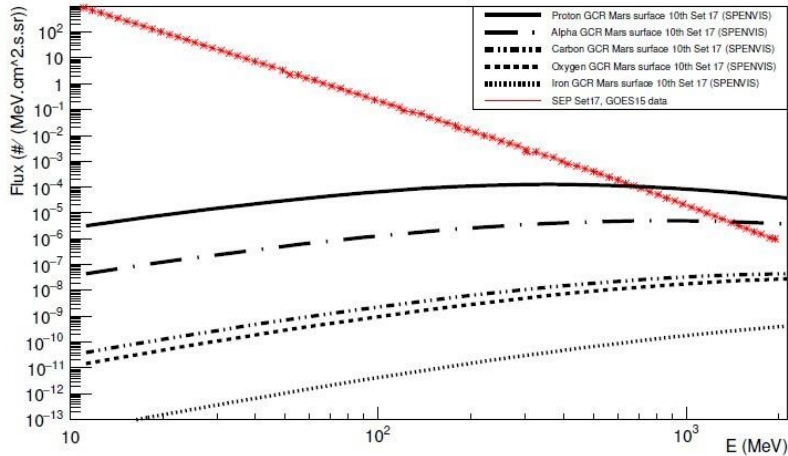
dMEREM



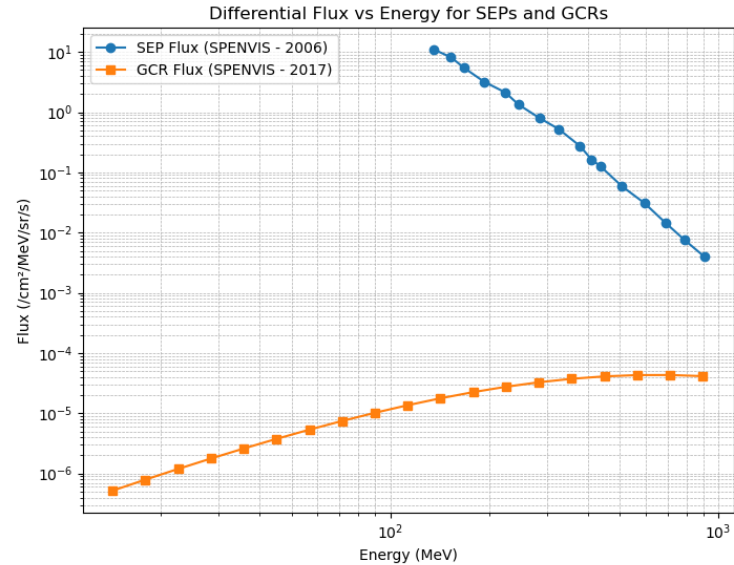
* If the Magnetic Field is on, the values and configuration are different

- ❑ Geometry Definition and Materials;
- ❑ Primary Particle Generation;
- ❑ Event Generation & Simulation;
- ❑ Physics Processes & Interactions;
- ❑ Sensitive Detectors & Scoring Mechanisms;
- ❑ Tracking & Data Collection;
- ❑ Output & Visualization.

Radiation Spectra Simulation



Differential flux as a function of kinetic energy (GCR)



Comparative flux spectra of GCRs and SEPs extracted from SPENVIS