



**LIP Summer Internship 2022**

# Probing the Cosmic Ray Composition with SWGO

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**Student:** Afonso Guerreiro

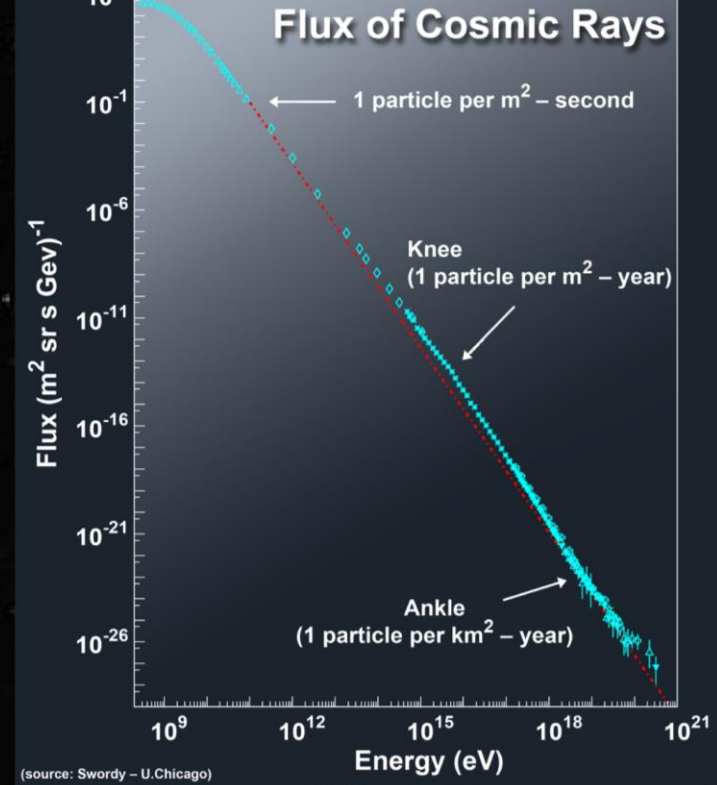
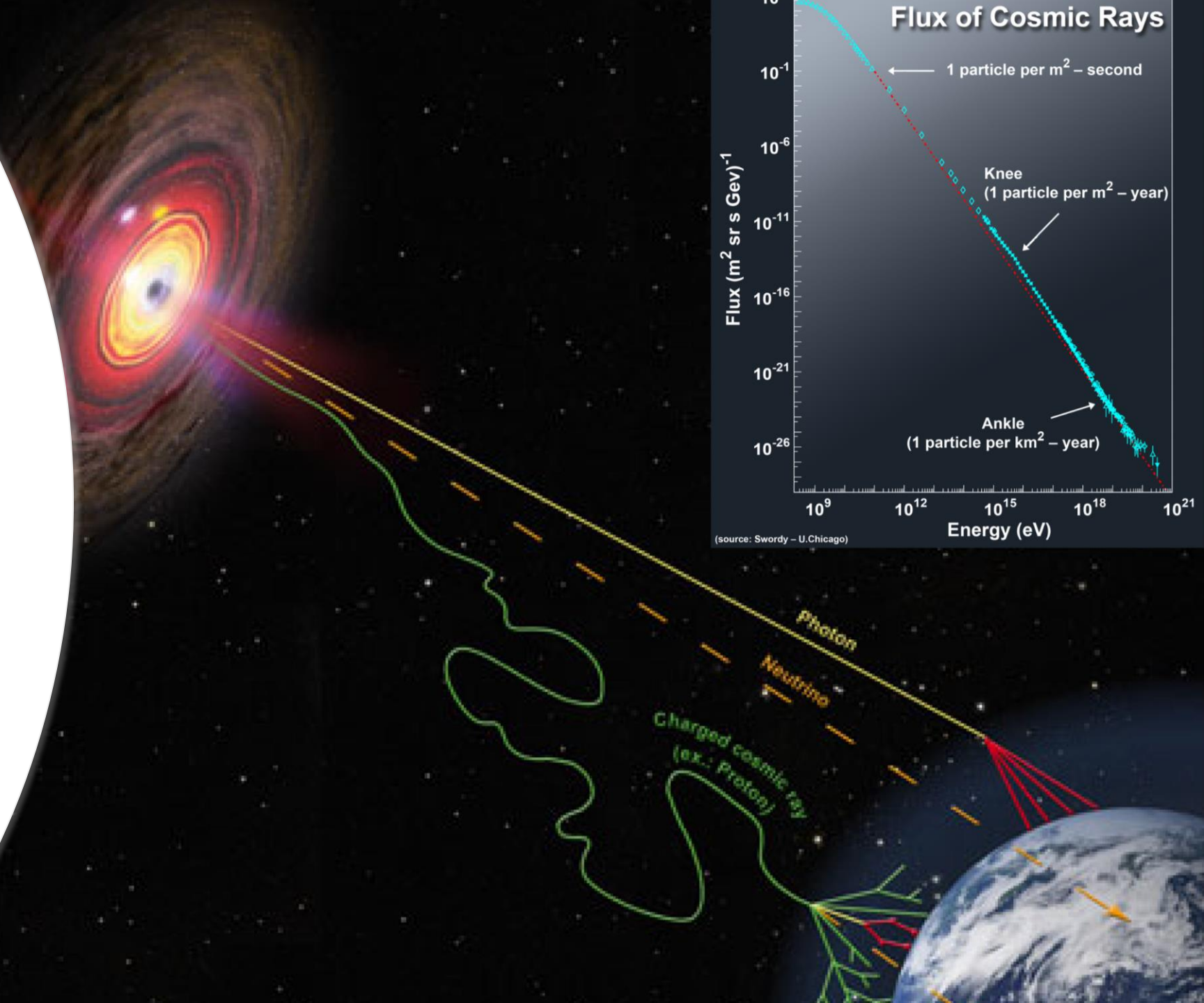
**Supervisors:** Prof. Rúben Conceição

Borja Serrano González



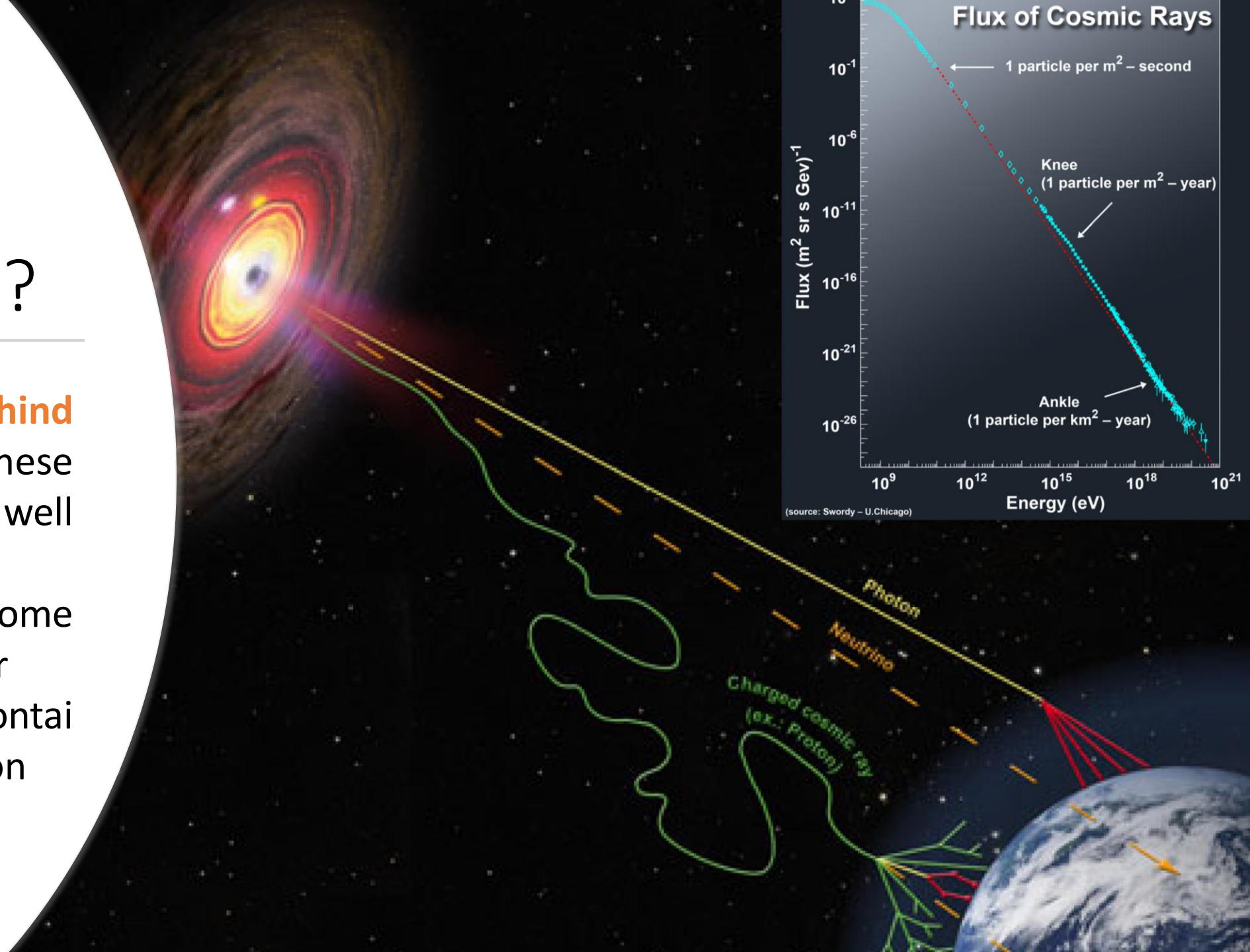
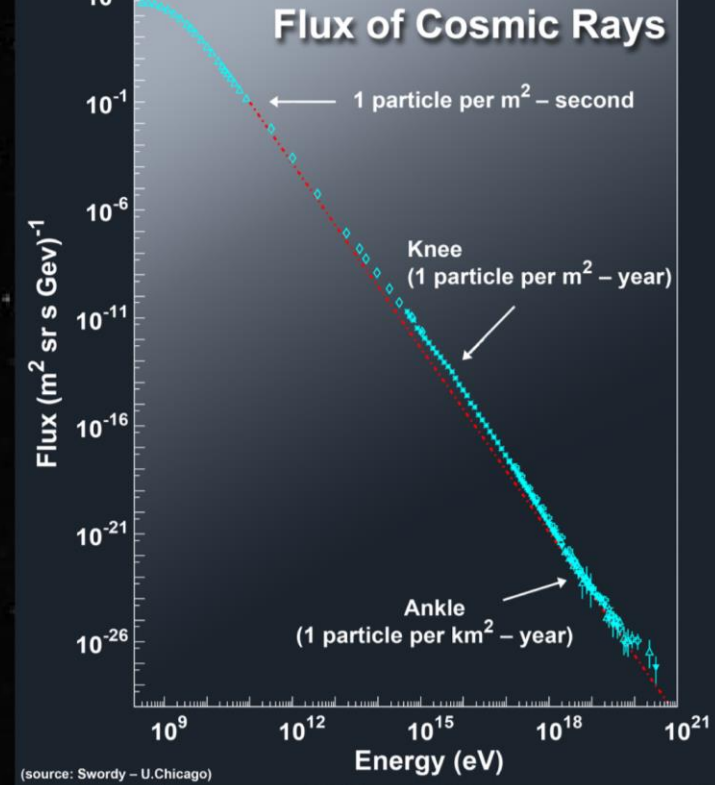
# What Are Cosmic Rays?

- Cosmic Rays are **high-energy particles** arriving from **outer space**
- Made mainly of **protons**, but also heavier nuclei
- Energies range from **1 GeV to  $10^8$  TeV**
- In the upper atmosphere, they collide, forming **particle showers**



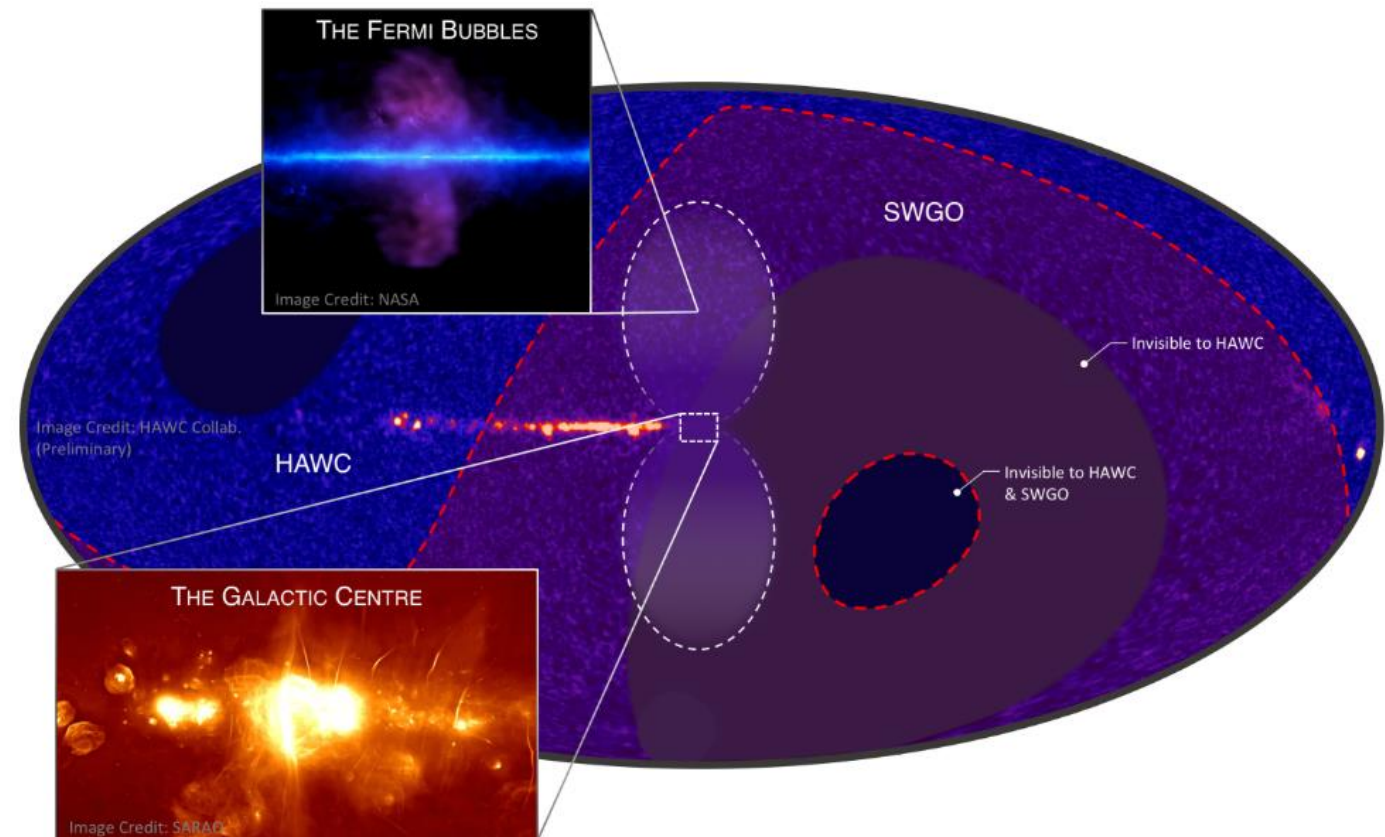
# Why study Cosmic Rays?

- The **mechanism behind the acceleration** of these particles is still not well understood
- Their **origins** in some cases are still not clear
- Their **composition** contains valuable information



# Southern Wide-field Gamma-ray Observatory (SWGGO)

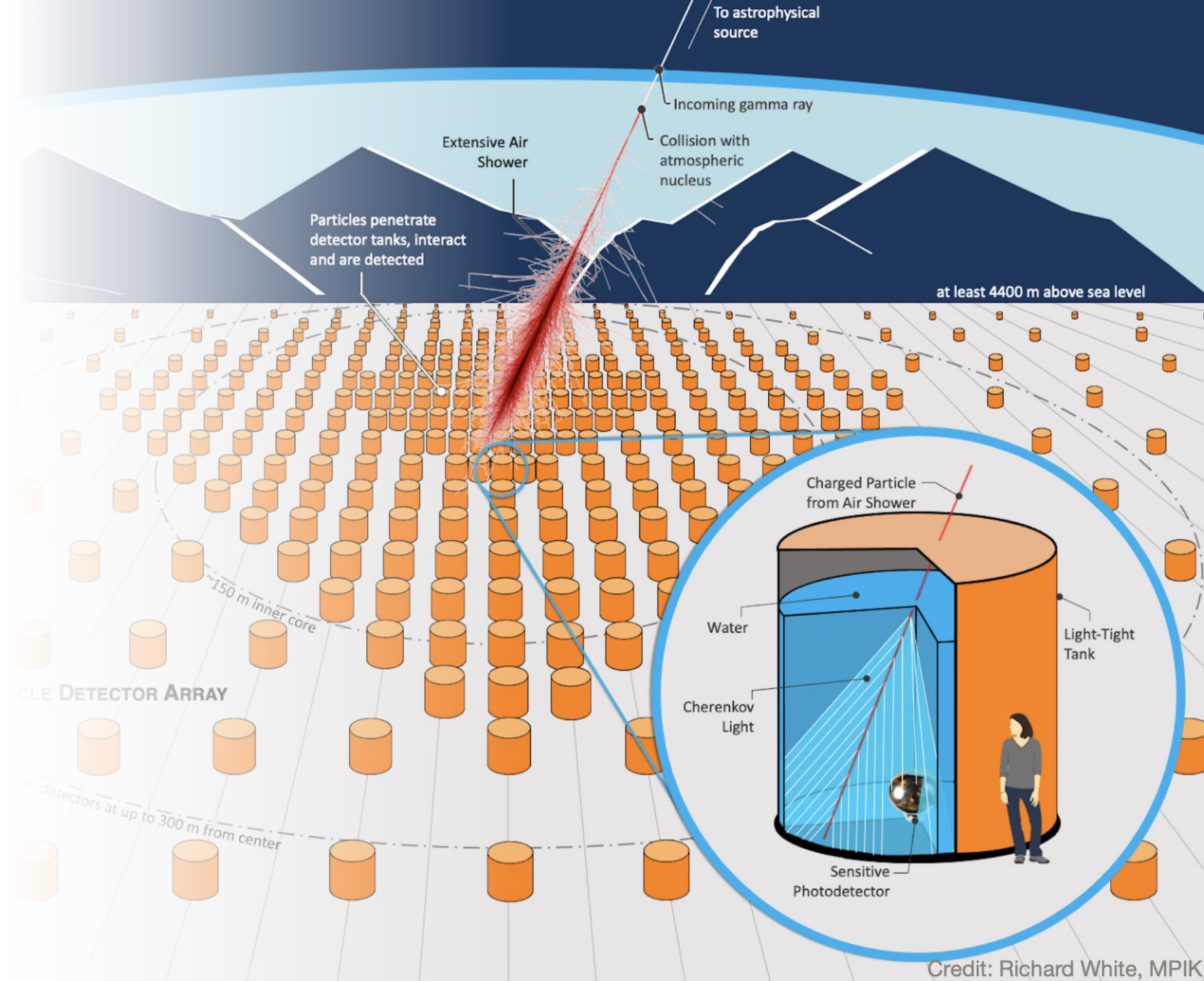
- A **gamma-ray observatory** based on ground-level particle detection
- Covering an energy range from **100s of GeV to 100s of PeV**
- Located in **South America** at a latitude between 10 and 30 degrees south
- At an **altitude of 4.4 km or higher**



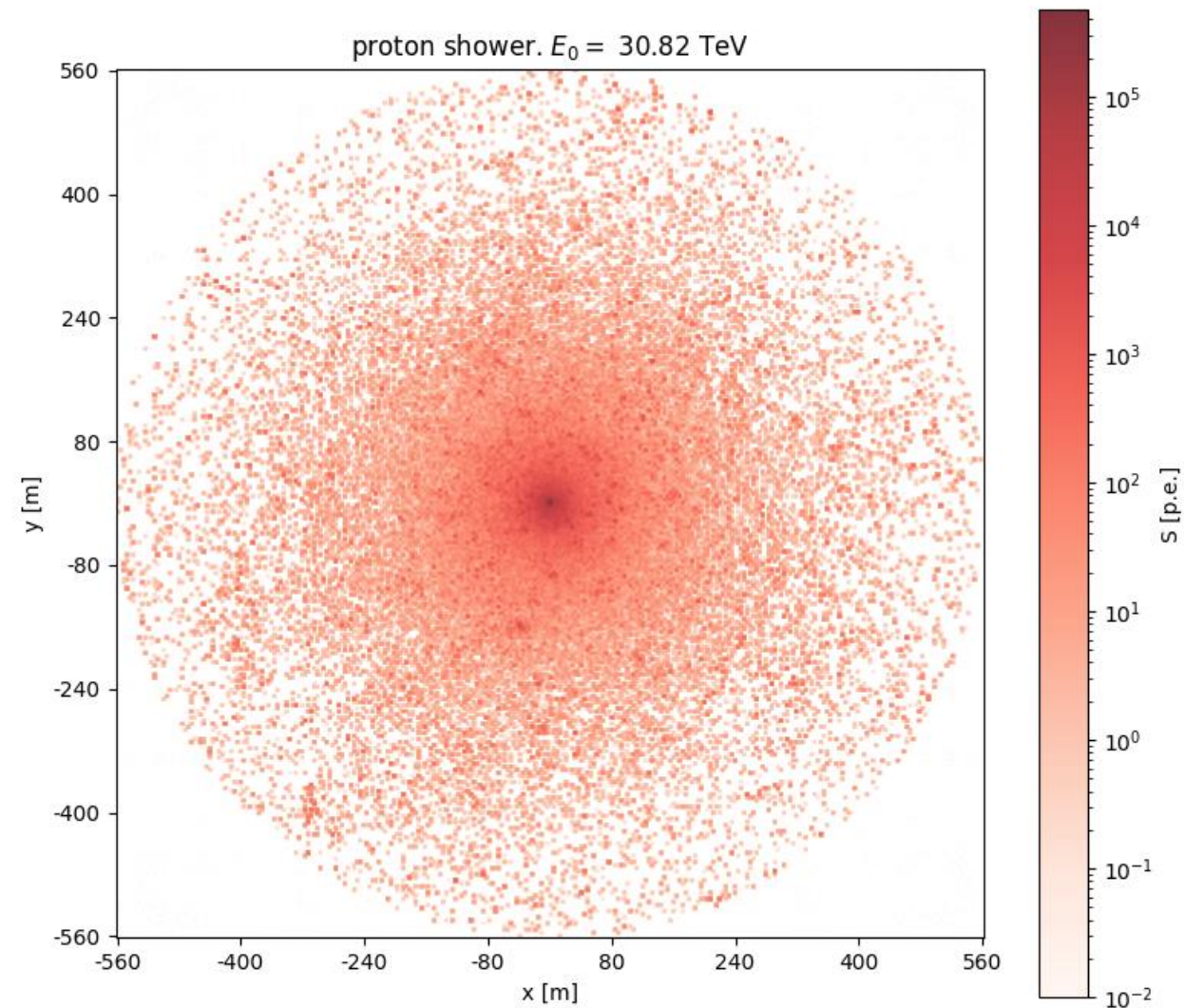


# Detection Process

1. The incoming **gamma ray** interacts with the atmosphere generating a particle shower
2. The resultant particles enter the detector producing **Cherenkov radiation**
3. The signal is picked up by **photo-multiplier tubes** in the station



# Simulating Cosmic Rays

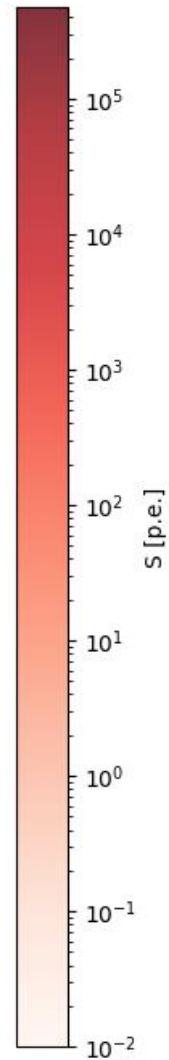
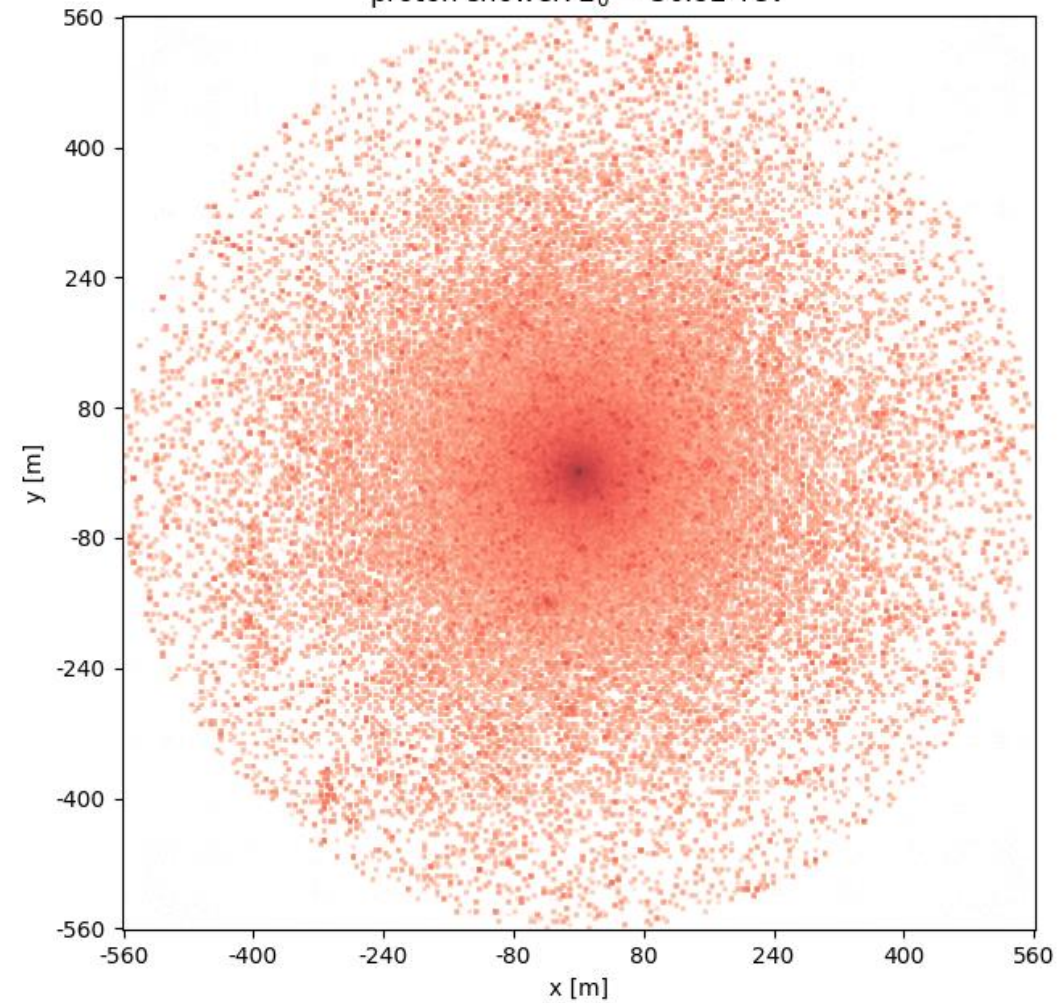


- The Simulation of extensive air showers was done resorting to **CORSIKA**
- An energy cut was used to make sure the reconstructed energies of the showers were comparable
- The particles collected at the ground were converted into a signal
- All showers are vertical

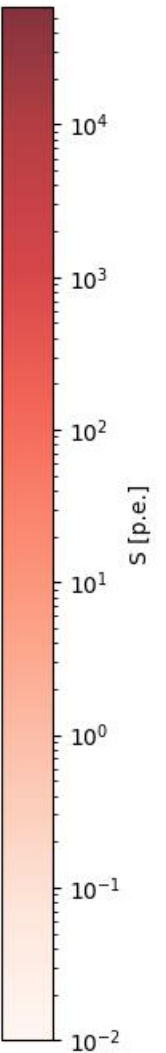
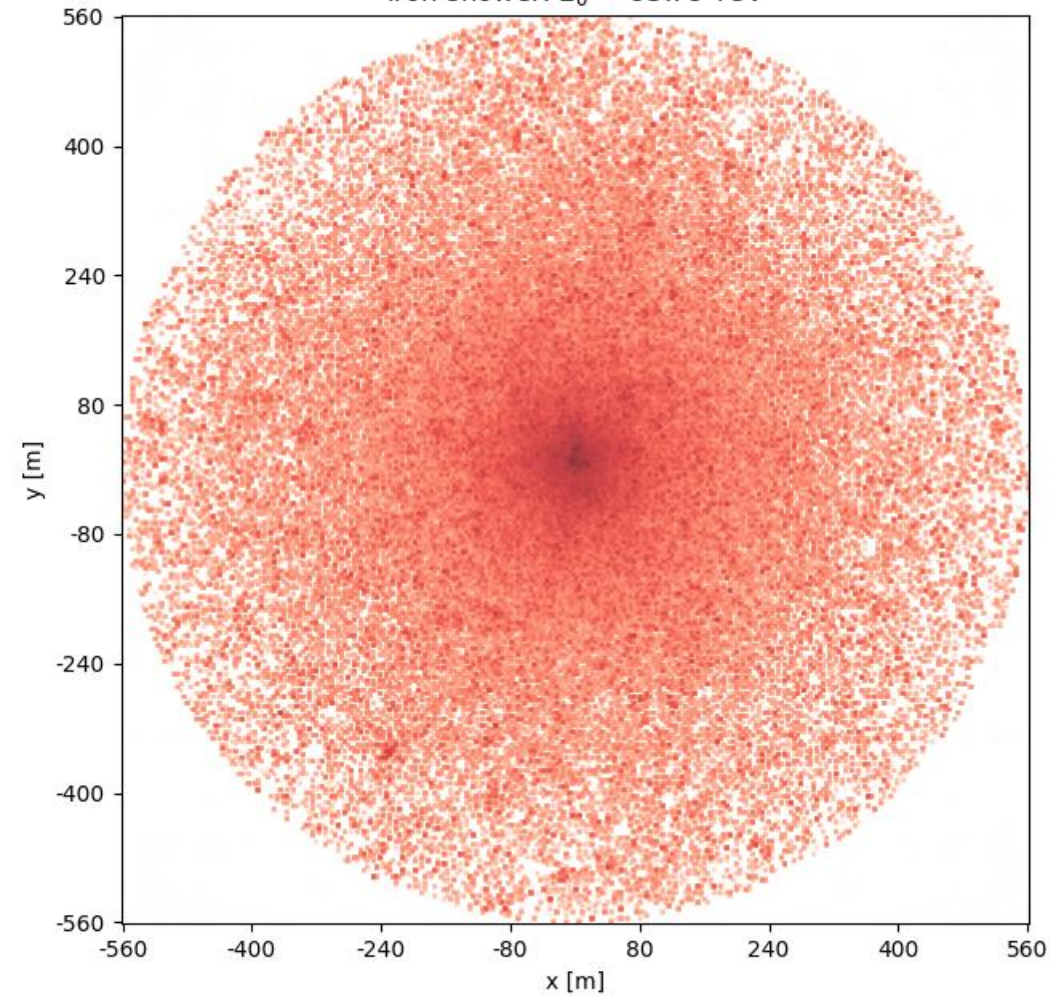


# Can we create an algorithm to distinguish proton showers from iron showers?

proton shower.  $E_0 = 30.82$  TeV

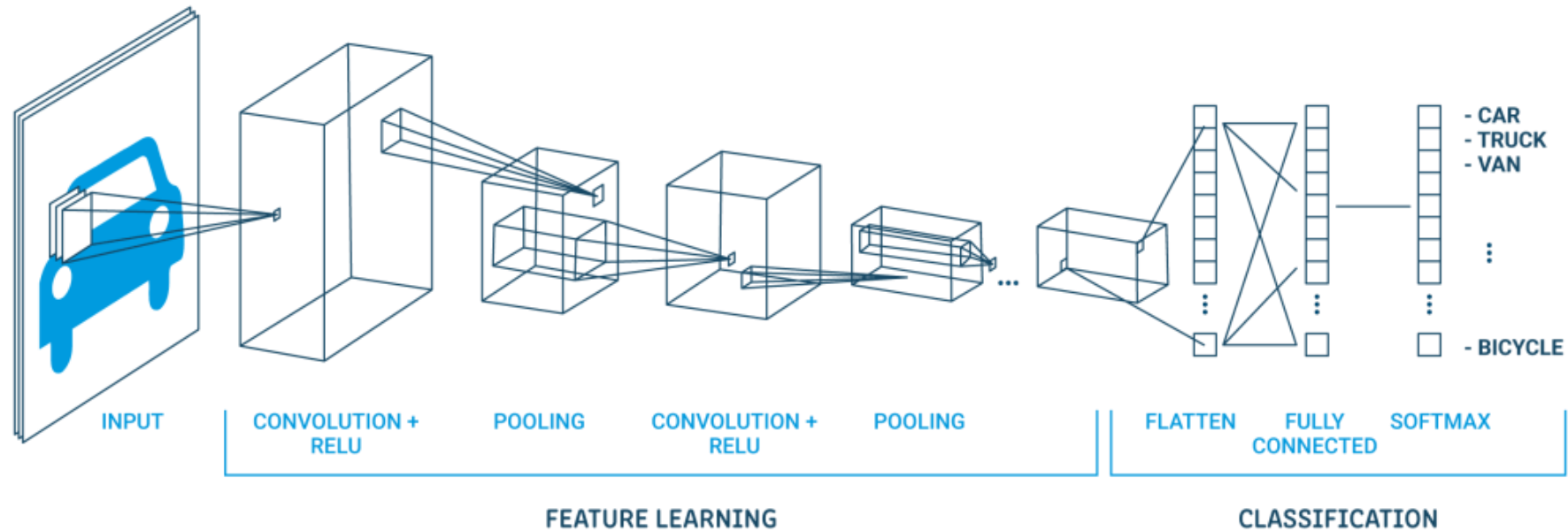


iron shower.  $E_0 = 93.78$  TeV



# Convolutional Neural Networks (CNNs)

A **Convolutional Neural Network (ConvNet/CNN)** is a Deep Learning algorithm that can take an image as input, assigning importance (learnable weights and biases) to various aspects/objects in the image, becoming able to differentiate one from the other.

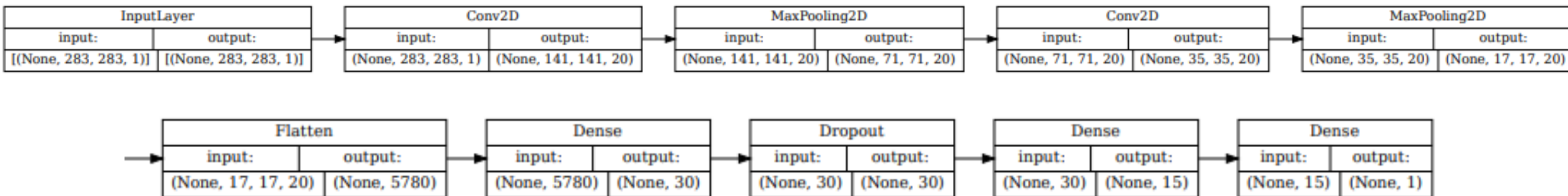




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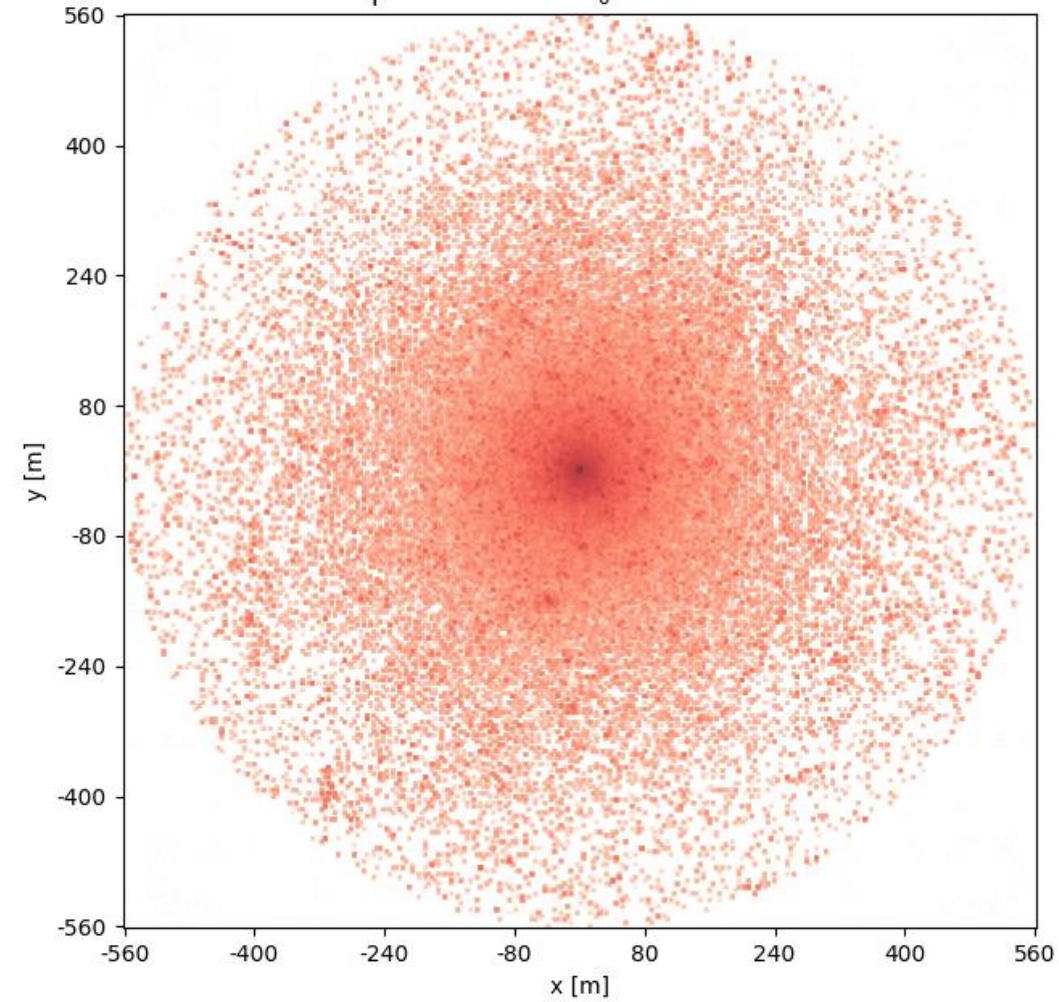
The **structure** of the Network used was



The **Fill Factor(FF)** is the ratio of the are occupied by detectors to its total area

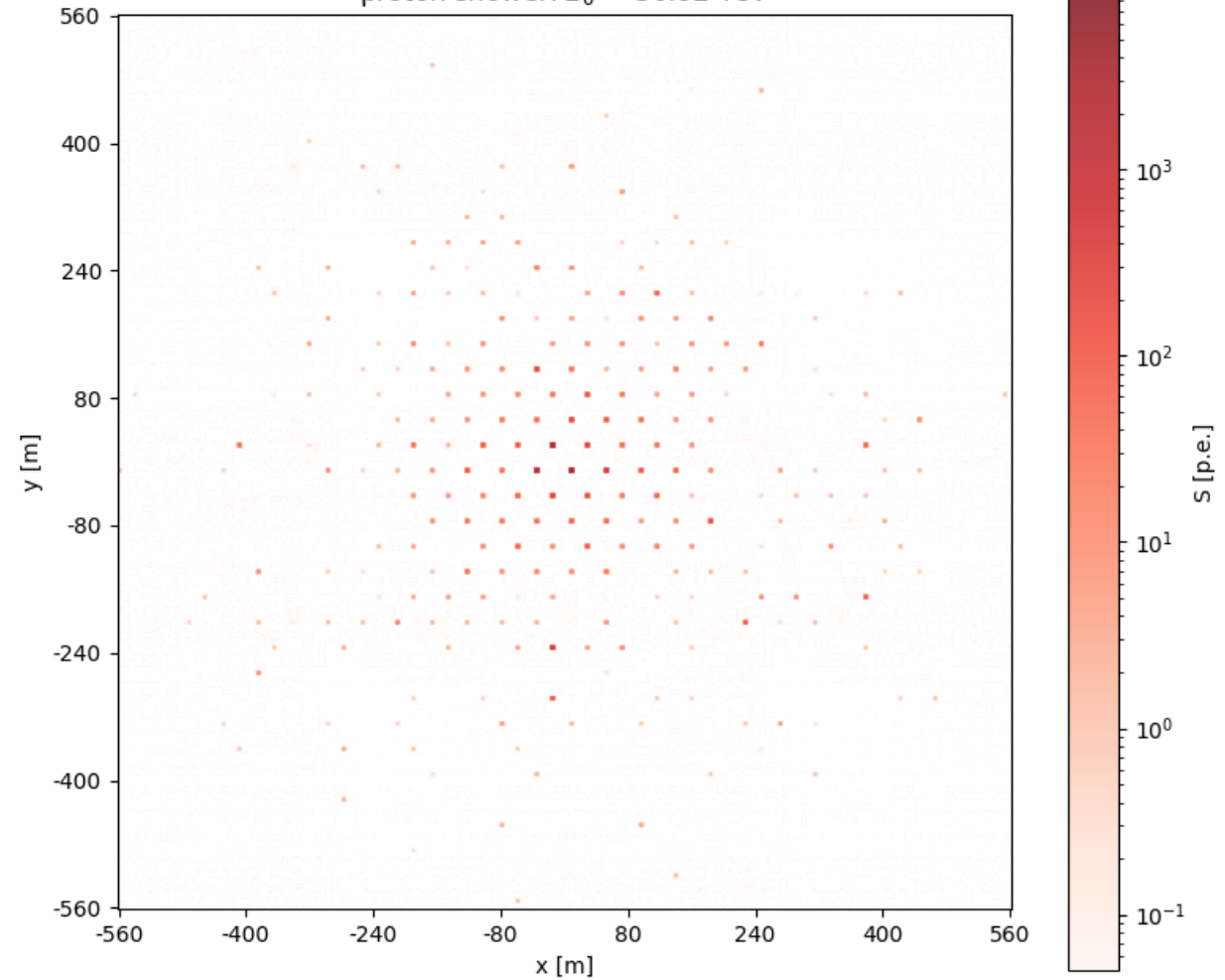
**FF=100%**

proton shower.  $E_0 = 30.82$  TeV



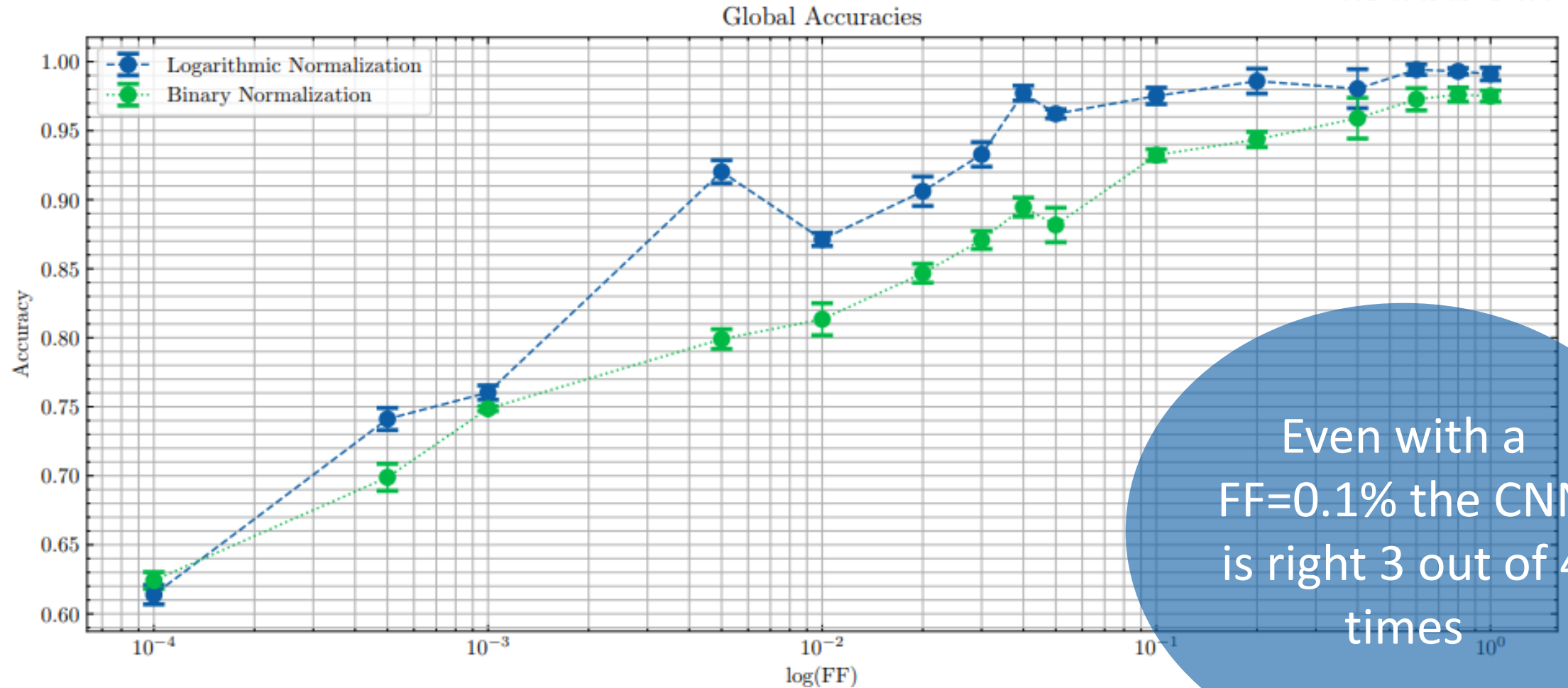
**FF=1%**

proton shower.  $E_0 = 30.82$  TeV





# Results



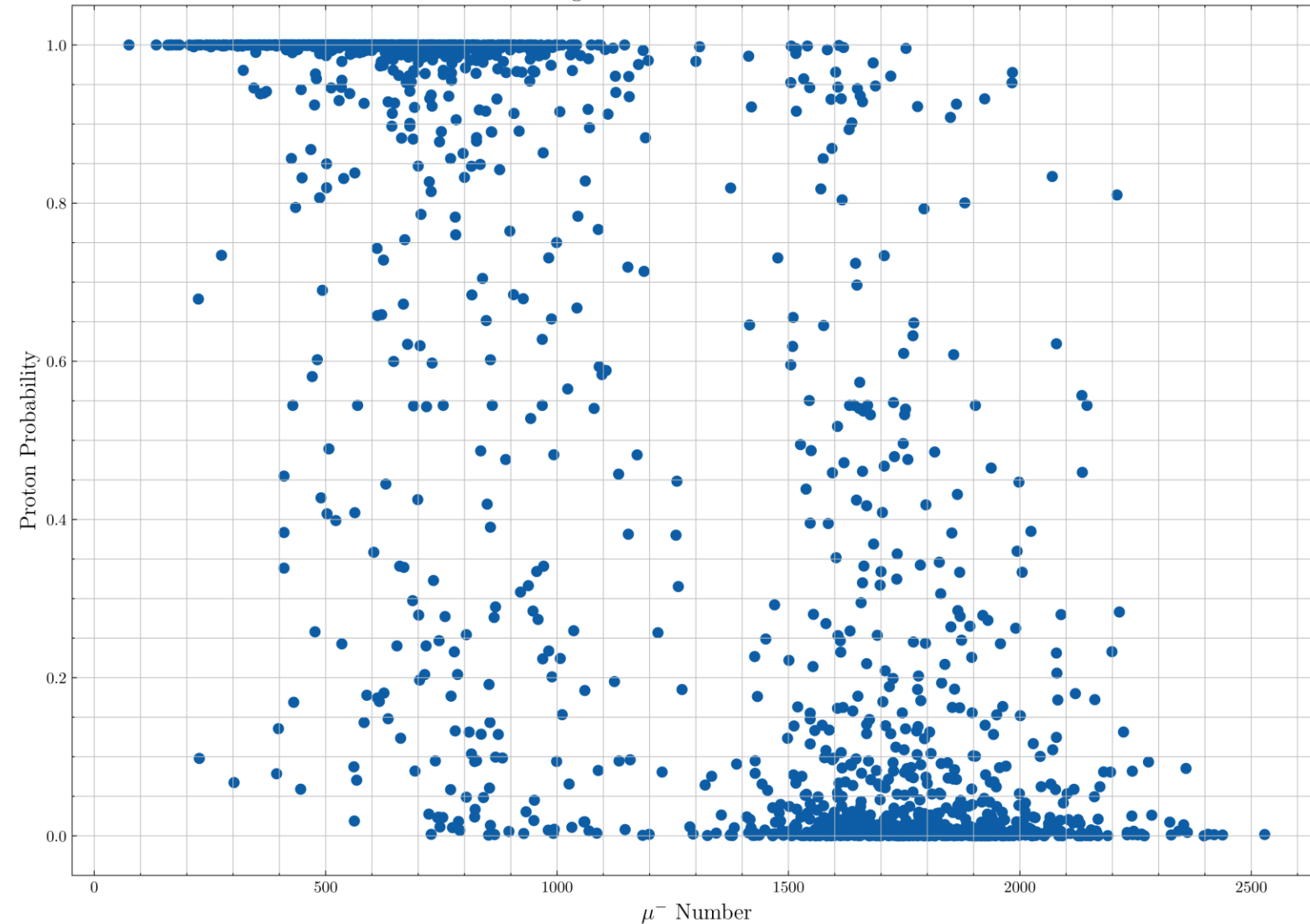
Even with a  
FF=0.1% the CNN  
is right 3 out of 4  
times

**Logarithmic Normalization**-> Used the Log of the signal normalized between 0 and 10

**Binary Normalization**-> Cell is either 1 (active) or 0 (inactive)

FF=1%

Logarithmic Normalization



- Number of muons very **sensitive to composition**
- The measurement of the number of muons is **very hard** to obtain
- **No correlation** between CNN prediction and number of muons
- CNN is picking **shower pattern** features sensitive to composition



# Conclusions

1

A **Machine Learning algorithm** can be used to distinguish between proton and iron showers

2

**Good Discrimination** can be attained until values of  $FF \sim 1\%$

3

The CNN is picking information on the **shower footprint** rather the shower muon content



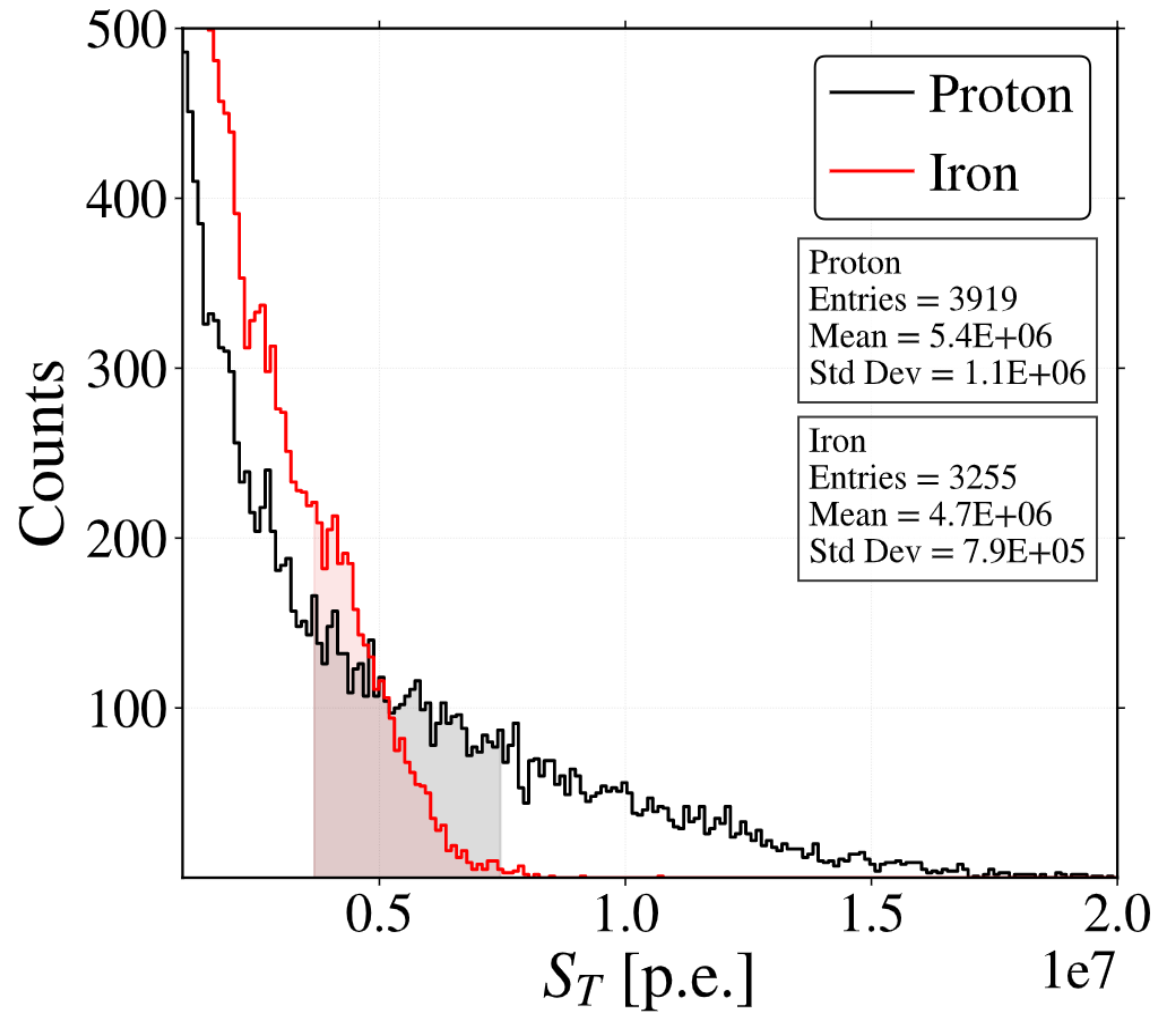
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**Thank you for  
you attention**





# Adicional Slides



- Shaded area corresponds to **selected events**
- Energies range from **40 to 60 TeV**