



From Particle Physics to Space

... and health

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Radiation Environment in the Solar System

Galactic Cosmic Rays low flux but highly penetrating **Protons & ions**

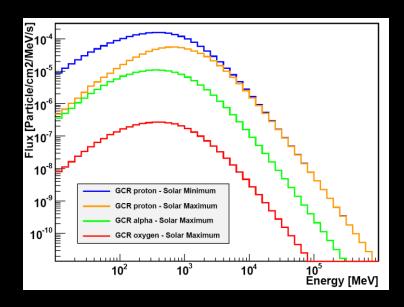
Solar Particle Events sporadic, intense & dangerous **Electrons, protons & ions**

Radiation Belts

high radiation dose **Electrons & protons**



"Galactic" Cosmic Rays (GCR)



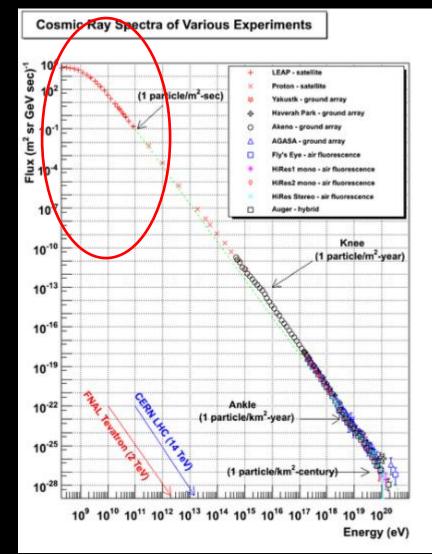
Low flux but highly penetrant

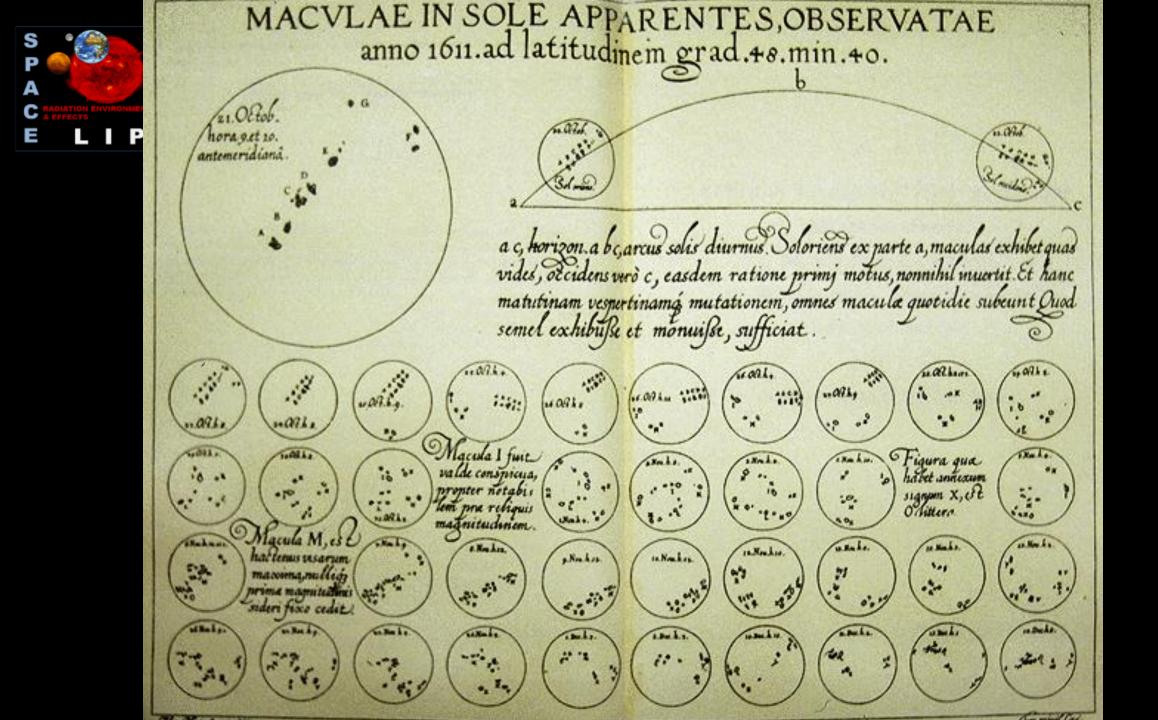
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- Protons and nuclei: energy spectra peak at ~1 GeV/n
- Solar cycle modulated flux inversely proportional to the Sun's activity
- E < 1 GeV/n: highly affected by solar activity</p>









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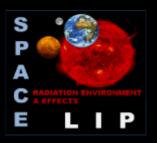
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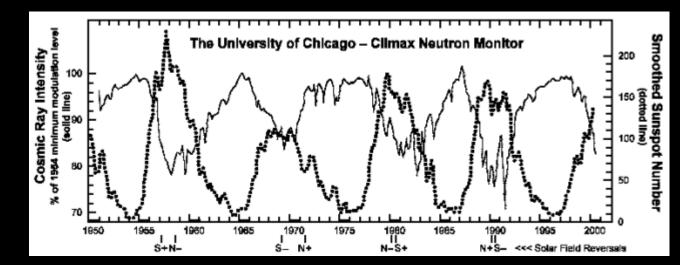
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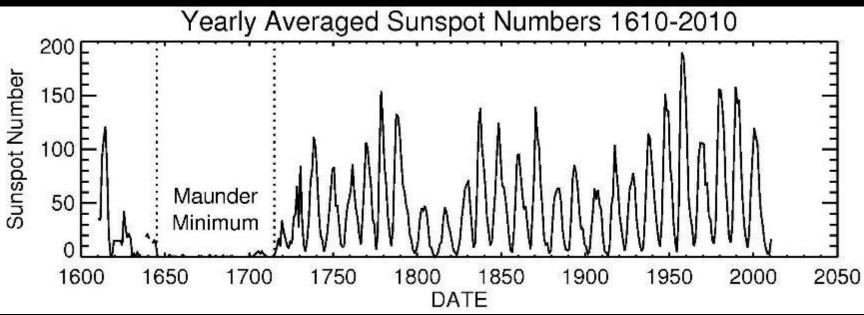


11 year Solar Cycle

Modulation with solar activity

- \star Solar Maximum: solar storms and SEP
- ★ Solar Minimum: more GCR

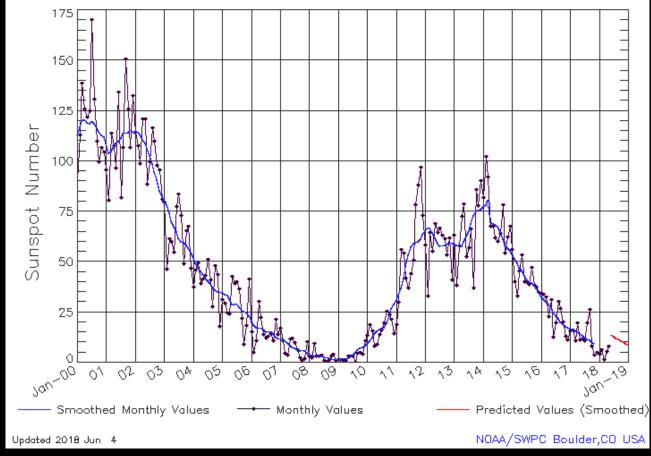




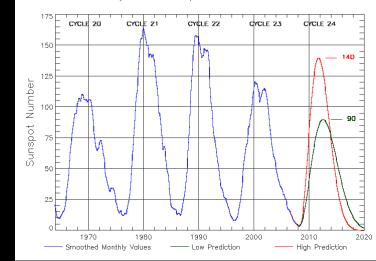


Solar Cycle 24

ISES Solar Cycle Sunspot Number Progression Observed data through May 2018



Solar Cycle 24 Sunspot Number Prediction



Solar Energetic Particle events

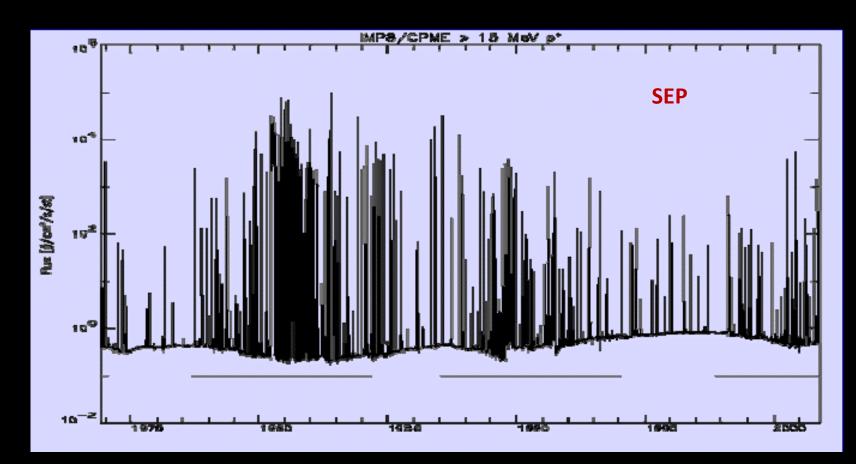
- P ★More frequent in "maximum" solar activity years
 - ★Highly unpredictable

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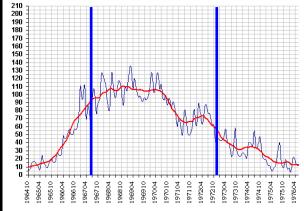
★ Design for by making statistical assessment

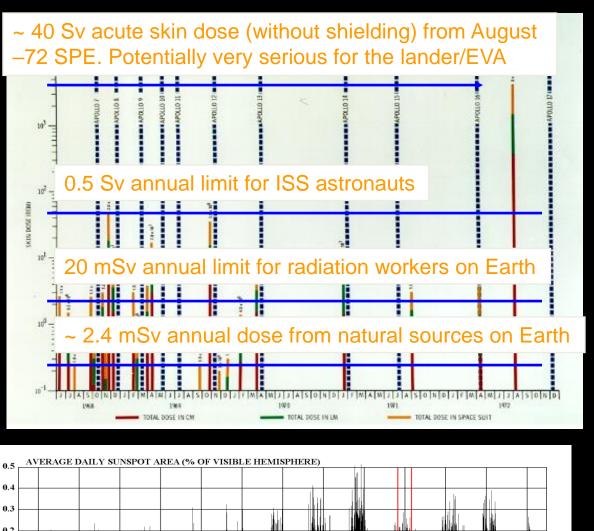




Apollo missions: Solar maximum







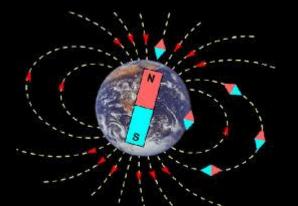
http://science.msfc.nasa.gov/ssl/pad/solar/images/bfly.gif

DATE



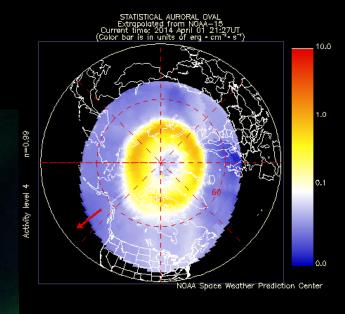
The Magnetosphere: an invisible shield

In Earth orbits the radiation belts containing trapped electrons and protons constitute the major radiation source.

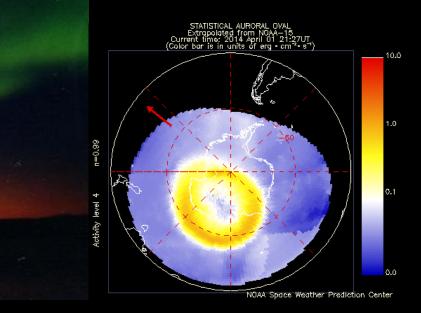




Auroras



Charged particles captured in the radiation belts excite N2 and O2 molecules that emit visible light while returning to the fundamental state.





Earth Radiation Belt Regions

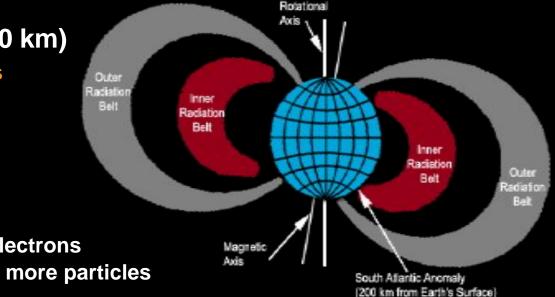
High radiation dose, electrons (<10 MeV) & protons (<250 MeV)

Inner belt (700-10000 km)

dominated by protons CRAND = Cosmic Ray Albedo Neutron Decay ~static E~100's MeV

Outer belt (~20000-70000 km)

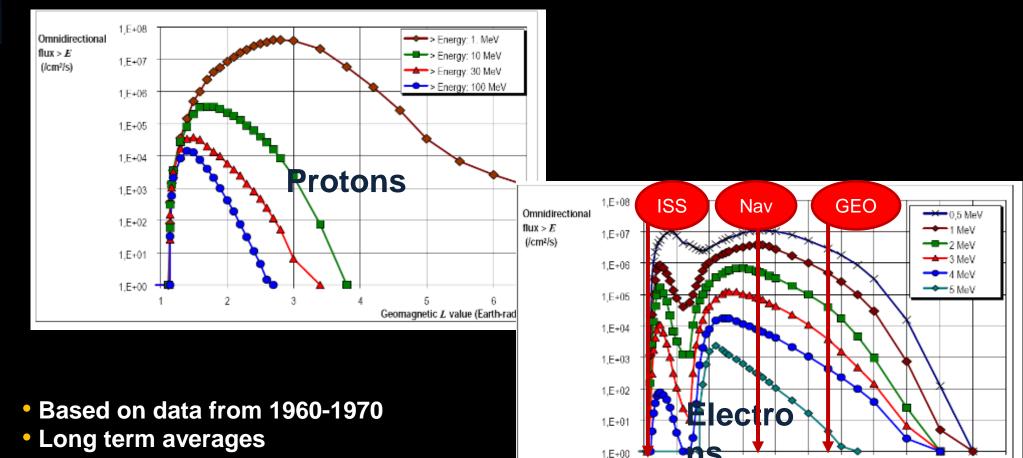
dominated by electrons Controlled by "storms" Very dynamic E~ MeV



Slot

low intensities of MeV electrons occasional injections of more particles





9

Geomagnetic L value (Earth-radii)

8

Δ

10

11

12

but : outer belt is very

stormy

S P A

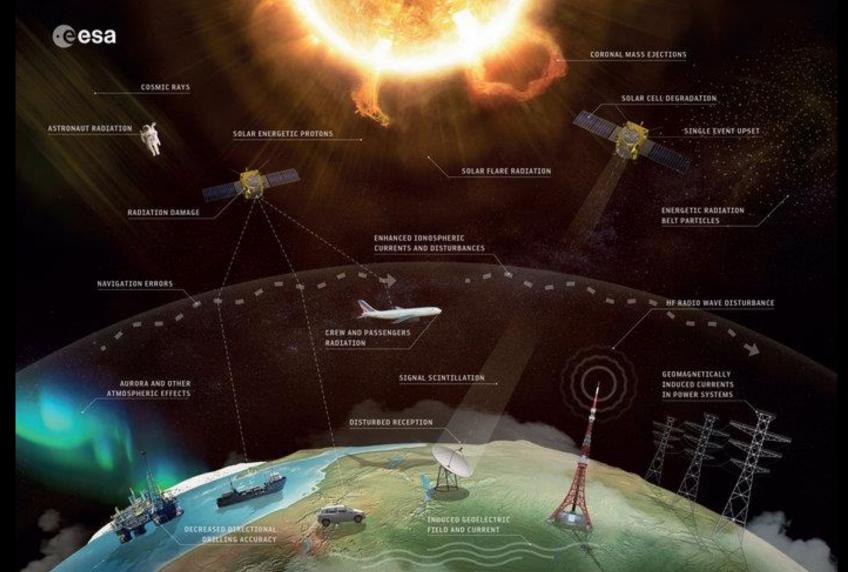
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LIP

ongoing work to update models

Radiation effects on Earth and in orbit

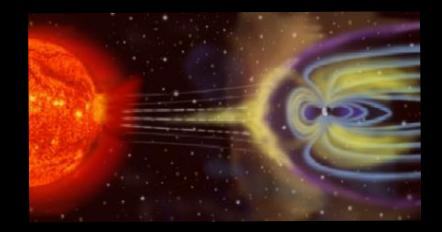


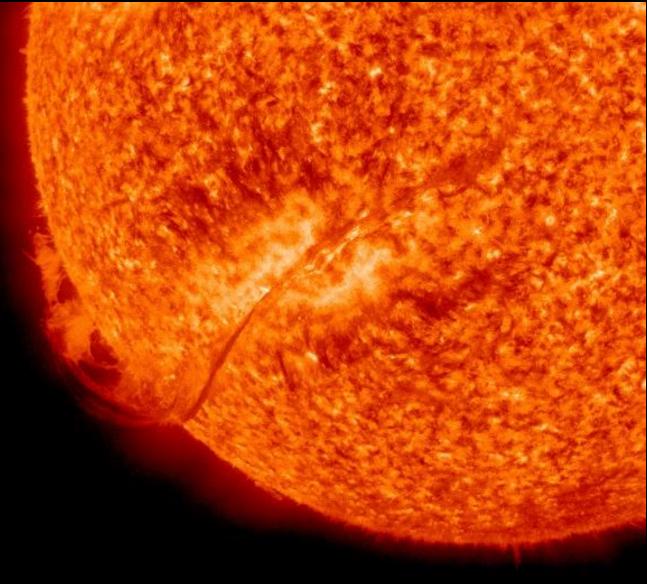
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In space, the radiation environment is responsible for spacecraft system, sub-system and component hazard and damage and it is also responsible by strict constrains on human space exploration.

At Earth's surface, the atmosphere in conjunction with the geomagnetic field provides considerable protection against cosmic rays and solar particle events.

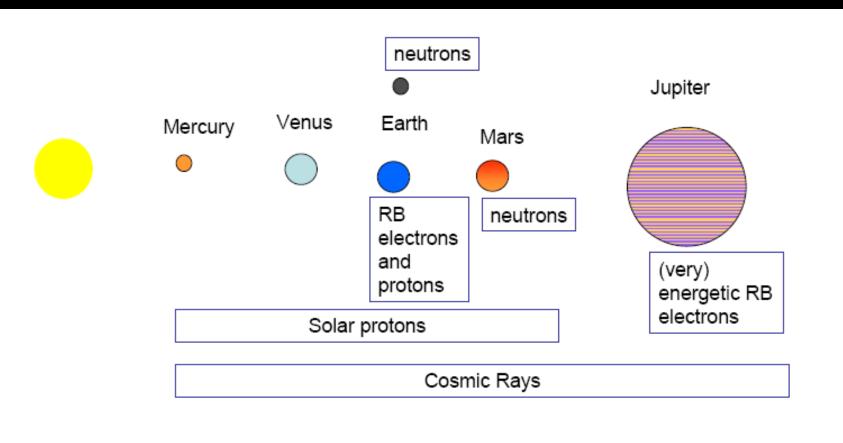




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Radiation Environment in Space



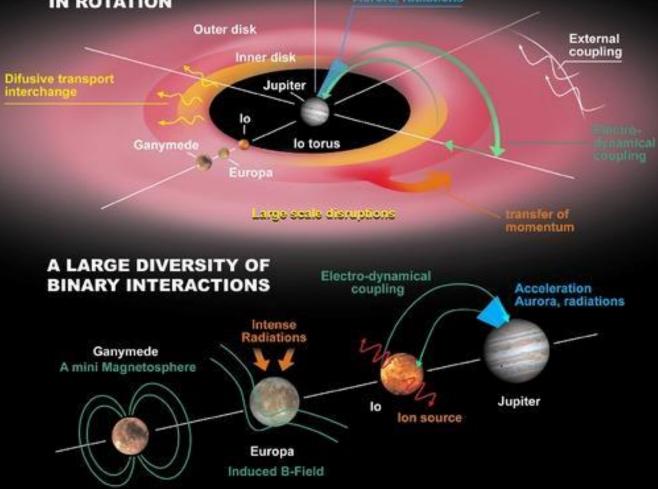
RB= radiation belt



Jovian System

Energetic Particle Environment

A GIANT SYSTEM Severe environment in terms of ionizing particles





Juice Mission Objectives

Study the emergence of habitable worlds around gas giants



Europa Ganymede

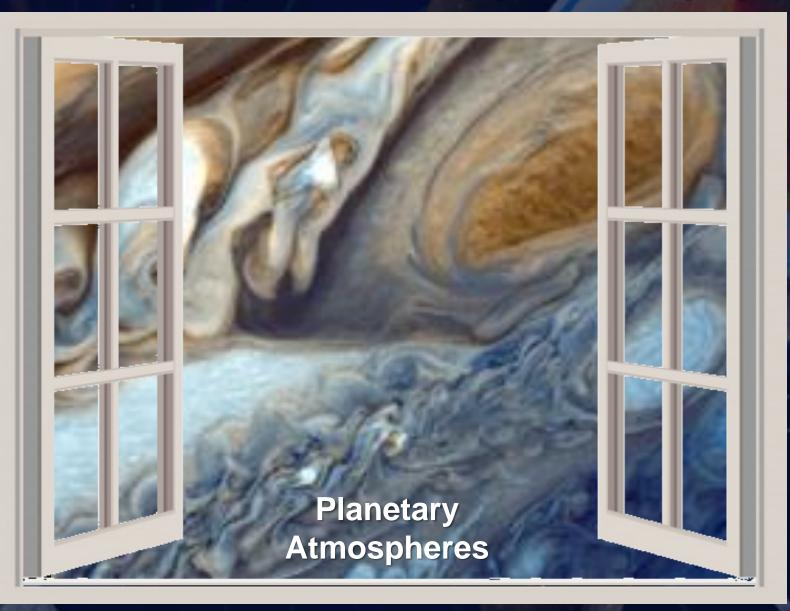
Callisto

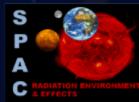
- Characterise the Jupiter Icy Moons: Ganymede, Europa and Callisto as planetary objects and potential habitats
- Explore the Jovian system as an archetype for gas giants



Juice mission

Implications for Astrophysics & Planetary Physics

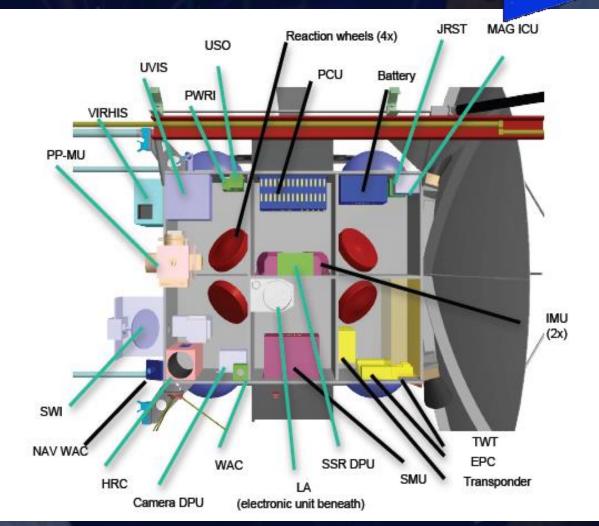




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Juice Scientific Payload

I P 11 instruments with total mass of ~100 kg



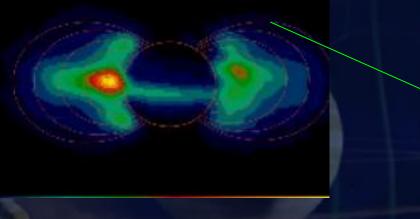
Possible configuration for Scientific Payload accomodation



Data and models

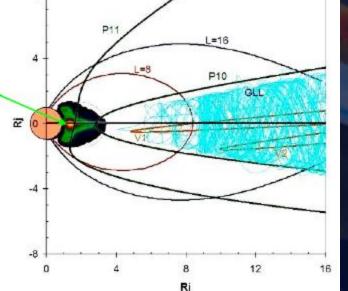
Synchrotron emission observations

obactived map actived miliz



Synchrotron emission predictions from Divine model

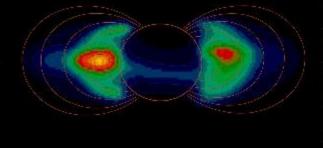
Revised Model



Limited data from Missions

- Pioneer 10-11 (1973, 1974)
- Voyager 1-2 (1979)
- Galileo (1989)
- Juno (NASA) 2017

Courtesy of H.B.Garret, Jet Prop.Lab

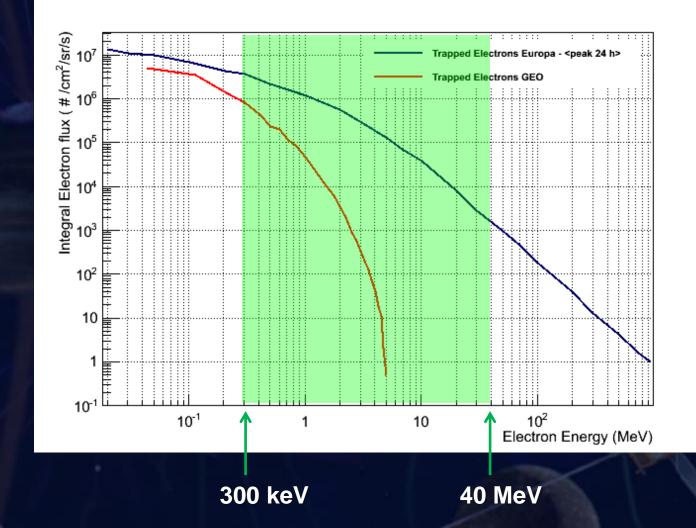


Models

- -Divine e- and proton models
- GIRE (Galileo Interim Radiation Environment Model)
- Salammbô(e-belt model from Jupiter surface to Europa)
- JOSE (based on Divine+GIRE+Salammbô+data)



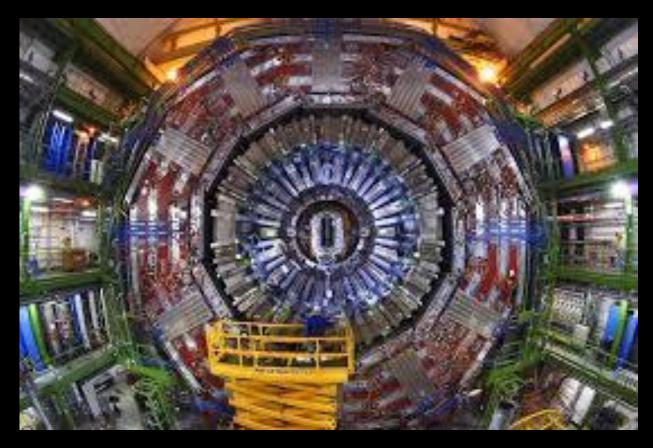
Very Hard Electron Spectrum





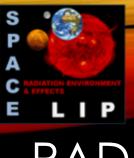
Radiation Monitors





Particle detectors in Space

Mass ~ 1 kg Power ~ 1 Watt Volume - 1 lt



RADIATION ENVIRONMENT MEASUREMENT

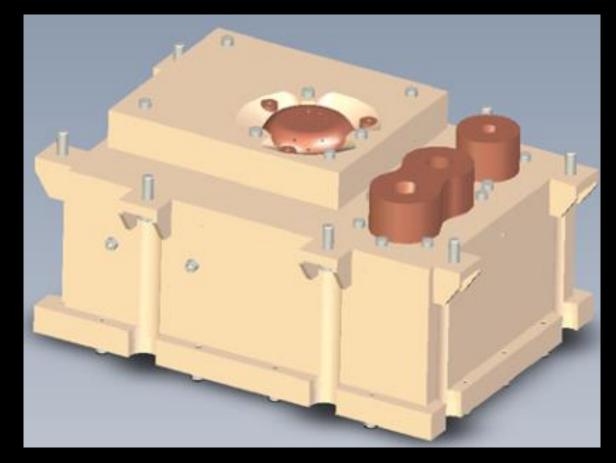
RADEM

Main Objectives

- Ensure mission safety
- Provide valuable scientific data

Requirements

- Electron detector
 - Spectral range 300 keV 40 MeV
 - o Peak Flux $10^9 \text{ e/cm}^2/\text{s}$
- Proton Detector
 - o Spectral range 5 MeV-250 MeV
 - o Peak Flux $10^9 \text{ e/cm}^2/\text{s}$
- Particle Separation
 - From Helium to Oxygen
- Dose determination
- Low mass
- Low power consumption







Electron Directionality Detector

JOVIAN RAD-HARD ELECTRON MONITOR PROTO-FLIGHT MODEL ESA/ESTEC CONTRACT 1-7560/13/NL/HB EFACEC,LIP, RUAG,PSI,IDEAS

170

LPA (degree)

Source State State

304-527 keV electrons

DATA: EPD measurements Galileo G29 encounter with Ganymede in December 2000

90

0

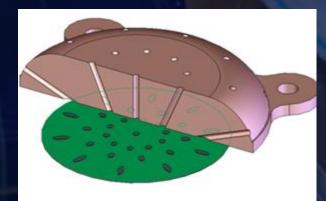
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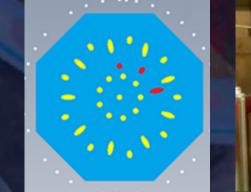


DD- THE RADEM DIRECTIONALITY DETECTOR

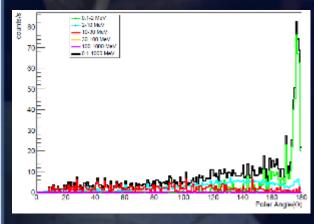
Design, testing and performance:

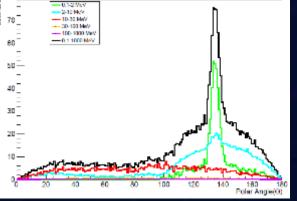
- Copper Collimator 28 directions
- Single 505µm Kapton Absorber
- Single Detection Plane
- 28 Si Diodes
- 3 Blind Sensors

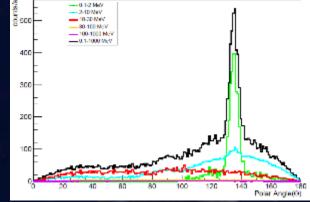


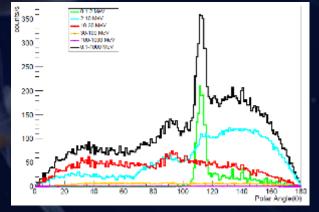












S P A C RADIATION INVERTION

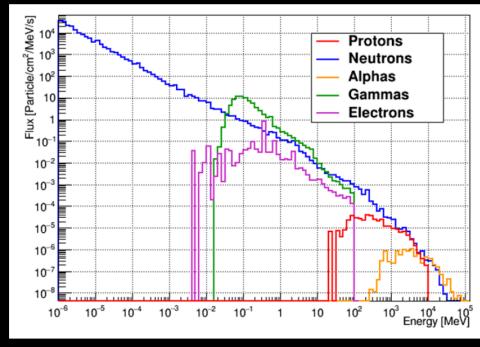
RADIATION ENVIRONMENT MODELING

MARSREM: THE MARS ENERGETIC RADIATION ENVIRONMENT MODELS

LIP developed **dMEREM**, a **Geant4 based model for the radiation environment on Mars**, **Phobos and Deimos**, including local treatment of surface topography and composition,

Inputs given as a function of latitude, longitude, in a 5 x 5 degree grid, and season.

Example of dMEREM results: Particle radiation arriving on Mars surface after GCR-alpha interaction with atmosphere and soil







...IT IS POSSIBLE !

It is possible to remain in Martian surface for some time with no serious risk for the astronauts!

For longer permanences shelters are required...

LIP ESA contracts

ESA JUICE mission

- RADEM Radiation Hard Electron Monitor
- EEE component testing for Jovian environment

ESA Mars Energetic Radiation Environment Models

Expert support to BERM and RADEM radiation monitors From Mercury to Jupiter

GEO Radiation Environment:

- Radiation Environment Measurement (MFS)
- EEE component test bed (CTTB)





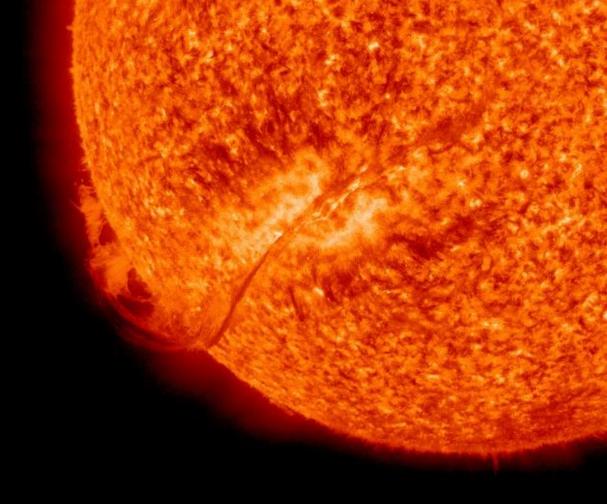


What can Space & Biomedical applications have in common?



RADIATION





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http://spaceweather.com/



The interplanetary travel

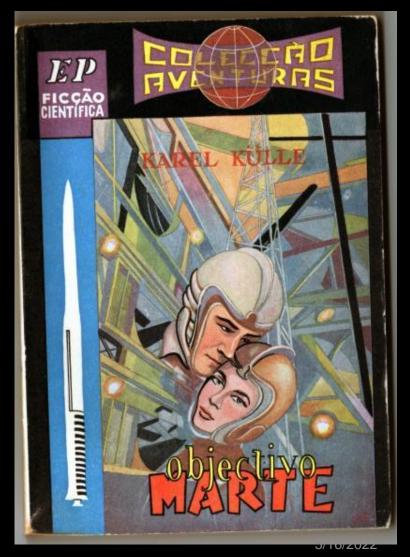
The most dangerous phase while travelling to Mars, from the point of view of the radiation hazard, is the interplanetary travel !

The biggest danger is the possibility of a SEP reaching the mission..

Mitigation Strategies:

- Shelters inside water compartments or other
- Faster propulsion system
- SEP Forecasting tools and alarms

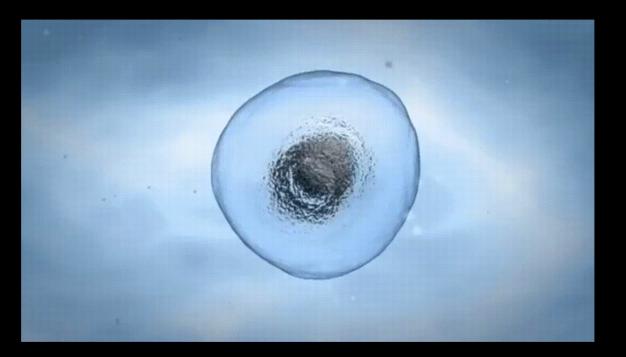
Radiopharmaceuticals ...





What is cancer?

Cancer is a group of diseases in which there is an abnormal and uncontrolled proliferation of cells that originated from a "normal" cell that mutated, giving rise to cells that have the ability to continue to multiply, spreading to other tissues and organs beyond those to which the original cell belonged.





Radiotherapy

Radiotherapy with radioactive sources or internal radiotherapy

• Brachytherapy:

Photons and electrons released from sources - "seeds" - placed close to the tissue to treat





Radiotherapy with external beams

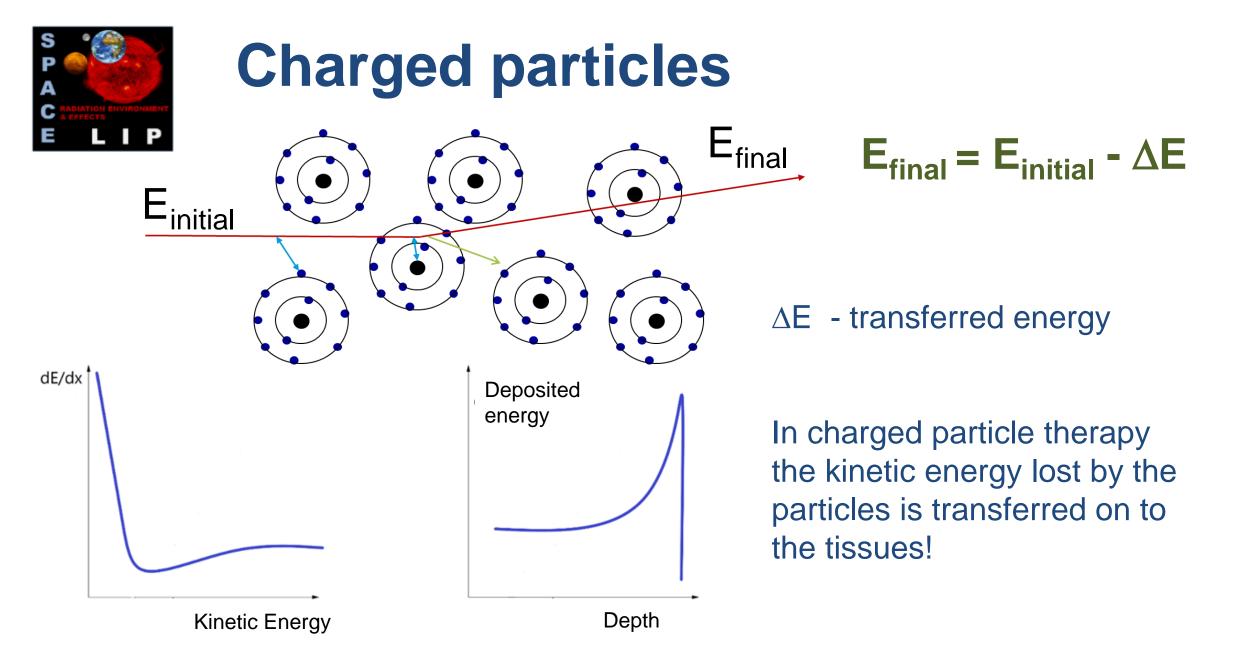
LINACs

electrons and photons

• Cyclotrons and synchrotrons protons and carbon ions









Therapy with hadrons

Proton/ion therapy

- energetic protons/ions are produced in an accelerator
- energetic protons/ions are directed at the tumor
- Varying the energy of the protons/ions results in good deep control
- Beam can be focused to the size of a pin
- less damage to healthy tissue than electron or gamma therapy



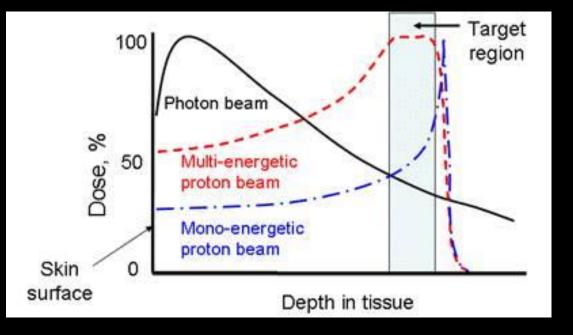


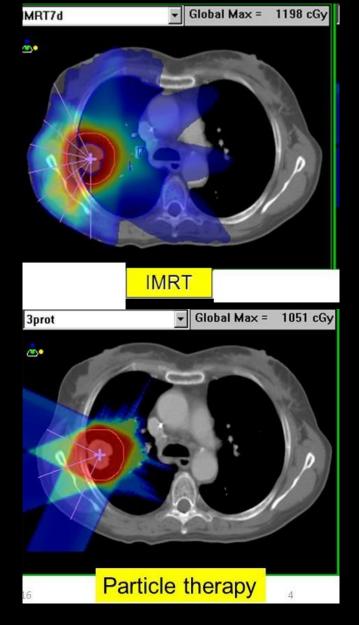


Therapy with hadrons

therapeutic application of accelerators:

particle-based radiotherapy, including protons and heavy ions







and there is more...

Questions today Theses tomorrow