Jet tomography from large to small systems

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jets as a tool to probe formation of QCD matter

The origin of complex matter



The origin of complex matter



How to probe matter?



A tool to probe the QGP formation



The State of the Art

Matter branch:

- the latter stages (th+ex)
- **G** the early stages (some th)
- main tool: hydrodynamics



compromises the success of the ongoing and future experimental programs

Jet branch:

- jets at latter stages (th+ex)
- **some** jet tomography

main tool: perturbative QCD

The State of the Art

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jets in non-equilibrium/evolving QCD matter (**fluctuating evolving** matter)



formation of QCD matter in smaller systems?

formation of complex QCD matter

Why now?



jets in hydrodynamic matter (e.g. our works during the last year)

evolving QCD matter at the LHC, RHIC, and EIC

collectivity in small systems vs. no jet quenching

a clear big problem in the field to be solved

the upcoming large and small system experiments (e.g. sPHENIX and O+O at the LHC)

the LHC Run 3 (2022 - 2025) the sPHENIX program at RHIC (2023 - 2025) the LHC Run 4 (2027+) the EIC experiment (2030+)

Jets in evolving matter

What do we have?

- Jets see the matter in HIC (and DIS) at multiple scales, and essentially X-ray it;
- Current theory is based on multiple simplifying assumptions: static matter, no fluctuations, etc;

What is missing?

- The coupling of jets to the flow, to the structure (matter anisotropy), to the transverse fields during initial stages, to the fluctuations, etc.
- An updated parton evolution equation needed for most modern simulations of jets in QCD matter;
- New jet observables sensitive to the medium evolution;
- Coupled simulations of matter and jets for quantitative phenomenology;

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- expertise at LIP

expertise

at LIP

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 - Coupled simulations of matter and jets for quantitative phenomenology;

R. Baier et al, NPB, 1997 B. G. Zakharov, JETP, 1997 R. Baier et al, NPB, 1998 M. Gyulassy et al, NPB, 2000 M. Gyulassy et al, NPB, 2001

Jet quenching formalisms



Jets in evolving matter



AS, M. Sievert, I. Vitev, PRD, 2021 J. Barata, AS, C. Salgado, PRD, 2022 C. Andres, F. Dominguez, AS, CS, arxiv, 2022 J. Barata, AS, X.-N. Wang, arxiv 2022 J. Barata, X. Mayo, AS, CS, 2022 (?)

Color potential

$$\begin{aligned} x_i & a_i(q) \\ q \\ & & \\ \hline q \\ \hline p_{s,i} & - q \\ \hline p_{s,i} & - q \\ \end{aligned}$$

$$A^{\mu a}(q) = \sum_i (ig \, t_i^a) \, e^{iq \cdot x_i} \, (2p_{s\,i} - q)_\nu \frac{-ig^{\mu\nu}}{q^2 - \mu_i^2 + i\epsilon} \, (2\pi) \, \delta\Big((p_{s\,i} - q)^2 - \overset{\mathsf{large}}{M^2}\Big) \, .$$

$$\bigvee (q^2) - \text{the Gyulassy-Wang potential}$$

$$iM_{1}(p) = \int \frac{d^{4}q}{(2\pi)^{4}} \left[ig t^{a}_{\text{proj}} A^{\mu a}_{\text{ext}}(q) (2p-q)_{\mu} \right] \left[\frac{i}{(p-q)^{2} + i\epsilon} \right] J(p-q)$$

$$gA^{\mu a}_{ext}(q) = \sum_{i} e^{iq \cdot x_{i}} t^{a}_{i} u^{\mu}_{i} v_{i}(q) (2\pi) \delta \left(q^{0} - \vec{u}_{i} \cdot \vec{q}\right)$$

$$\text{the fluid velocity}$$

Medium averaging



Gradient expansion



$$\mathcal{P}^{(0)}(\boldsymbol{r}, L; \boldsymbol{r}_0, 0) = e^{-\mathcal{V}(\boldsymbol{r})L} \delta^{(2)}(\boldsymbol{r} - \boldsymbol{r}_0)$$

$$\mathcal{V}(\boldsymbol{q}, z) \equiv -\mathcal{C} \rho(z) \left(\left| v(q_{\perp}^2) \right|^2 - \delta^{(2)}(\boldsymbol{q}) \int d^2 \boldsymbol{l} \left| v(l_{\perp}^2) \right|^2 \right)$$

$$\mathcal{P}^{(1)}(\boldsymbol{p},L;\boldsymbol{p}_{0},0) = \int d^{2}\boldsymbol{r} \, e^{-i(\boldsymbol{p}-\boldsymbol{p}_{0})\cdot\boldsymbol{r}} e^{-\mathcal{V}(\boldsymbol{r})L} \, \frac{u_{\alpha}}{E} \left[2L \, \boldsymbol{p}_{0\beta} \left(\mathcal{V}(\boldsymbol{r})\delta_{\alpha\beta} + \mathcal{V}_{\alpha\beta}(\boldsymbol{r}) \right) \right) \\ - iL \, \nabla_{\beta} \mathcal{V}_{\alpha\beta}(\boldsymbol{r}) + iL^{2} \left(\mathcal{V}(\boldsymbol{r})\delta_{\alpha\beta} + \mathcal{V}_{\alpha\beta}(\boldsymbol{r}) \right) \nabla_{\beta} \mathcal{V}(\boldsymbol{r}) \right]$$

$$\mathcal{V}_{\alpha\beta} = \mathcal{C} \rho \left[-\boldsymbol{q}_{\alpha} \boldsymbol{q}_{\beta} \frac{\partial v^2}{\partial q_{\perp}^2} - (2\pi)^2 \delta^{(2)}(\boldsymbol{q}) \frac{\delta_{\alpha\beta}}{2} \int \frac{d^2 \boldsymbol{l}}{(2\pi)^2} v(l_{\perp}^2)^2 \right]$$

$$E\frac{d\mathcal{N}}{d^2\boldsymbol{p}\,dE} = \int \frac{d^2\boldsymbol{p}_0}{(2\pi)^2}\,\mathcal{P}(\boldsymbol{p},L;\boldsymbol{p}_0,0)\left[1-\boldsymbol{u}\cdot(\boldsymbol{p}-\boldsymbol{p}_0)\frac{\partial}{\partial E}\right]E\frac{d\mathcal{N}^{(0)}}{d^2\boldsymbol{p}_0\,dE}$$

- The odd moments of this re-summed final distribution are proportional to the transverse flow velocity, while the even moments are unmodified;
- The initial and final distributions are not factorized anymore in coordinate space (due to the energy derivative);

$$E\frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = f(E)\delta^{(2)}(\mathbf{p}_0)$$

uniform matter

Eikonal approximation -- $E \rightarrow \infty$

$$\langle p_{\perp}^{2k} \boldsymbol{p} \rangle = \int \frac{d^2 \boldsymbol{p} \, d^2 \boldsymbol{r}}{(2\pi)^2} \, p_{\perp}^{2k} \boldsymbol{p} \, e^{-i\boldsymbol{p}\cdot\boldsymbol{r}} e^{-\mathcal{V}(\boldsymbol{r})L} = 0 + \mathcal{O}\left(\frac{\perp}{E}\right)$$

Opacity expansion -- $\chi \equiv C \frac{g^4 \rho}{4 \pi \mu^2} L \ll 1$

$$\langle p_{\perp}^{2k} \boldsymbol{p} \rangle \simeq -\frac{\boldsymbol{u}}{2E} \mathcal{C}\rho L \int \frac{d^2 \boldsymbol{p}}{(2\pi)^2} p_{\perp}^{2k+2} \left[E \frac{f'(E)}{f(E)} v(p_{\perp})^2 + p_{\perp}^2 \frac{\partial v^2}{\partial p_{\perp}^2} \right]$$



$$\langle \mathbf{p} \rangle \simeq 3 \, \chi \, \mathbf{u} \, \frac{\mu^2}{E} \log \frac{E}{\mu}$$

$$E\frac{d\mathcal{N}^{(0)}}{d^2\mathbf{p}_0 dE} = f(E)\delta^{(2)}(\mathbf{p}_0)$$

inhomogeneous matter

$$\frac{d\mathcal{N}}{d^{2}\boldsymbol{x}dE} \simeq \exp\left\{-\mathcal{V}\left(\boldsymbol{x}\right)L\right\} \left\{ \begin{bmatrix} 1 - \frac{iL^{3}}{6E}\boldsymbol{\nabla}\mathcal{V}\left(\boldsymbol{x}\right) \cdot \hat{\boldsymbol{g}} \,\mathcal{V}\left(\boldsymbol{x}\right) \end{bmatrix} \frac{d\mathcal{N}^{(0)}}{d^{2}\boldsymbol{x}dE} + \frac{iL^{2}}{2E} \,\hat{\boldsymbol{g}} \,\mathcal{V}\left(\boldsymbol{x}\right) \cdot \boldsymbol{\nabla} \frac{d\mathcal{N}^{(0)}}{d^{2}\boldsymbol{x}dE} \right\} \\ \hat{\boldsymbol{g}} \equiv \left(\boldsymbol{\nabla}\rho\frac{\delta}{\delta\rho} + \boldsymbol{\nabla}\mu^{2}\frac{\delta}{\delta\mu^{2}}\right) \\ \rho \sim T^{3}$$

$$\langle \mathbf{p} \, p_{\perp}^2 \rangle \simeq \chi^2 \, \frac{L \nabla T}{2T} \, \frac{\mu^4}{E} \left(\log \frac{E}{\mu} \right)^2$$

inhomogeneous matter

 $\uparrow \quad \uparrow \quad \uparrow$

$$\langle \mathbf{p} \, p_{\perp}^2 \rangle \simeq \chi^2 \, \frac{L \nabla T}{2T} \, \frac{\mu^4}{E} \left(\log \frac{E}{\mu} \right)^2$$

- Opacity $\chi \approx 4$
- $u \approx 0.7$ (about $\pi/4$ to z-axis)
- $\mu = gT$ with $g \approx 2$ and $T \approx 500 MeV$

$$\left\langle \frac{p_{\perp}}{E} \right\rangle \simeq 3 \chi \, \frac{u_{\perp}}{1 - u_z} \frac{\mu^2}{E^2} \log \frac{E}{\mu}$$

```
What jet energy corresponds to \langle \theta \rangle \approx 1^o?

Jet energies: E < 50 \text{ GeV}
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inhomogeneous matter

- Opacity $\chi \approx 4$
- $u \approx 0.7$ (about $\pi/4$ to z-axis)
- $\mu = gT$ with $g \approx 2$ and $T \approx 500 MeV$
- $L\nabla T > T$

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What jet energy corresponds to \langle "\theta" \rangle \approx 1^o?

Jet energies: E < 100 \text{ GeV}
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$$\left\langle \frac{\mathbf{p}}{E} \frac{p_{\perp}^2}{\mu^2} \right\rangle \simeq \chi^2 \frac{L\nabla T}{2T} \frac{\mu^2}{E^2} \left(\log \frac{E}{\mu} \right)^2$$

Gluon emission



Summary

- We have already constructed a generalization of the jet quenching theory which includes the effects of the medium evolution and structure;
- It still should be further improved to take into account the non-equilibrium dynamics and the medium response (and more differential properties: heavy flavor, spin effects, etc.)
- Now we can make the next step turning to the actual phenomenology, and seek for the relevant observables, sensitive to the medium evolution;
- Having all these elements will allow us to turn to coupled simulations of the matter and jet observables, which are needed for quantitative predictions/analysis;